MATERIALS SCIENCE

P. Sheng Introduction to Wave Scattering, Localization and Mesoscopic Phenomena



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P. Sheng

# Introduction to Wave Scattering, Localization and Mesoscopic Phenomena

Second Edition

With 72 Figures



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# Preface to the Second Edition

It has been ten years since the publication of the first edition. During this period there has been much progress in research on waves and localization. In particular, light localization has been observed in semiconductor powders, and random scatterings were found to simulate the effect of cavity in inducing lasing of light. These are in addition to the tremendous advances made in photonic and phononic crystals studies, not to mention the more recent developments in wave functional materials that involve negative index of refraction. There is also the pertinent development of metal-insulator transition observed in 2D electron gas, which seems to violate the scaling theory of localization. In preparing the second edition, therefore, the author is faced with the decision of whether to include these new materials. After much consideration, the choice is to stick with the original purpose of this book, i.e., rather than trying to be inclusive and complete, it is better to clarify a few essential points in a simple and coherent manner. In the case of metal-insulator transition in 2D electron gas the inclusion of this material may represent a major departure from the original elements, and therefore should wait for the emergence of a more definitive picture.

However, the second edition does contain changes and new materials. A new chapter is added (Chap. 4) to expand on what was only hinted at in the first edition. In the new chapter some of the effective medium theory applications are detailed. The extension of the effective medium theory to the intermediate frequency regime where the wavelength is comparable to the scattering feature size—the spectral function approach—is also described in some detail, together with its successful applications. Other than the addition of this new chapter the changes involve mainly the addition of subheadings, aimed at facilitating an easier read of the text. In this regard it should be pointed out that in Chap. 5, the full derivation of the diffusive wave characteristics is preceded by an added section on a simplified model, with the purposes of pointing out the intrinsic connection between the Ward identity and the diffusive pole, and the demonstration of inevitability of the diffusive pole through simple mathematics.

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In preparing this new edition, thanks are owed to Louiza Law, Chun-Fung Chan, and Carman Chan for transcribing and correcting the text, Milky Tang for redrawing the figures, to Xiaodun Jing for carrying out new calculations that underlie Figs. 4.13, 4.14, and 4.16 and to the many students and colleagues for their comments and encouragement.

Clear Water Bay, Hong Kong August 18, 2005 Ping Sheng

# Preface to the First Edition

Wave behavior in disordered media is an old subject that has undergone a tremendous transformation in the past thirty years. Initiated by Anderson's seminal paper in predicting the phenomenon of wave localization, this transformation was achieved through a general proliferation of research activities in the physics of disorder and inhomogeneous materials. As a result, compared with thirty years ago there is now a new picture of wave characteristics in disordered media. The purposes of this volume are to delineate a coherent outline of this picture and to make the relevant technical materials accessible to a larger audience than those specialized in this area of research. Consistent with this intention, the presentation of both the physics and the mathematical framework is done with a minimum of assumed knowledge on the readers' part about the subject matter and the relevant theoretical techniques. In order not to interrupt the main line of presentation, however, many of the mathematical details are given at the end of each chapter, in the form of problems and solutions. The style of presentation is that of a physicist, whereby physical picture and reasoning are emphasized over mathematical rigor, and the mathematics is developed only to the extent necessary to support the derivation of results. Therefore, although Green function is used throughout the book, the discussion of Green function techniques is kept to a minimum. Also associated with the style of presentation is the subjective choice of the topics, which reflects this author's attempt to give not just a cartoon picture of the subject on the one hand and to avoid writing a review on the other. The readers can judge for themselves the degree of success to which this balance has been achieved.

The topics covered in this book can be categorized under four phenomena/concepts: effective medium, wave diffusion, wave localization, and intensity fluctuations and correlations. An overview and some explanation of the unconventional features of wave behavior in disordered media are given in Chapter 1. Basics of Green functions are introduced in Chapter 2. The central themes of Chapter 3 are wave scattering and the concept of effective medium, which reconciles the physical appearance of a homogeneous medium with the ubiquitous existence of wave scattering. The limitation of the effective

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medium concept and what happens beyond that are discussed in some detail in Chapter 4. This chapter is perhaps the most mathematical one of the book, since it contains the derivation of the wave diffusion constant and the demonstration of Ward identity. A new experimental measurement technique diffusive wave spectroscopy—is also described. Chapters 2, 3, and 4 are essentially self-contained and may form an independent unit in a course on waves.

The coherent backscattering effect and its consequences, including wave localization, are treated in Chapters 5, 6, and 8. A physical description of the coherent backscattering is given in Chapter 5. The dynamical implications of the effect are described in Chapter 6, which also contains an alternative, more exact treatment of wave localization in one-dimensional systems. The full localization implications of coherent backscattering are presented in Chapter 8.

The phenomenological scaling theory of localization is presented in Chapter 7, which is relatively independent from the other chapters. The scaling theory illustrates how the essential qualitative characteristics of localization can be understood on the basis of a few crucial elements and offers a physical way to view the progression from an extended state to a localized one through the variation of sample size.

The effects of multiple scattering and localization are illustrated in Chapters 9 and 10 in two different contexts. In Chapter 9 the effects are viewed in bulk electronic systems through the temperature and magneticfield variations of conductances. In Chapter 10 the effects are viewed through "mesoscopic" samples where the intensity fluctuations resulting from wave interference are manifest. These fluctuations are shown to contain novel correlations as well as to give rise to electrical conduction phenomena not seen in conventional bulk samples.

Writing this book was made possible by a six-month sabbatical leave from the Exxon Research and Engineering Company and a visiting membership at the Institute for Advanced Study, Princeton, New Jersey, for the 1992-1993 academic year. For the support of these institutions I owe my sincere gratitude. Special acknowledgments must be made of my colleagues Minyao Zhou and Zhao Qing Zhang, who provided me with invaluable criticisms and corrected my many mistakes during the writing process. Minyao Zhou, in particular, is responsible for the numerical calculations necessary for many of the figures, some of them quite involved. To them I give my heartfelt thanks. Appreciation is also extended to Ms. Marianne Kane, who typed most of the book and was tireless in going through the many revisions, and to Ms. Dot Fazekas for general secretarial support throughout the writing process.

Finally, I express my warm thanks to my wife, Deborah Wen, and to my daughters, Ellen and Ada, not only for their encouragement and spiritual support, but also for accepting the many hours I spent away from the family in writing.

Clear Water Bay, Hong Kong March, 1994 Ping Sheng

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Waves are everywhere around us. We rely on light and sound to sense our immediate surroundings. Radio waves and microwaves are indispensable means of communication. Water waves are responsible for the ocean's perpetually dynamic image. Quantum waves associated with electrons and atoms, while not directly visible, are important in maintaining the structure and stability of solids. With such a ubiquitous presence, wave phenomena naturally occupy a central position in our study of the physical world. Indeed, for waves in simple systems and ordered structures, an extensive literature already exists. However, for the more difficult problem of waves in disordered media, i.e., multiply scattered waves, a coherent (but by no means complete) understanding has only recently emerged, and from what is already known the picture is very different from that we normally associate with waves. In particular, the possibility that a wave can become localized in a random medium is especially intriguing because localization involves a change in the basic wave character. A localized wave has no spatial periodicity or possibility for transport and thus requires a new theoretical framework for its description and understanding. The purposes of this volume are to delineate the main features of this emerging picture of wave behavior in disordered media and to introduce the theoretical techniques for describing these features. Mesoscopic phenomena, which are the natural manifestations of wave scattering and interference effects, are also treated. A brief sketch below of the prominent random-wave characteristics serves as both an introduction to the subject and a map to what follows.

# 1.1 Relevant Length Scales

In an infinite, uniform medium, a (plane) wave may be characterized by a frequency and a direction of propagation. In contrast, a wave cannot propagate freely in a disordered medium because of the many scatterings it encounters. There are two types of scattering. One type, inelastic scattering, alters both

the wave frequency and the propagation direction. Another type, elastic scattering, preserves the frequency but alters the propagation direction. This book is concerned mainly with the effect of elastic scattering. Accordingly, the term "incoherent" is defined to mean waves having different propagation directions but the same frequency.

The consequence of multiple elastic scatterings may be described in accordance with the scale of observation. There are two obvious yardsticks in the problem. One is the average size R of the inhomogeneous scatterers. If the density of scatterers is not too low, then the interscatterer separation is also on the same order as R. Another yardstick is the wavelength  $\lambda$ . The ratio between R and  $\lambda$  is an important parameter in determining the average distance of coherent propagation between two scatterings. That distance, usually called the mean free path, is the relevant length scale for separating the different regimes of wave phenomena. When  $\lambda \gg R$ , the scattering is weak and the mean free path is large ( $\gg R$ ) for classical waves, i.e., electromagnetic and elastic waves. In addition, the scatterers and their placement geometry are beyond the resolution limit of the wave. Therefore, on the local scale of one to two mean free paths or less, a disordered medium appears as a homogeneous effective medium to the probing wave. In fact, since all matter is discrete at the atomic level, our everyday understanding of a uniform homogeneous medium reflects this effective medium concept. The same effective medium characterization holds for the quantum wave at the local scale. However, on the scale of many mean free paths, the effective medium can no longer be a valid description; even if locally the scattering is weak, over long distances the scattering effect accumulates and a wave can still be significantly randomized. When that happens, the result - diffusive transport - is similar to that of a classical particle undergoing random Brownian motion. The same result holds for the case of  $\lambda \leq R$ , except that the onset of diffusive transport occurs at a scale comparable to R, and there is no longer a valid effective medium because the local microstructure can now be clearly resolved by the wave.

# 1.2 Diffusive Transport

The fact that wave transport can be diffusive has a prominent example in our everyday experience of heat conduction. From statistical mechanics, it is well known that heat in electrically insulating solids is carried by randomly scattered (mostly short-wavelength) elastic waves, called phonons. Since heat conduction in solids is known to be governed by the diffusion equation even in the absence of inelastic scattering (Sheng and Zhou 1991), one immediately concludes that randomly scattered elastic waves transport diffusively.

Despite such clear-cut examples, however, diffusive transport for waves still raises some important questions concerning basic principles. A basic property of the wave equation is that of causality, which means that the speed at which a wave can carry information is always finite. For the diffusion equation, however, causality is not valid because the speed of propagation for a disturbance can be infinite. In addition, directly related to causality is the phase information of a wave, which is completely absent in diffusion. Just from this simple consideration it is already clear that the diffusion description of multiply scattered waves cannot be completely accurate since, whatever its appearance, the multiply scattered wave is still a solution of the wave equation (albeit with random coefficients) and therefore must satisfy the basic properties of its solution. Thus for wave diffusion there should be some deviation from classical diffusion where the wave character asserts itself. In the past decades, this expectation was borne out by the experimental demonstration (Tsang and Ishimaru 1984; van Albada and Lagendijk 1985; Wolf and Maret 1985) of the *coherent backscattering effect*, or the weak-localization effect as it is sometimes called, which represents not only a deviation from classical diffusion, but also the precursor to wave localization.

# 1.3 Coherent Backscattering and the Approach to Localization

The coherent backscattering effect means what the name implies: After a wave is multiply scattered many times, its phase coherence is preserved in the direction opposite to its incident direction (backscattering direction), but not in other directions. The reason for this behavior is fully explained in a later chapter, but we may appreciate some of its consequences here. By preserving the coherence in the backscattering direction, the probability of backscattering is enhanced through constructive interference. This leads to a decrease in the diffusion constant from its classical value, because whatever the direction of the wave, the increased backscattering tends to drag it back as if the wave medium were more "viscous" than it should be classically.

From elementary kinetic theory, the diffusion constant may be expressed as D = (1/3)vl, where v is the velocity of the diffusing "particle" and l is its elastic collision mean free path. When D decreases, either l or v, or both, may be the cause. For wave diffusion, the mean free path may be measured as the distance a plane wave can penetrate into a scattering medium before it loses its phase front. The velocity v, on the other hand, is a problem because the usual wave speed, whether the phase velocity or the group velocity, is defined for wave states that have well-defined wave vectors. Definitions of phase velocity requires a constant-phase surface perpendicular to the wave vector, and the definition of the group velocity – the wave vector derivative of frequency requires both the wave vector and the dispersion relation between frequency and wave vector to be well defined. However, in a strongly scattering medium a wave vector cannot possibly describe a multiply scattered wave state because such a state has neither a unique direction of propagation nor a unique spatial periodicity for defining a wavelength. What, then, is the v in the wave diffusion constant?

The v relevant to the wave diffusion is the average speed at which the wave *energy* is transported locally. In the presence of strong scattering, especially resonant scattering, this transport velocity can be smaller than the free-space phase velocity for classical waves (van Albada et al. 1991), but it is always identical to the group velocity for the quantum wave associated with free electrons. This striking difference between the classical and the quantum waves has its origin in the different dispersion relations for the two kinds of waves and the manner in which the randomness enters the two types of wave equations. Therefore, through increasingly strong scattering the mean free path must decrease for the quantum case but both the transport speed and the mean free path can decrease for the classical waves.

The downward renormalization of the diffusion constant and the coherent backscattering effect itself have several important features. First, the decrease in the diffusion constant is proportional to the scattering strength. If the scattering is weak, the coherent backscattering effect is correspondingly weak so that it can be ignored in general. But if the backscattering is strong enough, the decrease in the diffusion constant can make it vanish, thus leading to wave *localization*.

Second, the coherent backscattering effect is fully operative only when the system is time-reversal invariant, meaning that macroscopically there should be no preferred direction of time established, for example, by a uniform average velocity of the scatterers or by the presence of a magnetic field. In the presence of effects that break time-reversal invariance, the correction to the wave diffusion constant is diminished.

Third, the coherent backscattering is fully effective only when all scatterings are elastic. In the presence of inelastic scattering, which is inevitable in real materials having various dissipation mechanisms, the coherent backscattering effect is again diminished. However, the manner in which this occurs is directly coupled to the next feature; that is, the magnitude of the coherent backscattering effect, manifested in the amount of (negative) correction to the diffusion constant, is a monotonically increasing function of the physical sample size. Where the sample size is infinite, the de facto "sample size" is set by the inelastic scattering rate, because over a distance where there are several inelastic scatterings, the coherent backscattering can no longer be operative, and that distance essentially becomes the limiting "sample size." Since electron inelastic scattering is generally temperature dependent, the combination of the last two features offers a means by which to observe the coherent backscattering effect (indirectly) in electronic systems through the temperature dependence of conductivity in disordered materials.

# 1.4 Sample Size Dependence

From the point of view of classical physics, the sample size dependence of the coherent backscattering effect is revolutionary because it makes the



Fig. 1.1. A schematic of the variation of the renormalized diffusion constant with sample size, i.e., the scale of observation. The sample size must be larger than the mean free path before diffusion can be observed. The three cases of weak scattering, strong scattering, and localization are shown.  $D^{(B)}$  denotes the classical Boltzmann value of the diffusion constant

renormalized diffusion constant no longer intensive, as would generally be expected, since the diffusion constant belongs to the same class of intensive quantities as density, electrical conductivity, and temperature. Figure 1.1 shows the renormalized diffusion constant schematically as a function of sample size for three cases.

In the weak scattering limit, the diffusion constant is independent of the sample size, as expected classically. When the scattering is strong the diffusion constant is renormalized downward as a function of the sample size, with an asymptotic value that can be significantly less than its classical value (which is nonetheless still observable at small sample size). When the asymptotic value of the renormalized diffusion constant vanishes, then by definition a localized state is created. Therefore, a pulse injected into a strongly scattering medium would evolve initially as in a uniform medium, then quickly make a transition into diffusive transport, accompanied by a gradual slowdown of diffusion over time. Localization occurs when the overall diffusion is stopped.

Two comments can be made concerning the sample size dependence of the coherent backscattering effect. First, on the most elementary level, the backscattering aspect of the effect is responsible for its sample size dependence, because larger samples are more opaque than smaller samples. Thus one can expect more backscattering and less transmission for the larger samples. The coherence part of the effect then enhances the backscattering intensity

from what is expected from ordinary diffusion. Through refinement, this simple viewpoint can be made quantitative, as will be seen in Chap. 8. Second, the size dependence is necessary if one looks ahead to localization. When a wave is trapped by randomness, there appears a new length scale – the localization length – which naturally introduces sample size dependence into the physical transport property of the system. Let us expand on this point below by using the quantum waves associated with the electronic system as an example.

If one measures the electrical conductivity of a sample in which the electrons are localized – a disordered insulator – the usual measurement at finite temperatures is expected to yield only the conductivity of the electrons thermally activated from their localized states to some higher-energy mobile states able to carry electrical current. Now imagine a thought experiment where the temperature is lowered to absolute zero so that all electrons are in their localized states. The electrical conductivity of a bulk insulating sample at zero degree is usually regarded as zero – or too small to be measured. Nevertheless, if the sample dimension is small enough so that it becomes comparable to the localization length – a possibility which is increasingly becoming a technological reality nowadays - then even an insulator can conduct some electricity. Conduction can occur because localization prevents the wave, in this case the quantum wave associated with an electron, only from moving outside the spatial domain defined by the localization length, but the electron can still be mobile inside the domain of localization. If the edge of the localized domain is not abrupt but has an exponential tail, then the ability of a localized electron to conduct electricity would decrease exponentially as a function of the sample size when its linear dimension increases beyond the localization length. The result is the sample size dependence of the transport characteristics expected from localized states. If one did not know about the coherent backscattering effect but wanted to invent a mechanism for localization, such a mechanism would need to incorporate some kind of sample size dependence as necessitated by its desired result. From this perspective, the sample size dependence is an essential and necessary attribute for a localization mechanism.

# 1.5 Localization and Scaling

Since localization is a major theme of this book, a brief digression on the development of the localization concept would be helpful. In the early days of solid-state physics, the recognition that electronic states in a periodic lattice form energy bands was a breakthrough that clarified a basic question about why some materials are electrically conducting and some are insulating. In the simplest version of the band picture, electrical conductors have a half-filled energy band whereas an insulator has filled bands. Mott took the conductor picture a step further and proposed that if the lattice constant of a half-filled band conductor can be continuously increased, then at some point

the conductor becomes an insulator. The rationale behind this hypothesis is that when the atoms are separated far enough, they will behave as individual neutral atoms instead of as a metal in which the conduction electrons can pass freely from one atomic site to the next. The basic physics of the socalled Mott transition is the Coulomb interaction between the electrons (Mott 1949, 1974). It does not involve disorder. In contrast, Anderson proposed in 1958 that electronic diffusion can vanish in a sufficiently random potential, in the absence of any electron–electron interaction (Anderson 1958). This proposal, together with Mott's previous works, formed the theoretical basis for the study of the metal-insulator transition in doped semiconductors. However, it was not until the late 1970s and the early 1980s that the Anderson localization was linked to the coherent backscattering effect and explained on that basis. At about the same time, the study of Anderson localization was extended to classical waves, which offer an advantage over the disordered electronic systems where the Mott mechanism and the Anderson mechanism are inseparable: In classical waves, the Anderson localization may be studied alone, without the additional complication of wave-wave interaction. In this volume, the term "localization" denotes the phenomenon only in the sense of the Anderson mechanism.

Although the coherent backscattering effect was important to the understanding of wave multiple scattering phenomena and localization, an overview of the localization phenomenon was actually first achieved through the different perspective of a phenomenological theory, the scaling theory of localization (Abrahams et al. 1979). The scaling theory is a scheme for interpolating between an extended wave state and a localized one. The term "scaling" is popular in physics nowadays. In the present context, it has the following connotations. First, scaling implies that the conclusions of the theory are independent of the many details of the physical model. For example, it is often the case that electron multiple scattering and localization are studied in the context of a lattice model of the solid atomic lattice. A scaling theory of localization would imply that the conclusions of the theory are independent of the type of lattice, be it simple cubic, body-centered cubic, or whatever. The conclusions are also independent of the type of random scatterings, the statistics of the randomness, etc. Therefore, scaling implies broad applicability: Because the theory depends on few essential physical quantities, one hopes to obtain a better overall picture of the phenomenon without being encumbered by details. Herein lies the attraction of a scaling theory.

Second, scaling here means literally changing the scale, or physical size, of the sample under consideration. The scaling theory of localization considers how a quantity, defined as the dimensionless conductance  $\gamma$ , varies under a sample size change. For an electronic system,  $\gamma$  is simply the ordinary conductance  $\Gamma$  divided by the quantum unit of conductance,  $e^2/h$ , where e denotes the electronic charge and h is Planck's constant. For classical waves,  $\gamma$  may be expressed alternatively as the ratio of two energy scales, which are defined and motivated in Chap. 8. How the dimensionless conductance  $\gamma$  varies with

sample size is very simple for electrical conductors (samples in which the wave states are extended). How  $\gamma$  varies with sample size for disordered insulators, in which the wave states are all localized, can also be inferred from the earlier discussion on the sample size dependence of the coherent backscattering effect. To these known facts the scaling theory adds the following assumption: The rule governing the variation of  $\gamma$  with sample size can depend on only one parameter, which is the value of  $\gamma$  itself. With this seemingly innocent proposition, the scaling theory reaches some startling conclusions, the most striking of which is the dependence of the localization phenomenon on the spatial dimensionality of a sample. In particular, the scaling theory tells us that regardless of how weak the randomness, all waves are localized in 1D or 2D samples of infinite extent. For 3D samples, the scaling theory predicts the possibility of coexistence for extended and localized wave states, where the extended states can exist in one or several frequency regimes and the localized states can exist in the other. The frequency that separates a localized regime from a neighboring extended regime is called the "mobility edge." In the jargon of the physics community, spatial dimension two is called the "marginal dimension" for localization. In order to understand this prediction, let us first clarify the meaning of spatial dimensionality as applied to physical samples, as well as the conditions under which the predictions apply.

# 1.6 Spatial Dimensionality in Localization and Diffusion

All physical objects must have finite cross-sections or thicknesses. Therefore, there cannot be true 2D or 1D samples of vanishing thickness or cross-section. Spatial dimensionality means that if wave propagation and scattering are allowed only in two (backward and forward) directions defined by a line, the sample is 1D; if propagation and scattering are allowed only in directions defined by a surface, the sample is 2D; and if they are allowed in all 3D space, the sample is 3D. The restriction on the direction(s) of wave propagation and scattering can be achieved by making the thickness, or cross-sectional dimension of a sample smaller than, or comparable to, the wavelength. For example, if a wave is confined inside a sample whose thickness is smaller than the wavelength (but larger than half the wavelength), then the excitation in the thickness direction must be a standing wave. In addition, all other standing wave states are higher in frequency, with frequency increments large enough to make them inaccessible (e.g., through thermal excitation). Therefore, all scattering and propagation are confined to the planar directions, and the sample may be described as 2D. Another possibility of observing 2D waves is found in interfacial excitations, where the wave amplitude decays exponentially away from the interface and the wave can propagate only along the interface, as the name implies. Similar reasoning applies to the description of a 1D sample.

On the basis of the scaling theory predictions, should all thin and wirelike objects be good electrical and thermal insulators? The answer depends on the conditions under which the measurements are made. The conclusions of the scaling theory are meant to apply only to samples at absolute zero temperature, so that finite-temperature effects, such as the activation of charge carriers and the inelastic scattering (that essentially limits the sample size as described before), are all absent. Therefore, the predicted 1D and 2D localization effects should not be ordinary, everyday experience. However, they should apply under laboratory conditions where the temperature of a sample is lowered to close to absolute zero and the sample dimensions are controlled to within the tolerance of being qualified as 1D or 2D.

A particularly intriguing prediction of the scaling theory is the inevitability of wave localization in 1D and 2D samples even at the weak scattering limit. To better understand this limiting case, let us first construct and analyze an "apparent" dilemma. Consider a 1D sample of finite extent L. The heterogeneties in the sample are characterized by a scale R which is assumed to be smaller than the wavelength  $\lambda$ . To be more specific, the wave is a classical pulse with a gentle envelope whose width is many wavelengths so that the additional frequency components are negligible. The question is: What happens when  $L \to \infty$  and  $\lambda \to \infty$ ? If  $\lambda \to \infty$  first, the previous discussion on effective medium shows that the 1D random medium should always appear homogeneous to the probing wave, with no net scattering in the static limit of  $\lambda/R \to \infty$ . This would remain true for every finite L as  $L \to \infty$ . Therefore, the end result of  $\lambda, L \to \infty$  should be an infinite (homogeneous) effective medium where there is no localization. Now consider the reverse order of taking the limits. By letting  $L \to \infty$  first, one can immediately invoke the prediction of the scaling theory – that all waves are localized in 1D samples of infinite extent – and this conclusion remains true for every  $\lambda$  as  $\lambda \to \infty$ . If a unique physical state is assumed for  $\lambda, L \to \infty$ , then the two contradicting conclusions yield a dilemma. The above description applies to a 1D sample, but similar reasoning leads to the same dilemma in 2D samples.

What can be the resolution of this dilemma? The answer is that both conclusions are correct, because even as  $L, \lambda \to \infty$ , the ratio  $L/\lambda$  still needs to be adjusted. There can indeed be different physical states depending on whether that ratio approaches 0 or  $\infty$ , which are the two possibilities probed by taking the limits in two different orders. But what does this imply physically for wave localization? If  $L \to \infty$  first, then as  $\lambda \to \infty$ , locally (inside the wave packet) the effective medium should be an increasingly good approximation, because scattering by classical waves diminishes as  $\lambda/R \to \infty$ . As the wave packet travels through the medium, the scattering is small at every instant of time. Therefore, if localization were to occur as predicted, the small scatterings at successive times *must accumulate* so that the net effect is large enough to localize. In other words, even if  $R/\lambda = \varepsilon$  is small, yet over large travel distances, measured in terms of  $1/\varepsilon$  (or even larger in 2D), there is still an order one effect, i.e., localization. If the travel distance is limited by the sample size L, then the localization limit can never be reached and one always obtains the effective medium limit instead. Therefore, the following physical picture emerges from

an analysis of this apparent dilemma: In the weak scattering limit, a localized wave can exhibit propagating behavior *locally*; through increasing sample size and the effect of accumulated scatterings, the wave transport character is altered progressively from propagating to diffusion to localized. Of course, one recognizes the coherent backscattering effect in this scattering accumulation process, especially its sample size dependence. But why should this accumulation process be especially effective only in 1D and 2D? What is so magical about the spatial dimensionality two? Some insight into this question may be obtained through the special character of diffusion and its interaction with the coherent backscattering effect.

The usual way to describe diffusion is through the net distance r traveled by a random walker at the end of a time period t. The diffusion relation is described by  $r^2 \propto t$ , independent of the spatial dimensionality where the random walker executes its "walks." That is, in time t the random walker covers an "area" proportional to  $r^2$ . If the random walker is limited to moving on a line or a flat surface, the trace of its path would appear dense, but if the random walker is a flying particle that can freely traverse the 3D space, then the trace of its path would appear to be flimsy, because an area does not go a long way toward covering a volume. While this description of diffusion may be somewhat simplistic, it can be made rigorous by adding qualifying conditions. Let us consider the path traversed by a random walker from t = 0onward. As  $t \to \infty$  in d = 1, 2 the random walker will visit the infinitesimal neighborhood of any given point with probability one, i.e., with certainty. In d = 3, however, the probability is infinitesimal that it would visit any given infinitesimal neighborhood. One way to appreciate that difference is by calculating the probability, at a time  $t_m > 0$  onward, that a random walker will return to the neighborhood of the origin, the point where the walker started its motion at t = 0. According to the solution of the diffusion equation, the probability density for the walker at distance r from the origin at time t is given by  $P(r,t) = (4\pi Dt)^{-d/2} \exp(-r^2/4\pi Dt)$ , where d denotes the spatial dimensionality of the random walk and D is the diffusion constant. The desired probability is therefore given by

$$\lim_{T \to \infty} \int_{t_m}^T P(0,t) \mathrm{d}t = \lim_{T \to \infty} \int_{t_m}^T \frac{\mathrm{d}t}{(4\pi Dt)^{d/2}}$$

For d = 1, 2 the integral diverges as  $T \to \infty$ , independent of  $t_m$ , implying that the random walker will certainly return to the neighborhood of the origin. This is intuitively plausible from the fact that the path of a random walker covers an "area." For d = 3, on the other hand, the integral is proportional to  $1/\sqrt{t_m}$ , so that as  $t_m$  increase, the probability decreases toward zero. Thus, in terms of this probability, there is a qualitative difference between diffusion in one and two dimensions and diffusion in three dimensions. If the probability of returning to the origin is heuristically equated with backscattering, then this special property of diffusion translates into more effective coherent backscattering in one and two dimensions, leading to localization. From this viewpoint, the marginal dimension of localization is a direct result of the exponent two in the diffusion relation  $r^2 \propto t$ .

Although the scaling theory gives a global view of what can happen in different spatial dimensions, the details of this explanation need to be filled in with results of calculations based on the mechanisms of diffusion plus coherent backscattering. This proves to be possible, because the correction to the diffusion constant introduced by the coherent backscattering is consistent with the scaling hypothesis. As a result, formulas for the localization length and its frequency dependence, plus explicit forms for the scaling functions, can all be derived in spatial dimensions one and two. In spatial dimension three, however, our knowledge of the subject is still incomplete. It is known that if randomness is introduced into a periodic structure or if the system is discrete, such as an electron in a disordered lattice, then the localized states and mobility edge(s) can exist. This is due to the fact that in the absence of randomness, periodic structures yield frequency bands, and near the band edges the Bragg scattering gives rise to standing waves, i.e., waves with zero group velocity. Such wave states are, in a sense, waiting to be localized as randomness is introduced. For a continuous random system without any longrange periodic correlation, however, quantum wave is known to localize at low energies, whereas classical wave localization is possible only if the scattering is strong enough. One possibility is resonant scattering, e.g., Mie resonance for electromagnetic wave scattering from small particles, which can give a much higher scattering cross-section than usual. But resonant scattering has the disadvantage of associated large absorption, which can mask the localization effect. Hence the most likely systems for classical wave localization are those with large index of refraction contrast between the scatterers and the matrix material in which they are dispersed (Sheng 1986). Indeed, light localization was observed in GaAs powders with an index of refraction value  $\sim 3.5$ . (Wiersma et al. 1997)

# 1.7 Mesoscopic Phenomena

Wave multiple scattering and localization can lead to various physical manifestations. For electronic systems, the direct proportionality relation between the conductivity and the diffusion constant is known as the "Einstein relation." Therefore, the observation of electronic conductivity is equivalent to observing electron diffusion. For classical phonons, heat conduction has already been mentioned as a manifestation of phonon diffusion. In dirty conductors, indirect observation of the coherent backscattering effect is contained in the measurements of anomalous temperature dependence of resistivity and the anomalous magnetoresistance, both in conducting films. However, these indirect measurements suffer from the ambiguity that a competing effect – that

of electron – electron interaction–can yield similar results, so there is no way to attribute the measured behavior entirely to the coherent backscattering effect. A more direct observation of the consequences of wave multiple scattering is via measurements on mesoscopic samples. Here "mesoscopic" denotes a sample size regime intermediate between the molecular and the bulk. More precisely, it means a sample size whose linear dimension is smaller than a "dephasing length,"  $\sqrt{D\tau_{\rm in}}$ , where  $\tau_{\rm in}$  is the inelastic collision time for the wave. In a mesoscopic sample, all collisions are elastic, so the effects of coherent backscattering and interference can be manifest and not be averaged out. For classical waves, the inelastic scattering rate is generally low and temperature insensitive. Wave-wave interaction is also negligible, in contrast to the interaction between electrons. Thus a mesoscopic sample in this case is a bulk sample, and the experiment can be done at room temperature. In fact, the cleanest quantitative demonstration of the coherent backscattering effect came from light scattering from bulk disordered systems. For electronic systems, on the other hand, the inelastic scattering length is generally small and inversely temperature dependent. Therefore, the observation of mesoscopic phenomena in electronic systems generally requires small samples whose linear dimensions are on the order of micrometers or smaller. In addition, the experiments must be performed at low temperatures to ensure that the inelastic scattering rate is sufficiently low. However, with today's microfabrication and cryogenic technologies these requirements are not difficult barriers, and the relentless push for miniaturization in the electronic industry means that the mesoscopic phenomena could very well serve as the basis for tomorrow's quantum devices.

Mesoscopic phenomena are manifest in the measurements of conventional quantities in mesoscopic samples. Electrical conductivity is an example. The definition of mesoscopic conductance is problematic at first sight because electrical conductance is always associated with dissipation, whereas in mesoscopic samples the absence of inelastic scattering means that there is no dissipation. Thus, there is a need to define what one means by conductance in a mesoscopic sample. By analogy with the resistance of a tunnel junction, Landauer has proposed that the conductance of a mesoscopic sample should be proportional to the wave transmission probability (Landauer 1957). In that case, dissipation occurs not inside the sample but in the leads connected to the sample, and through consideration of equilibrium with the leads an expression for the mesoscopic conductance can be derived.

Conductance of mesoscopic samples has many unconventional characteristics. Not only does it vary with sample size (due to the coherent backscattering effect), it can also fluctuate wildly from sample to sample, even if the samples were fabricated identically in the same batch. The conductivity can also show large fluctuations upon the application of a varying magnetic field. Moreover, as the sample size increases, these fluctuations do not decrease (as long as inelastic scattering is absent); i.e., the fluctuations are not averaged out as one would expect from additive noise. All these effects are basically due to the fact that even though the waves are multiply scattered, they can still exhibit phase interference and retain some long-range memory.

# 1.8 Localization vs. Confinement

The term localization has often been used in the literature to denote nonpropagating wave states which are not localized in the Anderson sense. In particular, wave confinement, e.g., standing waves formed inside a cavity or by totally reflecting interfaces, should be distinguished from the Anderson wave localization that has just been described. Wave confinement may involve walls made of material that has no wave state at the relevant frequency. Hence waves can only totally reflect from such interfaces, penetrating the interface only by an exponentially decaying evanescent tail. Spectral bandgaps in photonic or phononic crystals or crystalline electronic systems are typical examples which can form effective wave confining cavities or waveguides. Viewed in more general terms, the crucial difference in spectral bandgap confinement and Anderson localization is that a bandgap denotes a frequency regime which is empty of wave states, whereas a localized wave is a nonpropagating wave state. The two mechanisms can interact, nevertheless. A most interesting case is the disordered spectral bandgap systems in which the crystalline periodicity is perturbed by disorder as mentioned before. The sharp edge of the crystalline bandgaps would then be smeared out to form a transition region. In that transition regime there would be spatial regions that are deficit of wave states, so that instead of total confinement, the waves would be restricted in their propagation directions. That is, instead of propagating in straight lines, the waves would be traveling in a labyrinth. Another way of saying the same thing is that the total scattering is increased, leading to the enhancement of localization effect. That is why the bandedge states are easily localized.

# 1.9 Topics not Covered

The discussion thus far is a brief and qualitative guide to the contents of this volume. At this point it is perhaps also important to point out the relevant subjects that are left out. First is wave dissipation. In this volume, dissipation is treated only as a constraint, or limitation, on the effects due to elastic scattering. No attempt is made to examine either the mechanism of dissipation or the combined effect of both elastic scattering and dissipation. The second neglected topic is nonlinearity, which can couple waves of different frequencies. Since the magnitude of nonlinearity depends on the amplitude of a wave, its neglect means that we will be concerned only with small-amplitude waves, and waves of different frequencies will be treated as independent. The third neglected topic is wave–wave interaction. A well-known example in this regard is

the electron–electron interaction through their electrical charges. Here, overlooking wave–wave interaction implies that wave scattering is treated as a "single-body" problem, in contrast to the "many-body" problem where the presence of more than one electron changes the nature of the system. For electrons, interaction effects are important at low densities. At high densities, the Coulomb interaction effect is weakened by screening, resulting in an electron dressed in a cloud of screening charges that can be treated as an independent "quasiparticle." The neglect of electron–electron interaction here implies that whenever the term electron is used in this volume, it denotes a quasi-particle in the high-density limit. For classical waves, on the other hand, the interaction between electromagnetic waves is weak, and the same is true for elastic waves of small amplitudes. Therefore, their neglect is well justified.

The three topics not covered – dissipation, nonlinearity, and interaction – are interesting and rich subjects in themselves. They have been omitted solely to attain the simplicity and coherence made possible by focusing on the subjectively chosen main line of exposition. The author, rather than trying to be inclusive and complete, intends this volume to clarify a few essential points, leaving the readers to further explore this challenging field.

# Quantum and Classical Waves

# 2.1 Preliminaries

In a uniform medium, a wave at a fixed frequency  $\omega$  has the property that a snapshot of its amplitude would display spatial periodicity. The period, or wavelength  $\lambda$ , is related to  $\omega$  by the so-called dispersion relation, which is a basic property of the wave and the medium in which it propagates. For a quantum particle/wave, the dispersion relation reflects the definition of energy. That is, if by  $E = \hbar \omega$  one denotes the total energy of the quantum particle/wave, where  $\hbar$  is Planck's constant, and by  $p = \hbar k$  the momentum of the particle, where  $k = 2\pi/\lambda$  is the magnitude of the wave vector **k** that points at the direction of wave/particle propagation, then for a particle of mass m, E is the sum of kinetic and potential energies:

$$E = \hbar\omega = \frac{p^2}{2m} + V = \frac{\hbar^2 k^2}{2m} + V.$$
 (2.1)

Here V denotes the potential energy, taken to be a constant for the uniform medium. Equation (2.1) represents the relation between  $\omega$  and k for a quantum particle/wave. By representing the temporal oscillation of the wave by  $\exp(-i\omega t)$  and the spatial periodicity by  $\exp(i\mathbf{k}\cdot\mathbf{r})$ , E can be obtained from the wave amplitude function  $\phi = \exp[-i(\omega t - \mathbf{k}\cdot\mathbf{r})]$  by the operation of  $i\hbar\partial/\partial t$  on  $\phi$ , i.e.,

$$\mathrm{i}\hbar\frac{\partial\phi}{\partial t} = E\phi. \tag{2.2}$$

Also,  $p^2/2m$  can be obtained by the operation of  $-(\hbar^2/2m)\nabla^2$  on  $\phi$ , where  $\nabla^2 = \partial^2/\partial x^2 + \partial^2/\partial y^2 + \partial^2/\partial z^2$  is the Laplacian operator, i.e.,

$$-\frac{\hbar^2}{2m}\nabla^2\phi = \frac{\hbar^2k^2}{2m}\phi.$$
 (2.3)

Another way of stating (2.3) is that  $\exp(i\mathbf{k} \cdot \mathbf{r})$  is the eigenfunction of the Laplacian operator with the eigenvalue  $-k^2$ . From (2.1)–(2.3), one thus infers the Schrödinger equation,

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$$i\hbar\frac{\partial\phi}{\partial t} + \frac{\hbar^2}{2m}\nabla^2\phi - V\phi = 0, \qquad (2.4)$$

from the dispersion relation.

In contrast to the quantum wave, the classical electromagnetic and elastic waves have the dispersion relation

$$\omega = v|\mathbf{k}|,\tag{2.5}$$

where v is the wave speed. One way to get rid of the absolute value sign is to rewrite (2.5) as

$$\omega^2 = v^2 k^2. \tag{2.6}$$

By using the same time and space derivative operations on the wave amplitude function  $\exp[-i(\omega t - \mathbf{k} \cdot \mathbf{r})]$ , (2.6) implies the following scalar wave equation:

$$\frac{\partial^2 \phi}{\partial t^2} - v^2 \nabla^2 \phi = 0. \tag{2.7}$$

The scalar wave equation differs from the electromagnetic and elastic wave equations in that the true classical waves have vector character, i.e., different polarizations. For the electromagnetic wave, there are two polarizations transverse to  $\mathbf{k}$ , and for the elastic wave there is in addition a longitudinal polarization parallel to k. In a uniform, isotropic medium, classical waves of different polarizations are each described by a scalar wave equation, with a velocity v which can be different for the elastic longitudinal wave and the elastic transverse (shear) waves because they rely on different material properties for their propagation. Since different polarizations are decoupled from each other in a uniform, isotropic medium, the scalar wave equation is thus accurate for each polarization component of the classical waves. This is not true in a disordered medium. Because of the coupling between the different polarization components at a scattering interface (between materials of different wave properties), classical waves with different polarizations can be interconverted into each other. As a result, the scalar wave equation is no longer accurate in a disordered medium. However, in spite of such complexities arising from the polarization consideration, many of the basic classical wave characteristics are still qualitatively captured by the scalar wave equation. In particular, for the aspects of wave transport and localization considered in this volume, the scalar wave equation is adequate in illustrating the basic physics involved. For simplicity, the scalar wave would serve as our model for the description of classical waves.

## 2.1.1 Quantum and Classical Wave Equations

An important difference between the Schrödinger (quantum) wave equation and the scalar (classical) wave equation is the order of the time derivative. Whereas it is first order in the quantum case, it is second order in the classical case. This difference arises from the different dispersion relations for the two types of waves and implies that, whereas for the classical waves the speed of propagation is a constant of the wave medium, for a free quantum particle (V = 0) the wave speed, defined as dE/dp for the group velocity, is dependent on its momentum (just like a classical particle). Quantum waves of different frequencies thus travel with different speeds. This important difference between quantum and classical waves is reflected in many aspects of their behavior in disordered media.

For waves of a given energy or frequency, the quantum and classical wave equations can be put in the same form

$$\nabla^2 \phi + \kappa^2 \phi = 0, \tag{2.8}$$

where

$$\kappa^2 = \frac{2m(E-V)}{\hbar^2} \tag{2.9}$$

for the quantum case, and

$$\kappa^2 = \frac{\omega^2}{v^2} \tag{2.10}$$

for the classical case. Here a striking difference between the quantum wave and the classical waves is already apparent, because whereas E - V can be negative in (2.9),  $\omega^2/v^2$  is always positive. Therefore, classical waves have a quantum correspondence only when E > V.

In general,  $\kappa^2$  is a function of **r**, and it can be written as

$$\kappa^2 = \kappa_0^2 - \sigma(\mathbf{r}), \qquad (2.11)$$

where  $\kappa_0^2$  represents the uniform background and  $\sigma(\mathbf{r})$  represents the deviation from it. In the quantum case, if V is a constant, it can be absorbed into E as a redefinition of that quantity. Therefore, it is always possible to write

$$\kappa_0^2 = \frac{2m\omega}{\hbar} \tag{2.12}$$

and

$$\sigma(\mathbf{r}) = \frac{2mV(\mathbf{r})}{\hbar^2} \tag{2.13}$$

for the quantum wave. In the classical case, if  $v_0$  denotes the wave speed in the uniform medium, such as the speed of light in vacuum, and  $v = v_0/\sqrt{\varepsilon}$  is the speed inside a scattering object, then  $\varepsilon$  is generally known as the dielectric constant, where  $\varepsilon = n^2$ , *n* being the index of refraction. By expressing  $v^2 = v_0^2/\varepsilon(\mathbf{r})$ , where  $\varepsilon = 1$  for the background, then

$$\kappa_0^2 = \frac{\omega^2}{v_0^2},$$
 (2.14)

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and

$$\sigma(\mathbf{r}) = [1 - \varepsilon(\mathbf{r})] \frac{\omega^2}{v_0^2}.$$
(2.15)

The quantity  $\sigma(\mathbf{r})$  is seen to depend on  $\omega$  for classical waves but is independent of  $\omega$  for the Schrödinger wave. More precisely, the deviation from a uniform medium occurs *additively* with frequency for the quantum case, but *multiplicatively* with frequency for the classical case. One immediate result is that as  $\omega \to 0$ , classical waves always scatter weakly because  $\sigma(\mathbf{r})$  vanishes in that limit. The same is not true for the quantum case.

In Sect. 2.2 the behavior of quantum and classical waves in a uniform medium is examined under different spatial dimensionalities. The meaning of spatial dimensionality for a physical sample has been discussed in the last chapter. Whereas real 2D and 1D samples must have physical thicknesses and finite cross-sections, in the subsequent theoretical development 2D and 1D are treated as meaning true mathematical surfaces and lines. This is, of course, an approximation to reality, but one which is accurate for the essential physics of the lower dimensionalities.

# 2.2 Green Functions for Waves in a Uniform Medium

Green function is one of the most useful tools in physics because of the physical information it can provide about the system described by the differential equation. Simply put, Green function in the present context is just the solution to the wave equation with a specific initial source and a boundary condition at infinity. One might ask why that is more useful than just any solution  $\phi(t, \mathbf{r})$ to the homogeneous wave equation. The answer could be illustrated by the following analogy.

Suppose there are N mechanical players in a game whose rules, while well defined, are not known a priori to an outside observer. The game involves the exchange of some currency between the players (such as in Monopoly, except here the players have no free will), which we just call "money". The purpose of the observer is to figure out the game rules and predict what is going to happen. The observer can record the money each player has at every instant and how it varies with time. The recorded quantity, M(t, i), where i is the player index, may be compared to the wave function  $\phi(t, \mathbf{r})$ , since both give successive snapshots of the state of affairs. However, an intelligent observer might want to ask the question: What is the interrelationship between the amounts of money of any two players, if any? This important information cannot be obtained by simply looking at each player individually. Instead, one has to look at the *correlation* between pairs of players as a function of time separation, i.e., the correlation between M(t, A) and  $M(t + \Delta t, B)$ . If the observer has the power to do a controlled experiment on the players, this correlation may be obtained by (1) setting M = 0 for all players, (2) giving an amount  $M_0$  to A at t = 0, and (3) watching how much of  $M_0$  is transferred

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to B at a later time t. This is precisely the sort of information that a Green function would provide. By making A and B the same player, the Green function can also monitor the information contained in the wave function. Therefore it is more general. In the case of the wave equation, the precise knowledge of the Green function is sufficient to figure out all the rules of the system; i.e., it is equivalent to an exact solution of the problem. In the absence of exact solutions, which are often difficult beyond simple or periodic systems, Green function also provides a solution approach to the wave equation that is amenable to systematic perturbation analysis. Moreover, the meaning of these analyses can be interpreted physically. Therein lies the attractiveness of the Green function approach.

# 2.2.1 Green Function for the Uniform Medium

The specific initial condition for the Green function is that at t = 0, the system is excited by a pulse localized at  $\mathbf{r} = \mathbf{r}'$ ; i.e., the right-hand sides of (2.4) and (2.7) are replaced by  $\delta(t)\delta(\mathbf{r} - \mathbf{r}')$ , where  $\mathbf{r}'$  denotes the source position. The reason for such a source is that it can excite the natural resonances (or eigenfunctions) of all temporal and spatial frequency in the system (since the source contains all temporal and spatial frequencies), so that the subsequent development may contain information about them. A useful approach for analyzing the Green function is through its frequency components. The left-hand side of the wave equation for the frequency component  $\omega$  is given by (2.8), and the same frequency component on the right-hand side is simply  $\delta(\mathbf{r} - \mathbf{r}')$  because  $\delta(t)$  gives 1 as the amplitude for every frequency component. Equating both sides yields

$$(\nabla^2 + \kappa^2)G(\omega, \mathbf{r}, \mathbf{r}') = \delta(\mathbf{r} - \mathbf{r}'), \qquad (2.16)$$

where  $G(\omega, \mathbf{r}, \mathbf{r}')$  denotes the Green function in the frequency and spatial domains and  $\mathbf{r}'$  is used to mark the source point. The source point position is in general an important parameter for the Green function if the system is inhomogeneous, since the response to a point source can depend on its position. In addition to (2.16), a unique determination of  $G(\omega, \mathbf{r}, \mathbf{r}')$  also requires the boundary condition that  $G(\omega, \mathbf{r}, \mathbf{r}') \to 0$  as  $|\mathbf{r} - \mathbf{r}'| \to \infty$ , as well as the physical condition of causality; i.e., cause always precedes effect. In this chapter, only the Green function for the uniform medium is considered. The effects induced by  $\sigma(\mathbf{r})$  are treated in the subsequent chapters.

For the uniform medium case of  $\kappa = \kappa_0$ , the response of the system is independent of the source position, and the Green function depends only on the relative separation  $\mathbf{r} - \mathbf{r}'$  between the source and the detector. Therefore,

$$(\nabla^2 + \kappa_0^2)G_0(\omega, \mathbf{r} - \mathbf{r}') = \delta(\mathbf{r} - \mathbf{r}').$$
(2.17)

A simple solution approach to (2.17) is to first solve for the spatial frequency component  $G_0(\omega, \mathbf{k})$  by writing

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$$G_0(\omega, \mathbf{r} - \mathbf{r}') = \frac{L^d}{N} \int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^d} \exp[\mathrm{i}\mathbf{k} \cdot (\mathbf{r} - \mathbf{r}')] G_0(\omega, \mathbf{k})$$
(2.18)

and

$$\delta(\mathbf{r} - \mathbf{r}') = \int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^d} \exp[\mathrm{i}\mathbf{k} \cdot (\mathbf{r} - \mathbf{r}')], \qquad (2.19a)$$

where d denotes the spatial dimensionality,  $L^d$  is the sample volume  $[(2\pi)^d L^{-d}]$ may be regarded as the smallest volume unit  $(\delta k)^d$  in the wave vector space], and N is the number of states, or atoms, in the sample.  $L^d/N = (\delta r)^d$  is the smallest volume unit in the system. The notation here is based on viewing the continuum as the limit of infinitesimal discretization, so that both the continuum case and the lattice case may be treated on the same footing. The presence of N in (2.18) is due to the fact that a plane wave has magnitude  $\sqrt{N}$ . The convention in this book is to associate a factor 1/N with every **k**-integration (summation) to compensate for that fact. To be consistent,  $\delta(\mathbf{k} - \mathbf{k}')$ , as defined by

$$\int \mathrm{d}\mathbf{k}\,\delta(\mathbf{k}-\mathbf{k}')=1$$

has the following Fourier representation:

$$\delta(\mathbf{k} - \mathbf{k}') = \frac{1}{(2\pi)^d} \int d\mathbf{r} \exp[i(\mathbf{k} - \mathbf{k}') \cdot \mathbf{r}].$$
 (2.19b)

In (2.19a)

$$\int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^d}$$

denotes integration over all **k** vectors. For d = 1, that means

$$\int_{-\infty}^{\infty} \frac{\mathrm{d}k}{2\pi} \, .$$

For d = 2, that means

$$\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \frac{\mathrm{d}k_x \mathrm{d}k_y}{(2\pi)^2}$$

in Cartesian coordinates and

$$\int_0^{2\pi} \mathrm{d}\theta \int_0^\infty \frac{k \mathrm{d}k}{(2\pi)^2}$$

in circular (cylindrical) coordinates. For d = 3, that means

$$\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \frac{\mathrm{d}k_x \mathrm{d}k_y \mathrm{d}k_z}{(2\pi)^3}$$
in Cartesian coordinates and

$$\int_0^{2\pi} \mathrm{d}\phi \int_0^{\pi} \mathrm{d}\theta \int_0^{\infty} \frac{\sin\theta \, k^2 \mathrm{d}k}{(2\pi)^3}$$

in spherical coordinates. Substitution of (2.18) and (2.19a) into (2.17) gives

$$\frac{L^{d}}{N} \int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^{d}} \exp[\mathrm{i}\mathbf{k} \cdot (\mathbf{r} - \mathbf{r}')] [\kappa_{0}^{2}(\omega) - k^{2}] G_{0}(\omega, \mathbf{k}) 
= \int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^{d}} \exp[\mathrm{i}\mathbf{k} \cdot (\mathbf{r} - \mathbf{r}')].$$
(2.20)

Equating the amplitudes of each  $\mathbf{k}$  component on both sides yields

$$[\kappa_0^2(\omega) - k^2]G_0(\omega, \mathbf{k}) = \frac{N}{L^d},$$
(2.21)

or

$$G_0(\omega, \mathbf{k}) = \frac{1}{\kappa_0^2(\omega) - k^2} \frac{N}{L^d}.$$
(2.22)

The Green function  $G_0(\omega, \mathbf{k})$  gives the response of the homogeneous medium to a source excitation that has frequency  $\omega$  and spatial wave vector  $\mathbf{k}$ . It is noted that if the source  $\mathbf{k}$  and  $\omega$  coincide with the wave dispersion relation of the medium, i.e.,  $k^2 = \kappa_0^2(\omega)$ , then  $G_0$  diverges. In analogy with the response of a forced oscillator which has its own resonance frequencies, this divergent response is expected, since the dispersion relation essentially gives the condition of natural resonance(s) in a medium. On the other hand, when the source frequency and  $\mathbf{k}$  do not coincide with any natural mode of the system, (2.22) indicates that it can still excite many modes just by its proximity to the resonance condition, as in the case of a forced oscillator.

#### 2.2.2 Green Function and the Density of States

The divergence in  $G_0(\omega, \mathbf{k})$  causes a problem in the transformation back to the spatial domain. To fix this problem, it is generally the practice to add a small imaginary constant  $i\eta$  to the denominator, then to take the limit as  $\eta$  approaches zero. This simple procedure is illustrated in the solution to Problem 2.1 at the end of this chapter. The result is

$$\frac{L^d}{N}G_0^{\pm}(\omega, \mathbf{k}) = \lim_{\eta \to 0} \frac{1}{\kappa_0^2(\omega) - k^2 \pm i\eta} 
= P \frac{1}{\kappa_0^2(\omega) - k^2} \mp i\pi \delta[\kappa_0^2(\omega) - k^2].$$
(2.23)

Here the sign of the imaginary constant is important in determining the behavior of the Green function in real space, as will be seen later. The symbol Pmeans taking the principal value of  $[\kappa_0^2(\omega) - k^2]^{-1}$  as that quantity diverges, which is defined in the solution to Problem 2.1. The imaginary part of the Green function, on the other hand, is interesting because the delta function

picks out exactly the modes of the system. If one wants to count the number of modes per unit frequency range and per unit volume  $L^d/N$ , i.e., to get the density of states  $\rho_0(\omega)$ , it can be achieved by integrating the delta function over all possible free-space eigenstates that are characterized by the **k** vectors, i.e.,

$$\rho_0(\omega) = \frac{\mathrm{d}\kappa_0^2(\omega)}{\mathrm{d}\omega} \int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^d} \,\delta[\kappa_0^2(\omega) - k^2]. \tag{2.24}$$

Here the factor

$$\frac{\mathrm{d}\kappa_0^2(\omega)}{\mathrm{d}\omega} = \begin{cases} \frac{2m}{\hbar} & \text{quantum} \\ \frac{2\omega}{v_0^2} & \text{classical} \end{cases}$$
(2.25)

accounts for the fact that the delta function counts the states in units of  $\kappa_0^2$ , so it is required for conversion to states per unit frequency. It is shown in Problem 2.2 that  $\rho_0(\omega)$  is given by

$$\rho_0(\omega) = \begin{cases} \frac{m^{3/2}}{\sqrt{2}\hbar^{3/2}\pi^2}\sqrt{\omega} & 3D\\ \frac{m}{2\pi\hbar} & 2D\\ \frac{1}{\pi}\sqrt{\frac{m}{2\hbar}\frac{1}{\sqrt{\omega}}} & 1D \end{cases}$$
(2.26)

for the quantum case, and

$$\rho_{0}(\omega) = \begin{cases} \frac{\omega^{2}}{2\pi^{2}v_{0}^{3}} & 3D\\ \frac{\omega}{2\pi v_{0}^{2}} & 2D\\ \frac{1}{\pi v_{0}} & 1D \end{cases}$$
(2.27)

for the classical wave case. Here  $\omega$  is defined to be positive.

In (2.24), the integral may be re-expressed as

$$\int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^d} \,\delta[\kappa_0^2(\omega) - k^2] = -\frac{1}{\pi} \frac{L^d}{N} \int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^d} \mathrm{Im}G_0^+(\omega, \mathbf{k})$$
$$= -\frac{1}{\pi} \mathrm{Im}G_0^+(\omega, \mathbf{r} = \mathbf{r}'), \qquad (2.28)$$

where Im means taking the imaginary part of what follows. The last equality in (2.28) follows because at  $\mathbf{r} = \mathbf{r}'$ , the phase factor  $\exp[\mathbf{i}\mathbf{k} \cdot (\mathbf{r} - \mathbf{r}')]$  in the Fourier transform of  $G_0^+(\omega, \mathbf{k})$  becomes 1, so the real and imaginary parts of  $G_0^+(\omega, \mathbf{r}=\mathbf{r}')$  are just the **k**-integrals of the real and imaginary parts of  $G_0^+(\omega, \mathbf{k})$ . Substitution of (2.28) into (2.24) gives

$$\rho_0(\omega) = -\frac{\mathrm{d}\kappa_0^2(\omega)}{\mathrm{d}\omega} \frac{1}{\pi} \mathrm{Im} G_0^+(\omega, \mathbf{r} = \mathbf{r}').$$
(2.29)

The reason for expressing the density of states in terms of the imaginary part of the Green function (in the  $\omega$  and **r** representation) is the general validity

of the formula not only for waves in a uniform medium but also for waves in random media. In solution to Problem 2.3, it is shown that

$$\rho(\omega, \mathbf{r}) = -\frac{\mathrm{d}\kappa_0^2(\omega)}{\mathrm{d}\omega} \frac{1}{\pi} \mathrm{Im} G^+(\omega, \mathbf{r} = \mathbf{r}')$$
(2.30)

gives the "local density of states" at  $\mathbf{r}$  in the general case. The fact that the density of states can always be obtained this way is due to the Green function's general property that it diverges at the resonant modes of the system, and as a result its imaginary part offers a way of picking them out as seen from (2.23). Another general property of the Green function is that  $G^+(\omega, \mathbf{r}, \mathbf{r}')$  and  $G^-(\omega, \mathbf{r}', \mathbf{r})$  are complex conjugates of each other (note the interchange of  $\mathbf{r}$  and  $\mathbf{r}'$ ), and  $G_0^+$ ,  $G_0^-$  are complex conjugates in both the real space and the momentum space. These and other facts about the Green function are made clear in the solution to Problem 2.3.

#### 2.2.3 Real Space Green Functions

The Green function in the  $(\omega, \mathbf{k})$  domain is noted to have the same expression in different spatial dimensionalities. However,  $G_0^{\pm}(\omega, \mathbf{r} - \mathbf{r}')$  is different in different spatial dimensions. This can be seen by the Fourier transformation of  $G_0^{\pm}(\omega, \mathbf{k})$ . In 3D, we have

$$G_{0}^{\pm}(\omega, \mathbf{r} - \mathbf{r}') = \lim_{\eta \to 0} \frac{1}{4\pi^{2}} \frac{L^{a}}{N}$$

$$\times \int_{0}^{\infty} \frac{k^{2} dk}{\kappa_{0}^{2} - k^{2} \pm i\eta} \int_{-1}^{1} d(\cos \theta) \exp(ik|\mathbf{r} - \mathbf{r}'|\cos \theta) \frac{N}{L^{d}}$$

$$= \lim_{\eta \to 0} \frac{1}{4\pi^{2}|\mathbf{r} - \mathbf{r}'|} \frac{1}{i} \int_{-\infty}^{\infty} \frac{k \exp(ik|\mathbf{r} - \mathbf{r}'|)}{\kappa_{0}^{2} - k^{2} \pm i\eta} dk$$

$$= -\frac{\exp[\pm i(\kappa_{0} \pm i\eta)|\mathbf{r} - \mathbf{r}'|]}{4\pi|\mathbf{r} - \mathbf{r}'|} \qquad (2.31)$$

through contour integration in the upper complex k plane (so that the integrand vanishes in the limit of  $k \to +i\infty$ ).  $G_0^{\pm}(\omega, \mathbf{r} - \mathbf{r}')$  is seen to correspond to outgoing (+) and incoming (-) spherical waves. The sign of  $i\eta$ determines whether the wave is outgoing or incoming because  $G_0$  must vanish as  $|\mathbf{r} - \mathbf{r}'| \to \infty$ , so that  $\kappa_0 + i\eta$  must be outgoing whereas  $\kappa_0 - i\eta$  must be incoming. In 2D, similar contour integration and comparison with the integral representation of the Hankel function yield  $G_0^{\pm}$  as outgoing and incoming cylindrical waves

$$G_0^{\pm}(\omega, \mathbf{r} - \mathbf{r}') = \mp \frac{i}{4} H_0^{(1,2)}(\kappa_0 |\mathbf{r} - \mathbf{r}'|), \qquad (2.32)$$

where  $H_0^{(1,2)}$  denotes the zeroth-order Hankel function of the first (outgoing wave) kind and the second (incoming wave) kind. In 1D,  $G_0^{\pm}$  is given by

$$G_0^{\pm}(\omega, \mathbf{r} - \mathbf{r}') = \mp \frac{\mathrm{i}}{2\kappa_0} \exp(\pm \mathrm{i}\kappa_0 |\mathbf{r} - \mathbf{r}'|).$$
(2.33)

It is easy to check that the imaginary part of  $G_0^{\pm}(\omega, \mathbf{r}=\mathbf{r}')$ , when it is substituted into (2.29), gives the correct density of states. A further Fourier transform with respect to frequency  $\omega$  would yield Green function in the  $(t, \mathbf{r})$  domain. Since  $\kappa_0^2$  is linear in  $\omega$  for quantum waves and quadratic in  $\omega$  for classical waves, the Green functions for the two types of waves are expected to differ in the  $(t, \mathbf{r})$  domain (Economou 1983). However, they share the same causal characteristics: In the time domain  $G_0^+(t, |\mathbf{r} - \mathbf{r}'|) = 0$  for t < 0 and  $G_0^-(t, |\mathbf{r} - \mathbf{r}'|) = 0$  for t > 0. This fact is easy to see in the case of 3D classical waves, where the Fourier transform of the numerator of (2.31) gives  $\delta(|\mathbf{r} - \mathbf{r}'|/v_0 \mp t)$ . Thus,  $G_0^+$  is usually denoted the *advanced* Green function.

# 2.3 Waves on a Discrete Lattice

In the case of electrons in solids, the interaction of an electron with the solid lattice is known to play a dominant role in the electron's resulting behavior. It is therefore desirable that the lattice effect be taken into account first, before examining any effect due to disorder. Basic to the consideration of the lattice effect is the lattice potential. If the potential V at each site of an ordered, perfect lattice is given by a Coulomb or some form of attractive potential, then the solution of the Schrödinger equation for each site, considered in isolation from other sites, would result in many atomic energy levels. On adding to this problem the interaction between the sites, i.e., the effect of the lattice structure, the combined state of affair can become very complicated indeed. However, in many electronic problems of physical interest only particular one atomic state is important, i.e., the valence electronic state. If the energy separations between the valence state and other atomic states are large compared to the interaction energy between the sites, then the problem can be simplified by solution in two stages. The first stage considers only the atomic-level problem for each site, in isolation from other sites. The next stage considers the effect of lattice structure on each atomic level, without worrying about the intermixing of the atomic levels. This approach is the basis of the "tight-binding approximation" in which the lattice-structure effect is treated as a perturbation on the single-site atomic energy levels. In what follows, the description of the quantum Schrödinger wave is based on this tight binding approach, where the one relevant atomic level is assumed to be identified, and our attention is therefore focused on the effect of the lattice structure plus the randomness.

#### 2.3.1 Dispersion Relation

The question before us is: If  $\phi(t, \mathbf{l})$  stands for the wave amplitude of the relevant atomic level at lattice site  $\mathbf{l}$  and time t, then what is the effect of the lattice on its frequency component  $\phi(\omega, \mathbf{l})$ ? To answer this question, let

us start with the wave form in a perfect lattice,  $\exp(-i\omega t + i\mathbf{k} \cdot \mathbf{l})$ , and ask what general spatial operator would yield  $\exp(i\mathbf{k} \cdot \mathbf{l})$  as its eigenfunction. Since in continuum  $\exp(i\mathbf{k} \cdot \mathbf{r})$  is the eigenfunction of the Laplacian operator  $\nabla^2$ , the obvious choice in the present case is the Laplacian defined on the lattice, i.e., a discrete Laplacian  $\Delta^{(2)}$ . In 1D, where  $l = \ldots, -2, -1, 0, 1, 2, \ldots$  and the lattice constant is denoted by a, the operation of the discrete Laplacian is defined by

$$\Delta^{(2)}\phi(\omega,l) = \frac{1}{a^2} [\phi(\omega,l+1) + \phi(\omega,l-1) - 2\phi(\omega,l)].$$
(2.34)

Equation (2.34) expresses the operation of taking the discrete second derivative of  $\phi$ . For a simple square lattice in 2D and a simple cubic lattice in 3D, the operation of  $\Delta^{(2)}$  is a direct extension from 1D:

$$\Delta^{(2)}\phi(\omega,\mathbf{l}) = \frac{1}{a^2} \left[ -2d\phi(\omega,\mathbf{l}) + \sum_{\mathbf{n}} \phi(\omega,\mathbf{l}+\mathbf{n}) \right], \qquad (2.35)$$

where **n** denotes the vectors pointing from the lattice site **l** to its nearest neighbors. The sum therefore has four terms in 2D and six terms in 3D. By substituting  $\exp(-i\omega t + i\mathbf{k}\cdot\mathbf{l})$  for  $\phi(\omega,\mathbf{l})$  and similar forms for  $\phi(\omega,\mathbf{l}+\mathbf{n})$  into (2.35), we get

$$\Delta^{(2)}\phi(\omega,\mathbf{l}) = -e(\mathbf{k})\phi(\omega,\mathbf{l}), \qquad (2.36)$$

where  $e(\mathbf{k})$  is given by

$$e(\mathbf{k}) = \frac{2}{a^2} \sum_{\alpha=1}^{d} (1 - \cos k_{\alpha} a).$$
(2.37)

Here  $k_{1,2,3}$  means  $k_{x,y,z}$ . When  $k_{\alpha}a \to 0$ , i.e., in the long-wavelength limit,  $1 - \cos k_{\alpha}a \cong k_{\alpha}^2 a^2/2$ , and

$$e(\mathbf{k}) \cong \sum_{\alpha=1}^{d} k_{\alpha}^2 = k^2, \qquad (2.38)$$

which corresponds precisely to the negative of the eigenvalue for the Laplacian. Since  $\hbar^2 k^2/2m$  is the kinetic energy of a quantum wave/particle,  $\hbar^2 e(\mathbf{k})/2m^*$  may thus be identified as the similar quantity on a lattice. Here  $m^*$  is the "effective mass," to be defined later.

Equation (2.37) indicates significant differences between lattice and continuum waves. In the first place,  $e(\mathbf{k})$  is periodic in  $\mathbf{k}$ . This is due to the fact that lattice has a minimum length scale, a, which means that there cannot be a wave with wavelength less than 2a, or a  $\mathbf{k}$  with  $|k_{\alpha}| \geq \pi/a$ . When  $|k_{\alpha}|$  is formally extended beyond  $\pi/a$ , it becomes equivalent to a state with  $|k_{\alpha}| \leq \pi/a$ . A related second difference is that  $e(\mathbf{k})$  has a maximum, defined by  $4d/a^2$ . The eigenvalues of  $-\Delta^{(2)}$  thus form a band, extending from 0 to

 $4d/a^2$ . At the maximum, which occurs at  $|k_{\alpha}| = \pi/a$ , the dispersion relation gives a group velocity  $de(\mathbf{k})/dk_{\alpha} = 0$  in the direction of  $k_{\alpha}$ , which indicates a standing wave caused by Bragg scattering from the lattice structure. At frequencies beyond the maximum so that  $e(\mathbf{k}) > 4d/a^2$ , i.e., outside the band, it is clear that the only way for (2.37) to be satisfied is by switching  $k_{\alpha}$  from real to imaginary so that  $\cos(k_{\alpha}a)$  becomes  $\cosh(k_{\alpha}a)$ . However, a "state" which is associated with an imaginary k must decay exponentially toward infinity. These waves are called evanescent waves and are analogous to what happens in waveguides below the cutoff frequency. The third difference is that on a lattice, the "mass" of a quantum wave/particle is not necessarily the same as that in the continuum. Since in the long-wavelength limit the dispersion relation of the lattice wave has the same form as that in the continuum, the *effective mass*  $m^*$  on a lattice may be defined by the curvature of the kinetic energy expression around  $\mathbf{k} = 0$ :

$$\frac{1}{m^*} = \frac{1}{\hbar^2} \frac{\mathrm{d}^2}{\mathrm{d}k^2} (\text{kinetic energy})_{\mathbf{k}=0}.$$
 (2.39)

The kinetic energy is dependent on the strength of coupling between the neighboring sites. If the interaction between neighboring sites is weak, the electron tends to be bound at a given site, and the transfer between the sites would occur slowly, implying a large  $m^*$ . The effective mass  $m^*$  may therefore be regarded as a parameter for measuring the lattice effect on an electron.

The total energy, E, for a quantum wave on a lattice is

$$E = \frac{\hbar^2}{2m^*} e(\mathbf{k}) + \varepsilon_0 - \frac{\hbar^2}{2m^*} \left(\frac{2d}{a^2}\right), \qquad (2.40)$$

where  $\varepsilon_0$  is the energy of the atomic level, defined relative to potential V = 0. The fact that  $\varepsilon_0$  always defines the center of the energy band accounts for the last term, which is half of the energy bandwidth. Equation (2.40) relates  $\mathbf{k}$  and  $E = \hbar \omega$  for the quantum Schrödinger wave on a lattice. Just as in the case of waves in continuum, the corresponding wave equation may be inferred from the dispersion relation. By noting that  $e(\mathbf{k})$  is obtained as the eigenvalue of  $-\Delta^{(2)}$ , it is easy to see that the lattice wave equation is

$$(\beta \Delta^{(2)} + \kappa_0^2)\phi(\omega, \mathbf{l}) = 0,$$
 (2.41)

where  $\kappa_0^2$  is now renormalized with respect to  $\varepsilon_0 - (\hbar^2 d/m^* a^2)$ :

$$\kappa_0^2 = \frac{2m(E-\varepsilon_0)}{\hbar^2} + \frac{2\beta d}{a^2},\tag{2.42}$$

with

$$\beta = \frac{m}{m^*}.\tag{2.43}$$

So far we have discussed only quantum waves on a lattice. However, (2.41) is equally applicable to waves in a classical spring-and-mass lattice. This can

be seen by letting  $\beta = 1$  and  $\kappa_0^2 = \omega^2/v_0^2$ , where  $v_0^2 = Ka^2/m$  is the square of wave speed in the long-wavelength limit, K being the spring constant of the springs linking identical nearest-neighbor particles of mass m. Equation (2.41) then becomes

$$Ka^{2}\Delta^{(2)}\phi(\omega,\mathbf{l}) = m\omega^{2}\phi(\omega,\mathbf{l}).$$
(2.44)

The left-hand side of (2.44) is the negative of the force on the particle situated at **l**, where  $\phi$  denotes the displacement of the particle away from its equilibrium position. For example, in 1D the expression

$$K\{[\phi(\omega, l+1) - \phi(\omega, l)] - [\phi(\omega, l) - \phi(\omega, l-1)]\}$$

gives the force on the particle at l due to the net contraction/extension of the springs on its two sides. The right-hand side of (2.44) is the mass times acceleration of the particle (at frequency  $\omega$ ). Therefore, this is precisely the dynamical equation for the vibrational modes on an elastic lattice. One thus concludes that (2.41) is the general form for both the quantum and the classical waves on a lattice.

An alternative way of expressing (2.41) is to use a vector  $\mathbf{\Phi}$  to express the wave function, where each component of the vector corresponds to the value of the wave function at a site. The operation  $-\beta\Delta^{(2)}$  can then be expressed by a matrix  $\mathbf{D}$  whose diagonal elements are

$$(\mathbf{D})_{\mathbf{l},\mathbf{l}} = \frac{2\beta d}{a^2} \tag{2.45}$$

and whose off-diagonal elements are

$$(\mathbf{D})_{\mathbf{l},\mathbf{l}+\mathbf{n}} = -\frac{\beta}{a^2} \tag{2.46}$$

and zero otherwise. A succinct way of writing the matrix is by using Dirac's bra and ket notation, where  $|\phi\rangle = \Phi$  means the wave function expressed as a column vector and  $\langle \phi |$  means the same wave function but expressed as a row vector.  $\langle \phi | \phi' \rangle$  means the inner product (dot product between two vectors), and  $|\phi'\rangle \langle \phi |$  means the outer product. In Dirac's notation,

$$\mathbf{D} = \sum_{\mathbf{l}} \frac{2\beta d}{a^2} \left| \mathbf{l} \right\rangle \left\langle \mathbf{l} \right| + \sum_{\mathbf{l},\mathbf{n}} \left( -\frac{\beta}{a^2} \right) \left| \mathbf{l} \right\rangle \left\langle \mathbf{l} + \mathbf{n} \right|, \qquad (2.47)$$

where the second sum is over all the nearest neighbors of **l**, for all **l**, and  $|\mathbf{l}\rangle$  means a column vector which has all zero components except for the component **l**, which is 1. The same holds for row vector  $\langle \mathbf{l} |$ . The outer product  $|\mathbf{l}\rangle \langle \mathbf{l} + \mathbf{n} |$  is a matrix whose elements are all zero except the matrix element with value one at the row corresponding to **l** and the column corresponding to  $\mathbf{l} + \mathbf{n}$ . **D** is expressed in (2.47) as the sum of such matrices. Equation (2.41) can then be written as

$$\mathbf{D} \left| \phi \right\rangle = \kappa_0^2 \left| \phi \right\rangle. \tag{2.48}$$

#### 2.3.2 Random Discrete Lattices: the Anderson Model

There are two ways to introduce randomness into the lattice problem – either via random variation in the atomic energy level or via random variation in  $\beta$ . If  $\Delta \varepsilon_1$  denotes the random deviation (from  $\varepsilon_0$ ) of the atomic energy level at site l, then  $\Delta \varepsilon_1$  randomizes only the diagonal matrix elements. For the phonon case, the same effect is obtained by varying the mass of the particle at each site by  $\Delta m_1$ . When there is randomness,  $\kappa_0^2$  should be replaced by  $\kappa_0^2 - \sigma(1)$ in (2.41), where

$$\sigma(\mathbf{l}) = \begin{cases} \frac{2m\Delta\varepsilon_{\mathbf{l}}}{\hbar^2} & \text{quantum} \\ \frac{\omega^2}{v_0^2} \frac{\Delta m_{\mathbf{l}}}{m} & \text{classical.} \end{cases}$$
(2.49)

Since one is free to define  $\varepsilon_0$  and m as the mean values of the respective quantities,  $\Delta \varepsilon_1$  and  $\Delta m_1$  may be regarded as having zero mean:

$$\left\langle \sigma(\mathbf{l})\right\rangle_{c} = 0, \tag{2.50}$$

where  $\langle \rangle_c$  means configurational averaging, which is equivalent to averaging over all sites in the present case. By defining a new matrix **M** such that the diagonal elements consist of only the deviations, i.e.,

$$\mathbf{M} = \sum_{\mathbf{l}} \sigma(\mathbf{l}) \left| \mathbf{l} \right\rangle \left\langle \mathbf{l} \right| + \sum_{\mathbf{l},\mathbf{n}} \left( -\frac{\beta}{a^2} \right) \left| \mathbf{l} \right\rangle \left\langle \mathbf{l} + \mathbf{n} \right|, \qquad (2.51)$$

the equation

$$\left[\beta\Delta^{(2)} + \kappa_0^2 - \sigma(\mathbf{l})\right]\phi(\omega, \mathbf{l}) = 0 \tag{2.52}$$

may be written as

$$\mathbf{M} \left| \phi \right\rangle = q_0^2 \left| \phi \right\rangle, \tag{2.53}$$

where

$$q_0^2 = \kappa_0^2 - \frac{2\beta d}{a^2},$$
 (2.54a)

with

$$\kappa_0^2 = \frac{2m(E-\varepsilon_0)}{\hbar^2} + \frac{2\beta d}{a^2}$$

in the quantum case and

$$\kappa_0^2 = \frac{\omega^2}{v_0^2},\tag{2.54b}$$

 $\beta = 1$  in the phonon case. Equations (2.51) and (2.53), together with the condition (2.50), are known as the *Anderson model*. Here the disorder is denoted "diagonal randomness."

The second way to introduce randomness into the model is by varying  $\beta$ . That would randomize the off-diagonal elements as well as the diagonal elements of **D**. However, for the quantum case  $q_0 = 2m(E - \varepsilon_0)/\hbar^2$ , therefore  $\beta$ does not enter into the diagonal elements of **M**. Randomizing  $\beta$  thus constitutes an "off-diagonal randomness" model. For the phonon problem, on the other hand, varying  $\beta$  would randomize both the diagonal and off-diagonal elements in a correlated manner. In what follows, only the case of diagonal randomness is pursued because the randomness in  $\beta$  is known to have an effect on the scattering and localization behavior qualitatively similar to that in the diagonal randomness case. Studying one is therefore sufficient to illustrate the physics involved.

# 2.4 Lattice Green Functions

The Green function for the perfect lattice is defined by the solution to the following equation:

$$(\beta\Delta^{(2)} + \kappa_0^2)G_0(\omega, \mathbf{l} - \mathbf{l}') = \frac{1}{a^d}\delta_{\mathbf{l},\mathbf{l}'}, \qquad (2.55)$$

where  $\delta_{\mathbf{l},\mathbf{l}'}$  is the Kronecker delta defined by

$$\delta_{\mathbf{l},\mathbf{l}'} = \begin{cases} 1 & \text{if } \mathbf{l} = \mathbf{l}' \\ 0 & \text{otherwise.} \end{cases}$$
(2.56)

To solve for  $G_0(\omega, \mathbf{l} - \mathbf{l}')$ , we use the same approach as in the continuum case by first solving for the spatial frequency component  $G_0(\omega, \mathbf{k})$ , related to  $G_0(\omega, \mathbf{l} - \mathbf{l}')$  by

$$G_0(\omega, \mathbf{l} - \mathbf{l}') = \frac{L^d}{N} \int_{1BZ} \frac{d\mathbf{k}}{(2\pi)^d} \exp[i\mathbf{k} \cdot (\mathbf{l} - \mathbf{l}')] G_0(\omega, \mathbf{k}).$$
(2.57)

The Kronecker delta may also be expressed as

$$\frac{1}{a^d} \delta_{\mathbf{l},\mathbf{l}'} = \int_{1\mathrm{BZ}} \frac{\mathrm{d}\mathbf{k}}{(2\pi)^d} \exp[\mathrm{i}\mathbf{k} \cdot (\mathbf{l} - \mathbf{l}')].$$
(2.58)

Here the notation 1BZ means restricting the values of  $|k_x|, |k_y|, |k_z| \leq \pi/a$ , which is the first Brillouin zone for the simple cubic lattice in 3D and the simple square lattice in 2D. These are the only lattices considered throughout the succeeding chapters. By substituting (2.57) and (2.58) into (2.55), the operation of  $\Delta^{(2)}$  on exp[ik(l - l')] is seen to yield  $-e(\mathbf{k})$  as its eigenvalue. Therefore, by equating the left- and right-hand sides of each **k** component, one gets

$$[\kappa_0^2 - \beta e(\mathbf{k})]G_0(\omega, \mathbf{k}) = \frac{N}{L^d} = \frac{1}{a^d},$$
(2.59)

or

$$G_0(\omega, \mathbf{k}) = \frac{1}{\kappa_0^2 - \beta e(\mathbf{k})} \frac{1}{a^d}.$$
(2.60)

# 2.4.1 1D Lattice Green Function

Comparison with the continuum case, (2.22), shows the only difference to be the replacement of  $k^2$  by  $\beta e(\mathbf{k})$ , and  $L^d/N$  by  $a^d$ . In 1D, the calculation of  $G_0(\omega, l - l')$  from  $G_0(\omega, \mathbf{k})$  may be performed analytically. By substituting (2.60) into (2.57) and using the definition of  $e(\mathbf{k})$  (2.37), we get

$$G_{0}^{\pm}(\omega, l - l') = a \int_{-\pi/a}^{\pi/a} \frac{\mathrm{d}k}{2\pi} \frac{\exp[\mathrm{i}k(l - l')]}{\kappa_{0}^{2} - 2\beta(1 - \cos ka)/a^{2} \pm \mathrm{i}\eta} \frac{1}{a}$$
$$= \frac{a}{2\pi\beta} \int_{-\pi}^{\pi} \mathrm{d}(ka)$$
$$\times \frac{\exp[\mathrm{i}ka|l - l'|/a]}{2(\kappa_{0}^{2}a^{2}/2\beta - 1) + [\exp(\mathrm{i}ka) + \exp(-\mathrm{i}ka)]}, \quad (2.61)$$

where  $\pm i\eta$  is dropped in the second line of (2.61), but the constant  $(\kappa_0^2 a^2/2\beta - 1)$  should be remembered to have a  $\pm i\eta$  imaginary part. Also, since the integral is invariant with respect to the interchange of l and l' (because the ka integral is from  $-\pi$  to  $+\pi$ ), it must be a function of |l-l'|. By changing to the variable  $z = \exp(ika)$  and noting that d(ka) = dz/iz, the definite integral may be converted into a contour integral around the unit circle in the complex z-plane:

$$G_0^{\pm}(\omega, l - l') = \frac{a}{2\pi i\beta} \oint dz \frac{z^{|l-l'|/a}}{z^2 + 2uz + 1},$$
(2.62)

where  $u = (\kappa_0^2 a^2/2\beta - 1) \pm i\eta$ . The denominator of the integrand has two roots,

$$z_1 = -u + \sqrt{u^2 - 1}, \tag{2.63}$$

$$z_2 = -u - \sqrt{u^2 - 1}. \tag{2.64}$$

For  $|u| \leq 1$ ,  $|z_1| = |z_2| = 1$ , and the contour integration has to rely on  $\pm i\eta$  to give

$$G_0^{\pm}(\omega, l - l') = \frac{\mp a \mathbf{i}}{2\beta\sqrt{1 - u^2}} (-u \pm \mathbf{i}\sqrt{1 - u^2})^{|l - l'|/a}.$$
 (2.65a)

This form of  $G_0^{\pm}(\omega, l-l')$  may be put in a more illuminating form by defining a  $k_0$  as the solution to the equation  $\kappa_0^2 = \beta e(k_0)$ . Then from the definition of u and  $e(k_0)$  one gets  $-u = \cos k_0 a$ , so that

$$G_0^{\pm}(\omega, l - l') = \frac{\mp a \mathbf{i}}{2\beta \sin k_0 a} \exp(\pm \mathbf{i} k_0 |l - l'|)$$
$$= \frac{\mp \mathbf{i}}{(\partial \beta e(\mathbf{k})/\partial k)_{k_0}} \exp(\pm \mathbf{i} k_0 |l - l'|).$$
(2.65b)

Comparison with (2.33) shows remarkable similarity. The only difference is that the pre-exponential factor is now modified. The continuum form of the pre-exponential factor is recovered in the limit of  $k_0 \rightarrow 0$  and  $\beta = 1$ .

When u > 1, then  $|z_1| < 1$  and  $|z_2| > 1$ , which means only the pole at  $z = z_1$  is inside the contour. But if u < -1, the reverse is true. In either case,  $\pm i\eta$  does not play a role any more, and

$$G_0^{\pm}(\omega, l-l') = \frac{\pm a}{2\beta\sqrt{u^2 - 1}}(-u \pm \sqrt{u^2 - 1})^{|l-l'|/a},$$
(2.66)

where the + and - signs correspond to the cases u > 1 and u < -1, respectively. Since the right-hand side of (2.66) is a real number, it follows that the density of states is zero for frequencies that give |u| > 1. In this case  $G_0^{\pm}(\omega, l - l')$  is also noted to decay (or grow) geometrically (exponentially) as a function of |l - l'|.

To calculate the density of states from (2.29), (2.65), and (2.66), the substitution  $u = ma^2(E - \varepsilon_0)/\beta\hbar^2$  is used for the quantum case to obtain

$$\rho_{0}(\omega) = -\frac{1}{\pi} \frac{\mathrm{d}\kappa_{0}^{2}}{\mathrm{d}\omega} \mathrm{Im}G_{0}^{+}(\omega, \mathbf{l} = \mathbf{l}') \\
= \begin{cases} \frac{ma}{\pi\hbar\beta} \frac{1}{\sqrt{1 - (\hbar\omega - \varepsilon_{0})^{2}/(\beta\hbar^{2}/ma^{2})^{2}}} & |\hbar\omega - \varepsilon_{0}| \leq \frac{\beta\hbar^{2}}{ma^{2}} \\ 0 & |\hbar\omega - \varepsilon_{0}| > \frac{\beta\hbar^{2}}{ma^{2}}. \end{cases}$$
(2.67)

This is shown in Fig.2.1. Near the lower band edge  $\rho_0(\omega)$  is noted to have the same divergence as in a uniform medium. For the phonon case,  $u = (m\omega^2/2K) - 1$ ,  $\beta = 1$ , so that



Fig. 2.1. The electronic density of states for a 1D lattice, normalized to per site per frequency unit  $\beta\hbar/ma^2$ . Here  $u = (\hbar\omega - \varepsilon_0)/(\beta\hbar^2/ma^2)$ . The integrated density of states over the whole band is 1. The *solid curve* is described by the function  $(\pi\sqrt{1-u^2})^{-1}$ .



Fig. 2.2. The phonon density of states on a 1D lattice, normalized to per site per frequency unit  $\sqrt{K/m}$ . Here  $u = \omega/\sqrt{K/m}$ . The integrated density of states over the whole band is 1. The *solid curve* is described by the function  $(\pi\sqrt{1-(u/2)^2})^{-1}$ .

$$\rho_0(\omega) = \begin{cases} \frac{2}{\pi a} \frac{1}{\sqrt{(4K/m) - \omega^2}} & \omega \le 2\sqrt{K/m} \\ 0 & \omega > 2\sqrt{K/m}. \end{cases}$$
(2.68)

As  $\omega \to 0, \rho_0(\omega)$  for the continuum case is recovered exactly (by using the identity  $v_0 = a\sqrt{K/m}$ ), as expected. At the upper band edge the qualitative behavior is similar to that of the quantum case. This is shown in Fig. 2.2. For the sake of uniformity in notation the variable u is redefined here to mean the dimensionless frequency  $\omega/\sqrt{K/m}$ .

#### 2.4.2 2D Lattice Green Function

In 2D, the general Green function can no longer be calculated analytically. This is because a constant value of  $e(\mathbf{k})$  does not correspond to a fixed  $|\mathbf{k}|$ , i.e., the Brillouin zone is not circular, and as a result analytic calculation becomes difficult. However, the Green function for  $\mathbf{l} = \mathbf{l}'$ , and therefore the density of states, can still be explicitly evaluated in terms of elliptic integrals. Details of the evaluation are given in the solution to Problem 2.4. The answer is

$$\rho_0(\omega) = \frac{1}{2\pi^2\beta} \frac{\mathrm{d}\kappa_0^2(\omega)}{\mathrm{d}\omega} \mathbf{K} \left[ \sqrt{1 - \left(\frac{\kappa_0^2 a^2 - 4\beta}{4\beta}\right)^2} \right]$$
(2.69)

for  $|(\kappa_0^2 a^2 - 4\beta)/4\beta| < 1$ , and zero otherwise. Here **K** denotes the complete elliptic integral of the first kind (Arfken 1970). Written out explicitly, the 2D quantum case gives

$$\rho_0(\omega) = \frac{m}{\pi^2 \beta \hbar} \mathbf{K} \left[ \sqrt{1 - \left(\frac{m(\hbar \omega - \varepsilon_0)a^2}{2\beta \hbar^2}\right)^2} \right]$$
(2.70)



Fig. 2.3. The electronic density of states for the 2D square lattice, normalized to per site per frequency unit  $\beta\hbar/ma^2$ . Here  $u = (\hbar\omega - \varepsilon_0)/(\beta\hbar^2/ma^2)$ . The density of states integrates to 1. The *solid curve* is given by  $(1/\pi^2)\mathbf{K}[\sqrt{1-(u/2)^2}]$ . The divergence at the band center is logarithmic in nature.

for  $|\hbar\omega - \varepsilon_0| \leq 2\beta\hbar^2/ma^2(|u| < 2)$ , and zero otherwise. This is shown in Fig. 2.3. At the band edge, where the argument of **K** vanishes, the value of  $\mathbf{K}(0) = \pi/2$ , and the continuum density of states is recovered if  $\beta = 1$ . It is also interesting to note that at the center of the band  $\mathbf{K}(1)$  is logarithmically divergent. For the phonon case,

$$\rho_0(\omega) = \frac{\omega}{\pi^2 v_0^2} \mathbf{K} \left[ 2 \left( \frac{\omega}{\omega_0} \right) \sqrt{1 - \left( \frac{\omega}{\omega_0} \right)^2} \right], \qquad (2.71)$$

for  $0 \le \omega \le \omega_0 = 2\sqrt{2}\sqrt{K/m}$ , and zero otherwise. This is shown in Fig. 2.4. Again, at  $\omega$  close to zero the continuum density of states is recovered.

### 2.4.3 3D Lattice Green Function

In 3D, it is no longer possible to have an analytic density of states expression for the simple cubic lattice, even for the l = l' case. However, in the solution to Problem 2.5 it is shown that the density of states can nevertheless be expressed as an integral which has to be evaluated numerically:

$$\rho_0(\omega) = -\frac{1}{\pi} \frac{\mathrm{d}\kappa_0^2}{\mathrm{d}\omega} \mathrm{Im} G_0^+(\omega, \mathbf{l} = \mathbf{l}'), \qquad (2.72)$$

where

$$\operatorname{Im} G_0^{\pm}(\omega, \mathbf{l} = \mathbf{l}') = \frac{1}{2\pi^2 a\beta} \operatorname{Im} \int_0^{\pi} \mathrm{d}\theta \mu^{\pm}(\theta) \mathbf{K}[\mu^{\pm}(\theta)], \qquad (2.73)$$

$$\mu^{\pm}(\theta) = \frac{4\beta}{|\kappa_0^2 a^2 - 6\beta + 2\beta \cos \theta| \pm \mathrm{i}\eta}.$$
(2.74)



Fig. 2.4. The phonon density of states on a 2D square lattice, normalized to per site per frequency unit  $\sqrt{K/m}$ , plotted as a function of  $u = \omega/\sqrt{K/m}$ . The density of states integrates to 1, and the upper band edge is at  $\sqrt{8}\sqrt{K/m}$ . The *solid curve* is given by  $(u/\pi^2)\mathbf{K}[(u/\sqrt{2})\sqrt{1-(u^2/8)}]$ . The divergence in the density of states is logarithmic in nature.

In Figs. 2.5 and 2.6 the densities of states for the quantum case and the phonon case are plotted, respectively. The 3D density of states differs from those of 1D and 2D in that it has no divergences.

# 2.5 Treating Continuum Problems on a Lattice

In the prior discussion, it has been noted that the continuum limit can be achieved on a lattice at the  $k \to 0$  limit. Therefore, one can always use the lattice to simulate waves in continuum by making sure that the lattice spacing is much smaller than the wavelength so that the  $k \to 0$  limit is well approximated.

Technically, however, there is a limit to the usefulness of the lattice approach in its application to wave scattering problems. One potential complication has already been mentioned, namely that the internal degrees of freedom at each lattice site, e.g., the atomic energy levels, may intermix. If there are inhomogeneities which have many internal degrees of freedom and which intermix with the lattice effect, then the simplicity of the lattice model would be lost. Another technical complication can arise when the inhomogeneities are *correlated* from site to site. Thus, if the lattice approach is used to model light scattering from spherical particles whose diameters are on the order of the light wavelength, it is necessary to subdivide the interior of each particle into many lattice sites so that the wave behavior may be modeled correctly. However, in that case the inhomogeneous randomness must be correlated over the size of each sphere. Technically, correlation would present difficulty in a



Fig. 2.5. The electronic density of states for a 3D simple cubic lattice, normalized to per site per frequency unit  $\beta\hbar/ma^2$ . Here  $u = (\hbar\omega - \varepsilon_0)/(\beta\hbar^2/ma^2)$ . The area under the curve integrates to 1. The *solid curve* is given by the function  $-(1/\pi^3) \text{Im} \int_0^{\pi} d\theta \mu^+(u, \theta) \mathbf{K}[\mu^+(u, \theta)]$ , where  $\mu^+(u, \theta) = 2/(|u + \cos \theta| + i\eta)$ .



Fig. 2.6. The phonon density of states for a 3D simple cubic lattice, normalized to per site per frequency unit  $\sqrt{K/m}$ . Here  $u = \omega/\sqrt{K/m}$ , and  $\beta = 1$ . The integrated area under the curve is 1. The upper band edge is at  $\omega = \sqrt{12}\sqrt{K/m}$ . The solid curve here is given by  $-(u/\pi^3) \text{Im} \int_0^{\pi} d\theta \mu^+(u,\theta) \mathbf{K}[\mu^+(u,\theta)]$ , where  $\mu^+(u,\theta) = 4/(|u^2 - 6 + 2\cos\theta| + i\eta)$ .

lattice model, which is best suited to treat uncorrelated randomness that can vary independently from site to site. Even numerically, 3D problems with correlated randomness are still technically challenging.

However, the viewpoint – that the waves in continuum may be modeled by waves in a correlated lattice – is useful in illuminating one source of difference

between the scattering and localization behavior of classical waves and that of an electron on a lattice: The existence of resonances for scatterers in continuum and the inert, point-particle nature of the single-site scattering impurity in the lattice case. The equivalence between the continuum model and the correlated lattice model tells us that the scatterer resonances may be viewed as a consequence of correlation between the scattering sites on a lattice.

In what follows, the quantum case, pertaining to electrons in solids, will be treated by the lattice model, since this is fairly realistic. For classical waves the scalar wave equation with continuous spatial variables will be used as our model. Although the basic Green function formalism is common to the two cases, actual implementation can involve different approaches. In the case of short-wavelength phonons, it is already seen that the problem is qualitatively identical to that of the electron. Therefore, it will not be covered separately.

# 2.6 Problems and Solutions

## Problem 2.1 Show that

$$\lim_{\eta \to 0} \frac{1}{x \pm i\eta} = P\left(\frac{1}{x}\right) \mp i\pi\delta(x).$$

**Solution** By multiplying  $x \mp i\eta$  to both the numerator and denominator, one gets

$$\frac{1}{x \pm i\eta} = \frac{x \mp i\eta}{x^2 + \eta^2} = \frac{x}{x^2 + \eta^2} \mp i\frac{\eta}{x^2 + \eta^2}.$$
 (P2.1)

For the real part, taking the principal value means that in performing the integral over x one should first do the integral with a finite  $\eta$ , and then let  $\eta$  approach zero. That is,

$$P \int \frac{\mathrm{d}x}{x} = \lim_{\eta \to 0} \int \frac{x}{x^2 + \eta^2} \mathrm{d}x.$$
 (P2.2)

It is clear that if the integration limits are symmetrical about x = 0, the result would be zero since  $x/(x^2 + \eta^2)$  is an odd function of x. When the limits are not symmetrical about 0, the result would depend on the degree of asymmetry but would never diverge.

For the imaginary part, the integral over x is given by

$$\int_{-\infty}^{\infty} \frac{\eta}{x^2 + \eta^2} dx = \int_{-\pi/2}^{\pi/2} \frac{\eta^2 \sec^2 \theta d\theta}{\eta^2 (1 + \tan^2 \theta)} = \pi,$$
 (P2.3)

where the change of variable  $x = \eta \tan \theta$  is used. The final answer is noted to be independent of  $\eta$ . Therefore, when  $\eta$  approaches zero,  $\eta/(x^2 + \eta^2)$  becomes a narrower and taller Lorentzian, but its area,  $\pi$ , remains constant. In the

limit of  $\eta \to 0$ , That Lorenzian becomes a delta function, with a factor  $\pi$  to account for its integral value.

It should be noted here that a delta function always has the dimension that is the inverse of its argument's dimension. For example,  $\delta(k)$  has the dimension of length because k has the dimension of  $[length]^{-1}$ . The reason is that

$$\int_{-\infty}^{\infty} \delta(k) \mathrm{d}k = 1. \tag{P2.4}$$

Since the right-hand side is dimensionless,  $\delta(k)$  must have the dimension of length in order to cancel that of k.

**Problem 2.2** Derive the continuum density of states in 1D, 2D, and 3D uniform media for both the quantum wave and the classical scalar wave.

**Solution** Starting from (2.24), the main task is to evaluate the integral

$$\int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^d} \delta[\kappa_0^2(\omega) - k^2].$$

In 1D, we have

$$\int_{-\infty}^{\infty} \frac{\mathrm{d}k}{2\pi} \delta[\kappa_0^2(\omega) - k^2] = \frac{2}{2\pi} \int_0^{\infty} \mathrm{d}(k^2) \cdot \frac{\mathrm{d}k}{\mathrm{d}(k^2)} \delta[\kappa_0^2(\omega) - k^2].$$
(P2.5)

The integral immediately gives  $[2\pi\kappa_0(\omega)]^{-1}$ . In 2D, the integral may be written as

$$\frac{1}{4\pi} \int_0^\infty \mathrm{d}(k^2) \delta[\kappa_0^2(\omega) - k^2] = (4\pi)^{-1}.$$
 (P2.6)

In 3D, the integral is

$$\frac{1}{8\pi^3} \int_0^\infty 2\pi k \mathrm{d}(k^2) \delta[\kappa_0^2(\omega) - k^2] = \frac{\kappa_0(\omega)}{4\pi^2}.$$
 (P2.7)

By substituting these expressions into (2.24) and using (2.25) plus the respective  $\kappa_0(\omega)$  expressions for the quantum and the classical cases, (2.26) and (2.27) are obtained.

**Problem 2.3** Derive a general expression of (local) density of states for the scalar wave in an inhomogeneous medium. Express the diagonal Green function in terms of the density of states.

**Solution** The Green function for the general scalar wave equation is given by (2.16):

$$(\nabla^2 - \sigma(\mathbf{r}) + \kappa_0^2) G(\omega, \mathbf{r}, \mathbf{r}') = \delta(\mathbf{r} - \mathbf{r}').$$
(P2.8)

Although this equation is not solvable in general, one can nevertheless formally define the eigenfunctions of the operator  $\nabla^2 - \sigma(\mathbf{r})$  as  $\phi_n(\mathbf{r})$ , with eigenvalues  $-d_n$ :

$$\left[\nabla^2 - \sigma(\mathbf{r})\right]\phi_n(\mathbf{r}) = -d_n\,\phi_n(\mathbf{r}).\tag{P2.9}$$

Here n is the index for the eigenfunctions and their corresponding eigenvalues. It is noted that for the operator  $\nabla^2$ ,  $\phi_n(\mathbf{r})$  is just  $\exp(\mathbf{i}\mathbf{k}\cdot\mathbf{r})$ , with the eigenvalue  $-k^2$ . A formal solution of (P2.8) may be obtained by first expressing  $G(\omega, \mathbf{r}, \mathbf{r}')$  and  $\delta(\mathbf{r} - \mathbf{r}')$  as

$$G(\omega, \mathbf{r}, \mathbf{r}') = \frac{1}{N} \sum_{n} G(\omega, d_n) \phi_n(\mathbf{r}) \phi_n^*(\mathbf{r}'), \qquad (P2.10)$$

$$\delta(\mathbf{r} - \mathbf{r}') = (1/L^d) \sum_n \phi_n(\mathbf{r}) \phi_n^*(\mathbf{r}'), \qquad (P2.11)$$

where the summation over n is seen to correspond to summation over  $\mathbf{k}$  in the free-space case, i.e.,

$$\sum_{\mathbf{k}} = L^d \int \frac{\mathrm{d}k}{(2\pi)^d}.$$

Equation (P2.11) is the completeness condition of the eigenfunctions, guaranteed for the operator  $\nabla^2 - \sigma(\mathbf{r})$  if  $\sigma(\mathbf{r})$  is real. By substituting (P2.10) and (P2.11) into (P2.8) and equating the left-hand and right-hand sides for each component n, it is seen that

$$\frac{L^d}{N}G^{\pm}(\omega, d_n) = \lim_{\eta \to 0} \frac{1}{\kappa_0^2(\omega) - d_n \pm i\eta} 
= P \frac{1}{\kappa_0^2(\omega) - d_n} \mp i\pi \delta[\kappa_0^2(\omega) - d_n]. \quad (P2.12)$$

The imaginary part of  $G^+(\omega, \mathbf{r} = \mathbf{r}')$  is given by

$$\frac{L^d}{N} \operatorname{Im} G^+(\omega, \mathbf{r} = \mathbf{r}') = -\frac{\pi}{N} \sum_n \delta[\kappa_0^2(\omega) - d_n] |\phi_n(\mathbf{r})|^2.$$
(P2.13)

If one defines a local density of states as

$$\rho(\omega, \mathbf{r}) = \frac{\mathrm{d}\kappa_0^2(\omega)}{\mathrm{d}\omega} \frac{1}{L^d} \sum_n \delta[\kappa_0^2(\omega) - d_n] |\phi_n(\mathbf{r})|^2, \qquad (P2.14a)$$

then

$$\rho(\omega, \mathbf{r}) = -\frac{1}{\pi} \frac{\mathrm{d}\kappa_0^2(\omega)}{\mathrm{d}\omega} \operatorname{Im} G^+(\omega, \mathbf{r} = \mathbf{r}').$$
(P2.14b)

The volume-averaged density of states is then

$$\bar{\rho}(\omega) = \frac{1}{L^d} \int \rho(\omega, \mathbf{r}) \, \mathrm{d}\mathbf{r}$$

$$= \frac{\mathrm{d}\kappa_0^2(\omega)}{\mathrm{d}\omega} \frac{1}{L^d} \sum_n \delta[\kappa_0^2(\omega) - d_n] \frac{1}{L^d} \int |\phi_n(\mathbf{r})|^2 \mathrm{d}\mathbf{r}$$

$$= \frac{\mathrm{d}\kappa_0^2(\omega)}{\mathrm{d}\omega} \frac{1}{L^d} \sum_n \delta[\kappa_0^2(\omega) - d_n]. \qquad (P2.15)$$

For a uniform medium,  $|\phi_n(\mathbf{r})|^2 = |\mathbf{e}^{\mathbf{i}\mathbf{k}\cdot\mathbf{r}}|^2 = 1$ , so there is no distinction between the local density of states and the volume-averaged density of states. We also note here that from (P2.12) and (P2.10),

$$G^+(\omega, \mathbf{r}, \mathbf{r}') = [G^-(\omega, \mathbf{r}', \mathbf{r})]^*,$$

where \* means complex conjugation.

There is an important and general relation between the local density of states and the diagonal component of the Green function which can easily be shown as follows. From (P2.10), (P2.12), and (P2.14a), we have

$$G^{+}(\omega, \mathbf{r} = \mathbf{r}') = \frac{1}{L^{d}} \sum_{n} \frac{|\phi_{n}(\mathbf{r})|^{2}}{\kappa_{0}^{2}(\omega) - d_{n} + \mathrm{i}\eta}$$
$$= \frac{1}{L^{d}} \sum_{n} \int \mathrm{d}\kappa_{1}^{2} \frac{\delta[\kappa_{1}^{2} - d_{n}]|\phi_{n}(\mathbf{r})|^{2}}{\kappa_{0}^{2}(\omega) - \kappa_{1}^{2} + \mathrm{i}\eta}$$
$$= \int \mathrm{d}\omega' \frac{\rho(\omega', \mathbf{r})}{\kappa_{0}^{2}(\omega) - \kappa_{0}^{2}(\omega') + \mathrm{i}\eta},$$
(P2.16)

where we have set  $\kappa_1^2 = \kappa_0^2(\omega')$ . The real part of  $G^+$  may be obtained by taking the principal part of the integral. Equation (P2.16) is also known as the Kramers–Kronig relation. Its physical basis lies in the causal nature of the retarded Green function; i.e., response always follows the excitation. Translated into the frequency domain, causality means  $G^+$  is always analytic in the upper half of the complex  $\omega$  plane. This is clear from (P2.12), where the poles for  $G^+(\omega, d_n)$  occur at the lower half of the complex  $\omega$  plane. The relation (P2.16) is valid for both the continuum and the lattice Green functions.

**Problem 2.4** Derive an expression for the density of states on a 2D square lattice.

**Solution** From (2.37), (2.57), and (2.60), one can write for the 2D square lattice

$$G_{0}^{\pm}(\omega, \mathbf{l} = \mathbf{l}') = \frac{L^{2}}{N} \iint_{-\pi/a}^{\pi/a} \frac{\mathrm{d}k_{x} \,\mathrm{d}k_{y}}{(2\pi)^{2}} \\ \times \frac{1}{\kappa_{0}^{2} - (2\beta/a^{2})[(1 - \cos k_{x}a) + (1 - \cos k_{y}a)] \pm \mathrm{i}\eta} \frac{1}{a^{2}} \\ = \int_{-\pi}^{\pi} \frac{\mathrm{d}\theta_{x}}{2\pi} \int_{-\pi}^{\pi} \frac{\mathrm{d}\theta_{y}}{2\pi} \frac{1}{(\kappa_{0}^{2}a^{2} - 4\beta) + 2\beta(\cos \theta_{x} + \cos \theta_{y})}.$$
(P2.17)

In the last expression  $\theta_{x,y} = k_{x,y}a$ , and  $\pm i\eta$  is absorbed into  $\kappa_0^2$ . By a change of variables

$$\theta_1 = \frac{\theta_x + \theta_y}{2}, \qquad \theta_2 = \frac{\theta_x - \theta_y}{2},$$

one can write

$$G_0^{\pm}(\omega, \mathbf{l} = \mathbf{l}') = \frac{1}{\pi^2} \int_0^{\pi} \mathrm{d}\theta_1 \int_0^{\pi-\theta_1} \mathrm{d}\theta_2 \left| \frac{\partial(\theta_x, \theta_y)}{\partial(\theta_1, \theta_2)} \right| \\ \times \frac{1}{(\kappa_0^2 a^2 - 4\beta) + 4\beta \cos \theta_1 \cos \theta_2}.$$
(P2.18)

In obtaining (P2.18) the identity

$$\cos \theta_x + \cos \theta_y = \cos(\theta_1 + \theta_2) + \cos(\theta_1 - \theta_2) = 2\cos \theta_1 \cos \theta_2$$

has been utilized. The notation  $|\partial(\theta_x, \theta_y)/\partial(\theta_1, \theta_2)|$  denotes the Jacobian of transformation, and in this case it has the value 2. Since the integrand is an even function of  $\theta_1$  and  $\theta_2$ , it is written as an integral over only the first quadrant, with a factor 4 multiplied. The limits of integration come from the fact that the domain of integration is modified under the transformation as shown in Fig. P2.1. The shaded region denoted 1 is the region of integration in (P2.18). It is easy to show that the integral over the region denoted 2 is identical to the integral over region 1 under a transformation of variables  $\theta'_1 = \pi - \theta_1$  and  $\theta'_2 = \pi - \theta_2$ . Therefore, (P2.18) can be rewritten as

$$G_0^{\pm}(\omega, \mathbf{l} = \mathbf{l}') = \frac{1}{\pi} \int_0^{\pi} \mathrm{d}\theta_1 \left[ \frac{1}{2\pi} \int_{-\pi}^{\pi} \mathrm{d}\theta_2 \frac{1}{(\kappa_0^2 a^2 - 4\beta) + 4\beta \cos \theta_1 \cos \theta_2} \right].$$
(P2.19)

The integral inside the bracket (denoted below by the square brackets) may be transformed into a contour integral just as shown in the 1D case, (2.61) and (2.62). The result is, for  $|\kappa_0^2 a^2 - 4\beta| > 4\beta \cos \theta_1$ ,

$$\begin{bmatrix} = \frac{\operatorname{sgn}(\kappa_0^2 a^2 - 4\beta)}{[(\kappa_0^2 a^2 - 4\beta)^2 - (4\beta)^2 \cos^2 \theta_1]^{1/2}} \\ = \frac{1}{(\kappa_0^2 a^2 - 4\beta)(1 - h^2 \cos^2 \theta_1)^{1/2}}, \quad (P2.20)$$



Fig. P2.1. The integration domain for the variables  $\theta_x$ ,  $\theta_y$  is a square  $[-\pi, \pi] \times [-\pi, \pi]$ . The domain after the transformation into  $\theta_1$  and  $\theta_2$  is a smaller tilted square. The 1 and 2 denote the two triangular regions in the first quadrant referred to in the text.

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where

$$h = \frac{4\beta}{\kappa_0^2 a^2 - 4\beta}.\tag{P2.21}$$

On the other hand, if  $|\kappa_0^2 a^2 - 4\beta| < 4\beta \cos \theta_1$ , then

$$[] = \frac{\mp i}{|\kappa_0^2 a^2 - 4\beta| (h^2 \cos^2 \theta_1 - 1)^{1/2}}.$$
 (P2.22)

The real part of (P2.19) is therefore

$$G_0^{\pm}(\omega, \mathbf{l} = \mathbf{l}') = \frac{2}{\pi(\kappa_0^2 a^2 - 4\beta)} \int_{u(h)}^{\pi/2} \frac{\mathrm{d}\theta}{(1 - h^2 \cos^2 \theta)^{1/2}},$$
 (P2.23)

with the condition that  $|h \cos \theta| < 1$ , where u(h) = 0 if  $|h| \le 1$ , and  $u(h) = \cos^{-1}(1/|h|)$  if |h| > 1. Alternatively, the imaginary part of the Green function is given by

$$G_0^{\pm}(\omega, \mathbf{l} = \mathbf{l}') = \frac{\mp 2i}{\pi |\kappa_0^2 a^2 - 4\beta|} \int_0^{u(h)} \cdot \frac{\mathrm{d}\theta}{(h^2 \cos^2 \theta - 1)^{1/2}},$$
 (P2.24)

with the conditions that |h| > 1 and  $|h \cos \theta| > 1$ .

To evaluate (P2.23) and (P2.24) explicitly, the easiest case is  $|h| \leq 1,$  which yields

$$G_0(\omega, \mathbf{l} = \mathbf{l}') = \frac{2}{\pi(\kappa_0^2 a^2 - 4\beta)} \mathbf{K} \left(\frac{4\beta}{\kappa_0^2 a^2 - 4\beta}\right)$$
(P2.25)

for  $|\kappa_0^2 a^2 - 4\beta| > 4\beta$ , i.e., outside the energy band. The symbol **K** stands for the complete elliptic integral of the first kind (Arfken 1970), defined by (P2.34). In this case the Green function is completely real, and the density of states is therefore zero. If  $|h| \ge 1$ , i.e., inside the energy band, the real part of  $G_0(\omega, \mathbf{l} = \mathbf{l}')$  can be written from (P2.23) as

$$\operatorname{Re}[G_{0}(\omega, \mathbf{l} = \mathbf{l}')] = \frac{2}{\pi(\kappa_{0}^{2}a^{2} - 4\beta)} \int_{\cos^{-1}(1/|h|)}^{\pi/2} \frac{\mathrm{d}\theta}{(1 - h^{2}\cos^{2}\theta)^{1/2}} = \frac{\operatorname{sgn}(\kappa_{0}^{2}a^{2} - 4\beta)}{2\pi\beta} \int_{0}^{\pi/2} \frac{\mathrm{d}\phi}{(1 - h^{-2}\cos^{2}\phi)^{1/2}} = \frac{\operatorname{sgn}(\kappa_{0}^{2}a^{2} - 4\beta)}{2\pi\beta} \mathbf{K}\left(\frac{\kappa_{0}^{2}a^{2} - 4\beta}{4\beta}\right), \quad (P2.26)$$

for  $|\kappa_0^2 a^2 - 4\beta| \leq 4\beta$ . The change of variable that accomplished the task in (P2.26) is  $\cos \phi = |h| \cos \theta$ . For the imaginary part inside the energy band, one starts with (P2.24), from which one obtains

$$\operatorname{Im}[G_{0}^{\pm}(\omega, \mathbf{l} = \mathbf{l}')] = \mp \frac{1}{2\pi\beta} \int_{0}^{\cos^{-1}(1/|h|)} \frac{\mathrm{d}\theta}{(\cos^{2}\theta - h^{-2})^{1/2}}$$
$$= \mp \frac{1}{2\pi\beta} \int_{0}^{\pi/2} \frac{\mathrm{d}\phi}{[1 - (1 - h^{-2})\sin^{2}\phi]^{1/2}}$$
$$= \mp \frac{1}{2\pi\beta} \mathbf{K} \left[ \sqrt{1 - \left(\frac{\kappa_{0}^{2}a^{2} - 4\beta}{4\beta}\right)^{2}} \right]$$
(P2.27)

for  $|\kappa_0^2 a^2 - 4\beta| \leq 4\beta$ . The transformation that does the trick in (P2.27) is  $\sin \theta = (1 - h^{-2})^{1/2} \sin \phi$ . By a further transformation  $\phi' = \pi/2 - \phi$ , one obtains the last line of (P2.27). The density of states is given by

$$\rho_{0}(\omega) = -\frac{1}{\pi} \frac{\mathrm{d}\kappa_{0}^{2}(\omega)}{\mathrm{d}\omega} \mathrm{Im}[G^{+}(\omega, \mathbf{l} = \mathbf{l}')]$$
$$= \frac{1}{2\pi^{2}\beta} \frac{\mathrm{d}\kappa_{0}^{2}(\omega)}{\mathrm{d}\omega} \mathbf{K} \left[ \sqrt{1 - \left(\frac{\kappa_{0}^{2}a^{2} - 4\beta}{4\beta}\right)^{2}} \right].$$
(P2.28)

for  $|\kappa_0^2 a^2 - 4\beta| < 4\beta$ , and zero otherwise.

**Problem 2.5** Derive an expression for the diagonal elements of the Green function on a 3D simple cubic lattice.

Solution The diagonal elements of Green function are defined as those with  $\mathbf{l}=\mathbf{l}',$  or

$$G_{0}^{\pm}(\omega, \mathbf{l} = \mathbf{l}') = \int_{-\pi/a}^{\pi/a} \int_{-\pi/a}^{\pi/a} \int_{-\pi/a}^{\pi/a} \frac{\mathrm{d}k_{x}\mathrm{d}k_{y}\mathrm{d}k_{z}}{(2\pi)^{3}} \\ \times \frac{1}{\kappa_{0}^{2} - (2\beta/a^{2})(3 - \cos k_{x}a - \cos k_{y}a - \cos k_{z}a) \pm \mathrm{i}\eta} \\ = \frac{1}{2\pi a} \int_{-\pi}^{\pi} \mathrm{d}\theta_{z} \left[ \int_{-\pi}^{\pi} \frac{\mathrm{d}\theta_{y}}{2\pi} \int_{-\pi}^{\pi} \frac{\mathrm{d}\theta_{x}}{2\pi} \\ \times \frac{1}{(\kappa_{0}^{2}a^{2} - 6\beta + 2\beta\cos\theta_{z}) + 2\beta(\cos\theta_{x} + \cos\theta_{y})} \right], \quad (P2.29)$$

where  $\pm i\eta$  is absorbed into  $\kappa_0^2$ . From the previous problem, the double integral inside the square brackets can be performed just as in the 2D square lattice case, with the result that

$$G_0(\omega, \mathbf{l} = \mathbf{l}') = \frac{1}{2\pi^2 a\beta} \int_0^{\pi} \mathrm{d}\theta \,\mu(\theta) \,\mathbf{K}[\mu(\theta)], \qquad (P2.30)$$

with

$$\mu(\theta) = \frac{4\beta}{\kappa_0^2 a^2 - 6\beta + 2\beta \cos\theta}$$
(P2.31)

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when the integrand is real. When the integrand is imaginary, one has to remember the  $\pm i\eta$  in  $\kappa_0^2$ , and moreover, the Green function is now given by

$$G_0^{\pm}(\omega, \mathbf{l} = \mathbf{l}') = \frac{1}{2\pi^2 a\beta} \int_0^{\pi} \mathrm{d}\theta \mu^{\pm}(\theta) \mathbf{K}[\mu^{\pm}(\theta)], \qquad (P2.32)$$

where  $\mu^{\pm}(\theta)$  is given by

$$\mu^{\pm}(\theta) = \frac{4\beta}{|\kappa_0^2 a^2 - 6\beta + 2\beta \cos \theta| \pm i\eta}.$$
 (P2.33)

In all cases  $\mathbf{K}(\mu)$  stands for

$$\mathbf{K}(\mu) = \int_0^{\pi/2} \frac{\mathrm{d}\phi}{(1 - \mu^2 \cos^2 \phi)^{1/2}}.$$
 (P2.34)

# Wave Scattering and the Coherent Potential Approximation

## 3.1 An Overview of the Approach

This chapter begins the consideration of waves in random media. Before delving into more detailed considerations, however, a discussion of the overall approach would be helpful in pointing out the rationale of the subsequent development.

In a random medium, a complete solution of the wave equation is represented by the knowledge of  $G(\omega, \mathbf{r}, \mathbf{r}')$  for all values of  $\omega, \mathbf{r}, \mathbf{r}'$  in the presence of disorder  $\sigma(\mathbf{r})$ . Complications arise because the accurate solution of the wave equation is not generally possible in the presence of  $\sigma(\mathbf{r})$ . Moreover, there is also the problem of how to extract the desired information from  $G(\omega, \mathbf{r}, \mathbf{r}')$ even if it were known. Here a comparison with the classical random walk would be illuminating. Suppose the position of a random walker is known at every instant of time, given by  $\mathbf{r}(t)$ . This is analogous to the knowledge of the Green function for the wave equation and represents the complete solution of the random walk problem in principle. But the diffusive behavior, which results from the statistical character of  $\mathbf{r}(t)$ , is not directly evident from  $\mathbf{r}(t)$ . What is necessary to bring out the diffusive behavior is the evaluation of the moments of  $\mathbf{r}(t) - \langle \mathbf{r} \rangle_c$ ,  $\langle \mathbf{r}^2 \rangle_c$ , etc. – where the angular brackets with the subscript c denote averaging over different configurations (of the random perturbations  $\sigma(\mathbf{r})$  in the wave scattering case, and different random walk trajectories in the random walker case). Similarly, for the wave problem the first objective of our approach is the approximate evaluation of  $\langle G \rangle_c$ . Since  $\langle G \rangle_c$  contains much less detailed information than G, its calculation is simpler than that for G. For example, whereas G depends on the source position  $\mathbf{r}'$ ,  $\langle G \rangle_c$  depends only on the source–detector separation  $\mathbf{r} - \mathbf{r}'$  because after configurational averaging,  $\langle G \rangle_c$  can no longer depend on any particular  $\sigma(\mathbf{r})$  (since an averaged quantity cannot depend on the averaged variable, and in this case independence from  $\sigma(\mathbf{r})$  means all spatial positions are equivalent), and only  $\mathbf{r} - \mathbf{r}'$  plus the statistical properties of the  $\sigma(\mathbf{r})$  ensemble are relevant.

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#### 3.1.1 Information from the First Moment

The knowledge of  $\langle G \rangle_c$  can tell us the following. First, it gives the wave propagation characteristics in an averaged sense. This is what is meant by the "effective medium" as seen by a wave. Second, it gives the spatial scale beyond which this effective medium description can no longer be valid. Mathematically, this second point is manifest in the exponential decay of  $\langle G \rangle_c$  as a function of  $|\mathbf{r} - \mathbf{r}'|$ . However, it should be emphasized that the decay here does not indicate wave localization; just as in random walk  $\langle \mathbf{r} \rangle_c = 0$  does not imply that the random walker is localized at the origin. Rather, the exponential decay of  $\langle G \rangle_c$  means that the wave *coherence*, in the sense of a unique wave propagation direction and phase relation, is lost. This decay length is defined as the *mean free path*. For the random walk counterpart,  $\langle \mathbf{r} \rangle_c = 0$  implies the absence of ballistic motion.

# 3.1.2 Information from the Second Moment

Whereas  $\langle G \rangle_c$  gives us the character of the coherent part of the wave propagation in a random medium, the transport dynamics beyond the scale of mean free path are contained in the second moment,  $\langle GG^* \rangle_c$ , just as in the random walk case the diffusion dynamics emerge from the time dependence of  $\langle \mathbf{r}^2 \rangle_c$ . Therefore the next objective in our approach is the evaluation of  $\langle GG^* \rangle_c$ . The outcome of this calculation will show that at the long-time, long-propagationdistance limit, intensity transport is indeed diffusive in character, and the attendant diffusion constant may be explicitly calculated. It is in the context of diffusive transport that the coherent backscattering effect may be demonstrated as a correction to the diffusion constant. Our exposition will give a global view of the possible consequences of coherent backscattering through a description of the scaling theory of localization, which is then substantiated by showing consistency between the assumption of the scaling theory and the calculations involving  $\langle GG^* \rangle_c$ .

## 3.1.3 Information from Higher Moments

Since the intensity transport behavior is given by  $\langle GG^* \rangle_c$ , fluctuations in the intensity transport, which are observable in mesoscopic samples, have to be calculated from  $\langle GG^*GG^* \rangle_c$ . The fourth moment represents the intensity– intensity correlation and contains information concerning the long-range memory and phase interference effects that may be retrieved from intensity (or current) fluctuations, e.g., speckle patterns formed by light after passing through a random medium. Together, these moments of G are indicative of the special statistical character of wave transport in random media. Their evaluation will be the main technical task for this and subsequent chapters.

Wave transport characteristics obtained from the moments of G necessarily differ from those in a single configuration. Nevertheless, these moments can give a general sense of the single-configuration characteristics because in most physical situations the *ergodic hypothesis* is valid, which means that the configurationally averaged behavior may be equated to the infinite-time average of the single-configuration behavior. However, the detailed spatial and temporal variations of the wave field in a single configuration can never be retrieved from the moments of G. For such information there is no substitute for an exact solution to the problem.

# 3.2 Wave Scattering Formalism

A useful property of the Green function is that it can solve inhomogeneous differential equations of the form

$$(\nabla^2 + \kappa_0^2)\phi(\omega, \mathbf{r}) = f(\mathbf{r}), \qquad (3.1)$$

where  $f(\mathbf{r})$  is an inhomogeneous source term with arbitrary spatial dependence. For example, if instead of  $\delta(t)\delta(\mathbf{r} - \mathbf{r}')$  the excitation source is of the form  $\delta(t)f(\mathbf{r})$ , then (3.1) would result. We would like to show that

$$\phi(\omega, \mathbf{r}) = \phi_0(\omega, \mathbf{r}) + \int G_0(\omega, \mathbf{r} - \mathbf{r}') f(\mathbf{r}') d\mathbf{r}'$$
(3.2)

is the solution to (3.1). Here  $\phi_0$  represents the solution to the homogeneous equation [i.e., right-hand side of (3.1) equals zero]. By substituting (3.2) into (3.1), it is seen that since the operation of  $(\nabla^2 + \kappa_0^2)$  is on the **r** variable of  $G(\omega, \mathbf{r}, \mathbf{r}')$ , it can be performed inside the integral. This results in

$$\int \delta(\mathbf{r} - \mathbf{r}') f(\mathbf{r}') d\mathbf{r}' = f(\mathbf{r}), \qquad (3.3)$$

thus recovering the right-hand side of (3.1).

Let us now write the Green function equation in the presence of  $\sigma(\mathbf{r})$  as

$$(\nabla^2 + \kappa_0^2)G(\omega, \mathbf{r}, \mathbf{r}') = \delta(\mathbf{r} - \mathbf{r}') + \sigma(\mathbf{r})G(\omega, \mathbf{r}, \mathbf{r}').$$
(3.4)

By comparison with (3.1) and (3.2), one can write down the formal solution to (3.4) as

$$G(\omega, \mathbf{r}, \mathbf{r}') = \int d\mathbf{r}_1 G_0(\omega, \mathbf{r} - \mathbf{r}_1) [\delta(\mathbf{r}_1 - \mathbf{r}') + \sigma(\mathbf{r}_1) G(\omega, \mathbf{r}_1, \mathbf{r}')]$$
  
=  $G_0(\omega, \mathbf{r} - \mathbf{r}') + \int d\mathbf{r}_1 G_0(\omega, \mathbf{r} - \mathbf{r}_1) \sigma(\mathbf{r}_1) G(\omega, \mathbf{r}_1, \mathbf{r}'), \quad (3.5)$ 

where the homogeneous term  $\phi_0$  is eliminated by the boundary condition that G must vanish as  $|\mathbf{r} - \mathbf{r}'| \to \infty$ .

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#### 3.2.1 The Operator Notation

At this point it would be convenient to introduce the operator notation to rewrite (3.5). That is, if the space is regarded as finely discretized so that every location  $\mathbf{r}$  is associated with an index number and a small volume  $(\delta r)^d$ , where  $\delta r$  is the linear size of the discretized unit, e.g., the atomic unit cell, then the Green function  $G(\omega, \mathbf{r}, \mathbf{r}')$  has two indices, one associated with  $\mathbf{r}$ and the other one associated with  $\mathbf{r}'$ ; i.e., G may be regarded as a matrix. Similarly,  $G_0(\omega, \mathbf{r} - \mathbf{r}_1)$  is also a matrix, and  $\sigma(\mathbf{r}_1)$  is a diagonal matrix. By writing the second term on the right-hand side of (3.5) as

$$\sum_{\mathbf{r}_1}\sum_{\mathbf{r}_2} (\delta r)^{2d} G_0(\omega, \mathbf{r} - \mathbf{r}_1) \sigma(\mathbf{r}_1) [\delta_{\mathbf{r}_1, \mathbf{r}_2} / (\delta r)^d] G(\omega, \mathbf{r}_2, \mathbf{r}'),$$

where the quantity in square brackets is a discrete representation of  $\delta(\mathbf{r}_1 - \mathbf{r}_2)$ , one may regard the double integral as a double summation. The whole term is thus equivalent to the multiplication of three matrices

## $\mathbf{G}_{0}\mathbf{VG},$

where the elements of the  $\mathbf{V}$  matrix are given by

**V** is called the impurity potential operator, and in our convention it has the dimension  $[length]^{d-2}$ . Alternatively, one can use Dirac's bra and ket notation as described in the last chapter to write the matrix multiplication as

$$\begin{split} \sum_{\mathbf{r}_1} \sum_{\mathbf{r}_2} \left< \mathbf{r} \right| \mathbf{G}_0 \left| \mathbf{r}_1 \right> \left< \mathbf{r}_1 \right| \mathbf{V} \left| \mathbf{r}_2 \right> \left< \mathbf{r}_2 \right| \mathbf{G} \left| \mathbf{r}' \right> . \end{split}$$
 Here  $\left< \mathbf{r} \left| \mathbf{G}_0 \right| \mathbf{r}_1 \right> = G_0(\omega, \mathbf{r} - \mathbf{r}_1), \left< \mathbf{r}_2 \left| \mathbf{G} \right| \mathbf{r}' \right> = G(\omega, \mathbf{r}_2, \mathbf{r}'), \text{ and} \\ \sum_{\mathbf{r}} \left| \mathbf{r} \right> \left< \mathbf{r} \right| = \mathbf{I}, \end{split}$ 

the identity matrix. This should be clear since  $|\mathbf{r}\rangle \langle \mathbf{r}|$  means a matrix with only one nonzero element (with value 1) at the diagonal position corresponding to row and column associated with  $\mathbf{r}$ . The summation of such matrices fills the diagonal elements of the resulting matrix with 1, which is the identity matrix. In this notation  $\langle \mathbf{r} | \mathbf{r}' \rangle = \delta_{\mathbf{r},\mathbf{r}'}$ , and  $\delta(\mathbf{r},\mathbf{r}') = \delta_{\mathbf{r},\mathbf{r}'}/(\delta r)^d$ . Either way, (3.5) may be expressed succinctly as

$$\mathbf{G} = \mathbf{G}_0 + \mathbf{G}_0 \mathbf{V} \mathbf{G},\tag{3.6}$$

where the symbols stand for operators (e.g., matrices), and their ordering is important and cannot be altered at will. In the operator notation, (3.6) is valid regardless of whether  $\mathbf{G}$ ,  $\mathbf{G}_0$ , and  $\mathbf{V}$  are in the  $\mathbf{r}$  domain representation or in the  $\mathbf{k}$  domain representation. The manipulation of the formalism thus simplifies.

#### 3.2.2 The T Operator

An alternative way to express (3.6) is obtained by iterating on G:

$$\mathbf{G} = \mathbf{G}_0 + \mathbf{G}_0 \mathbf{V} [\mathbf{G}_0 + \mathbf{G}_0 \mathbf{V} (\mathbf{G}_0 + \mathbf{G}_0 \mathbf{V} \cdots)]$$
  
=  $\mathbf{G}_0 + \mathbf{G}_0 \mathbf{V} \mathbf{G}_0 + \mathbf{G}_0 \mathbf{V} \mathbf{G}_0 \mathbf{V} \mathbf{G}_0 + \cdots$   
$$\mathbf{G} = \mathbf{G}_0 + \mathbf{G}_0 \mathbf{T} \mathbf{G}_0, \qquad (3.7)$$

or

$$\mathbf{T} = \mathbf{V} + \mathbf{V}\mathbf{G}_0\mathbf{V} + \mathbf{V}\mathbf{G}_0\mathbf{V}\mathbf{G}_0\mathbf{V} + \dots$$
  
=  $\mathbf{V}(\mathbf{I} - \mathbf{G}_0\mathbf{V})^{-1} = (\mathbf{I} - \mathbf{V}\mathbf{G}_0)^{-1}\mathbf{V}$  (3.8)

is called the **T** matrix, or the scattering matrix. In (3.8) it is seen that if each **V** represents one scattering, then **T** includes all the multiple scatterings. The formal summation of the operator series may be expressed as the inverse of  $(\mathbf{I} - \mathbf{G}_0 \mathbf{V})$  (just as in the summation of the power series). In the form of the inverses, (3.6) can be put in another popular form which may be derived as follows. From (3.6) we have

$$(\mathbf{I} - \mathbf{G}_0 \mathbf{V})\mathbf{G} = \mathbf{G}_0. \tag{3.9}$$

Taking the inverse of both sides and then right-multiplying by  $(I - G_0 V)$  give

$$\mathbf{G}^{-1} = \mathbf{G}_0^{-1}(\mathbf{I} - \mathbf{G}_0 \mathbf{V}),$$

or

$$\mathbf{G}^{-1} = \mathbf{G}_0^{-1} - \mathbf{V}. \tag{3.10}$$

Knowledge of  $\mathbf{T}$  is sufficient to completely solve the general wave equation

$$[\nabla^2 + \kappa_0^2 - \sigma(\mathbf{r})]\phi(\omega, \mathbf{r}) = 0$$

in terms of the uniform-medium solutions  $\phi_0(\omega, \mathbf{r})$  and  $G_0(\omega, \mathbf{r} - \mathbf{r}')$ . This can be seen by rewriting the equation as

$$(\nabla^2 + \kappa_0^2)\phi(\omega, \mathbf{r}) = \sigma(\mathbf{r})\phi(\omega, \mathbf{r}).$$

Then from (3.2) the solution may be written in terms of the bra and ket notation as

$$\begin{aligned} |\phi\rangle &= |\phi_0\rangle + \mathbf{G}_0 \mathbf{V} |\phi\rangle = |\phi_0\rangle + \mathbf{G}_0 \mathbf{V} |\phi_0\rangle + \mathbf{G}_0 \mathbf{V} \mathbf{G}_0 \mathbf{V} |\phi_0\rangle + \cdots \\ &= |\phi_0\rangle + \mathbf{G}_0^+ \mathbf{T}^+ |\phi_0\rangle . \end{aligned}$$
(3.11)

The + superscripts on  $\mathbf{G}_0$  and  $\mathbf{T}$  in the last line of (3.11) are meant to select the physical solution branch where the scattering from an inhomogeneity is represented by an outgoing wave (from the inhomogeneity) rather than by an incoming wave, which would be selected by  $\mathbf{G}_0^-\mathbf{T}^-$ . 50 3 Wave Scattering and the Coherent Potential Approximation

## 3.2.3 Configurational Averaging

From (3.7), the configurationally averaged Green function is given by

$$\langle \mathbf{G} \rangle_{c} = \mathbf{G}_{0} + \mathbf{G}_{0} \langle \mathbf{T} \rangle_{c} \mathbf{G}_{0},$$
 (3.12)

where  $\mathbf{G}_0$  is independent of  $\sigma(\mathbf{r})$  and is therefore not affected by the average. From the **T** representation of (3.8), we have

$$\langle \mathbf{T} \rangle_{c} = \left\langle \mathbf{V} (\mathbf{I} - \mathbf{G}_{0} \mathbf{V})^{-1} \right\rangle_{c}.$$
 (3.13)

It is noted that  $\langle \mathbf{T} \rangle_c$  contains all the higher-order correlations of  $\mathbf{V}$ , such as  $\langle \mathbf{VG}_0 \mathbf{V} \rangle_c$  and  $\langle \mathbf{VG}_0 \mathbf{VG}_0 \mathbf{V} \rangle_c$ , and all the multiple scatterings caused by  $\mathbf{V}$ . Since in the real space representation  $\langle \mathbf{G} \rangle_c$  depends only on the spatial separation between the source and the receiver  $\mathbf{r} - \mathbf{r'}$ , its Fourier transform is a function of one  $\mathbf{k}$  only, just as for  $\mathbf{G}_0$ . The general validity of this statement is shown in Problem 3.1 at the end of this chapter. In Problem 3.2, it is shown that a convolution integral in real space, as symbolized by  $\mathbf{G}_0 \langle \mathbf{T} \rangle_c \mathbf{G}_0$ , means simple multiplication of the transformed quantities in  $\mathbf{k}$  space. Therefore from (3.12) it is thus clear that in the  $\mathbf{k}$  representation,  $\langle \mathbf{T} \rangle_c$  is also a function of one  $\mathbf{k}$  only, which implies a dependence on only the separation  $\mathbf{r} - \mathbf{r'}$  in real space.

#### 3.2.4 The Self-Energy

In the same way as the derivation of (3.10), from (3.12) one can define a "self energy"  $\Sigma$  operator as

$$\langle \mathbf{G} \rangle_{\mathbf{c}}^{-1} = \mathbf{G}_0^{-1} - \Sigma \frac{L^d}{N}, \qquad (3.14)$$

In the **k** representation, since both  $\langle \mathbf{G} \rangle_{c}$  and  $\mathbf{G}_{0}$  are functions of one **k** only,  $\Sigma$  is also a function of one **k** only, which means a dependence on  $\mathbf{r} - \mathbf{r}'$  in real space, just as for  $\langle \mathbf{T} \rangle_{c}$ . From (3.12) and (3.14),  $\Sigma$  is related to  $\langle \mathbf{T} \rangle_{c}$  by

$$\Sigma = \frac{N}{L^d} \left\langle \mathbf{T} \right\rangle_{\rm c} \left( \mathbf{I} + \left\langle \mathbf{T} \right\rangle_{\rm c} \mathbf{G}_0 \right)^{-1}. \tag{3.15}$$

 $\Sigma$  is denoted the self-energy operator, and (3.14) is known as the *Dyson equa*tion. In the **k** representation, (3.14) and (3.15) are simple algebraic equations because all the relevant operators are diagonal matrices. Comparison of (3.14) with (3.10) shows that in spite of the apparent similarity, the self-energy  $\Sigma$ is a very different object from the operator **V** of the exact Green function in a fixed configuration. In real space, whereas the latter is just the (local) perturbation  $\sigma(\mathbf{r})$ ,  $\Sigma$  represents a nonlocal operator as can be seen from the equation satisfied by  $\langle G \rangle_c$ . Since in the **k**-domain we have

$$\left[\kappa_0^2(\omega) - k^2 - \Sigma(\omega, \mathbf{k})\right] \langle G \rangle_{\rm c}(\omega, \mathbf{k}) = \frac{N}{L^d},$$

it follows that in real space,  $-k^2$  means  $\nabla^2$ , and multiplication in the **k**-domain means convolution in real space (see solution to Problem 3.2), so that

$$\begin{aligned} & \left[\kappa_0^2(\omega) + \nabla^2\right] \left\langle G \right\rangle_{\rm c} \left(\omega, \mathbf{r} - \mathbf{r}'\right) - L^{-d} \int \Sigma(\omega, \mathbf{r} - \mathbf{r}_1) \left\langle G \right\rangle_{\rm c} \left(\omega, \mathbf{r}_1 - \mathbf{r}'\right) \mathrm{d}\mathbf{r}_1 \\ &= \delta(\mathbf{r} - \mathbf{r}'). \end{aligned}$$

This equation reduces to the ordinary wave equation only if  $\Sigma(\omega, \mathbf{r} - \mathbf{r}_1) = \Sigma'(\omega)L^d\delta(\mathbf{r} - \mathbf{r}_1)$ . This turns out to be possible when the effective medium description is valid, as will be seen later.

# 3.3 Single Scatterer: the Lattice Case

Consider a single scatterer embedded in a uniform lattice. The Green function of the system is given by

$$\mathbf{g} = \mathbf{G}_0 + \mathbf{G}_0 \mathbf{t} \mathbf{G}_0, \tag{3.16}$$

where the lowercase  $\mathbf{g}$  and  $\mathbf{t}$  are used to denote the Green function and the scattering matrix in the presence of a *single* scatterer. In this case, an explicit expression for  $\mathbf{t}$  can be obtained.

In the lattice model, a single scatterer may be expressed by a deviation  $\Delta \varepsilon$  of the site energy at  $l_0$ . In the bra and ket notation, the operator V is given by

$$\mathbf{V} = \mathbf{v} = \sigma_0 a^d \left| \mathbf{l}_0 \right\rangle \left\langle \mathbf{l}_0 \right|, \qquad (3.17)$$

$$\sigma_0 = \frac{2m\Delta\varepsilon}{\hbar^2},\tag{3.18}$$

where  $\mathbf{v}$  denotes the perturbation due to a single scatterer and a is the lattice constant. The operator  $\mathbf{v}$  may be thought of as a matrix with only one nonzero element, which is at the diagonal position of column  $\mathbf{l}_0$  and row  $\mathbf{l}_0$ . From (3.8), one gets

$$\mathbf{t} = (\sigma_0 a^d) |\mathbf{l}_0\rangle \langle \mathbf{l}_0| + (\sigma_0 a^d)^2 |\mathbf{l}_0\rangle \langle \mathbf{l}_0| \mathbf{G}_0 |\mathbf{l}_0\rangle \langle \mathbf{l}_0| + \cdots$$

$$= (\sigma_0 a^d) |\mathbf{l}_0\rangle \langle \mathbf{l}_0| \{1 + \sigma_0 a^d G_0(\omega, \mathbf{l} = \mathbf{l}') + [\sigma_0 a^d G_0(\omega, \mathbf{l} = \mathbf{l}')]^2 + \cdots \}$$

$$= |\mathbf{l}_0\rangle \langle \mathbf{l}_0| \frac{\sigma_0 a^d}{1 - \sigma_0 a^d G_0(\omega, \mathbf{l} = \mathbf{l}')}.$$
(3.19)

Knowledge of **t** explicitly solves the Green function  $g(\omega, \mathbf{l}, \mathbf{l}')$  for the single-scatterer case:

$$g(\omega, \mathbf{l}, \mathbf{l}') = G_0(\omega, \mathbf{l} - \mathbf{l}') + G_0(\omega, \mathbf{l} - \mathbf{l}_0)G_0(\omega, \mathbf{l}_0 - \mathbf{l}') \\ \times \frac{\sigma_0 a^d}{1 - \sigma_0 a^d G_0(\omega, \mathbf{l} = \mathbf{l}')}.$$
(3.20)

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From (3.11), the wave function for the one-scatterer case may also be similarly expressed as

$$\phi(\omega, \mathbf{l}) = \exp(\mathbf{i}\,\mathbf{k}_0 \cdot \mathbf{l}) + \left[\frac{a^2\sigma_0 \exp(\mathbf{i}\,\mathbf{k}_0 \cdot \mathbf{l}_0)}{1 - \sigma_0 a^d G_0^+(\omega, \mathbf{l} = \mathbf{l}')}\right] a^{d-2} G_0^+(\omega, \mathbf{l} - \mathbf{l}_0), \qquad (3.21)$$

where the notation  $\mathbf{k}_0$  is used to denote a wave vector which satisfies the relation  $\kappa_0^2(\omega) = \beta e(\mathbf{k}_0)$ . Equation (3.21) has the following physical interpretation. The difference between the uniform medium and the single-scatterer solution is represented by the second term on the right-hand side of (3.21). That term shows the single scatterer acting as a point source whose strength is given by the expression inside the square brackets. That strength is essentially the probability amplitude whereby the wave in the uniform medium is diverted into a wave emanating isotropically from the scatterer, i.e., the scattering probability.

#### 3.3.1 Impurity Bound State

It is interesting that the scattering probability amplitude has a denominator which may vanish when

$$\frac{1}{a^d \sigma_0} = G_0^+(\omega, \mathbf{l} = \mathbf{l}').$$

This condition can never be satisfied for  $\omega$  values inside the energy band because  $G_0^+(\omega, \mathbf{l} = \mathbf{l}')$  has a nonzero imaginary part which gives the density of states inside the band. Outside the band, however, the condition can indeed be satisfied.

The denominator of the scattering cross-section arises physically from the scattered wave that is multiply scattered back by the lattice. Since the same denominator is in the expression for g, (3.20), its zero(s) thus represents the existence of resonance mode(s) of the system. Such state(s) must have an amplitude that is localized around  $\mathbf{l}_0$ , because if one looks at (3.21), the spatial dependence of the new state is determined by  $G_0^+(\omega, \mathbf{l} - \mathbf{l}_0)$ , and since for  $\omega$  outside the energy band  $G_0^+$  must decay exponentially, as discussed in the last chapter, the new resonance mode created by the single "impurity" is thus a state bound to the neighborhood of the impurity site. Whereas in 1D and 2D the creation of such bound state(s) always accompanies the presence of an impurity (see Figs. 3.1 and 3.2, where the real part of the Green function is seen to diverge at the band edges), in 3D the strength of the impurity perturbation must exceed a critical threshold, i.e.,

$$|\sigma_0 a^2| \ge \frac{1}{0.2519} = 3.97,$$
 (3.22)

before a bound state can be generated (see Fig. 3.3, where it is seen that the value of  $|\text{Re}(aG_0)| < 0.2519$  in the band gap).

In the solution to Problem 3.3, an expression for the total scattering acrosssection O is derived; given by



Fig. 3.1. The scattering cross-section (the *solid line*) for an impurity in a 1D lattice. The impurity strength is  $\sigma_0 a^2 = 2$ . The real part of the lattice Green function  $G_0 a^{-1}$  is shown by the *dashed lines*, with  $\beta = 1$ . The *solid line* is given by the function  $(-4\text{Imy}(u))/(\bar{k}_0 a|1-2y(u)|^2)$ , where  $y(u) = 1/2\sqrt{u^2-1}$ , and  $\bar{k}_0$  is given by (P3.18) with d = 1. The *dashed lines* are given by sgn(u)y(u), for |u| > 1. Here  $u = (\hbar\omega - \varepsilon_0)/(\beta\hbar^2/ma^2)$ . The resonant state outside the band (at  $G_0 a^{-1} = 0.5$ ) is too far removed from the band edge for any appreciable effect on the scattering cross-section.

$$O = \frac{-(\sigma_0 a^d)^2 \text{Im} G_0^+(\omega, \mathbf{l} = \mathbf{l}')}{\bar{k}_0 |1 - \sigma_0 a^d G_0^+(\omega, \mathbf{l} = \mathbf{l}')|^2},$$
(3.23)

where the definition of  $\bar{k}_0$  and the relation between O and the imaginary part of  $\mathbf{t}^+$ , known as the optical theorem, are given in the same problem solution. In Figs. 3.1, 3.2, and 3.3, the real part of  $G_0^+(\omega, \mathbf{l} = \mathbf{l}')$  is shown for d = 1, 2, and 3, respectively, together with plots of the scattering cross-sections for fixed values of  $\sigma_0 a^2$ . An interesting observation is that although the resonant scattering cannot occur inside the band, the scattering cross-section can be affected if the resonant impurity state is close to the band edge. This is clearly seen in Figs. 3.2 and 3.3, where a resonant impurity state near the upper band edge enhances the scattering cross-section just inside the band edge.

# 3.4 Single Scatterer: the Continuum Case

In the classical scalar wave case, the single scatterer is taken to be a sphere of radius R (a circle in 2D, a line in 1D) with dielectric constant  $\varepsilon \neq 1$ . Then

In the bra and ket notation, the impurity operator  $\mathbf{v}$  may be expressed as

$$\mathbf{v} = \sum_{\mathbf{r}} (\Delta r)^d \sigma(\mathbf{r}) \left| \mathbf{r} \right\rangle \left\langle \mathbf{r} \right|.$$
(3.25)



Fig. 3.2. The scattering cross-section (in units of *a*) for an impurity in a 2D square lattice (shown by the *solid line*), The impurity strength is  $\sigma_0 a^2 = 3$ . The real part of the lattice Green function is shown by the *dashed line*, with  $\beta = 1$ . The *solid line* is given by the function  $(-9\text{Im}y(u))/(\bar{k}_0a|1-3y(u)|^2)$ , where  $\bar{k}_0$  is given by (P3.18) with d = 2 and  $y(u) = (1/2\pi) \left\{ \text{sgn}(u) \mathbf{K}(u/2) - i\mathbf{K} \left[ \sqrt{1-(u/2)^2} \right] \right\}$ . The dashed line is given by the function  $(1/\pi u) \mathbf{K}(2/u)$  for |u| > 2 and  $(\text{sgn}(u)/2\pi) \mathbf{K}(u/2)$  for |u| < 2. Here  $u = (\hbar\omega - \varepsilon_0)/(\beta\hbar^2/ma^2)$ , and  $\mathbf{K}(u)$  denotes the complete elliptic integral of the first kind. Since the resonant impurity state (at  $\text{Re}(G_0) = 1/3$ ) is close to the upper band edge, the scattering cross-section is seen to be pulled up in its vicinity.



Fig. 3.3. The scattering cross-section (in units if  $a^2$ ) for an impurity in a 3D simple cubic lattice (shown by the *solid line*). The impurity strength is  $\sigma_0 a^2 = 5$ . The real part of the lattice Green function is shown by the *dashed line*, with  $\beta = 1$ . The *solid line* is given by the function  $(-25 \text{Im}y(u))/(\bar{k}_0 a|1 - 5y(u)|^2)$ , where  $y = a\beta G_0^+(\omega, \mathbf{l} = \mathbf{l}')$ , with  $G_0^+$  given by (P2.30)–(P2.33). The *dashed line* is given by Re[y(u)]. Here  $u = (\hbar\omega - \varepsilon_0)/(\beta\hbar^2/ma^2)$ . The resonant impurity state (at  $\text{Re}(G_0 a) = 0.2$ ), is close to the upper band edge. As a result, the scattering cross-section is enhanced in its vicinity.

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In contrast to the lattice case, **v** is now a diagonal matrix with elements nonzero for all  $|\mathbf{r}| \leq R$ . The **t** operator is thus

$$\mathbf{t} = \sum_{\mathbf{r}} (\Delta r)^{d} \sigma(\mathbf{r}) |\mathbf{r}\rangle \langle \mathbf{r}| + \sum_{\mathbf{r}} \sum_{\mathbf{r}_{1}} (\Delta r)^{2d} \sigma(\mathbf{r}) \sigma(\mathbf{r}_{1}) |\mathbf{r}\rangle \langle \mathbf{r}| \mathbf{G}_{0} |\mathbf{r}_{1}\rangle \langle \mathbf{r}_{1}| + \cdots$$
(3.26)

Except for the first term, the rest of the series consists of full matrices, since each  $|\mathbf{r}\rangle \langle \mathbf{r}_1|$  represents a nondiagonal entry at column  $\mathbf{r}_1$  and row  $\mathbf{r}$ . The explicit form for  $\mathbf{t}$  is thus difficult to obtain by using the series or the inversion of the matrix  $(\mathbf{I} - \mathbf{G}_0 \mathbf{v})$ . However, it is observed that physically, the multiple scattering is caused by the boundary of the scatterer with the uniform medium. The net effect of summing the series in (3.26) must therefore be equivalent to the solution of a boundary value problem, which is easily manageable.

## 3.4.1 Wave Function in the Single Scatterer Case

With this idea in mind, we write the wave function for the one-scatterer case by using (3.11):

$$\phi(\omega, \mathbf{r}) = \exp(i\kappa_0 \cdot \mathbf{r}) + \sum_{\mathbf{r}_1} \sum_{\mathbf{r}_2} \langle \mathbf{r} | \mathbf{G}_0^+ | \mathbf{r}_1 \rangle \langle \mathbf{r}_1 | \mathbf{t}^+ | \mathbf{r}_2 \rangle \exp(i\kappa_0 \cdot \mathbf{r}_2), \quad (3.27)$$

where  $\kappa_0$  denotes a wave vector with magnitude  $\kappa_0$ . In 3D,  $G_0^+(\omega, \mathbf{r} - \mathbf{r}_1) = -\exp(i\kappa_0|\mathbf{r} - \mathbf{r}_1|)/4\pi|\mathbf{r} - \mathbf{r}_1|$  [(2.31)]. Since  $\langle \mathbf{r}_1 | \mathbf{t}^+ | \mathbf{r}_2 \rangle$  is nonzero only for  $r_1, r_2 < R$ , for  $r \gg R$ ,  $|\mathbf{r} - \mathbf{r}_1|$  may be expanded as

$$|\mathbf{r} - \mathbf{r}_1| = (r^2 + r_1^2 - 2rr_1 \cos \theta)^{1/2} \cong r(1 - r_1 \cos \theta/r) = r - r_1 \cos \theta,$$

where  $\theta$  is the angle between  $\mathbf{r}_1$  and  $\mathbf{r}$ . That means for (3.27),

$$\begin{aligned} \phi(\omega, \mathbf{r}) &\cong \exp(\mathrm{i}\kappa_0 \cdot \mathbf{r}) \\ &- \frac{\exp(\mathrm{i}\kappa_0 r)}{4\pi r} \sum_{\mathbf{r}_1} \sum_{\mathbf{r}_2} \exp(-\mathrm{i}\kappa'_0 \cdot \mathbf{r}_1) \langle \mathbf{r}_1 | \mathbf{t}^+ | \mathbf{r}_2 \rangle \exp(\mathrm{i}\kappa_0 \cdot \mathbf{r}_2) \\ &= \exp(\mathrm{i}\kappa_0 \cdot \mathbf{r}) - \frac{\exp(\mathrm{i}\kappa_0 r)}{4\pi r} \langle \kappa'_0 | \mathbf{t}^+ | \kappa_0 \rangle, \qquad |\mathbf{r}| \gg R, \end{aligned}$$
(3.28)

where  $\kappa'_0$  is noted to have a different direction from  $\kappa_0$  in general. Here we have made the identification

$$\langle \mathbf{r}_2 \mid \boldsymbol{\kappa}_0 \rangle = \exp(i\boldsymbol{\kappa}_0 \cdot \mathbf{r}_2),$$
 (3.29)

and similarly for  $\langle \kappa'_0 | \mathbf{r}_1 \rangle = \exp(-i\kappa'_0 \cdot \mathbf{r}_1)$ .

# 3.4.2 A Digression on Notation

A brief digression on the convention to be followed is helpful at this point. Under the conventions of this book, each  $|\mathbf{k}\rangle$  or  $\langle \mathbf{k}|$  has magnitude  $\sqrt{N}$ , so that  $\langle \mathbf{k}' | \mathbf{k} \rangle = N \delta_{\mathbf{k},\mathbf{k}'}$  because 56 3 Wave Scattering and the Coherent Potential Approximation

$$\begin{aligned} \langle \mathbf{k}' \mid \mathbf{k} \rangle &= \sum_{\mathbf{r}} \langle \mathbf{k}' \mid \mathbf{r} \rangle \left\langle \mathbf{r} \mid \mathbf{k} \right\rangle = \frac{1}{(\delta r)^d} \int \mathrm{d}\mathbf{r} \exp[\mathrm{i}(\mathbf{k} - \mathbf{k}') \cdot \mathbf{r} \\ &= \frac{(2\pi)^d}{(\delta r)^d} \delta(\mathbf{k} - \mathbf{k}') = \delta_{\mathbf{k},\mathbf{k}'} \frac{L^d}{(\delta r)^d} = N \delta_{\mathbf{k},\mathbf{k}'}, \end{aligned}$$

where  $\delta(\mathbf{k} - \mathbf{k}') = \delta_{\mathbf{k},\mathbf{k}'}/(\delta k)^d = L^d \delta_{\mathbf{k},\mathbf{k}'}/(2\pi)^d$  since  $(\delta k)^d = (2\pi)^d L^{-d}$  as defined in the last chapter. Also,

$$\frac{1}{N}\sum_{\mathbf{k}}\left|\mathbf{k}\right\rangle\left\langle \mathbf{k}\right|=\mathbf{I},$$

and  $G_0(\omega, \mathbf{k}) = \langle \mathbf{k} | \mathbf{G}_0 | \mathbf{k} \rangle, \langle G \rangle_c(\omega, \mathbf{k}) = \langle \mathbf{k} | \langle \mathbf{G} \rangle_c | \mathbf{k} \rangle, G(\omega, \mathbf{k}, \mathbf{k}') = \langle \mathbf{k}' | \mathbf{G} | \mathbf{k} \rangle.$ 

# 3.4.3 Scattering Amplitude and the t Operator

In the literature, it is popular to define a scattering amplitude  $f(\kappa'_0, \kappa_0)$  as

$$\phi(\omega, \mathbf{r}) \underset{|\mathbf{r}| \to \infty}{\longrightarrow} \exp(\mathrm{i}\kappa_0 \cdot \mathbf{r}) + f(\kappa'_0, \kappa_0) \frac{\exp(\mathrm{i}\kappa_0 r)}{r^{(d-1)/2}}.$$
 (3.30)

Comparison with (3.28) gives, for d = 3,

$$f(\kappa'_{0},\kappa_{0}) = -\langle \kappa'_{0} | \mathbf{t}^{+} | \kappa_{0} \rangle / 4\pi.$$
(3.31)

In 2D,  $G_0^+(\omega, \mathbf{r} - \mathbf{r}_1)$  is given by  $-iH_0^{(1)}(\kappa_0|\mathbf{r} - \mathbf{r}_1|)/4$ . For  $|\mathbf{r}| \gg R$ ,  $H_0^{(1)}$  may be expanded to give

$$G_{0}^{+}(\omega, \mathbf{r} - \mathbf{r}_{1}) \cong -\frac{i}{4}\sqrt{\frac{2}{\pi\kappa_{0}}} \frac{\exp(i\kappa_{0}|\mathbf{r} - \mathbf{r}_{1}|)}{\sqrt{|\mathbf{r} - \mathbf{r}_{1}|}} \exp\left(\frac{-i\pi}{4}\right)$$

$$\cong -\frac{i}{4}\sqrt{\frac{2}{\pi\kappa_{0}}} \frac{\exp[i(\kappa_{0}r - \pi/4)]}{\sqrt{r}} \exp(i\kappa'_{0} \cdot \mathbf{r}_{1}).$$
(3.32)

By substituting (3.32) into (3.27), one gets for the 2D one-scatterer wave function

$$\phi(\omega, \mathbf{r}) \xrightarrow[|\mathbf{r}| \to \infty]{} \exp(\mathrm{i}\kappa_0 \cdot \mathbf{r}) - \frac{\mathrm{i}}{4} \sqrt{\frac{2}{\pi\kappa_0}} \langle \kappa'_0 | \mathbf{t}^+ | \kappa_0 \rangle \frac{\exp[\mathrm{i}(\kappa_0 r - \pi/4)]}{\sqrt{r}}, \quad (3.33a)$$

so that in 2D,

$$f(\kappa_0',\kappa_0) = -\frac{\mathrm{i}}{4}\sqrt{\frac{2}{\pi\kappa_0}} \langle \kappa_0' | \mathbf{t}^+ | \kappa_0 \rangle \exp\left(-\mathrm{i}\frac{\pi}{4}\right).$$
(3.33b)

In 1D,  $G_0^+(\omega, x - x_1) = -i \exp(i\kappa_0 |x - x_1|)/2\kappa_0$ . Expansion of  $|x - x_1|$  as  $|x| \to \infty$  gives

$$\phi(\omega, \mathbf{r}) \xrightarrow[|x| \to \infty]{} \exp(\mathrm{i}\kappa_0 x) - \frac{\mathrm{i}}{2\kappa_0} \langle \kappa'_0 | \mathbf{t}^+ | \kappa_0 \rangle \exp(\mathrm{i}\kappa_0 |x|), \qquad (3.34a)$$

and

$$f(\mathbf{\kappa}'_{0},\mathbf{\kappa}_{0}) = -\frac{\mathrm{i}}{2\kappa_{0}} \langle \mathbf{\kappa}'_{0} | \mathbf{t}^{+} | \mathbf{\kappa}_{0} \rangle. \qquad (3.34b)$$

It should be noted that  $\mathbf{t}^+$  has the dimension of  $[\text{length}]^{d-2}$ . Therefore comparison with (3.30) shows the scattering amplitude to have the dimension of length in 3D, the dimension of  $[\text{length}]^{1/2}$  in 2D, and dimensionless in 1D.

# 3.4.4 Scattering Amplitude and the Optical Theorem: 3D Case

From (3.28), (3.33), and (3.34), it is clear that the matrix elements of  $\mathbf{t}^+$  in the **k** representation may be obtained from the solution of the single-scatterer wavefunction, as shown in the solutions to Problems 3.4, 3.5, and 3.6 at the end of this chapter, for a spherical scatterer of radius R and  $v \neq v_0$ . The answer in 3D is

$$f(\kappa'_{0},\kappa_{0}) = -\frac{i}{\kappa_{0}} \sum_{n=0}^{\infty} D_{n} P_{n}(\cos\theta)(2n+1), \qquad (3.35)$$

where  $P_n(\cos \theta)$  is the *n*th Legendre polynomial,  $\theta$  is the angle between  $\kappa'_0$  and  $\kappa_0$ , *n* is the angular channel index, and

$$D_n = \frac{\kappa_0 j_n(\kappa R) j'_n(\kappa_0 R) - \kappa j'_n(\kappa R) j_n(\kappa_0 R)}{\kappa j'_n(\kappa R) h_n(\kappa_0 R) - \kappa_0 h'_n(\kappa_0 R) j_n(\kappa R)}$$
(3.36)

results from the solution of the boundary value problem. Here  $j_n$  and  $h_n$  denote the *n*th-order spherical Bessel function and spherical Hankel function of the first kind, respectively; prime means derivative with respect to the argument of the function; and  $\kappa = \omega/v$ ,  $\kappa_0 = \omega/v_0$ . In the limit of  $\kappa R, \kappa_0 R \rightarrow 0$ , f is dominated by  $D_0$ , whose leading order real and imaginary parts have the form

$$D_0 \simeq -\frac{\mathrm{i}}{3} \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right] (\kappa_0 R)^3 - \frac{1}{9} \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right]^2 (\kappa_0 R)^6 + \cdots \qquad (3.37)$$

As a result,

$$\langle \kappa'_0 | \mathbf{t}^+ | \kappa_0 \rangle = -4\pi f(\kappa'_0, \kappa_0) \cong \frac{4\pi}{3} \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right] \frac{(\kappa_0 R)^3}{\kappa_0} - \mathrm{i}\frac{4\pi}{9}$$
$$\times \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right]^2 \frac{(\kappa_0 R)^6}{\kappa_0} + \cdots . \quad (3.38)$$

It is noted that in the  $\kappa_0 R \to 0$  limit, the sphere would appear as a point scatterer relative to the wave, and the scattering is consequently isotropic, just as in the lattice case for a single impurity. Since  $\kappa$  and  $\kappa_0$  are both proportional to  $\omega$ , the scattering amplitude is noted to be proportional to  $\omega^2$  in the leading order. The scattering cross-section, which is just  $|f|^2$ , is thus proportional to  $\omega^4$  in the low-frequency limit. This dependence is well known as the signature of *Rayleigh scattering*. Also, (3.38) verifies an important identity relating the imaginary part of f to the total scattering cross-section O, which is just the angular integral of the leading part of  $|f|^2$ . In the present case, since the scattering is isotropic in the  $\omega \to 0$  limit, we have
$$O = \frac{4\pi}{9} \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right]^2 \frac{(\kappa_0 R)^6}{\kappa_0^2} = \frac{4\pi}{\kappa_0} \operatorname{Im} f(\kappa_0, \kappa_0).$$
(3.39)

Alternatively, (3.39) may be expressed in a more general form as

$$O = -\frac{1}{\kappa_0} \operatorname{Im} \langle \kappa_0 | \mathbf{t}^+ | \kappa_0 \rangle.$$
(3.40)

Equation (3.39) or (3.40) is known as the *optical theorem*. It relates the total scattering cross-section with the imaginary part of the *forward* scattering amplitude, i.e.,  $\kappa'_0 = \kappa_0$ . The forward direction is special because heuristically, what is lost from the original wave flux has to be accounted for by scattering. The optical theorem is just a precise statement of that fact, as shown in the solution to Problem 3.7. As (3.38) and (3.39) have demonstrated the validity of the optical theorem in the low-frequency limit, in Problems 3.4 and 3.5 it is shown that in 3D and 2D the optical theorem is valid for every angular momentum channel in the spherical scatterer case, and in 1D the theorem is trivially valid as shown below.

# 3.4.5 Scattering Amplitude and the Optical Theorem: 2D Case

For 2D the solution of  $\langle \kappa'_0 | \mathbf{t}^+ | \kappa_0 \rangle$  is given in the solution to Problem 3.5. The answer is

$$\langle \kappa'_0 | \mathbf{t}^+ | \kappa_0 \rangle = 4i \sum_{n=-\infty}^{\infty} D_n \exp(in\theta)$$
 (3.41)

with

$$D_n = \frac{\kappa J'_n(\kappa R) J_n(\kappa_0 R) - \kappa_0 J_n(\kappa R) J'_n(\kappa_0 R)}{\kappa_0 H'_n(\kappa_0 R) J_n(\kappa R) - \kappa H_n(\kappa_0 R) J'_n(\kappa R)}.$$
(3.42)

Here  $H_n$  stands for the Hankel function of the first kind, of order n, and  $J_n$  stands for the Bessel function of order n. At low frequencies,  $D_0$  dominates, and its leading terms in the expansion are given by

$$D_0 \simeq -i\frac{\pi}{4} \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right] (\kappa_0 R)^2 - \frac{\pi^2}{16} \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right]^2 (\kappa_0 R)^4 + \cdots . \quad (3.43)$$

The *t*-matrix element to the leading order is thus given by

$$\langle \kappa'_0 | \mathbf{t}^+ | \kappa_0 \rangle \cong \pi \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right] (\kappa_0 R)^2 - \mathrm{i} \frac{\pi^2}{4} \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right]^2 (\kappa_0 R)^4 + \cdots$$
(3.44)

Again, the scattering is seen to be isotropic because in the  $\kappa_0 R \to 0$  limit the scatterer appears as a point impurity to the incident wave. The scattering amplitude f to the leading order may be obtained by comparison with (3.33):

$$f(\kappa'_0,\kappa_0) \cong -i\frac{\sqrt{2\pi}}{4} \frac{(\kappa_0 R)^2}{\sqrt{\kappa_0}} \left[1 - \left(\frac{\kappa}{\kappa_0}\right)^2\right] \exp\left(-i\frac{\pi}{4}\right).$$
(3.45)

The total scattering cross-section, given by  $|f|^2$  integrated over all angles, is now

$$O = \frac{\pi^2}{4} \frac{(\kappa_0 R)^4}{\kappa_0} \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right]^2 = -\frac{1}{\kappa_0} \operatorname{Im} \langle \kappa_0 | \mathbf{t}^+ | \kappa_0 \rangle.$$
(3.46)

The optical theorem is seen to be satisfied, although its popular version, (3.39), is not satisfied here unless one gives a different definition for f in 2D. However, in the form of (3.40) it is always valid. Equation (3.46) also gives  $\omega^3$  as the frequency dependence for the 2D Rayleigh scattering cross-section in the  $\omega \to 0$  limit.

# 3.4.6 Scattering Amplitude and the Optical Theorem: 1D Case

In 1D, the solution to the scattering problem can be written down explicitly as shown in the solution to Problem 3.6. Since there are only two scattering directions, backward and forward, the answer may be expressed in terms of reflection and transmission coefficients  $\rho$  and  $\tau$ :

$$\langle -\kappa_0 | \mathbf{t}^+ | \kappa_0 \rangle = i 2\kappa_0 \rho, \qquad (3.47)$$

$$\langle \kappa_0 | \mathbf{t}^+ | \kappa_0 \rangle = i2\kappa_0(\tau - 1), \qquad (3.48)$$

where  $\rho$  and  $\tau$  are given by the following formulae:

$$\rho = -\mathrm{i} \frac{[1 - (\kappa/\kappa_0)^2] \sin 2\kappa R}{2(\kappa/\kappa_0) \cos 2\kappa R - \mathrm{i}[1 + (\kappa/\kappa_0)^2] \sin 2\kappa R} \exp(-2\mathrm{i}\kappa_0 R), \quad (3.49)$$

$$\tau = \frac{2(\kappa/\kappa_0)}{2(\kappa/\kappa_0)\cos 2\kappa R - i[1 + (\kappa/\kappa_0)^2]\sin 2\kappa R} \exp(-2i\kappa_0 R).$$
(3.50)

It is easily verified that  $|\rho|^2 + |\tau|^2 = 1$ , and the optical theorem is just a reflection of energy conservation in this case because the total scattering cross-section is given by  $O = |\tau - 1|^2 + |\rho|^2 = 2(1 - \text{Re}\tau)$ , whereas  $-\text{Im} \langle \kappa_0 | \mathbf{t}^+ | \kappa_0 \rangle / \kappa_0$  is simply  $-\text{Im}2i(\tau - 1) = 2(1 - \text{Re}\tau)$  and is therefore identical to the total scattering cross-section. Also, at low frequencies the Rayleigh scattering cross-section is noted to be proportional to  $\omega^2$ . The results in 1D, 2D, and 3D show that the frequency dependence of the Rayleigh scattering is  $\omega^{d+1}$  in general. Rayleigh scattering is general to classical waves, and its frequency dependence is a reflection of the classical wave dispersion relation and the fact that  $\sigma(\mathbf{r}) \propto \omega^2$ .

# 3.5 Infinite Number of Scatterers: the Effective Medium

When there are infinitely many scatterers, the T-matrix and the exact Green function are impossible to obtain accurately. However, it is noted that in the wave vector representation, the averaged Green function is given by

$$\langle G \rangle_{\rm c}(\omega, \mathbf{k}) = \frac{1}{\kappa_0^2(\omega) - \beta e(\mathbf{k}) - \Sigma(\omega, \mathbf{k})} \frac{N}{L^d},$$
 (3.51)

where  $\beta e(\mathbf{k}) = k^2$  for classical scalar waves in continuum. This form of  $\langle G \rangle_c$  follows from (3.14) and (2.22). It is clear from (3.51) that all nontrivial information about  $\langle G \rangle_c$  is contained in  $\Sigma(\omega, \mathbf{k})$ . If  $\Sigma$  turns out to be *independent* of  $\mathbf{k}$  in some frequency regime, then the effect of  $\Sigma(\omega)$  is just to renormalize  $\kappa_0^2(\omega)$ , the uniform medium property. This renormalized medium is called the effective medium.

### 3.5.1 Meaning of a k-Independent Self Energy

The meaning of self-energy's **k** independence is as follows. In general, **k** dependence reflects nonlocal correlations, because **k** implies spatial derivative. For example, consider the function  $F(\mathbf{r}) = 1$  for  $|\mathbf{r}| \leq R$  and zero otherwise (a sphere). In the Fourier transformed **k** domain,  $F(k) = 4\pi N(\sin kR - kR\cos kR)/(k^3L^d)$ . Here the **k** dependence of F(k) is the reflection of a finite sphere in real space. However, as  $kR \to 0$  it is easy to verify that F(k) becomes **k** independent, since relative to  $\lambda$  the sphere becomes a point, and a point has no length scale and no geometric structure. Indeed, the results for single scatterer in the preceding section show no **k** dependence in the *t*-matrix elements at the  $\kappa_0 R \to 0$  limit. If in addition the pointlike scatterers are situated randomly in space with no correlation between them, there is indeed a possibility that  $\Sigma$  could be **k** independent. From the discussion following (3.15), it is also clear that only when  $\Sigma$  has no **k** dependence would the wave equation be recovered. And in that case  $\Sigma(\omega)$  would combine directly with  $\kappa_0^2(\omega)$  to define the dispersion relation of the renormalized effective medium.

# 3.5.2 The Mean Free Path

Suppose the effective medium exists; then some of its properties can be easily inferred. First,  $\Sigma$  is likely to be frequency dependent, but it may not be the same frequency dependence as  $\kappa_0^2(\omega)$  for a uniform medium. For example, classical wave scattering from point scatterers gives rise to the Rayleigh frequency dependence of  $\omega^{d+1}$  in the low-frequency limit. Therefore, if randomly situated point scatterers can yield an effective medium, its frequency dependence should see a slight deviation from  $\kappa_0^2 \propto \omega^2$  at higher frequencies ( $\omega^2$  would

dominate over  $\omega^{d+1}$  at low frequencies for 2D and 3D). Second,  $\Sigma(\omega)$  is in general a complex number. Its imaginary part must be negative so that the imaginary part of  $\langle G^+ \rangle_c$  is negative. This is because  $-\text{Im} \langle G^+ \rangle_c /\pi$  is related to the density of states for the effective medium, which must be positive or zero. If the combination  $\kappa_0^2(\omega) - \Sigma(\omega)$  is viewed as a new  $[\kappa_e^*(\omega)]^2$  for the effective medium, then  $\kappa_e^*$  should have a positive imaginary part. For example, in 3D the real space Green's function of a continuous effective medium would have the form

$$\left\langle G^{+}\right\rangle_{c} = -\frac{\exp(\mathrm{i}\kappa_{e}^{*}|\mathbf{r}-\mathbf{r}'|)}{4\pi|\mathbf{r}-\mathbf{r}'|},\qquad(3.52)$$

just as in a uniform medium [see (2.31)]. If  $\kappa_{\rm e}^*$  has a positive imaginary part,  $\langle G^+ \rangle_{\rm c}$  decays exponentially. The meaning of this decay may be inferred from the optical theorem because the imaginary part of  $\Sigma(\omega)$  is directly related to the imaginary part of the averaged T matrix, which is related to the total scattering cross-section. Therefore, the spatial decay of  $\langle G^+ \rangle_{\rm c}$  is a result of the scattering loss. Because all scatterings considered are elastic, what is lost is the unique propagation direction and phase relation, i.e., the coherence, of a plane wave. The inverse of  $\mathrm{Im}\kappa_{\rm e}^*$  gives a length for such decay, which is generally denoted twice the mean free path l. The factor of 2 enters because l is defined as the *intensity* decay length for the coherent wave component. Inside the scale of a mean free path a propagating wave is expected to maintain its coherent behavior, but the dynamics beyond the scale of l have to be probed through the evaluation of  $\langle GG^* \rangle_{\rm c}$ , to be examined subsequently.

### 3.5.3 The Coherent Potential Approximation (CPA)

If the solution for an effective medium exists, i.e.,  $\Sigma(\omega, \mathbf{k}) = \Sigma(\omega)$ , an efficient approach to find such a solution is given by the coherent-potential approximation (CPA) (Lax 1951). The central idea of CPA is the simple observation that *relative* to the renormalized uniform medium with  $[\kappa_{\rm e}^*(\omega)]^2 = \kappa_0^2(\omega) - \Sigma(\omega)$ ,  $\langle \mathbf{G} \rangle_c$  is just a uniform-medium Green function, given by  $G_{\rm e}(\kappa_{\rm e}^*)$  in the krepresentation. Since  $\langle \mathbf{G} \rangle_c^{-1} = \mathbf{G}_{\rm e}^{-1} - \overline{\Sigma} L^d / N$ , that means  $\overline{\Sigma}$  must vanish in the renormalized effective medium. From (3.15),  $\overline{\Sigma} = 0$  implies

$$\left\langle \bar{\mathbf{T}} \right\rangle_{c} = 0$$
 (3.53)

in the effective medium. Here we have used an overbar to distinguish those self-energy and scattering matrix elements defined relative to the effective medium. In terms of  $\mathbf{G}_{\mathrm{e}}$ , the exact Green function is given by

$$\mathbf{G} = \mathbf{G}_{\mathrm{e}} + \mathbf{G}_{\mathrm{e}} \bar{\mathbf{T}} \mathbf{G}_{\mathrm{e}}, \qquad (3.54\mathrm{a})$$

so that

$$\langle \mathbf{G} \rangle_c = \mathbf{G}_{\mathrm{e}}.$$
 (3.54b)

The condition (3.53) means that in the effective medium, there is no scattering on average. However, scattering still exists, since  $\langle \bar{\mathbf{T}}\bar{\mathbf{T}}\rangle_c \neq 0$  in general. This point is important for understanding the origin of the diffusive wave transport behavior.

A theory would not be very useful if it can be expressed formally but its consequences can not be explicitly evaluated. In (3.53), the  $\bar{\mathbf{T}}$  matrix cannot be calculated in general for a disordered medium. However, the beauty of the CPA approach is that precisely in the  $\langle \bar{\mathbf{T}} \rangle_c \rightarrow 0$  limit, the theory becomes calculable because  $\langle \bar{\mathbf{T}} \rangle_c \rightarrow 0$  implies weak overall scattering. One way this may be achieved is for each individual scatterer to scatter weakly in the effective medium. The overall  $\bar{\mathbf{T}}$  matrix can then be written approximately as the sum of single scatterings from individual scatterers

$$\bar{\mathbf{\Gamma}} \cong \sum_{i} \bar{\mathbf{t}}_{i}, \qquad (3.55)$$

where i is the index for the scatterers. That is, when each scattering is weak, the multiple scattering terms can be ignored because they represent small numbers raised to some power >1. The accuracy of this approximation will be examined in the next section. With this approximation, the CPA condition becomes

$$\left\langle \bar{\mathbf{T}} \right\rangle_{\mathrm{c}} \cong \sum_{i} \left\langle \bar{\mathbf{t}}_{i} \right\rangle_{\mathrm{c}} = 0.$$
 (3.56)

The implications of the CPA condition are elucidated in the next chapter.

# 3.6 Accuracy of the CPA

The meaning of accuracy depends on the reference standard for comparison. Since the  $\mathbf{G}_{e}$  derived from the CPA is an approximation to  $\langle \mathbf{G} \rangle_{c}$ , its accuracy should be judged by comparison with  $\langle \mathbf{G} \rangle_{c}$ . In this context, the crucial approximation involved is (3.56), where  $\langle \mathbf{\bar{T}} \rangle_{c}$  is expressed as the sum of individual  $\langle \mathbf{\bar{t}}_{i} \rangle_{c}$ 's. In the solution to Problem 3.8, it is shown that the first few terms of  $\langle \mathbf{\bar{T}} \rangle_{c}$  may be written as

$$\left\langle \bar{\mathbf{T}} \right\rangle_{c} = \sum_{i} \left\langle \bar{\mathbf{t}}_{i} \right\rangle_{c} + \sum_{i} \sum_{j \neq i} \left\langle \bar{\mathbf{t}}_{i} \mathbf{G}_{e} \bar{\mathbf{t}}_{j} \right\rangle_{c} + \sum_{i} \sum_{j \neq i} \sum_{k \neq j} \left\langle \bar{\mathbf{t}}_{i} \mathbf{G}_{e} \bar{\mathbf{t}}_{j} \mathbf{G}_{e} \bar{\mathbf{t}}_{k} \right\rangle_{c} + \sum_{i} \sum_{j \neq i} \sum_{k \neq j} \sum_{l \neq k} \left\langle \bar{\mathbf{t}}_{i} \mathbf{G}_{e} \bar{\mathbf{t}}_{j} \mathbf{G}_{e} \bar{\mathbf{t}}_{l} \right\rangle_{c} + \cdots$$

$$(3.57)$$

For the second term, the fact that  $j \neq i$  means averaging over every subset of configurations with a fixed  $\overline{\mathbf{t}}_i$  would give zero under the CPA condition. The third term can have i = k, but there is always a single  $\overline{\mathbf{t}}$  left over to be separately averaged, thus also giving zero. Therefore, the first nonzero correction to the CPA occurs in the fourth-order term (i = k, j = l). The CPA is thus highly accurate. Physically, the higher moments missed by the CPA represent multiple scatterings from the same site(s), which could be important if there are local resonances or bound states.

Another inaccuracy of the CPA arises from its inability to capture fully the microstructure information, i.e., geometric correlations. In other words, microstructure information should supplement the CPA in order to obtain an adequate description of the effective medium properties. This important point is illustrated in Chap. 4.

# **Problems and Solutions**

**Problem 3.1** Show that  $\langle G \rangle_c$  is a function of  $\mathbf{r} - \mathbf{r'}$  in the real-space representation and of only one  $\mathbf{k}$  in the wave vector representation.

**Solution** The general Green function  $G(\omega, \mathbf{r}, \mathbf{r}')$  may be expressed in the wave vector representation as

$$G(\omega, \mathbf{r}, \mathbf{r}') = \frac{L^{2d}}{N^2 (2\pi)^{2d}} \int \int \mathrm{d}\mathbf{k} \,\mathrm{d}\mathbf{k}' G(\omega, \mathbf{k}, \mathbf{k}') \exp(\mathrm{i}\mathbf{k} \cdot \mathbf{r}) \exp(-\mathrm{i}\mathbf{k}' \cdot \mathbf{r}').$$
(P3.1)

By transforming  $\mathbf{r}$  and  $\mathbf{r}'$  to the center-of-mass coordinates  $R_{\rm c} = (\mathbf{r} + \mathbf{r}')/2$ and  $\Delta \mathbf{R} = (\mathbf{r} - \mathbf{r}')/2$ , (P3.1) becomes

$$G(\omega, \mathbf{r}, \mathbf{r}') = \frac{L^{2d}}{N^2 (2\pi)^{2d}} \int \int d\mathbf{k} \, d\mathbf{k}' G(\omega, \mathbf{k}, \mathbf{k}') \\ \times \exp[i(\mathbf{k} - \mathbf{k}') \cdot \mathbf{R}_c] \exp[i(\mathbf{k} + \mathbf{k}') \cdot \Delta \mathbf{R}].$$
(P3.2)

The action of averaging over configurations of  $\sigma(\mathbf{r})$  consists of two steps. The first is the replacement of  $G(\omega, \mathbf{k}, \mathbf{k}')$  by its averaged value,  $\langle G(\omega, \mathbf{k}, \mathbf{k}') \rangle_c$ . The second is to fix  $\Delta \mathbf{R}$  and let the source–detector pair be situated in all possible environments. Mathematically, this is easily carried out by the operation

$$\sum_{\mathbf{R}_{c}} G(\omega, \mathbf{r}, \mathbf{r'}) = \frac{N}{L^{d}} \int G(\omega, \mathbf{r}, \mathbf{r'}) d\mathbf{R}_{c}.$$
 (P3.3)

The integration over  $\mathbf{R}_{c}$  inside the integral of (P3.2) is seen to give

$$\frac{N}{L^d} \int \exp[i(\mathbf{k} - \mathbf{k}') \cdot \mathbf{R}_c] d\mathbf{R}_c = \frac{N}{L^d} (2\pi)^d \delta(\mathbf{k} - \mathbf{k}').$$
(P3.4)

By substituting (P3.4) into (P3.2), the integration over  $\mathbf{k}'$  can be immediately performed, with the result

$$\langle G(\omega, \mathbf{r}, \mathbf{r}') \rangle_{c} = \langle G \rangle_{c} (\omega, \mathbf{r} - \mathbf{r}')$$
  
=  $\frac{L^{d}}{N(2\pi)^{d}} \int d\mathbf{k} \langle G \rangle_{c} (\omega, \mathbf{k}, \mathbf{k}) \exp[i\mathbf{k} \cdot (\mathbf{r} - \mathbf{r}')].$ (P3.5)

The fact that  $\langle G \rangle_c$  depends on only one **k** (i.e., **k** = **k**') is clear from (P3.5), which proves our assertion.

**Problem 3.2** Show that the convolution integral in real space means simple multiplication in the wavevector space.

Solution A convolution integral is defined by

$$F(\mathbf{x}) = \frac{1}{L^d} \int G(\mathbf{x} - \mathbf{y}) H(\mathbf{y}) \mathrm{d}\mathbf{y}.$$
 (P3.6)

In Fourier components, each function may be written as

$$F(\mathbf{x}) = \frac{L^d}{N} \int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^d} F(\mathbf{k}) \exp(\mathrm{i}\mathbf{k} \cdot \mathbf{x}), \qquad (P3.7)$$

$$G(\mathbf{x} - \mathbf{y}) = \frac{L^d}{N} \int \frac{\mathrm{d}\mathbf{k}_1}{(2\pi)^d} G(\mathbf{k}_1) \exp[\mathrm{i}\mathbf{k}_1 \cdot (\mathbf{x} - \mathbf{y})], \qquad (P3.8)$$

$$H(\mathbf{y}) = \frac{L^d}{N} \int \frac{\mathrm{d}\mathbf{k}_2}{(2\pi)^d} H(\mathbf{k}_2) \exp(\mathrm{i}\mathbf{k}_2 \cdot \mathbf{y}).$$
(P3.9)

By substituting (P3.7)–(P3.9) into (P3.6) and performing the **y** integral, one gets the factor  $S(\mathbf{1}_{\mathbf{y}} = \mathbf{1}_{\mathbf{y}})/(2-\mathbf{y})^d/T^d$ 

$$\delta(\mathbf{k}_2-\mathbf{k}_1)(2\pi)^d/L^d$$

as the result. Carrying out the  $\mathbf{k}_2$  integration then gives

$$L^{d} \int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^{d}} F(\mathbf{k}) \exp(\mathrm{i}\mathbf{k} \cdot \mathbf{x})$$
  
=  $L^{d} \int \frac{\mathrm{d}\mathbf{k}_{1}}{(2\pi)^{d}} G(\mathbf{k}_{1}) H(\mathbf{k}_{1}) \exp(\mathrm{i}\mathbf{k}_{1} \cdot \mathbf{x}).$  (P3.10)

Thus, the Fourier amplitudes have the relation

$$F(\mathbf{k}) = G(\mathbf{k})H(\mathbf{k}), \qquad (P3.11)$$

i.e., simple multiplication. By interchanging  $\mathbf{k}$  and  $\mathbf{x}$ , it is clear that the converse is also true, i.e., convolution in the  $\mathbf{k}$  domain means simple multiplication in the spatial domain.

**Problem 3.3** Derive an expression for the total scattering cross-section of a single impurity in a lattice, and verify the optical theorem for that case.

**Solution** By starting from the  $t^+$  operator for the single impurity scatterer, (3.19),

$$\mathbf{t}^{+} = \left| \mathbf{l}_{0} \right\rangle \frac{\sigma_{0} a^{d}}{1 - \sigma_{0} a^{d} G_{0}^{+}(\omega, \mathbf{l} = \mathbf{l}')} \left\langle \mathbf{l}_{0} \right|,$$

it is easy to get

$$\langle \mathbf{k}'_{0} | \mathbf{t}^{+} | \mathbf{k}_{0} \rangle = \langle \mathbf{k}'_{0} | \mathbf{l}_{0} \rangle \frac{\sigma_{0} a^{d}}{1 - \sigma_{0} a^{d} G_{0}^{+}(\omega, \mathbf{l} = \mathbf{l}')} \langle \mathbf{l}_{0} | \mathbf{k}_{0} \rangle ,$$

$$= \exp[\mathbf{i}(\mathbf{k}_{0} - \mathbf{k}'_{0}) \cdot \mathbf{l}_{0}] \frac{\sigma_{0} a^{d}}{1 - \sigma_{0} a^{d} G_{0}^{+}(\omega, \mathbf{l} = \mathbf{l}')}, \quad (P3.12)$$

### 3.6 Accuracy of the CPA 65

where  $\mathbf{k}'_0$  and  $\mathbf{k}_0$  are wave vectors which satisfy the relations  $\kappa_0^2 = \beta e(\mathbf{k}_0)$ and  $\kappa_0^2 = \beta e(\mathbf{k}'_0)$ . The total scattering cross-section O is given by

$$O = \int_{\omega} |f|^2 \mathrm{d}\Omega = |f|^2 \int_{\omega} \mathrm{d}\Omega, \qquad (P3.13)$$

where  $|f|^2$  stands for the scattering cross-section, which is angle independent as can be seen from P3.12, and the integral is over all solid angles but conditioned on the availability of states at the given frequency  $\omega$ . The expression for f in the scalar wave case has been given in the text by (3.38) (3D), (3.45) (2D), and (3.47)–(3.50) (1D). In the long-wavelength limit they should agree with the similar quantities in the lattice case. That means we can write

$$f(\mathbf{k}'_{0}, \mathbf{k}_{0}) = \begin{cases} -\frac{1}{4\pi} \langle \mathbf{k}'_{0} | \mathbf{t}^{+} | \mathbf{k}_{0} \rangle & 3D \\ -\frac{i}{4} \sqrt{\frac{2}{\pi \bar{k}_{0}}} \exp(-i\frac{\pi}{4}) \langle \mathbf{k}'_{0} | \mathbf{t}^{+} | \mathbf{k}_{0} \rangle & 2D \\ -\frac{i}{2 \bar{k}_{0}} \langle \mathbf{k}'_{0} | \mathbf{t}^{+} | \mathbf{k}_{0} \rangle & , 1D \end{cases}$$
(P3.14)

where  $\bar{k}_0 = \kappa_0/\sqrt{\beta}$  in the long-wavelength limit. The general definition of  $\bar{k}_0$  is given below. To do the angular integral at a constant frequency, it is noted that whereas in the continuum case the phase space is isotropic, in the lattice case this is no longer true since the Brillouin zone is not spherical. Therefore, the angular integral should be expressed in terms of an integral over **k** (at fixed  $\omega$ ) as

$$\int_{\omega} \mathrm{d}\Omega = \begin{cases} \frac{2\beta}{\bar{k}_0} \int \mathrm{d}\mathbf{k}\,\delta[\kappa_0^2 - \beta e(\mathbf{k})] & 3\mathrm{D} \\ 2\beta \int \mathrm{d}\mathbf{k}\,\delta[\kappa_0^2 - \beta e(\mathbf{k})] & 2\mathrm{D} \\ 2\bar{k}_0\beta \int \mathrm{d}\mathbf{k}\,\delta[\kappa_0^2 - \beta e(\mathbf{k})]. & 1\mathrm{D} \end{cases}$$
(P3.15)

The validity of these expressions may easily be verified in the case when  $\kappa_0^2 \to 0$  so that the delta function is operative in the regime of  $|\mathbf{k}| \to 0$ . Then  $\beta e(\mathbf{k}) = \beta k^2$ ,  $\bar{k}_0 = \kappa_0/\sqrt{\beta}$ , and by separating d**k** into  $d\Omega \cdot k^{d-1}dk$ , the integral is seen to yield  $4\pi$ ,  $2\pi$ , 2 in 3D, 2D, 1D, respectively. The factor 2 is present because  $e(\mathbf{k})$  is a function of  $k^2$  (from  $\cos ka$ ). By combining (P3.12), (P3.13), (P3.14), (P3.15) and recognizing

$$\mathrm{Im}G_0^+(\omega, \mathbf{l} = \mathbf{l}') = -\pi \int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^d} \delta[\kappa_0^2 - \beta e(\mathbf{k})],$$

one immediately gets for all spatial dimensionalities,

$$O = -\frac{(\sigma_0 a^d)^2 \text{Im} G_0^+(\omega, \mathbf{l} = \mathbf{l}')}{\bar{k}_0 |1 - \sigma_0 a^d G_0^+(\omega, \mathbf{l} = \mathbf{l}')|^2}.$$
 (P3.16a)

From (P3.12) it is easily seen that

$$O = -\frac{\operatorname{Im} \langle \mathbf{k}_0 | \mathbf{t}^+ | \mathbf{k}_0 \rangle}{\bar{k}_0}, \qquad (P3.16b)$$

which is precisely the optical theorem.

To define  $\bar{k}_0$ , we note that an alternative definition of the angular integral is

$$\int_{\omega} \mathrm{d}\Omega = \frac{1}{\bar{k}_0^{d-1}} \int_{\omega} \mathrm{d}S, \qquad (P3.17a)$$

where dS represents a constant frequency surface integration. But

$$\int_{\omega} \mathrm{d}S = \int \mathrm{d}S \int \mathrm{d}k_{\perp} \left| \nabla_k \beta e(\mathbf{k}) \right| \delta[\kappa_0^2 - \beta e(\mathbf{k})], \qquad (P3.17b)$$

where  $k_{\perp}$  is the wave vector perpendicular to the constant-frequency surface. Since a given value of  $\beta e(\mathbf{k})$  defines the constant-frequency surface and  $\nabla_k \beta e(\mathbf{k})$  gives the direction perpendicular to it,  $dk_{\perp} |\nabla_k \beta e(\mathbf{k})|$  is essentially  $d[\beta e(\mathbf{k})]$ , so the integration over values of  $\beta e(\mathbf{k})$  in (P3.17b) is recognized to give 1 from the delta function. Since  $dSdk_{\perp} = d\mathbf{k}$ , a comparison between (P3.17a), (P3.17b) and (P3.15) thus gives

$$\bar{k}_0 = \frac{1}{2} \frac{\int \mathrm{d}\mathbf{k} \left| \nabla_k \beta e(\mathbf{k}) \right| \delta[\kappa_0^2 - \beta e(\mathbf{k})]}{\beta \int \mathrm{d}\mathbf{k} \delta[\kappa_0^2 - \beta e(\mathbf{k})]}.$$
(P3.18)

The right-hand side of (P3.18) is clearly an average of  $|\nabla_k \beta e(\mathbf{k})|$  over the states of constant frequency, which is the physical meaning of  $\bar{k}_0$ .

**Problem 3.4** Solve the scalar-wave scattering problem of a plane wave incident on a sphere with radius R and wave speed  $v \neq v_0$ . Find the expression for the scattering amplitude in the limit of  $\kappa_0 R \to 0$ , and verify the optical theorem for each angular momentum channel.

**Solution** For the scalar wave equation  $(\nabla^2 + \kappa^2)\phi = 0$ , where  $\kappa^2 = \omega^2/v^2$  inside the sphere, the solution may be expressed in terms of an expansion:

$$\phi^{<}(\mathbf{r}) = \sum_{n=0}^{\infty} A_n j_n(\kappa r) P_n(\cos\theta), \qquad r \le R.$$
(P3.19)

Outside the sphere,  $\phi^{>}(\mathbf{r})$  is a sum of the incident plane wave and the scattered wave, with  $\kappa^{2} = \kappa_{0}^{2} = \omega^{2}/v_{0}^{2}$ :

$$\phi^{>}(\mathbf{r}) = \exp(\mathrm{i}\kappa_0 \cdot \mathbf{r}) + \sum_{n=0}^{\infty} D_n h_n^{(1)}(\kappa_0 r) P_n(\cos\theta) \,\mathrm{i}^n(2n+1), \qquad r > R.$$
(P3.20)

Here  $j_n$  and  $h_n^{(1)}$  are the spherical Bessel function and the spherical Hankel function of the first kind, respectively. If the scalar wave equation is valid everywhere in space, then the boundary conditions have to be the continuity of  $\phi$  and  $\partial \phi / \partial r$  across the interface. This is because the jump in the value of  $\kappa$  from  $\omega/v$  inside the sphere to  $\omega/v_0$  outside the sphere implies the second

(normal) derivative of  $\phi$  must have the same jump so that the scalar wave equation holds across the interface. It follows that the normal derivative of  $\phi$  and  $\phi$  itself must both be continuous because they are, respectively, the first and second integrals of the jump. The two equations that result from the application of the boundary conditions are

$$A_n j_n(\kappa R) P_n(\cos \theta) = \mathbf{i}^n (2n+1) P_n(\cos \theta) [j_n(\kappa_0 R) + D_n h_n(\kappa_0 R)], \quad (P3.21)$$
$$\frac{\kappa}{\kappa_0} A_n j'_0(\kappa R) P_n(\cos \theta) = \mathbf{i}^n (2n+1) P_n(\cos \theta) [j'_n(\kappa_0 R) + D_n h'_n(\kappa_0 R)]. \quad (P3.22)$$

Here the prime indicates derivative with respect to the argument of the function, and the superscript (1) has been dropped from the spherical Hankel function. In obtaining the two equations, an expansion of the plane wave has been used:

$$\exp(\mathrm{i}\kappa_0 \cdot \mathbf{r}) = \sum_{n=0}^{\infty} \mathrm{i}^n (2n+1) P_n(\cos\theta) j_n(\kappa_0 r), \qquad (P3.23)$$

and the orthogonality of the different angular momentum channels has been utilized. By dividing (P3.22) by (P3.21) so as to eliminate  $A_n$ , one immediately gets

$$D_n = \frac{\kappa_0 j_n(\kappa R) j'_n(\kappa_0 R) - \kappa j'_n(\kappa R) j_n(\kappa_0 R)}{\kappa j'_n(\kappa R) h_n(\kappa_0 R) - \kappa_0 h'_n(\kappa_0 R) j_n(\kappa R)}.$$
(P3.24)

From (3.35), we can write

$$f(\mathbf{\kappa}'_0, \mathbf{\kappa}_0) = \sum_n f_n(\mathbf{\kappa}'_0, \mathbf{\kappa}_0), \qquad (P3.25)$$

where

$$f_n(\kappa'_0,\kappa_0) = -\frac{\mathrm{i}}{\kappa_0} D_n P_n(\cos\theta)(2n+1)$$
(P3.26)

is the scattering amplitude for the nth angular momentum channel. It is easy to show that

$$2\pi \int_{-1}^{1} |f_n|^2 d\cos\theta = \frac{4\pi}{\kappa_0^2} |D_n|^2 (2n+1) = O_n.$$
 (P3.27)

From (P3.24) and the definition

$$\left\langle \mathbf{\kappa}'_{0}\right|\mathbf{t}^{+}\left|\mathbf{\kappa}_{0}\right\rangle _{n}=-4\pi f_{n}(\mathbf{\kappa}'_{0},\mathbf{\kappa}_{0}),\tag{P3.28}$$

it is easy to see that, for  $\kappa'_0 = \kappa_0$ ,  $P_n(\cos \theta) = 1$  and

$$\frac{-\mathrm{Im}\langle \mathbf{\kappa}'_0 | \mathbf{t}^+ | \mathbf{\kappa}_0 \rangle_n}{\kappa_0} = \frac{4\pi}{\kappa_0^2} (2n+1) [-\mathrm{Imi}D_n].$$
(P3.29)

Since  $h_n = j_n + in_n$ , where  $n_n$  is the spherical Neumann function, and the numerator of  $D_n$  is noted to be real, it follows that

$$-\mathrm{Imi}D_n = -\mathrm{Re}D_n = |D_n|^2.$$
(P3.30)

Therefore,

$$O_n = \frac{-\mathrm{Im} \langle \kappa_0 | \mathbf{t}^+ | \kappa_0 \rangle_n}{\kappa_0}, \qquad (P3.31)$$

i.e., the optical theorem is valid for every angular channel.

To derive an expression for the scattering amplitude in the limit of  $\kappa_0 R \rightarrow 0$ , it is necessary only to consider  $D_0$ , because physically, the scattering should be isotropic for a point scatterer. By expanding all functions to the leading orders of  $\kappa_0 R$  and  $\kappa R$ , we have

$$j_{0}(x) \approx 1 - \frac{x^{2}}{6},$$
  

$$j_{0}'(x) \approx -\frac{x}{3} + \frac{x^{3}}{30},$$
  

$$h_{0}(x) \approx -\frac{i}{x} \left( 1 + ix - \frac{x^{2}}{2} - i\frac{x^{3}}{6} \right),$$
  

$$h_{0}'(x) \approx \left( \frac{1}{x} + i\frac{1}{x^{2}} \right) \left( 1 + ix - \frac{x^{2}}{2} - i\frac{x^{3}}{6} \right).$$
 (P3.32)

Collecting all the leading-order real and imaginary parts of the  ${\cal D}_0$  expansion gives

$$D_0 = -\frac{i}{3} (\kappa_0 R)^3 \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right] - \frac{1}{9} (\kappa_0 R)^6 \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right]^2.$$
(P3.33)

Therefore, the scattering amplitude to the leading order is given by

$$f(\kappa'_0,\kappa_0) = -\frac{1}{3} \frac{(\kappa_0 R)^3}{\kappa_0} \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right].$$
(P3.34)

Problem 3.5 Give the solution of the previous problem in two dimensions.Solution In 2D, the expansions are

$$\phi^{<}(\mathbf{r}) = \sum_{m=-\infty}^{\infty} A_m J_m(\kappa r) \exp(\mathrm{i}m\,\theta), \quad r \le R.$$
(P3.35)  
$$\phi^{>}(\mathbf{r}) = \exp(\mathrm{i}\kappa_0 \cdot \mathbf{r}) + \sum_{m=-\infty}^{\infty} \mathrm{i}^m D_m H_m^{(1)}(\kappa_0 r) \exp(\mathrm{i}m\,\theta)$$
$$= \sum_{m=-\infty}^{\infty} \mathrm{i}^m \exp(\mathrm{i}m\,\theta) [J_m(\kappa_0 r) + D_m H_m^{(1)}(\kappa_0 r)], \quad r > R.$$
(P3.36)

Here  $J_m$  and  $H_m^{(1)}$  are the Bessel function and the Hankel function of first kind, respectively, and we have utilized the 2D expansion formula for  $\exp(i\kappa_0 \cdot \mathbf{r})$ . The definitions of  $\theta$ ,  $\kappa$ , and  $\kappa_0$  are the same as in Problem 3.4. The boundary

conditions at r = R, continuity of  $\phi$  and  $\partial \phi / \partial r$ , may be combined to give the continuity of  $(\partial \phi / \partial r) / \phi = \partial (\ln \phi) / \partial r$ , which can be written as

$$\frac{\kappa J'_m(\kappa R)}{\kappa_0 J_m(\kappa R)} = \frac{J'_m(\kappa_0 R) + D_m H'_m(\kappa_0 R)}{J_m(\kappa_0 R) + D_m H_m(\kappa_0 R)}$$
(P3.37)

for each angular channel. Here we have dropped the superscript (1) on the Hankel function.  $D_m$  is thus given by the expression

$$D_m = \frac{\kappa J'_m(\kappa R) J_m(\kappa_0 R) - \kappa_0 J_m(\kappa R) J'_m(\kappa_0 R)}{\kappa_0 H'_m(\kappa_0 R) J_m(\kappa R) - \kappa H_m(\kappa_0 R) J'_m(\kappa R)}.$$
(P3.38)

The scattering amplitude is given by

$$f(\kappa'_0,\kappa_0) = \sum_{m=-\infty}^{\infty} f_m(\kappa'_0,\kappa_0), \qquad (P3.39)$$

$$f_m(\kappa'_0,\kappa_0) = \sqrt{\frac{2}{\pi\kappa_0}} D_m \exp(\mathrm{i}m\,\theta) \exp\left(-\mathrm{i}\frac{\pi}{4}\right). \tag{P3.40}$$

Thus

$$O_m = \int_0^{2\pi} |f_m|^2 \mathrm{d}\theta = \frac{4}{\kappa_0} |D_m|^2.$$
(P3.41)

From (P3.14) and the condition of  $\theta = 0$  for forward scattering, we have

$$\left\langle \kappa_{0}^{\prime}\right|\mathbf{t}^{+}\left|\kappa_{0}\right\rangle_{m}=4\mathrm{i}D_{m},\tag{P3.42}$$

from which it immediately follows that

$$O_m = -\frac{\operatorname{Im} \langle \kappa_0 | \mathbf{t}^+ | \kappa_0 \rangle_m}{\kappa_0}, \qquad (P3.43)$$

because  $-\text{Imi}D_m = -\text{Re}D_m = |D_m|^2$  due to the fact that  $H_m^{(1)}(x) = J_m(x) + iN_m(x)$ , where  $N_m(x)$  is the Neumann function.

For the  $\kappa_0 R \to 0$  limit, only the m = 0 term has to be expanded due to the consideration of isotropic scattering. Retaining the leading orders of real and imaginary parts means expanding  $J_0$  and  $H_0$  to the leading orders of their arguments:

$$J_0(x) = 1 - \frac{x^2}{4},$$
  

$$J'_0(x) = -\frac{x}{2},$$
  

$$H_0(x) = i\frac{2}{\pi}\ln x + 1,$$
  

$$H'_0(x) = i\frac{2}{\pi}\frac{1}{x} - \frac{x}{2}.$$
(P3.44)

Collecting terms together yields

$$D_0 = -i\frac{\pi}{4}(\kappa_0 R)^2 \left[1 - \left(\frac{\kappa}{\kappa_0}\right)^2\right] - \frac{\pi^2}{16}(\kappa_0 R)^4 \left[1 - \left(\frac{\kappa}{\kappa_0}\right)^2\right]^2.$$
 (P3.45)

It follows that to the leading order the scattering amplitude is

$$f(\kappa'_0,\kappa_0) = -\frac{\mathrm{i}}{4}\sqrt{\frac{2\pi}{\kappa_0}}(\kappa_0 R)^2 \left[1 - \left(\frac{\kappa}{\kappa_0}\right)^2\right] \exp\left(-\mathrm{i}\frac{\pi}{4}\right).$$
(P3.46)

Problem 3.6 Give the solution of Problem 3.4 in one dimension.

**Solution** In 1D, the solution expansion can be written as a function of the coordinate x:

$$\phi^{<}(x) = A \exp(i\kappa x) + B \exp(-i\kappa x) \quad |x| \le R,$$

$$\phi^{>}(x) = \exp(i\kappa_0 x) + \rho \exp(-i\kappa_0 x), \quad x < -R$$
(P3.47)

$$\rho^{>}(x) = \exp(i\kappa_0 x) + \rho \exp(-i\kappa_0 x), \quad x < -R$$
  
=  $\tau \exp(i\kappa_0 x). \qquad x > R.$  (P3.48)

Here 
$$\rho$$
 denotes the reflection coefficient and  $\tau$  the transmission coefficient.

The boundary conditions at  $x = \pm R$ , the continuity of  $\phi$  and  $d\phi/dx$ , give

$$A\exp(-i\kappa R) + B\exp(i\kappa R) = \exp(-i\kappa_0 R) + \rho\exp(i\kappa_0 R), \qquad (P3.49)$$

$$\frac{\kappa}{\kappa_0} [A \exp(-i\kappa R) - B \exp(i\kappa R)] = \exp(-i\kappa_0 R) - \rho \exp(i\kappa_0 R), \quad (P3.50)$$

$$A\exp(i\kappa R) + B\exp(-i\kappa R) = \tau \exp(i\kappa_0 R), \qquad (P3.51)$$

$$\frac{\kappa}{\kappa_0} [A \exp(i\kappa R) - B \exp(-i\kappa R)] = \tau \exp(i\kappa_0 R).$$
(P3.52)

After some algebra,  $\rho$  and  $\tau$  are solved to yield the expressions given by (3.49) and (3.50). In 1D,  $\rho$  and  $\tau$  together constitute the two scattering channels, and the optical theorem in that case has been explicitly demonstrated in the text.

In the  $\kappa_0 R \to 0$  limit, the forward scattering amplitude,  $\tau - 1$ , can be written as

$$\tau - 1 = -i\kappa_0 R \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right] - (\kappa_0 R)^2 \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right]^2.$$
(P3.53a)

The backward scattering amplitude is

$$\rho = i\kappa_0 R \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right] + (\kappa_0 R)^2 \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right]^2.$$
(P3.53b)

The total scattering cross-section in the  $\kappa_0 R \to 0$  limit is therefore

$$O = 2(\kappa_0 R)^2 \left[ 1 - \left(\frac{\kappa}{\kappa_0}\right)^2 \right]^2.$$
(P3.54)

**Problem 3.7** Show the optical theorem to be a statement of flux conservation.

**Solution** Since only elastic scattering is considered, flux conservation here is the same as energy conservation. The starting point of the consideration is the expression

$$\phi(\mathbf{r}) \underset{r \to \infty}{\longrightarrow} \exp(\mathrm{i}\kappa_0 \cdot \mathbf{r}) + f(\theta) \frac{\mathrm{e}^{\mathrm{i}\kappa_0 r}}{r^{(d-1)/2}}, \qquad (P3.55)$$

where for simplicity the scattering is assumed to be uniform in the azimuthal direction. The plane wave  $\exp(i\kappa_0\cdot \mathbf{r})$  may be written as

$$\exp(\mathbf{i}\kappa_0 \cdot \mathbf{r}) = \begin{cases} \sum_{n=0}^{\infty} \mathbf{i}^n (2n+1) j_n(\kappa_0 r) P_n(\cos\theta) & 3\mathbf{D} \\ \sum_{m=-\infty}^{\infty} \mathbf{i}^m \exp(\mathbf{i}m\theta) J_m(\kappa_0 r) & 2\mathbf{D}. \end{cases}$$
(P3.56)

In the 1D case, because the optical theorem is already explicitly demonstrated to be equivalent to energy conservation, it does not need to be considered further. By using the asymptotic expression for  $j_n(\kappa_0 r)$  and  $J_m(\kappa_0 r)$  as  $r \to \infty$ ,

$$j_n(\kappa_0 r) \underset{r \to \infty}{\longrightarrow} \frac{1}{2} \left[ (-\mathbf{i})^{n+1} \frac{\exp(\mathbf{i}\kappa_0 r)}{\kappa_0 r} + \mathbf{i}^{n+1} \frac{\exp(-\mathbf{i}\kappa_0 r)}{\kappa_0 r} \right], \tag{P3.57}$$

$$J_m(\kappa_0 r) \xrightarrow[r \to \infty]{} \frac{1}{\sqrt{2\pi}} \left[ (-\mathrm{i})^m \frac{\exp\left(\mathrm{i}\left(\kappa_0 r - \frac{\pi}{4}\right)\right)}{\sqrt{\kappa_0 r}} + \mathrm{i}^m \frac{\exp\left(-\mathrm{i}\left(\kappa_0 r - \frac{\pi}{4}\right)\right)}{\sqrt{\kappa_0 r}} \right],\tag{P3.58}$$

and substituting them into (P3.56), one obtains

$$\exp(\mathbf{i}\kappa_{0}\cdot\mathbf{r}) \xrightarrow[r\to\infty]{} \frac{\exp(\mathbf{i}\kappa_{0}r)}{\mathbf{i}\kappa_{0}r} \sum_{n=0}^{\infty} \frac{2n+1}{2} P_{n}(\cos\theta) -\frac{\exp(-\mathbf{i}\kappa_{0}r)}{\mathbf{i}\kappa_{0}r} \sum_{n=0}^{\infty} (-1)^{n} \frac{2n+1}{2} P_{n}(\cos\theta) \quad \text{3D,} \exp(\mathbf{i}\kappa_{0}\cdot\mathbf{r}) \xrightarrow[r\to\infty]{} \sqrt{\frac{2\pi}{\kappa_{0}r}} \exp\left[\mathbf{i}\left(\kappa_{0}r-\frac{\pi}{4}\right)\right] \sum_{m=-\infty}^{\infty} \frac{\exp(\mathrm{i}m\theta)}{2\pi} +\sqrt{\frac{2\pi}{\kappa_{0}r}} \exp\left[-\mathbf{i}\left(\kappa_{0}r-\frac{\pi}{4}\right)\right] \sum_{m=-\infty}^{\infty} (-1)^{m} \frac{\exp(\mathrm{i}m\theta)}{2\pi} \quad \text{2D.}$$

The angular summations in (P3.59) are angular delta functions. This can be shown by their action on some arbitrary angular function  $F(\theta)$ , which may be expanded as

$$F(\theta) = \begin{cases} \sum_{n=0}^{\infty} A_n P_n(\cos \theta) & \text{3D} \\ \sum_{m=-\infty}^{\infty} A_m \exp(im\theta) & \text{2D.} \end{cases}$$
(P3.60)

Integrating  $F(\theta)$  with one of the angular summations gives

$$\int_{-1}^{1} F(\theta) \sum_{n=0}^{\infty} (\pm 1)^{n} \frac{2n+1}{2} P_{n}(\cos \theta) \, \mathrm{d}(\cos \theta)$$
  
=  $\sum_{n,m} A_{m}(\pm 1)^{n} \frac{2n+1}{2} \int_{-1}^{1} P_{n}(x) P_{m}(x) \, \mathrm{d}x$   
=  $\sum_{m} A_{m}(\pm 1)^{m} = \sum_{m} A_{m} P_{m}(\pm 1)$   
=  $F(0)$  if +1  
=  $F(\pi)$  if -1. (P3.61)

That means

$$\sum_{n=0}^{\infty} (\pm 1)^n \frac{2n+1}{2} P_n(\cos \theta) = \delta(\cos \theta \mp 1).$$
 (P3.62)

Similar reasoning gives

$$\sum_{m=-\infty}^{\infty} (\pm 1)^m \frac{\exp(\mathrm{i}m\theta)}{2\pi} = \begin{cases} \delta(\theta) & \text{if } +1\\ \delta(\theta-\pi) & \text{if } -1. \end{cases}$$
(P3.63)

Therefore, in the limit of  $r \to \infty$ ,  $\phi(\mathbf{r})$  has the following form:

$$\phi(\mathbf{r}) \xrightarrow[r \to \infty]{} \frac{\exp(i\kappa_0 r)}{i\kappa_0 r} \delta(\cos \theta - 1) + f(\theta) \frac{\exp(i\kappa_0 r)}{r} - \frac{\exp(-i\kappa_0 r)}{i\kappa_0 r} \delta(\cos \theta + 1), \quad 3D$$

$$\phi(\mathbf{r}) \xrightarrow[r \to \infty]{} \sqrt{\frac{2\pi}{\kappa_0 r}} \exp\left[i\left(\kappa_0 r - \frac{\pi}{4}\right)\right] \delta(\theta) + f(\theta) \frac{\exp(i\kappa_0 r)}{\sqrt{r}} + \sqrt{\frac{2\pi}{\kappa_0 r}} \exp\left[-i\left(\kappa_0 r - \frac{\pi}{4}\right)\right] \delta(\theta - \pi). \quad 2D$$
(P3.64)

They have the simple interpretation that if one draws a large sphere (a circle in 2D) of radius  $R_0$  centered at the scatterer, then the incoming beam is represented by the  $\theta = \pi$ ,  $\exp(-i\kappa_0 r)/r^{(d-1)/2}$  term in (P3.64). The other two terms represent the outgoing part of the wave. In steady state, they must balance. That means in 3D,

$$\int_{-1}^{1} \left| \frac{\exp(i\kappa_0 R_0)}{i\kappa_0 R_0} \delta(\cos\theta - 1) + f(\theta) \frac{\exp(i\kappa_0 R_0)}{R_0} \right|^2 R_0^2 d(\cos\theta)$$
$$= \int_{-1}^{1} \left| \frac{\exp(-i\kappa_0 R_0)}{i\kappa_0 R_0} \delta(\cos\theta + 1) \right|^2 R_0^2 d(\cos\theta).$$
(P3.65)

If  $f(\theta) = 0$  (no scattering), the terms without  $f(\theta)$  are seen to exactly balance. That means, after squaring, the two terms involving  $f(\theta)$  must cancel each other:

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$$0 = \int_{-1}^{1} |f(\theta)|^2 d(\cos \theta) - \frac{2}{\kappa_0} \text{Im} f(0)$$
  
=  $\frac{O}{2\pi} + \frac{1}{2\pi\kappa_0} \text{Im} \langle \kappa_0 | \mathbf{t}^+ | \kappa_0 \rangle.$  (P3.66)

Equation (P3.66) is exactly the optical theorem. Similarly in 2D,

$$0 = \int_{0}^{2\pi} |f(\theta)|^{2} d\theta + f(0) \exp\left(i\frac{\pi}{4}\right) \sqrt{\frac{2\pi}{\kappa_{0}}} + f^{*}(0) \exp\left(-i\frac{\pi}{4}\right) \sqrt{\frac{2\pi}{\kappa_{0}}}$$
$$= O + \frac{1}{\kappa_{0}} \operatorname{Im}\left\langle\kappa_{0}\right| \mathbf{t}^{+} \left|\kappa_{0}\right\rangle, \qquad (P3.67)$$

which is again the optical theorem. This proves our assertion.

**Problem 3.8** Express the **T** operator in terms of the  $\mathbf{t}_i$ 's for the individual scatterers.

Solution From (3.8), it is only one step to see that  $\mathbf{T} - \mathbf{V}\mathbf{G}_0\mathbf{T} = \mathbf{V}$ , so

$$\mathbf{T} = \mathbf{V}(\mathbf{I} + \mathbf{G}_0 \mathbf{T}). \tag{P3.68}$$

Since

$$\mathbf{V} = \sum_{i} \mathbf{v}_{i},\tag{P3.69}$$

where  $\mathbf{v}_i$ 's are the scattering operators for individual scatterers, it follows that one can rewrite (P3.68) as

$$\mathbf{T} = \sum_{i} \mathbf{R}_{i},\tag{P3.70}$$

where

$$\mathbf{R}_{i} = \mathbf{v}_{i}(\mathbf{I} + \mathbf{G}_{0}\mathbf{T}) = \mathbf{v}_{i}\left(\mathbf{I} + \mathbf{G}_{0}\sum_{j}\mathbf{R}_{j}\right).$$
 (P3.71)

By subtracting  $\mathbf{v}_i \mathbf{G}_0 \mathbf{R}_i$  from both sides of this equation, it is seen that

$$(\mathbf{I} - \mathbf{v}_i \mathbf{G}_0) \mathbf{R}_i = \mathbf{v}_i \left( \mathbf{I} + \mathbf{G}_0 \sum_{j \neq i} \mathbf{R}_j \right),$$
 (P3.72)

or

$$\mathbf{R}_{i} = (\mathbf{I} - \mathbf{v}_{i}\mathbf{G}_{0})^{-1}\mathbf{v}_{i}\left(\mathbf{I} + \mathbf{G}_{0}\sum_{j\neq i}\mathbf{R}_{j}\right),$$

$$= \mathbf{t}_{i}\left(\mathbf{I} + \mathbf{G}_{0}\sum_{j\neq i}\mathbf{R}_{j}\right)$$

$$= \mathbf{t}_{i} + \mathbf{t}_{i}\mathbf{G}_{0}\sum_{j\neq i}\mathbf{t}_{j}\left(\mathbf{I} + \mathbf{G}_{0}\sum_{k\neq j}\mathbf{R}_{k}\right)$$

$$= \mathbf{t}_{i} + \sum_{j\neq i}\mathbf{t}_{i}\mathbf{G}_{0}\mathbf{t}_{j} + \sum_{j\neq i}\sum_{k\neq j}\mathbf{t}_{i}\mathbf{G}_{0}\mathbf{t}_{j}\mathbf{G}_{0}\mathbf{t}_{k} + \cdots.$$
(P3.73)

The second line of the above equation is noted to follow from (3.8). Substitution of (P3.73) into (P3.70) yields

$$\mathbf{T} = \sum_{i} \mathbf{t}_{i} + \sum_{i} \sum_{j \neq i} \mathbf{t}_{i} \mathbf{G}_{0} \mathbf{t}_{j} + \sum_{i} \sum_{j \neq i} \sum_{k \neq j} \mathbf{t}_{i} \mathbf{G}_{0} \mathbf{t}_{j} \mathbf{G}_{0} \mathbf{t}_{k} + \cdots$$
(P3.74)

Each succeeding index value is noted to be forced to differ from the one before it. However, there is no restriction on two indices not next to each other. Therefore, i = k is allowed in the third term.

# 4.1 Coherence and Homogenization

An important point which emerges from the treatment of wave scattering is the exponential decay of wave coherence over a length scale denoted the mean free path, determined by the imaginary part of the self energy. From the definition of self energy, (3.15), and the optical theorem, (3.40), it is seen that the magnitude of the mean free path is inversely correlated with the total scattering cross-section. If the scattering is weak, the waves can maintain their coherence over a long distance. In other words, the inhomogeneous medium would appear locally homogeneous to the propagating waves. In deriving the CPA condition, it has been further pointed out that a requisite condition for the CPA to be valid is the k independence of the self energy. As k implies spatial derivative (for example, Laplacian operator  $\nabla^2$  becomes  $-k^2$  in the wavevector representation), i.e., nonlocality in real space representation, the k independence therefore implies the absence of nonlocality. This can indeed be realized when both the size of the scatterers and their geometric correlation length are small compared to the wavelength, i.e., in the long wavelength regime, so that the structure of inhomogeneous medium is beyond the resolution limit of the wave. The inhomogeneous medium is therefore perceived as a uniform effective medium.

While it is known that all materials consist of hierarchical levels of discrete particles, i.e., molecules, atoms, electrons, nucleons, and elementary particles; the effective medium provides a physical basis for the perceived continuum that exists in our everyday reality.

In wave equation (3.4) the choice of  $\kappa_0$ , the reference homogeneous medium, is rather arbitrary. CPA gives a condition for determining an optimal  $\kappa_0$ , based on the requirement of zero overall scattering. In this chapter, we shall see that the CPA condition yields an approach for determining the effective medium properties from those of the inhomogeneous constituents plus their statistical characteristics: relative volume concentrations, geometric shapes, and correlations. In this context it should be noted that there is

an extensive literature on effective medium theories (Milton 2002), as they constitute an essential part of the study on composites, approached generally by using the mathematical technique of homogenization. From the present perspective this subject arises naturally as the coherent wave characteristics at the long wavelength limit.

In this chapter, we also show that a generalization of the CPA condition from zero overall scattering to minimum scattering, via the *spectral function approach*, can yield interesting information about quasimodes – those coherent waves which survive only for a short propagation distance or time – which manifest themselves only when the wavelength becomes comparable to or smaller than the size of the scatterers, i.e., in the intermediate frequency regime.

While coherent wave properties are obviously determined locally on a scale less than a few mean free paths, they are nevertheless important for setting the parameter values in the diffusion and localization regimes, where the scales involved are much larger than the mean free path. For example, the wave diffusion constant is determined by the coherent wave propagation velocity times the mean free path, both of which can be evaluated from the CPA or the spectral function approach. In wave localization it will be seen that both the criterion for determining the localization threshold, as well as the magnitude of the wave localization length, require inputs from the coherent wave properties for their evaluation. These associations show that while wave diffusion and localization are global wave characteristics, they are intimately related to the local properties of the coherent waves.

# 4.2 CPA: The Anderson Model

In the Anderson model, randomness is induced by random deviations of the atomic energy level at site **l**, denoted by  $\Delta \varepsilon_{\mathbf{1}}$ , with  $\langle \Delta \varepsilon_{\mathbf{1}} \rangle = 0$ ,  $\Delta \varepsilon_{\mathbf{1}}$  and  $\Delta \varepsilon_{\mathbf{m}}$  independent for  $\mathbf{l} \neq \mathbf{m}$ , i.e.,  $\langle \Delta \varepsilon_{\mathbf{1}} \Delta \varepsilon_{\mathbf{m}} \rangle = 0$ , and  $\Delta \varepsilon_{\mathbf{1}}$  has a flat distribution with width W:

$$D(\Delta \varepsilon_{1}) = \begin{cases} \frac{1}{W} & -W/2 \le \Delta \varepsilon_{1} \le W/2 \\ 0 & \text{otherwise.} \end{cases}$$
(4.1)

Here  $\Delta \varepsilon'_1$ s are all defined relative to the uniform medium. The equation defining the Green function is

$$\left[\left(\beta\Delta^{(2)} + \kappa_0^2(\omega) - \sigma(\mathbf{l})\right)\right] G^+(\omega, \mathbf{l}, \mathbf{l}') = \frac{1}{a^d} \delta_{\mathbf{l}, \mathbf{l}'},\tag{4.2}$$

where  $\sigma(\mathbf{l}) = 2m\Delta\varepsilon_{\mathbf{l}}/\hbar^2$ . To the leading order, the **T** operator in this case is just the addition of all the single-scatterer  $\mathbf{t}'s$ , each one given by (3.19):

$$\mathbf{T}^{+} \cong \sum_{\mathbf{l}} \mathbf{t}_{\mathbf{l}}^{+} = \sum_{\mathbf{l}} \frac{\sigma(\mathbf{l})a^{d}}{1 - \sigma(\mathbf{l})a^{d}G_{0}^{+}(\omega, \mathbf{l} = \mathbf{l}')} \left| \mathbf{l} \right\rangle \left\langle \mathbf{l} \right|.$$
(4.3)

The operation of configurational averaging  $\mathbf{t}_{\mathbf{l}}^+$  involves fixing the site  $\mathbf{l}$  and averaging over the  $\mathbf{t}^+$  matrix that appears on that site for all configurations. But that is equivalent to averaging over all possible values of  $\sigma(\mathbf{l})$ , i.e.,

$$\left\langle \mathbf{t}_{\mathbf{l}}^{+}\right\rangle_{c} = \left[\int \mathrm{d}\sigma D(\sigma) \frac{\sigma a^{d}}{1 - \sigma a^{d} G_{0}^{+}(\omega, \mathbf{l} = \mathbf{l}')}\right] \left|\mathbf{l}\right\rangle \left\langle \mathbf{l}\right|, \tag{4.4a}$$

where the label on  $\sigma(\mathbf{l})$  is dropped, and  $D(\sigma)$  denotes the distribution of the  $\sigma$  values. It becomes obvious that the value in the square bracket should be identical for every site **l** after configurational averaging. Therefore

$$\langle \mathbf{T}^+ \rangle_{\rm c} \cong \left[ \int \mathrm{d}\sigma D(\sigma) \frac{\sigma a^d}{1 - \sigma a^d G_0^+(\omega, \mathbf{l} = \mathbf{l}')} \right] \mathbf{I},$$
 (4.4b)

where  $\mathbf{I} = \Sigma_{\mathbf{l}} |\mathbf{l}\rangle \langle \mathbf{l}|$  is the identity operator. In accordance with the CPA recipe, we next try to make  $\langle \bar{\mathbf{T}}^+ \rangle_c$  vanish through shifting by an amount  $\Sigma^+$ . Since the shift has to be performed without altering the original problem as given by (4.2), the shift in  $\kappa_0^2$  must be accompanied by a similar shift in  $\sigma(\mathbf{l})$ , which means that the scattering potential also has to be renormalized relative to the effective medium. That is, the shift is simply the operation of rewriting the wave equation (4.2) as

$$\left[\beta\Delta^{(2)} + (\kappa_0^2(\omega) - \Sigma^+) - (\sigma(\mathbf{l}) - \Sigma^+)\right] G^+(\omega, \mathbf{l}, \mathbf{l}') = \frac{1}{a^d} \delta_{\mathbf{l}, \mathbf{l}'}.$$

With this shift, the CPA condition is given by [noting the  $\omega$  dependence of  $G_0^+$  is always implicitly defined by  $\kappa_0^2(\omega)$ ]:

$$\int d\sigma D(\sigma) \frac{\sigma - \Sigma^+}{1 - (\sigma - \Sigma^+) a^d G_0^+(\kappa_0^2 - \Sigma^+, \mathbf{l} = \mathbf{l}')} = 0.$$
(4.5)

To express this condition in a more useful form, it is noted that the integrand may be written as

$$D(\sigma) \left[ -1 + \frac{(a^d G_{\rm e}^+)^{-1}}{(a^d G_{\rm e}^+)^{-1} - (\sigma - \Sigma^+)} \right] \frac{1}{(G_{\rm e}^+)a^d},$$

which means

$$1 = \int \mathrm{d}\sigma \frac{D(\sigma)}{1 - (\sigma - \Sigma^+) a^d G_{\mathrm{e}}^+},\tag{4.6}$$

where  $G_{\rm e}^+ = G_0^+(\kappa_0^2 - \Sigma^+, \mathbf{l} = \mathbf{l}')$ . From (4.5) and (4.6) the CPA condition can be rewritten as

$$\int \mathrm{d}\sigma D(\sigma) \frac{\sigma}{1 - (\sigma - \Sigma^+) a^d G_0^+ (\kappa_0^2 - \Sigma^+, \mathbf{l} = \mathbf{l}')} = \Sigma^+.$$
(4.7)

This is the equation one can use to calculate  $\Sigma^+$  numerically by the method of iterative substitution; i.e., pick a value of  $\Sigma^+$ , calculate the left-hand side, and use that value as the new  $\Sigma^+$  to calculate the left-hand side until consistency is achieved. Usually a few iterations suffice in practice. For the flat distribution, (4.7) becomes

$$\frac{1}{2\sigma_{\rm w}} \int_{-\sigma_{\rm w}}^{\sigma_{\rm w}} \mathrm{d}\sigma \frac{\sigma}{1 - (\sigma - \Sigma^+) a^d G_0^+(\kappa_0^2 - \Sigma^+, \mathbf{l} = \mathbf{l}')} = \Sigma^+, \tag{4.8a}$$

where

$$2\sigma_{\rm w} = \frac{2mW}{\hbar^2}$$

is the total width of the distribution. The spatial dimensionality of the problem is contained in the form of the Green function  $G_0^+(\kappa_0^2 - \Sigma^+, \mathbf{l} = \mathbf{l}')$ . Since  $G_0^+$ is complex, the solution  $\Sigma^+$  is in general a complex number. The amount of randomness in the model is governed by  $\sigma_w$ . When  $\sigma_w = 0$ ,  $\Sigma^+ = 0$  necessarily, and the uniform medium is recovered.

# 4.2.1 Efficient Green Function Evaluation from DOS

For the 3D CPA, the evaluation of the Green function is perhaps the most time-consuming part of the numerical calculation. A useful and efficient approach to the calculation of the 3D Green function is to compile a numerical table of the density of states for the ordered lattice and store it in computer memory. The diagonal Green function can then be obtained from (P2.16), which involves only one integration. That is,

$$G_0^+(\kappa_0^2 - \Sigma^+, \mathbf{l} = \mathbf{l}') = \int \frac{\rho_0(\omega')}{\kappa_0^2 - \Sigma^+ - \kappa_0^2(\omega')} \,\mathrm{d}\omega'.$$
(4.8b)

The integration is over a finite domain of  $\omega'$ , since  $\rho_0(\omega')$  is nonzero only inside the band. Moreover, the imaginary part of  $\Sigma^+$  regularizes the singularity that appears in the integrand, thus making the numerical calculation quick and accurate.

Once  $\Sigma^+$  is obtained from (4.7), the mean free path l can also be evaluated because  $-\text{Im}\Sigma^+$  is proportional to a scattering time. However, in the lattice model the expression for l is not as simple as that for the continuum case, where  $G_e^+$  can be explicitly written down as in (3.52). In order to be consistent with the later development on wave diffusion, the expression for l is given by (5.110) in Chap. 5. In Figs. 4.1, 4.2, and 4.3 the density of states for the disordered system [which can be obtained from the imaginary part of the Green function as evaluated from (4.8b)] and its corresponding mean free path are plotted for 1D, 2D, and 3D, respectively, with the randomness parameter  $\sigma_{\rm w}a^2 = 0.5$ . It is interesting to observe that in the discrete lattice case, the CPA equation has solution for all frequencies inside the band. Comparison with Figs. 2.1, 2.3, and 2.5 shows that the effect of randomness is to both broaden the band and round off the singularities in the density of states.



Fig. 4.1. The density of states per site per unit frequency  $\hbar/ma^2$  (solid line) and the mean free path l/a (dashed line) for a 1D lattice with site randomness  $\sigma_w a^2 = 0.5$ .  $\beta$  is taken to be 1. The density of states should be compared with Fig.2.1 for the perfect lattice case. Here the dashed line is given by (5.110), with d = 1, and the solid line is given by  $-(2/\pi) \text{Im} F[u, \Sigma(u)a^2]$ , where  $F = (\text{sgn}[\text{Re}f])/(2\sqrt{f^2-1})$ ,  $f = u - \Sigma a^2/2$ ,  $u = (\hbar\omega - \varepsilon_0)/(\hbar^2/ma^2)$ , and  $\Sigma(u)$  is the solution to the CPA equation (4.8a) with the 1D Green function.



Fig. 4.2. The density of states per site per unit frequency  $\hbar/ma^2$  (solid line) and the mean free path l/a (dashed line) for a 2D square lattice with site randomness specified by  $\sigma_w a^2 = 0.5$ .  $\beta$  is taken to be 1. The density of states should be compared with Fig. 2.3 for the perfect lattice case. Here the dashed line is given by (5.110) with d = 2 and the solid line is given by  $-(2/\pi) \text{Im}F[u, \Sigma(u)a^2]$ , where  $F = (1/2\pi) \int_0^{\pi} d\theta(\text{sgn}[\text{Re}f])/(\sqrt{f^2 - 4\cos^2\theta}), f = u - \Sigma a^2/2, u = (\hbar\omega - \varepsilon_0)/(\hbar^2/ma^2)$ and  $\Sigma(u)$  is the solution to the CPA equation (4.8a) with the 2D Green function. The dip in the mean free path at the band center is due to the high density of scatterers at u = 0. The sharp decrease of the mean free path near the band edge, on the other hand, is attributable to the vanishing wave (group) velocity.



Fig. 4.3. The density of states per site per unit frequency  $\hbar/ma^2$  (solid line) and the mean free path l/a (dashed line) for a 3D simple cubic lattice with site randomness specified by  $\sigma_w a^2 = 0.5$ .  $\beta$  is taken to be 1. The density of states should be compared with Fig. 2.5 for the perfect lattice case. Here the dashed line is given by (5.110) for d = 3, and the solid line is given by  $-(2/\pi)\text{Im}F[u, \Sigma(u)a^2]$ , where  $f = (1/2\pi^2) \int_0^{\pi} d\theta_1 \int_0^{\pi} d\theta_2(\text{sgn}[\text{Re}f])/(\sqrt{f^2(\theta_1) - 4\cos^2\theta_2}), f = u - \Sigma(u)a^2 + \cos\theta_1,$  $u = (\hbar\omega - \varepsilon_0)/(\hbar^2/ma^2)$ , and  $\Sigma(u)$  is the solution to the CPA equation (4.8a) with the 3D Green function. The small value of l/a at the band center is directly correlated with the high density of states around u = 0.

# 4.3 CPA: The Classical Waves

For classical scalar waves, the scattering problem can be formulated from the discrete lattice point of view, where each scattering particle consists of infinitely many sites whose properties are correlated. The  $\mathbf{t}_i$  for a single particle i is now given by

$$\mathbf{t}_i = \mathbf{v}_i (\mathbf{I} - \mathbf{G}_0 \mathbf{v}_i)^{-1}.$$

where

$$\mathbf{v}_{i} = \sum_{\mathbf{r}} \left(\delta r\right)^{d} \sigma_{i}\left(\mathbf{r}\right) \left|\mathbf{r}\right\rangle \left\langle \mathbf{r}\right|, \qquad \mathbf{r} \in \text{inside particle } i.$$

A straightforward generalization of the CPA condition, (4.8), is to treat a *single particle* in the same manner as a single site in the Anderson model. The resulting condition is the matrix equation

$$\Sigma = \frac{1}{2\sigma_{\rm w}} \int_{-\sigma_{\rm w}}^{\sigma_{\rm w}} \mathrm{d}\sigma \,\mathbf{v}_i \,(\sigma) [\mathbf{I} - (\mathbf{v}_i \,(\sigma) - \Sigma a^d) \tilde{\mathbf{G}}_{\rm e}^+]^{-1}, \tag{4.9}$$

where we have used a function,

$$\tilde{\mathbf{G}}_{e}^{+}(\omega,\mathbf{r},\mathbf{r}') = H\left(\mathbf{r}\right)\mathbf{G}_{e}^{+}(\omega,\mathbf{r}-\mathbf{r}')H\left(\mathbf{r}'\right), \qquad (4.10a)$$

in the equation. H is a particle function defined as

$$H(\mathbf{r}) = \begin{cases} 1 & \text{for } \mathbf{r} \text{ inside the particle} \\ 0 & \text{else} \end{cases}$$
(4.10b)

to restrict the action of the Green function to within a single particle.

From (4.9) it is difficult to see that  $\Sigma$  should be proportional to the identity operator **I**, which is required if a single value of  $\Sigma$  were to satisfy the whole matrix equation. However, it turns out that there is indeed a low-frequency regime where the effective medium description holds. To demonstrate that, it is much simpler to look at the CPA condition in the k representation. Therefore we will not use (4.9) for our treatment of the CPA in the classical wave case. But the formal CPA condition, (4.9), is useful in demonstrating the Ward identity, shown in Chap. 5.

Since the CPA condition is

$$\bar{\Sigma}^{+} \cong \frac{N}{L^{d}} \left\langle \bar{\mathbf{T}}^{+} \right\rangle_{c} \cong \sum_{i} \frac{N}{L^{d}} \left\langle \bar{\mathbf{t}}_{i}^{+} \right\rangle_{c} = 0, \qquad (4.11a)$$

and since the configurationally averaged T-matrix is a function of only one **k** as discussed before, it follows immediately that it is only the *forward scattering* amplitudes that enters the CPA condition in the k representation(because  $\mathbf{k}' = \mathbf{k}$  implies forward scattering). Therefore,

$$\sum_{i} \frac{N}{L^{d}} \left\langle \bar{\mathbf{t}}_{i}^{+} \right\rangle_{c} = \frac{N}{L^{d}} \sum_{i} \frac{1}{N} \sum_{\kappa_{e}} \left\langle \kappa_{e} \right| \bar{\mathbf{t}}_{i}^{+} \left| \kappa_{e} \right\rangle \left| \kappa_{e} \right\rangle \left\langle \kappa_{e} \right| = 0, \quad (4.11b)$$

where  $\kappa_{\rm e}$  is a wave vector whose magnitude equals  $\sqrt{\kappa_0^2 - \text{Re}\Sigma^+}$ , which is the real part of  $\kappa_{\rm e}^*$  when  $\kappa_{\rm e}l \gg 1$ . By summing over the scatterer's index *i* and classifying the different scatterers into different species, one obtains the CPA condition as

$$\sum_{m} n_m \left\langle \mathbf{\kappa}_{\mathrm{e}} \right| \mathbf{\bar{t}}_m^+ \left| \mathbf{\kappa}_{\mathrm{e}} \right\rangle = 0, \qquad (4.12)$$

where  $n_m$  means the number density of the mth species of scatterers. In connection with (4.12) it should be pointed out that the real part of  $\langle \kappa_{\rm e} | \bar{\mathbf{t}}_m^+ | \kappa_{\rm e} \rangle$  is dominant at the long-wavelength limit, as seen by the solution to the single-scatterer problem. Therefore, it is usually that part, rather than the imaginary part, which enters the CPA equation. Because the real part does not represent the total scattering cross-section, as does the imaginary part, its sign can change from one species to the next, thus allowing the solution of (4.12) to exist.

## 4.3.1 The Role of Microstructure

The effective medium property of a continuous random medium is closely linked to its microstructure. Here the term microstructure refers to the volume fractions of the components plus the shapes and geometric arrangement

of the inhomogeneities. While the effective medium generally exists in the low-frequency limit where the individual scatterers are beyond the wave's resolution, the effective medium parameters can nevertheless be affected by presence or absence of some particular type of *statistical correlation* in a random composite that is represented by its microstructure. Such a statistical correlation generally extends throughout the sample and can thus affect its macroscopic properties. Therefore, an effective medium is uniquely determined only if the material properties of the components *and* their microstructure are both specified.

A useful approach to capturing the microstructural information of a random medium is through the concept of "structural unit(s)." This is illustrated in two commonly encountered microstructures, schematically depicted in Figs. 4.4a, b. In Fig. 4.4a, the two components of a random medium are symmetrical in the statistical sense; i.e., an interchange of the two components results in the same type of medium with interchanged volume fractions. This is denoted the symmetric microstructure. The same is not true for the microstructure depicted in Fig. 4.4b, where one component is always the dispersed phase in the matrix of the other. This is denoted the dispersion microstructure, like that of a colloid. For the symmetrical microstructure, there are two structural units, each approximated by a sphere (a circle in 2D, a line in 1D) of one material component. The volume fractions of the two units are p and 1 - p. For the dispersion microstructure, the dashed lines in the



Fig. 4.4. (a) The symmetric microstructure and its two structural units. On the right, the effective medium is indicated by the *cross-hatched region*. (b) The dispersion microstructure. In this case the *dashed lines* indicate a schematic division of the composite into structural units. On the right, the effective medium is indicated by the *cross-hatched region*.

left panel of Fig. 4.4b divide the medium into similar structural units, each consisting of a particle of dispersed phase coated by a layer of the matrix phase. By approximating the structural unit as a coated sphere, the coating layer thickness  $h = R_0 - R$  may be determined by the local volume fraction of the dispersed components:

$$h = R_0 - R = \left( (p')^{-1/d} - 1 \right) R, \tag{4.13}$$

where  $R_0$  is the radius of the total unit, R is the radius of the dispersed particle, and p' is its local volume fraction, which can fluctuate from one location to the next.

#### 4.3.2 The Symmetric Microstructure

To calculate the forward scattering amplitudes of a two-component composite in the symmetric microstructure, each structural unit, characterized by  $\kappa_m^2 = \varepsilon_m \kappa_0^2$  and radius  $R_m$  (m = 1, 2), is individually embedded in an effective medium with  $\kappa_e^2 = \bar{\varepsilon} \kappa_0^2$ . Since the calculation of the scattering amplitude has already been done in the last chapter, the answer may be directly quoted here:

$$\langle \kappa_{\mathbf{e}} | \, \overline{\mathbf{t}}_{m}^{+} | \kappa_{\mathbf{e}} \rangle = \begin{cases} \frac{4\pi i}{\kappa_{\mathbf{e}}} \sum_{n=0}^{\infty} \bar{D}_{n}^{(m)}(2n+1) & 3\mathrm{D}, \\ 4\mathrm{i} \sum_{n=-\infty}^{\infty} \bar{D}_{n}^{(m)} & 2\mathrm{D}, \\ 2\mathrm{i}\kappa_{\mathbf{e}}(\bar{\tau}^{(m)}-1) & 1\mathrm{D}, \end{cases}$$
(4.14)

where  $\theta$  has been set to 0 for the 3D and 2D cases, and  $\bar{D}_n^{(m)}$  and  $\bar{\tau}^{(m)}$  are given by (3.36), (3.42), and (3.50), respectively, for 3D, 2D, and 1D, in which  $\kappa_0$  is replaced by  $\kappa_e$ , and  $\kappa$  is replaced by  $\kappa_m$ .Combining (4.12) and (4.14) gives the CPA equation for the symmetric microstructure.

#### The Scalar Wave CPA for the Symmetric Microstructure

In the  $\kappa_{\rm e}R_m \to 0$  limit, the CPA equation for a two-component composite may be written down explicitly to the leading order in  $(\kappa_{\rm e}R_m)^d$ :

$$C_d \sum_{m=1}^{2} n_m (\kappa_{\rm e} R_m)^d \kappa_{\rm e}^{2-d} \left[ 1 - \left(\frac{\kappa_m}{\kappa_{\rm e}}\right)^2 \right] = 0, \qquad (4.15)$$

where  $C_d = 2, \pi$ , and  $4\pi/3$  for 1D, 2D, and 3D, respectively. Since  $C_d n_m R_m^d = p_m$ , the volume fraction of component m, we get

$$\bar{\varepsilon} = p\varepsilon_1 + (1-p)\varepsilon_2 \tag{4.16}$$

in all spatial dimensions. Alternatively, one can express the CPA equation in terms of the wave speeds  $\bar{v}$ ,  $v_1$  and  $v_2$ :

$$\frac{1}{\bar{v}^2} = \frac{p}{v_1^2} + \frac{1-p}{v_2^2}.$$
(4.17a)

### Wood's Formula and Bubbly Liquid

An important application of the effective medium equation is the prediction of sound speed in immiscible fluid – fluid mixtures, when the inhomogeneities are small compared to the wavelength. In this case, the sound speed in a uniform medium is given by  $v^2 = B/\rho$ , where B is the bulk modulus of the fluid and  $\rho$  denotes its mass density. The problem of acoustic wave scattering from a spherical inclusion of one fluid in the matrix of another involves boundary conditions which are slightly different from what have been used so far. Namely, if  $\phi$  = pressure P, then instead of  $\partial \phi/\partial r$  being continuous across a spherical interface, it should be the continuity of  $\rho^{-1}\partial \phi/\partial r$ . In the solution to Problem 4.1, this slight change is shown to result in the CPA equation

$$\frac{1}{\bar{B}} = \frac{p}{B_1} + \frac{1-p}{B_2},\tag{4.17b}$$

i.e.,  $v^2$  in (4.17a) is replaced by B. Equation (4.17b) is known as Wood's formula in the literature. A particularly interesting application of (4.17b) is to a bubbly liquid. Since B for a gas is on the order of its pressure, at atmospheric pressure this implies three to four orders of magnitude difference with the B of a liquid. Therefore, even the presence of 1% gas bubbles in a liquid would lower  $\bar{B}$  by two orders of magnitude from that of the liquid. This effect is analogous to squeezing a sandwich made of two steel plates with a sheet of foam rubber in between. The effective  $\bar{B}$  in that case is clearly dominated by that of the fluid by the presence of 1% gas [see (4.30e)], the acoustic wave speed in a bubbly liquid can thus be drastically reduced from the sound speed of the pure liquid. This effect has indeed been observed extensively. While not so dramatic, (4.17b) is also reasonably accurate for predicting the  $\bar{B}$  of mixtures consisting of solid particles dispersed in a fluid.

For higher frequencies, the CPA conditions (4.12) and (4.14) have to be solved numerically. Past experience has shown that a unique solution can generally be found up to the frequency where  $\kappa_e R_m$  is on the order of 1 but cannot be found beyond that because once the wave is able to discern the structure of individual scatterers, the effective medium description breaks down (see Sect. 4.5). Therefore, for classical waves in continuous random media the effective medium exists only when the condition  $\lambda \gg R$  is satisfied. It should be noted that sometimes for R very small, this condition can be satisfied at high frequencies, e.g., optical frequencies. Therefore, instead of stating the effective medium condition as being the low-frequency limit, it is perhaps more accurate to state it as being the *long-wavelength limit*, where "long" is measured in units of R.

# Mean Free Path Evaluation

The mean free path may also be explicitly calculated in the long-wavelength limit. In an effective medium, the scattering vanishes to the first order, so the averaged Green function may be written as

$$\langle G \rangle_{\rm c} = \frac{1}{\kappa_{\rm e}^2(\omega) - k^2 - \bar{\Sigma}(\omega)} \, \frac{N}{L^d},\tag{4.18}$$

where  $\operatorname{Re}\bar{\Sigma}(\omega) \cong NL^{-d} \langle \bar{T} \rangle_{c}(\omega) \cong 0$  to the leading order. However  $\langle \bar{\mathbf{t}} \rangle_{c}$  does not vanish to all orders. The mean free path information is contained in the next higher-order term of the  $\kappa_{e}R_{m}$  expansion of  $\bar{\Sigma}(\omega)$ , which is exactly the leading order *imaginary* part of  $\bar{\Sigma}(\omega)$ , given by

$$-\mathrm{Im}\bar{\Sigma}(\omega) = \Delta \cong -\mathrm{Im}\left[\sum_{m=1}^{2} n_m \langle \kappa_{\mathrm{e}} | \,\bar{\mathbf{t}}_m^+ \, | \,\kappa_{\mathrm{e}} \rangle\right]. \tag{4.19a}$$

By defining l as

$$\frac{1}{l} = 2 \mathrm{Im} \kappa^* = 2 \mathrm{Im} \sqrt{\kappa_{\mathrm{e}}^2 + i\Delta} \simeq \frac{\Delta}{\kappa_{\mathrm{e}}} = \sum_m n_m \bar{O}_m, \qquad (4.19\mathrm{b})$$

where the last equality is obtained through the optical theorem, with O defined as the total scattering cross-section evaluated in the effective medium. By gathering the results on the leading imaginary term of the scattering crosssection given in the last section, one gets

$$l = \frac{A_d \bar{\varepsilon}^2}{\kappa_{\rm e} [p(\kappa_{\rm e} R_1)^d (\bar{\varepsilon} - \varepsilon_1)^2 + (1 - p)(\kappa_{\rm e} R_2)^d (\bar{\varepsilon} - \varepsilon_2)^2]},$$
(4.20)

where  $A_d = 3$ ,  $4/\pi$ ,1 in 3D, 2D, 1D, respectively;  $\kappa_{\rm e}^2 = \bar{\varepsilon}\kappa_0^2$ ; and  $\bar{\varepsilon}$  is given by (4.16). It is noted that  $l \sim \kappa_{\rm e}^{-(d+1)} \sim \omega^{-(d+1)}$ , which is precisely the inverse frequency dependence of Rayleigh scattering. Therefore,  $l \to \infty$  as  $\omega \to 0$ , which implies that the effective medium is accurately valid at the  $\omega \to 0$  limit since not only  $\langle \bar{\mathbf{T}} \rangle_c = 0$ , but  $\bar{\mathbf{T}} = 0$  as well, so that the coherent wave propagation holds throughout the sample. Also, from (4.20) the mean free path is seen to be dependent on the particle size. If  $R_1 = R_2 = R$ , then

$$l = \frac{A_d \bar{\varepsilon}^2}{\kappa_{\rm e}^{d+1} R^d p (1-p) (\varepsilon_1 - \varepsilon_2)^2},\tag{4.21}$$

where (4.16) has been used. The  $R^d$  dependence emerges because the strength of Rayleigh scattering is sensitive to the ratio between the wavelength and the particle size.

#### The Electromagnetic CPA for the Symmetric Microstructure

For an important class of classical waves – electromagnetic waves – the effective medium equation described above has to be modified for the following two reasons. First, owing to the vector character of the transverse electromagnetic wave, its 2D and 3D scattering amplitudes in the  $\kappa_e R \rightarrow 0$  limit are always proportional to  $\cos \theta$ , i.e., dipolar in character, because the physical quantity here is the electric field **E**, which polarizes the scatterer in the scattering process. Since the projection of a vector on the radial coordinate **r** always involves an angle,  $P_1(\cos \theta) = \cos \theta$  is the lowest-order, non-isotropic term that enters. Second, in the electromagnetic equation the electric field may be written as  $\mathbf{E} = -\nabla \phi$  at the low-frequency limit. As a result, it follows the from the Maxwell equations that

$$\nabla \cdot \varepsilon \mathbf{E} = -\nabla \cdot \varepsilon \nabla \phi = 0$$

for systems without free charges. Thus across an interface it is the *normal* derivative of  $\phi$  times  $\varepsilon$  that is continuous. This can be seen from the usual arguments leading to the electromagnetic boundary conditions.

To incorporate these changes, we utilize the single-scatterer solutions worked out in the last chapter, modify them slightly to take into account the new boundary condition, and expand the n = 1 term ( $m = \pm 1$  in 2D) instead of the n = 0 term, in the  $\kappa_e R_m \to 0$  limit. This has been worked out in the solution to Problem 4.2. For d = 3 and 2, the resulting CPA equation is

$$p\frac{\bar{\varepsilon}-\varepsilon_1}{(d-1)\bar{\varepsilon}+\varepsilon_1} + (1-p)\frac{\bar{\varepsilon}-\varepsilon_2}{(d-1)\bar{\varepsilon}+\varepsilon_2} = 0.$$
(4.22a)

Equation (4.22a) is known in the literature as Bruggeman's symmetric effective medium equation (Bruggeman 1935). Also shown in the solution to Problem 4.2 is that in 1D, the modifications result in no change to the final answer, so the CPA equation remains the same as the n = 0 scalar case. Through similar considerations leading to (4.20), the electromagnetic mean free path in a symmetric-microstructure composite is obtained as

$$l_{em} = \begin{cases} \kappa_{e}^{-1} \left[ p(\kappa_{e}R_{1})^{3} \left( \frac{\bar{\varepsilon} - \varepsilon_{1}}{2\bar{\varepsilon} + \varepsilon_{1}} \right)^{2} + (1 - p)(\kappa_{e}R_{2})^{3} \left( \frac{\bar{\varepsilon} - \varepsilon_{2}}{2\bar{\varepsilon} + \varepsilon_{2}} \right)^{2} \right]^{-1} & \text{3D} \\ (2\pi\kappa_{e})^{-1} \left[ p(\kappa_{e}R_{1})^{2} \left( \frac{\bar{\varepsilon} - \varepsilon_{1}}{2\bar{\varepsilon} + \varepsilon_{1}} \right)^{2} + (1 - p)(\kappa_{e}R_{2})^{2} \left( \frac{\bar{\varepsilon} - \varepsilon_{2}}{2\bar{\varepsilon} + \varepsilon_{2}} \right)^{2} \right]^{-1} & \text{2D} \end{cases}$$

$$\left( \left( 2\pi\kappa_{\rm e}\right)^{-1} \left[ p(\kappa_{\rm e}R_1)^2 \left( \frac{\overline{\varepsilon} + \varepsilon_1}{\overline{\varepsilon} + \varepsilon_1} \right) + (1-p)(\kappa_{\rm e}R_2)^2 \left( \frac{\overline{\varepsilon} + \varepsilon_2}{\overline{\varepsilon} + \varepsilon_2} \right) \right] \quad \begin{array}{c} \text{2D,} \\ (4.22b) \end{array} \right)$$

where  $\bar{\varepsilon}$  is the solution to (4.22a).

For finite  $\kappa_{\rm e}R$  the electromagnetic wave scattering can have channels other than the one presented, which pertains to electric dipole scattering. Next in the order of importance is the magnetic dipole scattering (Bohren and Huffman 1983). When  $\kappa_{\rm e}R$  exceeds a certain value, the magnetic dipole scattering becomes dominant. However, the electric dipole scattering always dominates in the  $\kappa_{\rm e}R \rightarrow 0$  limit.

### Percolation Threshold in the Symmetric Microstructure

It should be noted that the Rayleigh frequency dependence and the particle size dependence of  $l_{em}$  are the same here as for the scalar wave scattering. However, the CPA condition (4.22a) gives a very different effective medium than that for the scalar wave CPA equation (4.16). In the solution to Problem 4.3, it is shown that (4.22a) predicts a *percolation threshold* when the material properties of the two components have infinite contrast, e.g., when  $\varepsilon_1 = 1$  and  $\varepsilon_2 = 0$ , the effective medium  $\overline{\varepsilon} = 0$  below a threshold  $p_c = 1/3$  in 3D ( $p_c = 1/d$ in general). This property has a simple geometric interpretation. Suppose a random medium is composed of an electrically conducting component whose material property is represented by  $\varepsilon_1 = 1$  and an insulating component whose material property is represented by  $\varepsilon_2 = 0$ . The electromagnetic boundary condition  $\bar{\varepsilon}(\partial \phi/\partial n) = \varepsilon_2(\partial \phi_2/\partial n)$ , where  $\partial/\partial n$  means the derivative normal to the interface, implies  $\partial \bar{\phi} / \partial n = 0$  at the interface with component 2; i.e., no electrical current can flow into component 2. The electrical current is therefore confined to component 1. In order for  $\bar{\varepsilon} \neq 0$ , it is clear that the  $\varepsilon_1$  component must be geometrically connected to form an infinite conducting network that spans the sample. Statistically, this proves to be possible only if p is greater than a threshold value  $p_{\rm c}$ , called the percolation threshold. In contrast, the CPA condition (4.16) gives no percolation threshold behavior because for the scalar wave, even if the electromagnetic boundary condition were applied so that  $\partial \phi / \partial n$  vanishes at the interface with component 2, the physical quantity in this case is  $\phi$ , which is still continuous across the interface. In contrast,  $\mathbf{E} = -\nabla \phi$  is the physical quantity in the electromagnetic case. The lesson here is that the (long-wavelength) expansion of the n = 0 term gives an effective medium theory for  $\phi$ , but the n = 1 term expansion gives an effective medium theory for  $\nabla \phi$ , with a percolation threshold.

#### The Elastic CPA for the Symmetric Microstructure

For completeness, the CPA equations for determining the elastic response of two-component composites in the symmetric microstructure case are given below. In isotropic elastic composites, there are three effective medium parameters to be determined: the two elastic moduli  $\beta$ ,  $\mu$  associated with the longitudinal and transverse (shear) waves, respectively, plus the mass density. The three relevant equations are

$$\frac{1}{\bar{\beta}} = \frac{p}{\beta_1 + \frac{4}{3}(\bar{\mu} - \mu_1)} + \frac{1 - p}{\beta_2 + \frac{4}{3}(\bar{\mu} - \mu_2)},$$
(4.23a)

$$\frac{1}{\bar{\mu} + H(\bar{\beta}, \bar{\mu})} = \frac{p}{\mu_1 + H(\bar{\beta}, \bar{\mu})} + \frac{1 - p}{\mu_2 + H(\bar{\beta}, \bar{\mu})},$$
(4.23b)

$$\bar{\rho} = p\rho_1 + (1-p)\rho_2,$$
(4.23c)

where

$$H(\bar{\beta},\bar{\mu}) = \frac{\bar{\mu}(9\beta - 4\bar{\mu})}{6\bar{\beta} + 4\bar{\mu}}.$$
(4.23d)

These coupled equations were first derived by Berryman (1979). It should be noted that in the elastic solid–liquid composite case, if p denotes the volume fraction of solid, then there is a rigidity percolation threshold at p = 0.6, and for  $0.4 \le p \le 0.6$  both components have continuous networks. That is, the composite is elastically bicontinuous. If the liquid is conducting and solid nonconducting, we note that in the present case the liquid percolation threshold, 0.4, is different from the conduction percolation threshold predicted by the Bruggeman (symmetric microstructure) equation. This difference in the percolation threshold values arises because the effective medium equations cannot accurately predict the threshold critical behavior, due to their inherent mean-field nature. But combining the predictions of (4.22) and (4.23), it can be stated that in the symmetric microgeometry, the two components are electrically and elastically bicontinuous in the range of  $0.4 \le p \le 0.6$ . This information is useful to the effective medium modeling of sedimentary rocks, shown below.

# 4.3.3 The Dispersion Microstructure

In the case of the dispersion microstructure, the same mathematical considerations are applied to a single type of structural unit, shown in Fig. 4.4b. In a random medium, the local volume fraction of the dispersed component can fluctuate from one region to the next. Therefore the CPA condition (4.12) should be expressed as

$$\int_{0}^{1} \mathcal{D}(p') \left\langle \kappa_{\mathrm{e}} \right| \overline{\mathbf{t}}^{+} \left| \kappa_{\mathrm{e}} \right\rangle_{p'} \mathrm{d}p' = 0, \qquad (4.24)$$

where  $\mathcal{D}(p')$  gives the distribution of local volume fraction p', and the forward scattering amplitude (noted to depend on p') is given in the solution to Problem 4.4.

### The Scalar Wave CPA for the Dispersion Microstructure

In the long-wavelength limit, (4.24) becomes

$$\bar{\varepsilon} = p\varepsilon_1 + (1-p)\varepsilon_2 \tag{4.25}$$

for 1D, 2D, and 3D, where p is the globally averaged volume fraction of the dispersed component 1. This equation is noted to be identical to that of the scalar wave CPA for the symmetric microstructure case. However, the mean free path expression is different:

$$l = \frac{B_d \bar{\varepsilon}}{(\kappa_e R_0)^d \kappa_e} \left\{ \int_0^1 \mathcal{D}(p') [\bar{\varepsilon} - (1 - p')\varepsilon_2 - p'\varepsilon_1]^2 \mathrm{d}p' \right\}^{-1}.$$
 (4.26)

Here  $B_d = 3\varepsilon_2$ ,  $4\varepsilon_2/\pi$ ,  $\bar{\varepsilon}$  for d = 3, 2, 1, respectively. In (4.24) and (4.26), a suitable form of the distribution function may be taken to be

$$\mathcal{D}(p')dp' = \mathcal{D}[\eta(p')]\frac{d\eta}{dp'}dp' = \frac{dp'}{\sqrt{2\pi}s_0p'(1-p')}\exp[-(\eta-\eta_0)^2/2s_0^2], \quad (4.27a)$$

where

$$\eta(p') = \ln \frac{p'}{1 - p'}.$$
(4.27b)

The two parameters  $\eta_0$  and  $s_0$  may be determined from the average volume fraction p and the variance of the local compositional fluctuations. In applying (4.25) and (4.26) to acoustic wave propagation in colloidal systems,  $\varepsilon$  should be replaced by the inverse of the bulk modulus as discussed earlier. Also, if the dispersed particles have a well-defined size, then  $R_0^d \approx R^d/p$  in (4.26) so that one may use the size of the dispersants directly. However, if the size of the particles varies over a large range, then  $R_0^d$  means its average value, i.e., the *dth* moment of the distribution function for  $R_0$ .

#### The Electromagnetic CPA for the Dispersion Microstructure

The electromagnetic case yields a different CPA equation for the dispersion microgeometry than that for the symmetric microgeometry in 3D and 2D. The details are worked out in the solution to Problems 4.5. The answer is

$$\int_0^1 \mathcal{D}(p') \,\mathcal{Q}(p') \mathrm{d}p' = 0, \qquad (4.28a)$$

where

$$\mathcal{Q}(p') = \frac{(\bar{\varepsilon} - \varepsilon_2)[\varepsilon_1 + (d-1)\varepsilon_2] + p'[\bar{\varepsilon} + (d-1)\varepsilon_2](\varepsilon_2 - \varepsilon_1)}{[\varepsilon_1 + (d-1)\varepsilon_2][\varepsilon_2 + (d-1)\bar{\varepsilon}] + (d-1)p'(\bar{\varepsilon} - \varepsilon_2)(\varepsilon_2 - \varepsilon_1)}.$$
(4.28b)

The corresponding mean free path expression is given by

$$l_{\rm em} = \frac{B'_d \varepsilon_2 / \bar{\varepsilon}}{\kappa_{\rm e} (\kappa_{\rm e} R_0)^d} \left[ \int_0^1 \mathcal{D}(p') \mathcal{Q}^2(p') \mathrm{d}p' \right]^{-1}, \qquad (4.28c)$$

where the  $B'_d = 1, 2/\pi$  for d = 3, 2, respectively. When,  $\mathcal{D}(p') = \delta(p' - p)$ , (4.28a) reduces to

$$\frac{\bar{\varepsilon} - \varepsilon_2}{\bar{\varepsilon} + (d-1)\varepsilon_2} - p \frac{\varepsilon_1 - \varepsilon_2}{\varepsilon_1 + (d-1)\varepsilon_2} = 0, \qquad (4.29)$$

which is known as the Maxwell-Garnett theory in the literature (Maxwell-Garnett 1904). It was first derived through the Clausius–Mossotti relation. In the present derivation, much more emphasis is given to the role of the microstructure.

In 1D, both the CPA equation and the expression for  $l_{em}$  are shown to be identical to (4.25) and (4.26), respectively.

Comparison of (4.29) and (4.22a) shows differences in their predictions, the most significant of which is the absence of a percolation threshold for  $\bar{\varepsilon}$  in the dispersion microgeometry. This is understandable because the dispersed phase is always prevented from forming a connected network by the presence of the coating, and therefore at the infinite contrast limit  $\bar{\varepsilon}$  reflects the property of the matrix phase (the coating component) only.

### Dielectric Anomaly in the Dispersion Microstructure

There is an optical consequence from (4.29), known as the dielectric anomaly (Cohen et al. 1973; Abeles et al. 1976), which has been widely observed in metal-insulator composites. If one considers a metal sphere coated by a layer of insulating dielectric (coated sphere being much smaller than the relevant wavelength), it is relatively straightforward to see that the elementary Mie resonance of the metal sphere, generally in the ultraviolet frequency range, would be modified and shifted downward in frequency (to the optical visible regime) by the presence of the dielectric coating, in the manner of adding an capacitive element to the resonance. This is precisely the dielectric anomaly phenomenon, which manifests itself as a resonance in the optical dielectric constant of the metal-insulator composites (see Fig. 4.7). Thus thinner the dielectric coating (increasing p and thus larger the capacitance), lower the frequency at which the dielectric anomaly would occur (Cohen et al. 1973). The dielectric anomaly is absent from the predictions of symmetric microgeometry effective medium equation (4.22a).

It should also be noted that in the limit of  $p \to 0$ , (4.22a) and (4.29) have the same expansion to first order in  $\Delta p$ :

$$\frac{\bar{\varepsilon}}{\varepsilon_2} \cong 1 + d\Delta p \frac{\varepsilon_1 - \varepsilon_2}{\varepsilon_1 + (d-1)\varepsilon_2} + \cdots$$

This fact is useful in the dilute limit.

#### The Elastic CPA for the Dispersion Microstructure

For elastic composites with the 3D dispersion microstructure, the relevant equations for the elastic constants are as follows:

$$\bar{\beta} = \frac{\beta_2 [4(\mu_2 - \mu_1) + 3\beta_1]}{3p\beta_2 + (1 - p)[4(\mu_2 - \mu_1) + 3\beta_1]} - \frac{4}{3}(\mu_2 - \bar{\mu}), \qquad (4.30a)$$

$$A\left(\frac{\bar{\mu}}{\mu_2}\right)^2 + B\left(\frac{\bar{\mu}}{\mu_2}\right) + C = 0, \qquad (4.30b)$$

where

$$A = -160(3k_2 + 2)R_{10}p^{10/3} - 2000R_7p^{7/3} + 10080R_5p^{5/3} -2000(3k_2^2 - 4k_2 + 3)R_3p - 80(2k_2 + 3)R_0 B = -80(3k_2 - 8)R_{10}p^{10/3} + 4000R_7p^{7/3} - 20160R_5p^{5/3} +1500(5k_2^2 - 14k_2 + 8)R_3p - 30(k_2 - 16)R_0 C = 80(9k_2 - 4)R_{10}p^{10/3} - 2000R_7p^{7/3} + 10080R_5p^{5/3} -250(27k_2^2 - 52k_2 + 24)R_3p + 10(19k_2 - 24)R_0$$

with

$$\begin{split} R_{10} &= (k_5 - 1)(9k_2k_5 - 4k_5 + 6k_2 + 4)(38k_2k_3k_5 - 48k_3k_5 \\ &+ 57k_2k_5 - 75k_5 - 38k_2k_3 - 57k_3 + 48k_2 + 72) \\ R_7 &= (9k_2k_5 - 4k_5 + 6k_2 + 4)(54k_2^2k_3k_5^2 - 104k_2k_3k_5^2 \\ &+ 48k_3k_5^2 + 81k_2^2k_5^2 - 156k_2k_5^2 + 72k_5^2 + 3k_2^2k_3k_5 \\ &- 35k_2k_3k_5 + 9k_3k_5 - 90k_2^2k_5 + 252k_2k_5 - 144k_5 \\ &- 57k_2^2k_3 + 76k_2k_3 - 57k_3 + 72k_2^2 - 96k_2 + 72), \\ R_5 &= (k_2 - 1)^2(k_5 - 1)(9k_2k_5 - 4k_5 + 6k_2 + 4) \\ &(16k_3k_5 + 24k_5 + 19k_3 - 24), \\ R_3 &= R_5/(k_2 - 1)^2, \end{split}$$

 $R_0 = (9k_2k_5 - 4k_5 + 6k_2 + 4)^2(16k_3k_5 + 24k_5 + 19k_3 - 24).$ 

In the above,  $k_2 = \beta_2/\mu_2$ ,  $k_3 = \beta_1/\mu_1$ , and  $k_5 = \mu_2/\mu_1$ . Equation (4.30a) was first derived as a bound for the bulk modulus of composites (Kerner 1956; Hashin and Shtrikman 1961), and (4.30b) was first obtained by Christensen and Lo (1979). A particular simple equation results from (4.30) when the matrix is liquid, i.e.,  $\mu_2 = 0$ . In that case the effective medium  $\bar{\mu} = 0$  also, since in the dispersion microgeometry the matrix component is the only continuous phase. From (4.30a), we see that for solid (or fluid) component dispersed in a liquid matrix,

$$\frac{1}{\overline{\beta}} = \frac{p}{\beta_1} + \frac{1-p}{\beta_2}.$$
(4.30c)

Equation (4.30c) is the generalized form of the Wood's formula (4.17b), now derived via the route of elastic CPA. It is seen that the simple form of the formula is retained even when the dispersed particles are solid rather than liquid.

It should be noted that there is another regime in the limit of  $\mu_2 \to 0$ . The reason is that the shear wave's wavevector  $k_s = \omega/\sqrt{\mu_2/\rho_2}$  can have two different limits when both  $\omega \to 0$  and  $\mu_2 \to 0$ , depending on which limit is taken first. If the  $\omega \to 0$  limit were taken first, then  $k_s a \to 0$ , a being the typical feature size, and we have the result shown above. However, if

the  $\mu_2 \to 0$  limit were taken first, then  $k_{\rm s}a \to \infty$  (while the longitudinal wavevector  $k_p = \omega/\sqrt{\beta_2/\rho_2}$  is still small so that  $k_pa \sim 0$ ), and the effective medium in that regime is very different (Berryman 1980):

$$\frac{1}{\bar{\beta}} = \frac{1-p}{\beta_2} + \frac{p}{\beta_1 - \frac{4}{3}\mu_1},$$
(4.30d)

$$\frac{\bar{\rho}}{\rho_2} = \frac{(2+p)\rho_1 + (1-p)\rho_2}{2(1-p)\rho_1 + (1+2p)\rho_2}.$$
(4.30e)

In (4.30d), if  $\mu_1 = 0$  (i.e., both components are fluids), then we recover (4.30c). But (4.30e) is quite different. Physically, this can be understood from the fact that when  $\mu_2 \rightarrow 0$  first, the liquid can be partially decoupled from the solid inclusions if the impedance mismatch between the solid and the fluid is large (Mei et al. 2006). The effective mass engaged by wave motions can thus differ from the arithmetic average value. This difference can be especially significant when there is large contrast in the components' mass densities. However, at  $p \sim 0$  (4.30e) becomes  $\bar{\rho}/\rho_2 \cong 1 + 3p(\rho_1 - \rho_2)/(2\rho_1 + \rho_2)$ , so the effect of significantly lower sound speed in bubbly liquid is still valid. In general, (4.30e) is the mass density formula for the fluid matrix case, whereas (4.23c) is the mass density expression for the solid matrix case.

# 4.4 Effective Medium Modeling of Inhomogeneous Materials

The application of effective medium theories is illustrated below in modeling the properties of two inhomogeneous material systems: granular metals and sedimentary rocks (perhaps one of the most prevalent forms of earthly matters). In both cases the two basic microstructures (symmetric and dispersion) are utilized as starting points, with subsequent theoretical frameworks based on physical inputs.

#### 4.4.1 Granular Metals

In many metal-insulator composite films both the phenomena of conduction percolation threshold and optical dielectric anomaly were observed. Hence neither the symmetric microgeometry nor the dispersion microgeometry can give an adequate description of the experimental results. For granular metals (GM), usually formed by co-sputtering or co-evaporation of metal and insulator, there has been extensive characterization of the films and measurements of their electrical and optical properties. Typical feature sizes in granular metal films can range from a few nanometers to tens of nanometers or submicrons. In many cases the feature sizes are small compared to the optical wavelength. Hence optical characteristics may be predicted using the effective medium approach. The GM effective medium equation(s) was formulated to overcome the defects of both the symmetric and the dispersion theories so that its predictions can give realistic account of the experimental observations (Sheng 1980).

#### GM Microstructure and its Effective Medium Predictions

In the GM microgeometry, there are two possible structural units when the composite film is formed through surface diffusion of sputtered or evaporated molecules: Component 1 forms the grain and component 2 the coating, denoted a type-1 unit; or component 2 forms the grain and component 1 the coating, denoted a type-2 unit. This is illustrated in Fig. 4.5. The relative probability of occurrence for the two units can be estimated by counting the number of equally possible final configurations (corresponding to different positions of the grain inside the region). By assuming the grains to be spherical, the number of configurations for type-1 unit is proportional to  $v_1 = (1-p^{1/3})^3$ , the free volume (up to a constant multiplicative factor) accessible by the grain center of mass inside the spherical region. The same reasoning leads to  $v_2 = [1 - (1-p)^{1/3}]^3$  for the type-2 unit. Thus the relative probability of occurrence for type-1 unit is  $f = v_1/(v_1 + v_2)$ , and that for type-2 unit is 1-f. Following the recipe for formulating the (electromagnetic) effective medium theory, the relevant equation is thus given by

$$fD_1 + (1 - f)D_2 = 0, (4.31)$$

where  $D_{1(2)}$  denotes the dipole moment induced on structural unit 1(2) by a static unit electric field when it is embedded in an effective medium with dielectric constant  $\bar{\varepsilon}$ . Equation (4.31) follows from the fact that in the long wavelength limit, the electromagnetic forward scattering amplitude is proportional to the induced dipole moment.

For f = 1, (4.31) can be reduced to the Maxwell-Garnett equation relevant to the dispersion microgeometry. In the present case, if we wish to consider the shape of the grains to deviate from spherical, then the general expressions for the orientationally averaged dipole moment of coated confocal spheroidal particles are

$$D_{1} = \frac{2}{3}D[\bar{\varepsilon}, \varepsilon_{1}, \varepsilon_{2}, p, A(\alpha, u), B(\alpha)] + \frac{1}{3}D[\bar{\varepsilon}, \varepsilon_{1}, \varepsilon_{2}, p, 3 - 2A(\alpha, u), 3 - 2B(\alpha)],$$
(4.32a)



Fig. 4.5. The two structural units for the GM microgeometry at the same value concentration of component 1 (*shaded*).
$$D_{2} = \frac{2}{3}D[\bar{\varepsilon}, \varepsilon_{2}, \varepsilon_{1}, 1 - p, A(\beta, w), B(\beta)] + \frac{1}{3}D[\bar{\varepsilon}, \varepsilon_{2}, \varepsilon_{1}, 1 - p, 3 - 2A(\beta, w), 3 - 2B(\beta)].$$

$$(4.32b)$$

Here  $\alpha$  and  $\beta$  are the ratio between the minor (major) and major (minor) axes of the elliptic cross-sections for the type-1 and type-2 oblate (prolate) spheroidal units, respectively,  $u = (p/\alpha)^{1/3}$ ,  $w = [(1-p)/\beta]^{1/3}$ , and D, A, and B have the following functional forms:

$$D(\bar{\varepsilon}, x, y, \theta, A, B) = \frac{[A\bar{\varepsilon} + (3 - A)y](y - x)\theta + [Bx + (3 - B)y](\bar{\varepsilon} - y)}{A(3 - A)(\bar{\varepsilon} - y)(y - x)\theta + [Bx + (3 - B)y][Ay + (3 - A)\bar{\varepsilon}]}, \quad (4.32c)$$

$$A(\gamma,\vartheta) = \frac{3}{2(1-\gamma^2)\vartheta^3} \left[ \frac{1}{\sqrt{1-\gamma^2}} \tan^{-1} \frac{\vartheta\sqrt{1-\gamma^2}}{\sqrt{s^2+\gamma^2\vartheta^2}} - \vartheta(s^2+\gamma^2\vartheta^2) \right]_{s\to0} (4.32d)$$
$$B(\gamma) = A(\gamma,\gamma^{-1/3}). \tag{4.32e}$$

In (4.32d), s is the solution of the equation  $(s^2 + \vartheta^2)^2(s^2 + \gamma^2\vartheta^2) = 1$ , and  $\tan^{-1}iz \rightarrow i \tanh^{-1}z$  when  $\gamma > 1$ .

The predictions of (4.32) have similarities to both the symmetric and dispersion microgeometries. As shown in Fig. 4.6, the DC conductivity exhibits a percolation threshold as in the symmetric case.

However, the (grain shape dependent) curvature in the conductivity vs. metal fraction p and the percolation threshold value make the present theory distinct from Bruggeman's theory, while at the same time offers a better account for the experimental data. At optical frequencies the present theory has features, such as the absorption peaks seen in Fig. 4.7, which are common in the dispersion microgeometry case as manifestation of the dielectric anomaly phenomenon. However, both the position and width of the peak, as well as the infrared behaviors are different. Overall, the present theory offers a unified DC and optical account of the experimental observations.

# 4.4.2 Sedimentary Rocks

For sedimentary rocks, the distribution of the feature sizes (of the scatterers) is not a peaked function. Rather it was found that for such systems the distribution can be very broad, following a power-law form over an extended size regime (Wong et al. 1986; Thompson et al. 1987). The *differential effective medium* (DEM) *microgeometry* is applicable to such media (Sen et al. 1981, Sheng and Callegari 1984). The basis of the DEM microgeometry is the *hierarchical* nature of the microstructure, shown schematically in Fig. 4.8. In this case mathematically one starts with a matrix component 2, adds a



Fig. 4.6. Normalized conductivity plotted as a function of metal volume fraction p for samples of W-Al<sub>2</sub>O<sub>3</sub> cermets (Abeles et al. 1975). Solid lines are from the predictions of the theory based on the GM microgeometry (Sheng 1980). Dashed line denotes Bruggeman's effective medium prediction

small fraction of the smallest-sized particles of component 1, and *regards* the resulting material as the new homogeneous matrix component 2', into which is added a small fraction of the next smallest-sized particles of component 1. This process of homogeneization is iterated for progressively larger particles of component 1, until the desired volume fraction p is reached. This process is shown schematically in Fig. 4.9. Such a hierarchical DEM microgeometry can yield a consistent description for the electrical and elastic properties of sedimentary rocks, shown below.

#### Mathematical Formulation of DEM

The basic mathematical process for obtaining the DEM effective dielectric constant in the two component case is to start with a homogeneous component 2 and replace a small fraction of it by component 1. The resulting effective medium property can be calculated as:

$$F(\bar{\varepsilon}, 1 - \Delta p, \varepsilon_2, \Delta p, \varepsilon_1) = 0, \qquad (4.33)$$

where F = 0 represents the effective medium condition for either the symmetric or the dispersion microgeometry case (the two are identical in the limit of



Fig. 4.7. Optical transmission as a function of light wavelength for a series of Au-SiO<sub>2</sub> composites (Cohen et al. 1973). The curves are displaced with respect to one another for clarity. The theoretical curves are normalized to the experimental values at the wavelength value of  $0.3 \,\mu\text{m}$ . The theoretical values of metal volume fraction p are labeled to the right of pairs of curves, whereas the experimental values of p and the film thickness are given above each pair of curves. The dip and peak in each curve (for  $p \leq 65\%$ ) are the manifestation of the dielectric anomaly phenomenon (from Sheng 1980).



Fig. 4.8. A schematic picture of the hierarchical microgeometry (applicable to the DEM), in which the large grains are surrounded by smaller grains, which in turn are surrounded by even smaller grains. In application to sedimentary rocks, the matrix is taken to be "cement," whose properties are obtainable from the symmetric microgeometry.



Fig. 4.9. A schematic diagram showing the iteration steps in the differential effective medium scheme. The starting medium (*white*), together with a small fraction of the added component (*solid circles*), are homogenized to a new starting medium (*light gray*). This step is then repeated.

 $\Delta p \to 0)$  . Differentiation of (4.33) yields the differential parameter changes induced by the replacement step:

$$\left(\frac{\partial F}{\partial \bar{\varepsilon}}\right)_{p=0} \frac{\mathrm{d}\bar{\varepsilon}}{\mathrm{d}p} = -\left(\frac{\partial F}{\partial p}\right)_{p=0},\tag{4.34}$$

or

$$\mathrm{d}\bar{\varepsilon} = -\frac{\left(\frac{\partial F}{\partial p}\right)_{p=0}}{\left(\frac{\partial F}{\partial \bar{\varepsilon}}\right)_{p=0}}\mathrm{d}p.$$
(4.35)

As this process is iterated, the relative concentration of component 1 will be increased. However, in any given iteration, the replacement of the homogeneous medium at that point by component 1 will increase the actual amount of component 1 by only

$$dp_{actual} = (1 - p_{actual})dp, \qquad (4.36)$$

since the composite already has a fraction  $p_{\text{actual}}$  of component 1. Provided the condition (4.36) is noted, at any given replacement step (4.33) is always valid if  $\varepsilon_2$  is replaced by  $\bar{\varepsilon}' = \bar{\varepsilon} - d\bar{\varepsilon}$ . Thus the final differential equation for the DEM microgeometry is

$$\mathrm{d}\bar{\varepsilon} = -\frac{\left(\frac{\partial F}{\partial p}\right)_{p=0}}{\left(\frac{\partial F}{\partial \bar{\varepsilon}}\right)_{p=0}} \frac{\mathrm{d}p}{(1-p)},\tag{4.37}$$

where we have used (4.36) but dropped the subscript "actual." For the purpose of modeling sedimentary rocks, in the following the implications of DEM will be pursued only in 3D.

#### **Electrical DEM Equation**

By using either the Bruggeman's (symmetric microstructure) effective medium equation or the Maxwell-Garnett (dispersion microstructure) equation for F, (4.37) can be readily integrated with the initial condition  $\bar{\varepsilon} = \varepsilon_2$  at p = 0:

$$1 - p = \left(\frac{\bar{\varepsilon} - \varepsilon_1}{\varepsilon_2 - \varepsilon_1}\right) \left(\frac{\varepsilon_2}{\bar{\varepsilon}}\right)^{1/3}.$$
(4.38)

In particular, if we let  $\varepsilon_1 = 0$ , then (4.38) acquires the simple form  $\bar{\varepsilon} = \varepsilon_2 (1-p)^{3/2}$ .

#### **Elastic DEM Equations**

The DEM effective elastic moduli equations can be similarly obtained, e.g., from (4.23):

$$\mathrm{d}\bar{\beta} = F^{\beta}(\bar{\beta}, \bar{\mu}, \beta_1, \mu_1) \frac{\mathrm{d}p}{(1-p)}, \qquad (4.39a)$$

$$d\bar{\mu} = F^{\mu}(\bar{\beta}, \bar{\mu}, \beta_1, \mu_1) \frac{dp}{(1-p)},$$
 (4.39b)

where

$$F^{\beta} = 3\bar{\beta} \frac{8(\mu_1 - \bar{\mu})[4\bar{\mu}(\mu_1 - \bar{\mu}) + \bar{\beta}(\mu_1 + 2\bar{\mu}) - 3\bar{\mu}\beta_1] + 3\bar{\beta}(\bar{\beta} - \beta_1)(2\mu_1 + 3\bar{\mu})}{(4\mu_1 - 4\bar{\mu} - 3\beta_1)(4\bar{\mu}\mu_1 + 6\bar{\beta}\mu_1 - 4\bar{\mu}^2 + 9\bar{\mu}\bar{\beta})}$$
(4.39c)

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$$F^{\mu} = \frac{15\beta\bar{\mu}(\mu_1 - \bar{\mu})}{4\bar{\mu}\mu_1 + 6\bar{\beta}\mu_1 - 4\bar{\mu}^2 + 9\bar{\mu}\bar{\beta}}.$$
 (4.39d)

The initial conditions are  $\bar{\beta} = \beta_2$ ,  $\bar{\mu} = \mu_2$  at p = 0. Numerical integration is necessary to yield predictions in the elastic case.

# Unified Electric and Elastic Description by Three-Component DEM

An important application of the DEM effective medium is the unified description of electric and elastic properties of sedimentary rocks. It is widely known that in spite of the apparent randomness in their microstructures, sedimentary rocks' electrical and elastic characteristics display approximate correlations with porosity  $\phi$  that have come to be known, respectively, as the Archie's law (Archie 1942) and the Wyllie equation (Wyllie et al. 1956). Archie's law states that the electrical conductivity  $\bar{\sigma}$  of a fluid-filled sandstone may be expressed as  $\bar{\sigma} = \sigma_f \phi^m$ , where  $\sigma_f$  denotes the fluid conductivity, and  $m \approx 2$ . Viewed in the framework of the percolation theory, Archie's law implies that sandstones are inhomogeneous composites with a microstructure that insures the persistence of its pore connectivity to very low porosities. The nonlinear exponent value of  $m \cong 2$  is also suggestive that only a subset of the pore space actually contributes to conduction (for example, at  $\phi = 0.1$ only 10% of the pore volume is effective), i.e., the pore diameters are larger than the connecting throats that control the conduction. Wyllie equation, on the other hand, is given by  $\bar{v}_{\rm p}^{-1} = (1-\phi)v_{\rm solid}^{-1} + \phi v_{\rm fluid}^{-1}$ , where  $\bar{v}_{\rm p}$  denotes the longitudinal p-wave velocity. It is not fortuitous that the DEM microgeometry can give a good account of the observed correlations. The diversity in the feature size, extending over several decades and following a fractal distribution, has been documented (Thompson et al. 1987). As shown later, Archie's law and Wyllie equation are the manifestations of this salient characteristic of the sedimentary rocks.

To apply the DEM equations, it is important to note that a feature of the DEM microgeometry is that the starting medium will always remain the connected "matrix" medium (Yonezawa and Cohen 1983), in which the successively larger particles of another component are dispersed. Hence for application to rocks the initial medium should be bicontinuous so that the resulting medium can exhibit elastic rigidity on the one hand, and also be electrically conducting through the pore fluid on the other. A three-component DEM theory was proposed to satisfy these requirements (Sheng 1990) in which the starting phase comprises a mixture of fluid and cement, with a pore fluid fraction of  $0.4 \le \phi_0 \le 0.6$ , i.e., in the range of bicontinuity.

#### Electrical Three-Component DEM Equation—Archie's Law

The electrical property (Archie's law) may be obtained by considering a generalization of the Bruggeman's equation (symmetry microgeometry) for

spheroidal grains with depolarization coefficients  $L_1 = L$  and  $L_2 = L_3 = (1-L)/2$ :

$$F(\bar{\varepsilon}, 1-p, \varepsilon_2, p, \varepsilon_1) = \sum_{i=1}^{3} \left[ \frac{(1-p)(\bar{\varepsilon}-\varepsilon_2)}{(1-L_i)\bar{\varepsilon}+L_i\varepsilon_2} + \frac{p(\bar{\varepsilon}-\varepsilon_1)}{(1-L_i)\bar{\varepsilon}+L_i\varepsilon_1} \right] = 0.$$
(4.40)

Here p denotes the volume fraction of particles, to be distinguished from that of cement. In accordance with the simple procedure outlined above, the electrical DEM equation is obtained by integrating (4.37) with the initial condition  $\bar{\varepsilon} = \varepsilon_2$  at p = 0, and a starting porosity  $\phi_0$  in the bicontinuous fluid-cement composite:

$$\frac{\phi}{\phi_0} = \left[\frac{\varepsilon_2}{\bar{\varepsilon}}\right]^a \left[\frac{\varepsilon_2 + \gamma\varepsilon_1}{\bar{\varepsilon} + \gamma\varepsilon_1}\right]^b \left[\frac{\bar{\varepsilon} - \varepsilon_1}{\varepsilon_2 - \varepsilon_1}\right],\tag{4.41a}$$

where the parameters  $a, b, \gamma$  are given by

$$a = 3L(1-L)/(3L+1),$$
 (4.41b)

$$b = 2(3L - 1)^2 / [(5 - 3L)(1 + 3L)], \qquad (4.41c)$$

$$\gamma = (1+3L)/(5-3L). \tag{4.41d}$$

In the three-component model, the starting component  $\varepsilon_2$  is obtained from (4.40) with fraction  $(1 - \phi_0)$  of cement  $(\varepsilon_c)$ ,  $(\phi_0)$  of fluid  $(\varepsilon_f)$ , and L = 1/3:

$$\varepsilon_2 = \frac{1}{4} \left[ N + \sqrt{N^2 + 8\varepsilon_c \varepsilon_f} \right], \qquad (4.42a)$$

where

$$N = (2 - 3\phi_0)\varepsilon_{\rm f} - (1 - 3\phi_0)\varepsilon_{\rm c}.$$
 (4.42b)

To make the connection with Archie's law, we focus on the imaginary part of the complex dielectric constant by setting  $\bar{\varepsilon} = \bar{\sigma}$ ,  $\varepsilon_{\rm f} = \sigma_{\rm f}$ , and  $\varepsilon_{\rm c} = \varepsilon_1 = 0$  in (4.41a) and (4.42) to obtain

$$\bar{\sigma} = C\sigma_{\rm f}\phi^m,\tag{4.43a}$$

$$m = (5 - 3L)/[3(1 - L^2)],$$
 (4.43b)

$$C = (3\phi_0 - 1)/2\phi_0^m. \tag{4.43c}$$

Equation (4.43a) is in the form of generalized Archie's law. Requiring m = 2 yields L = 0.73 (oblate spheroid with an aspect ratio of 4.5), which in turn gives a = 0.186, b = 0.314, and  $\gamma = 1.132$ . In addition, C = 1 implies  $\phi_0 = 0.5$ .

#### **Elastic Three-Component DEM Equations**

With the values of L and  $\phi_0$  fixed by Archie's law, the same model can be used to predict the speed of elastic waves. The differential equations are

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$$(1-p)\frac{\mathrm{d}\beta}{\mathrm{d}p} = g^{\beta},\tag{4.44a}$$

$$(1-p)\frac{\mathrm{d}\bar{\mu}}{\mathrm{d}p} = g^{\mu},\tag{4.44b}$$

where

$$g^{\beta,\mu} = \bar{\beta} \frac{\sum_{n,l=0}^{4} \left( A_{n,l}^{\beta,\mu} \mu_1 + B_{n,l}^{\beta,\mu} \beta_1 \right) \left( \frac{\bar{\mu}}{\mu_1} \right)^n \left( \frac{\bar{\beta}}{\bar{\mu}} \right)^l}{\sum_{n,l=0}^{4} \left( C_{n,l} \mu_1 + D_{n,l} \beta_1 \right) \left( \frac{\bar{\mu}}{\mu_1} \right)^n \left( \frac{\bar{\beta}}{\bar{\mu}} \right)^l}.$$
 (4.44c)

In (4.44c) the coefficients are dependent on L. For L = 0.73 (as fixed by Archie's law) these coefficient values have been numerically evaluated (Sheng 1990). Equations (4.44a), (4.44b) are to be integrated with the condition  $\bar{\beta} = \beta_2$ ,  $\bar{\mu} = \mu_2$  at p = 0. However, just as in the electrical case, here  $\beta_2$ ,  $\mu_2$  represent the effective moduli of the 50–50 fluid–cement composite, to be obtained as the solutions to the following coupled equations:

$$\frac{1}{\beta_2} = \frac{(1-\phi_0)}{\beta_c + \frac{4}{3}(\mu_2 - \mu_c)} + \frac{\phi_0}{\beta_f + \frac{4}{3}\mu_2},$$
(4.45a)

$$\frac{1}{\mu_2 + J} = \frac{(1 - \phi_0)}{\mu_c + J} + \frac{\phi_0}{J}.$$
(4.45b)

Here the moduli with subscript "c" denote those for the cement,  $\beta_{\rm f}$  is the longitudinal modulus for the fluid,  $\mu_{\rm f} = 0$ , the value of  $\phi_0 = 0.5$ , and  $J = \mu_2(9\beta_2 - 4\mu_2)/(6\beta_2 + 4\mu_2)$ .

#### Comparison with Data

Figure 4.10 shows a data plot (Pickett 1963) for sandstones (open squares) and dolomites (solid circles). While there is obvious scatter in the data, the main trend seen in the data is associated with porosity variation and its associated microstructures changes.

Instead of integrating (4.44) starting at p = 0 ( $\phi = \phi_0$ ) by choosing the values of cement, solid, and fluid moduli, it would minimize the amount of adjustable parameters by starting at a measured data point  $(v_p^{-1} = 58.2 \,\mu\,s/ft, v_s^{-1} = 92 \,\mu\,s/ft)$  for the sandstones and use the values of  $\rho_1 = 2.65 \,\mathrm{g \, cm^{-3}}, \beta_1 = 1 \times 10^{12} \,\mathrm{dynes \, cm^{-2}}, \mu_1 = 0.45 \times 10^{12} \,\mathrm{dynes \, cm^{-2}}$  corresponding to those for the  $\alpha$ -quartz, plus  $\phi = 8\%$  to integrate forward in porosity (to  $\phi = \phi_0 = 0.5$ ). The result is the solid line that goes through the sandstone data, obtained by the formula  $v_p^{-1} = \sqrt{\bar{\rho}/\bar{\beta}}, v_s^{-1} = \sqrt{\bar{\rho}/\bar{\mu}}$ . Here  $\bar{\rho} = [\phi + 2.65(1 - \phi)] \,\mathrm{g \, cm^{-3}}$ , i.e., the fluid density is taken to be that of water, and  $\phi = \phi_0(1 - p)$ . Not only is the calculated line almost indistinguishable from the least square fit to the data, but the range of porosity values used in the calculation (8–21%) also coincides well with those measured (5–30%). The cement moduli thus inferred are:  $\beta_c = 1.24 \times 10^{12} \,\mathrm{dynes \, cm^{-2}}$ 



Fig. 4.10. Comparison between the  $V_{\rm p}^{-1}$  vs.  $V_{\rm s}^{-1}$  data on sandstones and dolomites (Pickett 1963) with the predictions of the theory based on the DEM microgeometry (Sheng 1990). Here open squares denote the sandstone data, and solid circles denote the dolomite data.

 $\mu_{\rm c} = 0.23 \times 10^{12} \text{ dynes cm}^{-2}$ , with  $\beta_{\rm f} = 2.6 \times 10^{10} \text{ dynes cm}^{-2}$  corresponding to that of water. Similar treatment has resulted in the solid line for the dolomites.

Our results thus imply a definite correlation between the longitudinal and shear wave velocities with the rock porosity. This correlation is displayed in Fig. 4.11, where the corresponding Wyllie equation is shown by the dashed lines. It is seen that the Wyllie equation tends to consistently overpredict the porosity. Such deviation was indeed empirically observed (Schlumberger 1979).

Also, the Wyllie equation completely breaks down for the velocities of the empty-frame (vs. fluid saturated) rocks, whereas our calculations show the p-wave velocity in the empty-frame case to be only 5–15% lower than that in the fluid-saturated case, again in good agreement with the experiments (Gregory 1977).

It is also clear that the data contain large scatter. This implies factors, other than microstructural variations associated with porosity, to be affecting the velocities. Degree of compaction, clay content, gas saturation, etc., are all possible sources for observed scatter in the data.

# 4.5 The Spectral Function Approach: Coherent Quasimodes

When the wavelength becomes comparable to the typical feature size of a random medium, the coherent potential approach breaks down. This can appear



Fig. 4.11. Theoretically calculated porosity vs. interval travel time curves for both the longitudinal and the shear waves. (a) Sandstones, (b) dolomites (Sheng 1990).

in various forms in the solution of the relevant CPA equations. It is notable that the wavelength does not appear in the effective medium equations. That is because they are obtained in the limit of  $ka \rightarrow 0$ . When ka is small but no longer close to zero, wavelength dependence can appear in the forward scattering amplitudes. When  $ka \sim 1$ , solution(s) of such CPA equation(s) (involving the wavelength-dependent amplitudes) can suddenly disappear, or jump to some clearly unphysical branch, with multiple solutions appearing simultaneously. In other words, the coherent wave modes, well defined in the  $ka \sim 0$  limit, no longer appear so when  $ka \geq 1$ .

Physically, when  $ka \ge 1$  a wave's mean free path is much decreased from its value in the  $ka \sim 0$  limit. That means the coherent wave modes acquire finite lifetimes, i.e., they are at best quasimodes in the intermediate wavelength regime. Another aspect of the  $ka \ge 1$  regime is the increase in wave resolution, so that the various features of the random medium become "visible" to a wave. Mathematically, this aspect is usually manifest in the appearance of k dependence in the self energy, implying nonlocality and correlation as discussed earlier. It is thus not surprising that the CPA breaks down.

But in spite of the finite lifetime and the emergence of nonlocality, coherent waves do not disappear abruptly when  $ka \geq 1$ . They still persist in this

regime, associated with many novel and interesting phenomena. To describe them requires an approach different from the CPA, but which should be able to recover the CPA in the long wavelength limit. In other words, a more generalized approach is needed.

#### Generalization from the CPA

An important step in the derivation of the CPA equation is the recognition that the zeroes in the denominator of the Green function represent excitations (modes). The analytic property of the Green function dictates these zeroes to be equally marked by Green function's imaginary part, the spectral function, as delta functions. In the spectral function approach (Jing et al. 1991, 1992), detailed below, we shift our attention from the zeroes in the Green function denominator, which yielded the CPA, to the peaks in the spectral function. In the long wavelength regime, the two are the same but the denominator's zeroes offered a more direct and precise condition for formulating the CPA equation. In the intermediate wavelength regime, however, the finite lifetime of the coherent quasimodes means that the delta functions in the spectral function are broadened into finite peaks. The positions of the peaks are no longer necessarily associated with the zeroes in the denominator (of the real part of the Green function). Even the existence of the zeroes may be in question. Hence the breakdown of CPA does not mean the disappearance of the coherent modes in this regime, except now the spectral function offers a better and more direct way to identify and track these quasimodes.

Based on the above, it is seen that a point-by-point calculation of the spectral function in the  $\omega - k$  plane would yield a 2D map, in which not only the dispersions of the coherent quasimodes may be identified through the peak positions and their lifetimes obtained as the widths of the peaks, but also the complications arising from the **k** dependence of the self energy (giving rise to the possibility of multiple modes) may be circumvented. The latter is because in this method, it is no longer required to find solutions to equations. Instead, one is using an "experimental" approach to find the necessary information. Indeed, the spectral function is well known as a tool for identifying excitations in complex systems, e.g., angle-resolved photo-emission spectra. The difference here is that a theoretical simulation, if successful in obtaining agreement with experiments, can yield the relevant underlying physics of the phenomena.

Direct, accurate calculation of the spectral function is not trivial. What is proposed here is based on approximations which have proved to be reasonably accurate in CPA, extended to the present approach.

# 4.5.1 Formulation

The averaged Green function is in the form (3.51):

$$\langle G \rangle_{\rm c}(\omega, \mathbf{k}) = \frac{\text{const.}}{\frac{\omega^2}{v_{\rm e}^2} - k^2 - \text{Re}\Sigma(\omega, \mathbf{k}) - i\text{Im}\Sigma(\omega, \mathbf{k})}.$$
(4.46)

An important point to note is that there is a degree of freedom in the choice of  $v_{\rm e}$ , the wave speed of the reference medium, and the choice of  $v_{\rm e}$  directly affects the value of the self energy  $\Sigma$ . CPA is one scheme which "optimizes" the choice of  $v_{\rm e}$  by making  $\Sigma \cong 0$  to the leading order in ka, so that  $\langle G \rangle_c \cong$  $1/[(\omega^2/v_{\rm CPA}^2) - k^2]$ , recognized to be the form for a homogeneous medium with a renormalized wave speed.

As stated, the focus of the spectral function approach is on the imaginary part of  $\langle G \rangle_c$ :

$$-\mathrm{Im} \langle G \rangle_{\mathrm{c}} = \mathrm{const.} \frac{-\mathrm{Im}\Sigma}{\left[ \left( \frac{\omega^2}{v_{\mathrm{e}}^2} - k^2 - \mathrm{Re}\Sigma \right)^2 + \left( \mathrm{Im}\Sigma \right)^2 \right]}.$$
 (4.47)

In the long wavelength limit where the CPA is valid, the choice of  $v_e = v_{CPA}$  makes  $\text{Re}\Sigma = 0$  to the leading order in ka. But  $-\text{Im}\Sigma$ , being higher order in ka, is always a small positive number that is proportional to the scattering cross-section. If we take that into account, then

$$-\mathrm{Im} \langle G \rangle_{c} = \mathrm{const.} \frac{-\mathrm{Im}\Sigma}{\left[\frac{\omega^{2}}{v_{\mathrm{CPA}}^{2}} - k^{2}\right]^{2} + (\mathrm{Im}\Sigma)^{2}}, \qquad (4.48)$$

which is precisely in the form const.  $\pi \delta \left[ (\omega^2/v_{\rm CPA}^2) - k^2 \right]$  when  $-{\rm Im}\Sigma \to 0$ . The fact that  $-{\rm Im}\Sigma \to 0$  is guaranteed by Rayleigh scattering, which states that the scattering cross-section is proportional to  $(ka)^4 \sim 0$  in the long wavelength limit.

With the transition into the intermediate wavelength regime, the right hand side of (4.47) would change from a sharply peaked delta function to a function just with peak(s), signifying quasimodes. Accompanying this transition would be the fact that Re $\Sigma$  may not be zero anymore, regardless of the choice of  $v_{\rm e}$ , and  $-\text{Im}\Sigma$  may not be small. Hence the evaluation of the full  $-\text{Im} \langle G \rangle_c$  is the task at hand.

To obtain a point-by-point map of  $-\text{Im} \langle G \rangle_c$ , we first utilize the freedom in choosing  $v_e$  by setting  $v_e = \omega/k$  for every point on the  $(\omega, k)$  plane. Thus (4.47) becomes

$$-\mathrm{Im} \langle G \rangle_{\mathrm{c}} = \mathrm{const.} \frac{-\mathrm{Im}\Sigma'}{\left[ \left( \mathrm{Re}\Sigma' \right)^2 + \left( \mathrm{Im}\Sigma' \right)^2 \right]} \propto \mathrm{Im} \frac{1}{\Sigma'}, \qquad (4.49)$$

where  $\Sigma'$  is noted to be evaluated at a reference medium with wave speed  $\omega/k$ . Second, we note that in the map of the spectral function  $-\text{Im} \langle G \rangle_c$ , clearly the points corresponding to the local maxima, indicating quasimodes, are more important than those where the value of the function is small. From (4.49), the local maxima correspond with points where the real and imaginary parts of the self energy is small. Therefore, by sacrificing accuracy at those places where one does not care anyway, a reasonable approximation is to take (see (3.15), (3.57))

$$\Sigma' \cong \langle T \rangle_{c}|_{\frac{\omega}{k}} \cong \sum_{i} < \mathbf{t}_{i} >_{c}|_{\frac{\omega}{k}} = -4\pi \sum_{i} f_{i} (0)|_{\frac{\omega}{k}}, \qquad (4.50)$$

where the index *i* runs over all the scatterers, and the condition that the quantity is evaluated at the reference medium wave speed  $\omega/k$  is explicitly noted. In this way, one obtains from (4.49) and (4.50) a simple expression for the spectral function S:

$$-\mathrm{Im} \langle G \rangle_{\mathrm{c}} \propto S = -\mathrm{Im} \frac{1}{\sum_{i} f_{i} (0)|_{\frac{\omega}{k}}}.$$
(4.51)

Here we have reduced the problem again to the calculation of the forward scattering amplitudes. To go a step further would require the specification of the actual problem to be solved.

# 4.5.2 Coherent Acoustic Waves in the Intermediate Wavelength Regime

Through the applications of the spectral function approach to two examples, we wish to illustrate the richness of the coherent wave phenomena in the intermediate wavelength regime. The first is on the origin of *two* acoustic modes in colloidal suspensions, in the frequency regime where the wavelength is comparable to or smaller than the size of the suspended particles. This is surprising at first sight since a colloid is still a fluid, so one would generally expect only one longitudinal acoustic mode. The second case is on the explanation of the experimental observation of very low group velocity in a glass beads–water composite, again in the intermediate wavelength range, with a value which is 60% of the lowest wave speed of the constituents. Such a low value is in clear violation of the CPA lower bound and therefore can not possibly occur in the long wavelength limit.

In order to carry out quantitative calculations, it is necessary to consider the classical wave equations, i.e., elastic and electromagnetic wave equations, and their solutions in the scattering geometry. Since both examples involve acoustic waves in colloidal or unconsolidated particulate solid, we consider only the dispersion microgeometry for a solid sphere coated with a liquid layer, embedded in an effective medium fluid reference medium. It will be seen that the use of dispersion microgeometry's structural unit is important in obtaining the good agreement with experiments, because the wave solution in a coated sphere already takes into account (a significant part of) the multiple scattering that arise from the structural correlation between the dispersed particles and the matrix.

#### Acoustic Wave Scattering by a Coated Solid Sphere

Consider an acoustic plane wave in a fluid effective medium incident on a fluid-coated solid sphere with inner radius a and outer radius b. The relevant wave equation is given by

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$$\frac{\partial^2 \mathbf{u}}{\partial t^2} = c_{\rm t}^2(\mathbf{r}) \nabla^2 \mathbf{u} + [c_{\rm p}^2(\mathbf{r}) - c_{\rm t}^2(\mathbf{r})] \nabla (\nabla \cdot \mathbf{u})$$
(4.52)

in each of the three homogeneous regions with  $c_{\rm t} = v_{\rm t}, c_{\rm p} = v_{\rm p}$  for the shear wave and longitudinal wave velocities, respectively, inside the solid particle  $(r < a); c_{\rm t} = 0, c_{\rm p} = v_{\rm f}$  inside the fluid coating (a < r < b); and  $c_{\rm t} = 0$   $c_{\rm p} = \omega/k$  in the embedding matrix medium (r > b). Here **u** denotes displacement, expressible in terms of the scalar potentials  $\varphi, \xi$  as

$$\mathbf{u} = \nabla \varphi + \nabla \times \nabla \times (\mathbf{r}\xi), \tag{4.53}$$

where  $\varphi$ ,  $\xi$  must satisfy the scalar wave equations for the longitudinal and shear waves, respectively:

$$\nabla^2 \varphi + \frac{\omega^2}{c_{\rm p}^2} \varphi = 0, \qquad (4.54a)$$

$$\nabla^2 \xi + \frac{\omega^2}{c_{\rm t}^2} \xi = 0. \tag{4.54b}$$

The boundary conditions at each interface are the continuity of normal displacement and normal stress, plus vanishing tangential stress at the solid– liquid interface r = a. Here we take the liquid to be inviscid. Solution of the scattering problem (Jing et al. 1992) yields the forward scattering amplitude given by

$$f(0) = \frac{2\pi}{k} \sum_{\ell} (2\ell + 1)F_{\ell}, \qquad (4.55)$$

where

$$F_{\ell} = \frac{[k\rho_{\rm f}j'_{\ell}(kb)j_{\ell}(k_{\rm f}b) - k_{\rm f}\rho_{\rm e}j_{\ell}(kb)j'_{\ell}(k_{\rm f}b)] + [k\rho_{\rm f}j'_{\ell}(kb)h_{\ell}(k_{\rm f}b) - k_{\rm f}\rho_{\rm e}j_{\ell}(kb)h'_{\ell}(k_{\rm f}b)]A_{\ell}}{[k\rho_{\rm f}h'_{\ell}(kb)j_{\ell}(k_{\rm f}b) - k_{\rm f}\rho_{\rm e}h_{\ell}(kb)j'_{\ell}(k_{\rm f}b)] + [k\rho_{\rm f}h'_{\ell}(kb)h_{\ell}(k_{\rm f}b) - k_{\rm f}\rho_{\rm e}h_{\ell}(kb)h'_{\ell}(k_{\rm f}b)]A_{\ell}}$$

$$(4.56a)$$

Here k is the coordinate value on the  $(\omega, k)$  plane,  $\rho_e$  is given by (4.30e) with  $\rho_2 = \rho_f$ ,  $\rho_1 = \rho_s$ ,  $k_f = \omega/v_f$ ,  $j_\ell$  and  $h_\ell$  are the spherical Bessel and spherical Hankel functions of order  $\ell$ , respectively, and prime indicates derivative with respect to the argument of the function. In (4.56a),  $A_\ell$  is given by

$$A_{\ell} = -\frac{U^{(1)}\beta_2 + \ell(\ell+1)\alpha_1 V^{(1)}}{U^{(3)}\beta_2 + \ell(\ell+1)\alpha_1 V^{(3)}},$$
(4.56b)

with

$$\alpha_1 = 2 \left[ \frac{j_{\ell}'(k_{\rm p}a)}{k_{\rm p}a} - \frac{j_{\ell}(k_{\rm p}a)}{(k_{\rm p}a)^2} \right], \qquad (4.56c)$$

$$\beta_2 = \frac{[\ell(\ell+1)-2]j_\ell(k_{\rm t}a)}{(k_{\rm t}a)^2} + j_\ell''(k_{\rm t}a), \qquad (4.56d)$$

$$U^{(1)} = Q_{\rm o} \frac{\left[-j_{\ell}^{\prime\prime}(k_{\rm p}a) + Q_1 j_{\ell}(k_{\rm p}a)\right] j_{\ell}^{\prime}(k_{\rm f}a)}{k_{\rm f}a} - \frac{j_{\ell}^{\prime}(k_{\rm p}a) j_{\ell}(k_{\rm f}a)}{k_{\rm p}a}, \qquad (4.56e)$$

$$V^{(1)} = Q_{\rm o} \frac{[j_{\ell}(k_{\rm t}a)/k_{\rm t}a]' j_{\ell}'(k_{\rm f}a)}{k_{\rm f}a} + \frac{j_{\ell}(k_{\rm f}a)j_{\ell}(k_{\rm t}a)}{(k_{\rm t}a)^2}, \tag{4.56f}$$

where  $k_{\rm p} = \omega/v_{\rm p}, k_{\rm t} = \omega/v_{\rm t}, Q_{\rm o} = 2\rho_{\rm e}v_{\rm t}^2/\rho_{\rm f}v_{\rm f}^2, Q_1 = (v_{\rm p}^2 - 2v_{\rm t}^2)/2v_{\rm t}^2$ , and  $U^{(3)}$ and  $V^{(3)}$  have the same forms as  $U^{(1)}$  and  $V^{(1)}$ , respectively, except  $j_\ell(k_{\rm f}a)$ is replaced by  $h_\ell(k_{\rm f}a)$  and  $j'_\ell(k_{\rm f}a)$  is replaced by  $h'_\ell(k_{\rm f}a)$ . In (4.55), a cutoff in the summation over  $\ell$  at  $\ell \sim 20$  is generally sufficiently accurate. Apart from a constant factor, the spectral function S is then given by (4.51) and (4.55).

#### 4.5.3 Novel Acoustic Modes in Colloidal Suspensions

It is the conventional wisdom that an isotropic, homogeneous elastic solid has one longitudinal and two shear modes, and a fluid has only one longitudinal acoustic mode due to the vanishing shear modulus. However, in colloidal suspensions of polymethylmethacrylate (PMMA) spheres (370 nm in diameter) dispersed in fluid, it was found that whereas in the long wavelength limit there is indeed only one acoustic modes, once  $ka \ge 1$  there can be two distinct longitudinal modes whose characteristics vary with the concentration of the spheres (Liu et al. 1990). These experimental results, obtained through Brillouin scattering, directly challenged our usual understanding of one of the most basic topics in classical physics.

# **Brillouin Scattering Data**

Brillouin scattering is based on the scattering of light by tiny periodic dielectric constant variations induced by the density change  $(\nabla \cdot \mathbf{u})$  associated with the coherent acoustic modes. In Brillouin scattering experiments, the PMMA microspheres were dispersed in an optically index-matched mixture of dodecane and carbon disulphide, so as to eliminate multiple light scattering that might arise from the solid-fluid interfaces. The angle between the incident and the detected laser beams determines the (scattering) wavevector k of the acoustic excitation through momentum conservation. Since the amplitude of the weak optical gratings generated by the coherent acoustic modes varies time-harmonically (with periodicity  $\omega$ ), the scattered light can thus be shifted in frequency from the Rayleigh peak by  $\pm \omega$  (or some integral number times  $\omega$ ). The same conclusion can be reached through the consideration of energy conservation, i.e., via emission or absorption of phonons with energy  $\hbar\omega$ ,  $\hbar$ being the Planck's constant. By measuring the frequency shift as a function of k, one thus obtains the dispersion map of the coherent acoustic modes.

Data obtained through Brillouin scattering are shown in Fig. 4.12 for two different solid sphere concentrations: p = 0.38 and 0.51. Here the solid circles and open squares represent the data points. While the units on both axes are



Fig. 4.12. Measured dispersion curves for PMMA colloidal suspensions at two concentrations. The diameter of the solid spheres is d = 370 nm. The two solid straight lines are the dispersion curves for the longitudinal wave in the solid phase and the compressional wave in the liquid phase, as noted (Liu et al. 1990).

dimensionless, one can deduce from the wave velocity and the size of the microspheres that the relevant (angular) frequencies are in the 5–60 GHz range.

Two branches of the dispersion map are clearly seen for kd > 2(d = 2a)being the diameter of the solid spheres), i.e., in the intermediate wavelength regime. At volume concentration p of solid spheres less than 20%, there is only one mode whose dispersion relation deviates only very slightly from that of the liquid. Relative to the dispersion relations of the solid (that makes up the spheres) and the liquid, shown by the solid lines, the high-frequency mode has a velocity intermediate between the two, and it increases with increasing p. These tendencies are in accord with general intuition. However, the highfrequency mode is clearly separated from the low-frequency mode by a gap. Moreover, there seem to exist smaller, higher-frequency gaps in the dispersion relation. The dispersion relation for the lower-frequency mode is seen to fall below that of the liquid, with a propagation velocity lower than that of the liquid. As p increases, the lower frequency mode actually slows down further, opposite to the trend seen in the higher-frequency mode. Also, at higher sphere concentrations the dispersion relation is seen to develop a dip at  $kd \sim 2\pi$ . These novel features are the subject for theoretical explanation.

#### Predictions from the Spectral Function Approach

For inputs to the point-by-point spectral function calculations, we used the parameter values  $v_{\rm p} = 2.7 \times 10^5 {\rm cm \, s^{-1}}$ ,  $v_{\rm t} = 1.1 \times 10^5 {\rm cm \, s^{-1}}$  for PMMA,  $v_{\rm f} = 1.2 \times 10^5 {\rm cm \, s^{-1}}$ ,  $\rho_{\rm s} = \rho_{\rm f} = 1 {\rm g \, cm^{-3}}$  and the fluid coating thickness is

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Fig. 4.13. The magnitude of the spectral function, delineated in color, plotted as a function of  $\omega d/v_{\rm f}$  and kd. Here red indicate high and blue indicate low, with yellow as intermediate. Experimental data are shown as open circles, with bars indicating their frequency widths. The arrows on the sides indicate the frequency positions of the scattering cross-section peaks for a single sphere immersed in liquid. Here p = 0.38 (Jing et al. 1991, 1992).

determined from the relation  $p = a^3/b^3$ , where a = 185 nm. In Figs. 4.13 (p = 0.38) and 4.14 (p = 0.51), the colors in the two figures indicate the magnitude of the calculated spectral function, with red indicating high and blue indicating low, with yellow being intermediate.

Two bands of ridges are clearly seen in red and yellow. They correspond remarkably well to the experimentally observed dispersion relations for the two acoustic modes (here each data point has an error bar attached). It should be stressed that the calculation has no adjustable parameters. While not shown, the calculations also indicate that for volume fractions below p = 0.2 the two dispersion relations merge and become indistinguishable from the liquid dispersion relation, in exact agreement with the experimental observation.

#### Origin of the Two Acoustic Modes

The arrows in Figs. 4.13 and 4.14 indicate the frequency positions of the peaks in the calculated scattering cross-section for a single sphere immersed in liquid. They are the key to understanding the physical origin of the two modes. For the high-frequency mode, the positions of the arrows correspond directly with the gaps in the calculated dispersion relations. The first gap is clearly



Fig. 4.14. Same as Fig. 4.13, except p = 0.51. The lower frequency branch is seen to deviate further from the dispersion of liquid, with an asymptotic slope approaching the speed of the Stoneley wave (Jing et al. 1991, 1992).

seen in the experimental results for both concentrations and the second gap and hint of a third are also visible; although for the higher gaps their frequency positions do not agree with theory as well. This indicates that the high-frequency mode (between the gaps) results from the *antiresonance* condition for the single-sphere scattering. However, the fact that the minima are nonzero implies that the excitations are indeed quasimodes with finite lifetimes.

For the low-frequency mode, the frequency positions of the scattering peaks are noted to coincide with the peaks of the spectral function, and in the p=0.38 case the calculated gap in the dispersion is seen to correspond reasonably well with the experimentally observed one. The low-frequency mode is therefore the direct result of single-sphere resonances, in contrast to the high-frequency mode. The minima of scattering, as evidenced by the maxima of the spectral function, are the consequence of *coupling* between neighboring spheres and the resulting split of a single resonance mode into two, with a (nonzero) scattering minimum in between (in terms of the frequency).

The relevant single-sphere resonances discussed above depend crucially on the shear modulus of the solid sphere. They are analogous to the "surface modes" of small metallic spheres that may be viewed as an extension of the surface plasmon to the spherical geometry. Here the relevant flat-surface excitation is the Stoneley mode (Brekhovskikh 1980) for the solid–liquid interface (which vanishes if the solid shear modulus is zero). It should be noted that

being an interfacial mode, the Stoneley wave speed (~  $0.86 \times 10^5 \text{cm s}^{-1}$ ) is lower than that of the shear wave speed in solid or the longitudinal wave speeds in either the solid or liquid. It is the lowest of all the relevant wave speeds. A valid physical picture that emerges from the spectral function calculations is that of surface Stoneley modes coupled through their decaying portions in the liquid. Thus when the concentration of the solid spheres is increased, the interfacial area increases as well, and we expect the lower-frequency mode speed to move further away from that of the liquid and approach that of the Stoneley wave. Since the low-frequency mode depends on the coupling between the neighboring spheres, as concentration decreases the Stoneley mode becomes localized on individual spheres, no longer observable as a propagating wave. This explains the fact that at p < 0.2 there is only one mode.

It is indeed remarkable that the spectral function approach, with its simplifying assumptions, can capture the essential physics of the observed physical phenomenon. Its success lies in the use of the coated sphere as the scattering unit. Here the use of a slightly more complex structural unit (as compared to a bare sphere) makes it possible to obtain the physically important effects from the *single scattering* calculations repeated at every  $(\omega, k)$  point.

#### 4.5.4 Group Velocity in Strongly Scattering Media

It is well known that a wave is usually characterized by two velocities: the phase velocity  $\omega/k$  and the group velocity  $v_{\rm g} = {\rm d}\omega(k)/{\rm d}k$  at which a pulse or signal propagates. However, the meaning of the group velocity is strictly well defined only when the dispersion of the medium is not too large. This can be seen from the fact that if the index of refraction is frequency-dependent, i.e.,  $n(\omega)\omega = v_{\rm o}k$ , then  $v_{\rm g} = v_{\rm o}/[n(\omega) + \omega({\rm d}n/{\rm d}\omega)]$ . Since close to an optical resonance the index of refraction can vary violently as a function of frequency, its derivative can be very large, or negative, or anything in between. That means the group velocity calculated in this fashion can be almost any value. Nevertheless, even in a highly dispersive medium, a signal can still propagate, and the determination of its speed of propagation is a classic problem, having been recognized in the work of Sommerfeld (1914) and Brillouin (1960). One manifestation of this problem is the description of wave propagation through strongly scattering, inhomogeneous composite systems in the intermediate wavelength regime, where the excitation of internal modes within the scatterers leads to strong resonant scattering. This scattering leads to a large attenuation of the coherent wave component, and a correspondingly large dispersion. Such dispersion may result in unphysical values of the calculated group velocity. However, even in the most strongly scattering medium, a coherent, ballistic portion of the transmitted pulse can always be measured if one uses a sufficiently thin sample. This ballistic pulse consists of both the unscattered and forward scattered portions of the wave pulse. Hence a group velocity should be measurable.

# Measurement of the Coherent Wave Velocity

The velocity at which the ballistic pulse propagates is an intriguing problem because, on the one hand, it does not seem possible for the pulse to travel through the medium without being affected by scattering, yet on the other hand, if it is scattered, temporal coherence between the incident and transmitted pulses may be destroyed. If the pulse is not scattered, its velocity should be intermediate between the velocities of the individual components of the medium. By contrast, if the scatterers do affect the ballistic propagation, the resulting dispersion may become very large, making the group velocity inadequate to describe ballistic pulse propagation.

The relevant experiment (Page et al., 1996) was carried out using ultrasonic pulses incident on disc-shaped slabs of monodisperse glass beads, randomly packed in water at a glass volume fraction of 63% and contained between two 1.5 mm thick sheets of polystyrene. The thickness of the samples ranged from 2 to 5 mm, each immersed in water. The ultrasound was varied between 1 and 5 MHz, and the pulse typically consisted of ~10 oscillations.

A 25-mm-diameter flat receiving transducer was used on the other side of the sample to detect the transmitted signal. Because the detector averages the instantaneous pressure of the acoustic wave over the front face of the transducer, the scattered sound was effectively canceled as a result of its random phase fluctuations in the plane of the detector, leaving only the wave that was spatially and temporally coherent across the entire surface of the detector. The phase cancelation was further improved by translating the sample and averaging the transmitted field, thereby ensuring the scattered background was eliminated. The phase velocity was determined from the propagation time of the individual oscillations near the center of the pulse, and the correct correspondence between the oscillations of the incident and transmitted pulses was identified by ensuring consistency of the phase velocity between samples of several different thicknesses. The group velocity was determined by matching the centers of the transmitted and incident wave forms and measuring the propagation time.

## **Observed Low Group Velocity**

Shown in Figure 4.15 are the data gathered in the above manner. Significant velocity variation is seen as a function of frequency. The phase velocity drops to a minimum value of  $1.25 \,\mathrm{km \, s^{-1}}$  at  $\omega a/v_{\rm f} = k_{\rm f}a \sim 3$ , where the wave velocity in water is  $v_{\rm f} = 1.5 \,\mathrm{km \, s^{-1}}$ . The group velocity also displayed two minima, with the one at  $k_{\rm f}a \sim 2$  having a value of  $0.85 \,\mathrm{km \, s^{-1}}$ . Remarkably, for most of the frequencies at which the measurements were made, both velocities are substantially below the speed of sound in water, the longitudinal wave speed in glass  $(5.6 \,\mathrm{km \, s^{-1}})$ , the shear wave speed in glass  $(3.4 \,\mathrm{km \, s^{-1}})$ , or the Stoneley wave speed at the glass–water interface  $(1.5 \,\mathrm{km \, s^{-1}})$ . Such values of the velocities are outside the theoretical bounds predicted by the



Fig. 4.15. (a) Phase and (b) group velocities of acoustic waves in a strongly scattering medium consisting of glass beads in water. The data (*symbols*) are compared with the predictions of the spectral function approach (*solid lines*) (Page et al. 1996).

effective medium theories in the long wavelength limit. In fact, for all the effective medium theories the effective wave velocities in the composite medium must be intermediate between those of the components. However, from the values of  $k_{\rm f}a$  for the current experiment, it is clear that the observed velocity dispersion belongs to the intermediate wavelength regime, for which the effective medium theories are inapplicable.

#### Predictions from the Spectral Function Approach

Figure 4.16 shows the calculated dispersion map by using the spectral function approach (Page et al. 1996), in which a = 0.25 mm, b is determined from the relation  $a^3/b^3 = 0.63$ ,  $\rho_s = 2.6 \text{ g cm}^{-3}$ ,  $\rho_f = 1 \text{ g cm}^{-3}$ , and the relevant elastic constants were obtained from the velocities given above. Here the use of the coated sphere as the structural unit might seem a bit stretched, since the glass spheres were tightly packed. However, it should be noted that the spheres were only in frictional contact, and most of the ultrasound propagation between the spheres has to go through a fluid layer. Hence the use of the coated sphere is justified. Again color is used to delineate the height of the spectral function, with red being high and blue being low. The dispersion curve, defined by the peaks, is shown by the black dashed line. The white



Fig. 4.16. Magnitude of the spectral function, delineated in color, plotted as a function of  $\omega a/v_{\rm f}$  and ka. Here *red* indicates high and *blue* indicates low. The dispersion curve, defined by the peaks, is shown by *black dashed line*. The *white dashed line* is the dispersion curve for water. The theoretical calculations involve no adjustable parameters (Page et al. 1996).

dashed line indicates the dispersion relation for water. The calculations involve no adjustable parameters. In contrast to the phenomenon presented in the previous section, for the parameters used here the low-frequency coupled Stoneley mode does not exist, mainly because of the high rigidity of the glass. This fact shows that the existence of two acoustic modes in colloids requires the relevant material parameters to fall in certain regimes.

The phase velocity can be calculated directly from the dispersion relation, by dividing the frequency by the associated wavevector on the dispersion curve. The group velocity was calculated by numerically differentiating the dispersion curve. These are shown as the solid lines in Fig. 4.15. Excellent quantitative agreement is seen, capturing the full dispersion in the group velocity observed experimentally.

#### Physical Picture: The Renormalized Effective Medium

The physical picture that emerges is that when the scattering is strong, each scatterer must sense the scattered waves from the other particles. As a result, the embedding medium becomes renormalized, an effect that is analogous to the shifting of two quantum mechanical resonances when they are strongly coupled. The amount of this renormalization depends on the strength of scattering; the larger the scattering, the larger the adjustment. As the result of

this renormalization, the material properties of the effective medium matrix approach those of the scatterers, and the individual scattering resonances inevitably become leaky and weakened as the effective contrast between a scatterer and the embedding medium diminishes. It is precisely this effective renormalization that is sensed by the coherent group velocity. In our calculation, the coated sphere also possesses these scattering resonances, and these are modified by the coupling to the embedding medium. Our spectral approach identifies, for a wave of frequency  $\omega$ , the wavevector k that allows it to propagate through the medium with the least scattering. Hence, the physical origin of the remarkably low velocities of ballistic propagation lies in the renormalization of the effective medium by strong resonant scatterings, as correctly described by the theoretical model.

#### 4.5.5 Accuracy of the Spectral Function Approach

In the above we see that the spectral function approach can very well describe the wave phenomena in the intermediate wavelength regime, where either  $ka \geq 1$  or the wave absorption is strong so that the CPA breaks down. However, at the same time we also note that the present model tends to underestimate the linewidth of the quasimodes, i.e., overestimate the relevant mean free path in the novel acoustic modes in colloids, for example, where the error bars on the data are considerably larger than the width as seen from the colored dispersion map. The broadening of the linewidth could be accounted for by a better evaluation of the multiple scattering between the particles, as illustrated (Jing et al. 1992) through cluster calculations. Such considerations also lead to the limit of the spectral function approach when the scattering is so strong that  $k\ell \approx 1$ , the so-called the Ioffe–Regel criterion for wave localization, or when  $ka \gg 1$  so that the geometric optics limit is approached. In the latter case wave coherence is local in character and its effect may be treated perturbatively.

# **Problems and Solutions**

**Problem 4.1** Determine the CPA (effective medium) equation for acoustic wave propagation in immiscible fluid–fluid mixtures (emulsions), where the droplet size is small compared to the wavelength.

Solution For acoustic waves in fluid, the relevant equations are

$$\rho \frac{\partial^2 \mathbf{u}}{\partial t^2} = -\nabla P, \qquad (P4.1)$$

$$\frac{\partial P}{\partial t} = -B \frac{\partial (\nabla \cdot \mathbf{u})}{\partial t}.$$
(P4.2)

Here P denotes pressure,  $\rho$  the fluid mass density, **u** is the displacement, and B is the bulk modulus. The first equation is recognized to be Newton's equation

of force density = (mass density) × (acceleration). The second equation is the time derivative of the equation of state for a fluid:  $P = -B\nabla \cdot \mathbf{u}$ , where  $\nabla \cdot \mathbf{u}$  represents the fractional volume change at a point. To obtain a wave equation from the two equations, one first divides both sides of (P4.1) by  $\rho$ , and then takes the divergence of both sides. That gives

$$\frac{\partial^2 \nabla \cdot \mathbf{u}}{\partial t^2} = -\nabla \cdot \frac{1}{\rho} \nabla P. \tag{P4.3}$$

Substituting (P4.2) into (P4.3) and recognizing B to be time independent, one gets

$$\frac{\partial^2 P}{\partial t^2} - B\nabla \cdot \frac{1}{\rho} \nabla P = 0.$$
(P4.4)

This equation reduces to the scalar wave equation, with  $\phi = P$ , if  $\rho$  is a constant. In that case the speed of sound is identified as  $v = \sqrt{B/\rho}$ . For inhomogeneous fluid systems, however,  $\rho$  may have different values for different constituents. The scalar wave equation is valid in each homogeneous region, but one of the boundary conditions across the interfaces must be modified. Instead of requiring the normal derivative of P to be continuous across the interface, it is  $\rho^{-1}$  times the normal derivative. This can easily be seen from (P4.4) by the fact that since P (and therefore  $\partial^2 P/\partial t^2$ ) is continuous across the boundary, the term  $B\nabla \cdot \rho^{-1}\nabla P$  must also be continuous. As a result,  $\nabla \cdot \rho^{-1}\nabla P$  can, at worst, only have a jump across the interface to compensate for the jump in B. Due to the fact that the integral of a jump is continuous (with a discontinuity in the slopes), it follows that  $\rho^{-1}\nabla P$  must be continuous across the interface. Physically, the continuity of  $\rho^{-1}\nabla P$  implies displacement continuity across the interface.

With the modification of the boundary condition, the first step in the derivation of an effective medium equation is the consideration of the scattering solution for a single scatterer in the  $\kappa_0 R \rightarrow 0$  limit. In 3D, that means a modification of (P3.22) with the replacement of the  $\kappa/\kappa_0$  factor by  $(\kappa/\kappa_0)(\rho_0/\rho)$ . Tracing the effect of this change through (P3.24), we get a new expression for  $D_n$ :

$$D_n = \frac{\kappa_0 \rho j_n(\kappa R) j'_n(\kappa_0 R) - \kappa \rho_0 j'_n(\kappa R) j_n(\kappa_0 R)}{\kappa \rho_0 j'_n(\kappa R) h_n(\kappa_0 R) - \kappa_0 \rho h'_n(\kappa_0 R) j_n(\kappa R)},$$
(P4.5)

where  $\rho_0$  is the mass density of the uniform fluid medium. In the  $\kappa_0 R \to 0$ limit, only the  $D_0$  expansion has to be considered. By using the identities  $\kappa = \omega/\sqrt{B/\rho}$  and  $\kappa_0 = \omega/\sqrt{B_0/\rho_0}$ , the leading-order terms of the expansion are

$$D_0 \simeq -\frac{i}{3} (\kappa_0 R)^3 \left( 1 - \frac{B_0}{B} \right) - \frac{(\kappa_0 R)^6}{9} \left( 1 - \frac{B_0}{B} \right)^2.$$
(P4.6)

Comparison with (3.37) tells us that in order to obtain the CPA equation in the  $\kappa_{\rm e}R_m \to 0$  limit (m = 1, 2 for the two components), it is necessary only to replace the factor  $[1 - (\kappa_m/\kappa_{\rm e})^2]$  in (4.15) by  $(1 - \bar{B}/B_m)$ , which gives

$$\frac{1}{\bar{B}} = \frac{p}{B_1} + \frac{1-p}{B_2}.$$
(P4.7)

By denoting the effective mass density as  $\bar{\rho}$ , the magnitude of the effective wave vector is given by

$$\kappa_{\rm e} = \omega / \sqrt{\bar{B}/\bar{\rho}},\tag{P4.8}$$

and the mean-free path l is given by (4.20) with  $\varepsilon$  replaced by 1/B, i.e.,

$$l = \frac{3}{\kappa_{\rm e} \left[ p(\kappa_{\rm e}R_1)^3 \left( 1 - \frac{\bar{B}}{B_1} \right)^2 + (1-p)(\kappa_{\rm e}R_2)^3 \left( 1 - \frac{\bar{B}}{B_2} \right)^2 \right]}.$$
 (P4.9)

The CPA equations remains the same as (P4.7) for 2D and 1D, and the mean free path can be similarly obtained from (4.20). In the solution to Problem 4.4 it is shown that the CPA equation is also the same for the dispersion microgeometry. However, the mean free path expression is different. For application to colloids or mixtures with a well-defined size for the dispersed particles, the mean free path expression of the dispersion microgeometry should be more accurate.

The effective medium equation (P4.7) is noted to have a simple physical derivation in the static limit, where  $\nabla P = 0$  [from (P4.1)], so that P is a constant throughout space. From the equation of state, the fractional volume change is given by

$$\nabla \cdot \mathbf{u} = -\frac{P}{B}$$

The (volume) averaged fractional volume change is therefore given by

$$\langle \nabla \cdot \mathbf{u} \rangle_V = \frac{1}{L^d} \int \nabla \cdot \mathbf{u} \, \mathrm{d}\mathbf{r} = -P \frac{1}{L^d} \int \frac{\mathrm{d}\mathbf{r}}{B(\mathbf{r})}$$
$$= -P \left[ \frac{p}{B_1} + \frac{1-p}{B_2} \right].$$
(P4.10)

Expressing  $\langle \nabla \cdot \mathbf{u} \rangle_V = -P/\bar{B}$  yields (P4.7) directly.

**Problem 4.2** Derive the CPA (effective medium) equation for electromagnetic waves in a composite with the symmetric microstructure, where the scale of the inhomogeneities is small compared to the wavelength.

**Solution** In the Lorentz gauge, the electric field **E** is expressed as (Jackson, 1999)

$$\mathbf{E} = -\nabla\phi - \frac{1}{c}\frac{\partial\mathbf{A}}{\partial t},\tag{P4.11}$$

where  $\phi$  and **A** are the scalar and vector potentials, respectively, and c is the speed of light. In free space,  $\phi$  and every component of **A** satisfy a scalar wave

equation. In the quasistatic limit where the wavelength is large compared to the size of the scatterers, one may approximate  $\mathbf{E}$  by  $-\nabla\phi$  because  $c^{-1}\partial \mathbf{A}/\partial t$  yields a factor  $i\omega/c = i\kappa$ , and if R is the length unit, then  $c^{-1}\partial \mathbf{A}/\partial t$  is on the order of  $\kappa R \sim 0$ .

The boundary condition for **E** is dictated by the well-known electromagnetic boundary condition of normal  $\varepsilon \nabla \phi$  being continuous across the interface, instead of  $\nabla \phi$  by itself. Together with the fact that the leading order scattering term in the  $\kappa R \to 0$  limit must be the n = 1 term as discussed in the text, in 3D the changes in the solution to  $D_1$ , (P3.22), are contained in the substitution of  $\kappa/\kappa_0$  by  $(\kappa/\kappa_0)^3$  (due to the presence of  $\varepsilon$  in the boundary condition, and  $\varepsilon = (\kappa/\kappa_0)^2$ ). Equation (P3.24) is therefore modified to yield for n = 1,

$$D_{1} = \frac{\kappa_{0}^{3} j_{1}(\kappa R) j_{1}'(\kappa_{0} R) - \kappa^{3} j_{1}'(\kappa R) j_{1}(\kappa_{0} R)}{\kappa^{3} j_{1}'(\kappa R) h_{1}(\kappa_{0} R) - \kappa_{0}^{3} h_{1}'(\kappa_{0} R) j_{1}(\kappa R)},$$
(P4.12)

For  $x \to 0$ ,  $j_1(x) \cong (x/3) - (x^3/30)$ ,  $h_1(x) \cong (-i/x^2) + (x/3)$ . By substituting these small-parameter expansions into (P4.12),  $D_1$  becomes

$$D_1 \cong \frac{\mathrm{i}}{3} (\kappa_0 R)^3 \left(\frac{1-\varepsilon}{2+\varepsilon}\right) - \frac{(\kappa_0 R)^6}{9} \left(\frac{1-\varepsilon}{2+\varepsilon}\right)^2.$$
(P4.13)

The scattering amplitude is therefore [from (3.35)]

$$f(\kappa'_0,\kappa_0) = \frac{(\kappa_0 R)^3}{\kappa_0} \left(\frac{1-\varepsilon}{2+\varepsilon}\right) \cos\theta + i \frac{(\kappa_0 R)^6}{3\kappa_0} \left(\frac{1-\varepsilon}{2+\varepsilon}\right)^2 \cos\theta.$$
(P4.14)

For the CPA equation, only the forward-scattering ( $\theta = 0$ ) amplitude is important. So in (4.15), the factor  $[1 - (\kappa_m/\kappa_e)^2]$  should be replaced by

$$\left[\frac{1-(\varepsilon_m/\bar{\varepsilon})}{2+(\varepsilon_m/\bar{\varepsilon})}\right] = \frac{\bar{\varepsilon}-\varepsilon_m}{2\bar{\varepsilon}+\varepsilon_m}$$

which directly yields (4.22a) with d = 3 for the symmetric microstructure. In the 2D case, we have

$$D_1 = \frac{\kappa^3 J_1'(\kappa R) J_1(\kappa_0 R) - \kappa_0^3 J_1(\kappa R) J_1'(\kappa_0 R)}{\kappa_0^3 H_1'(\kappa_0 R) J_1(\kappa R) - \kappa^3 H_1(\kappa_0 R) J_1'(\kappa R)}.$$
 (P4.15)

By using the small-parameter expansions  $J_1(x) \cong (x/2) - (x^3/16)$  and  $H_1 \cong -(i2/\pi x) + (x/2)$ , (P4.15) becomes

$$D_1 \simeq \frac{\mathrm{i}\pi}{4} (\kappa_0 R)^2 \left(\frac{1-\varepsilon}{1+\varepsilon}\right) - \frac{\pi^2}{4} (\kappa_0 R)^4 \left(\frac{1-\varepsilon}{1+\varepsilon}\right)^2.$$
(P4.16)

In terms of the  $t^+$ -matrix element, we have

$$\langle \kappa'_0 | \mathbf{t}^+ | \kappa'_0 \rangle = -2\pi (\kappa_0 R)^2 \left( \frac{1-\varepsilon}{1+\varepsilon} \right) \cos \theta - i2\pi^2 (\kappa_0 R)^4 \left( \frac{1-\varepsilon}{1+\varepsilon} \right)^2 \cos \theta.$$
(P4.17)

For the CPA equation, only the forward-scattering amplitude is important, so the factor  $[1 - (\kappa_m/\kappa_e)^2]$  in (4.15) should be replaced by

$$\left[\frac{1-(\varepsilon_m/\bar{\varepsilon})}{1+(\varepsilon_m/\bar{\varepsilon})}\right] = \frac{\bar{\varepsilon}-\varepsilon_m}{\bar{\varepsilon}+\varepsilon_m}.$$
(P4.18)

This directly yields (4.22a) with d = 2 for the symmetric microstructure.

For the 1D case, if the wave propagation direction is along the x direction, then **E** is along either the y or the z direction. The continuity of the tangential component of **E** across the interface means the CPA equation can be simply derived in the static limit as follows.

Since **E** is constant in each uniform segment and is continuous across the interface, **E** must be a constant throughout the sample. Therefore, from the equation  $\mathbf{D} = \varepsilon \mathbf{E}$ , we get

$$\langle \mathbf{D} \rangle_V = \frac{1}{L} \int \mathbf{D}(x) \, \mathrm{d}x = \frac{1}{L} \int \varepsilon(x) \, \mathrm{d}x \cdot \mathbf{E} = [p\varepsilon_1 + (1-p)\varepsilon_2] \mathbf{E}.$$
 (P4.19)

Equation (4.16) results by letting  $\langle \mathbf{D} \rangle_V = \bar{\varepsilon} \mathbf{E}$ .

**Problem 4.3** Solve the 3D and 2D electromagnetic CPA equations for the symmetric microstructure and show that in the limit of infinite contrast between the two components, there is a percolation threshold.

Solution The electromagnetic CPA for the symmetric microstructure is

$$p\frac{\bar{\varepsilon}-\varepsilon_1}{(d-1)\bar{\varepsilon}+\varepsilon_1} + (1-p)\frac{\bar{\varepsilon}-\varepsilon_2}{(d-1)\bar{\varepsilon}+\varepsilon_2} = 0.$$
(P4.20)

This may be rewritten as a quadratic equation for  $\bar{\varepsilon}$ :

$$(d-1)\overline{\varepsilon}^2 + [(dp - (d-1))\varepsilon_2 + (1-dp)\varepsilon_1]\overline{\varepsilon} - \varepsilon_1\varepsilon_2 = 0.$$
(P4.21)

The solution is

$$\bar{\varepsilon} = \frac{-q + \sqrt{q^2 + 4(d-1)\varepsilon_1\varepsilon_2}}{2(d-1)},\tag{P4.22}$$

$$q = [dp - (d-1)]\varepsilon_2 + (1 - dp)\varepsilon_1.$$
 (P4.23)

Only the + sign solution is chosen because that is the branch yielding the correct limits of  $\bar{\varepsilon} = \varepsilon_1$  at p = 1 and  $\bar{\varepsilon} = \varepsilon_2$  at p = 0. In the limit of infinite contrast, we can let  $\varepsilon_1 = 1$  and  $\varepsilon_2 = 0$ . Then

$$q = (1 - dp)$$

and

$$\bar{\varepsilon} = \begin{cases} \frac{-(1-dp)+|1-dp|}{2(d-1)} = 0, \quad p < \frac{1}{d} \\ \frac{d}{d-1}\left(p-\frac{1}{d}\right), \qquad p > \frac{1}{d}. \end{cases}$$
(P4.24)

It is seen that in 3D,  $p_c = 1/3$  (or 2/3 if  $\varepsilon_2 = 1$ ,  $\varepsilon_1 = 0$ ), and in 2D,  $p_c = 1/2$ .

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	2	111111111
← p' -•   ←	- 1-p'	

Fig. P4.1. The structural unit for the 1D dispersion microgeometry. The effective medium is indicated by the *cross-hatched region*. Intrinsically there is no difference between the two types of microstructure in 1D. However, the unit taken here should give a better description than the two-unit approach because the short-range correlation between the two components is explicitly taken into account. While the CPA equation remains the same, the expression for l is different.

**Problem 4.4** Derive the CPA equation(s) for the dispersion microgeometry in 3D, 2D, 1D and give the respective expressions for the mean free path in the long-wavelength limit.

**Solution** The basic structural unit of the dispersion microgeometry is a coated sphere in 3D and a coated circle in 2D. The structural unit in 1D is taken to be two neighboring components 1 and 2 as shown in Fig. P4.1.

To obtain the CPA equations in 3D and 2D, it is necessary to first solve the single-scatterer problem where the structural unit is embedded in a homogeneous effective medium as shown in Figure 4.4. The radius of the outer sphere (circle) is denoted by  $R_0$  and that of the inner sphere (circle) by R. The material constant of the dispersed component, represented by the inner sphere (circle), is characterized by  $\kappa_1 = \omega/v_1$ , and that of the matrix component is characterized by  $\kappa_2 = \omega/v_2$ .  $R_0$  and R are related by the local volume fraction p' of the dispersed component:

$$p' = \left(\frac{R}{R_0}\right)^d.$$
 (P4.25)

The effective medium is characterized by  $\kappa_{\rm e} = \omega/v_{\rm e}$ . We define  $\varepsilon_1 = (\kappa_1/\kappa_0)^2$ ,  $\varepsilon_2 = (\kappa_2/\kappa_0)^2$ , and  $\bar{\varepsilon} = (\kappa_{\rm e}/\kappa_0)^2$  with respect to a homogeneous reference medium where  $\kappa_0 = \omega/v_0$ .

In 3D, the solutions in the three regions  $r \leq R$ ,  $R < r \leq R_0$ ,  $r > R_0$  can be written as

$$\phi^{(1)} = \sum_{n=0}^{\infty} A_n j_n(\kappa_1 r) P_n(\cos \theta), \qquad r \le R \qquad (P4.26)$$

$$\phi^{(2)} = \sum_{n=0}^{\infty} \left[ B_n j_n(\kappa_2 r) + C_n n_n(\kappa_2 r) \right] P_n(\cos \theta), \quad R < r \le R_0 \qquad (P4.27)$$

$$\phi^{(3)} = \exp(i\kappa_{e} \cdot \mathbf{r}) + \sum_{n=0}^{\infty} i^{n}(2n+1)D_{n}h_{n}^{(1)}(\kappa_{e}r)P_{n}(\cos\theta)$$
  
$$= \sum_{n=0}^{\infty} i^{n}(2n+1)P_{n}(\cos\theta)[j_{n}(\kappa_{e}r) + D_{n}h_{n}^{(1)}(\kappa_{e}r)]. \quad r > R_{0}$$
(P4.28)

Using the condition of  $(\partial \phi / \partial r) / \phi$  being continuous across interfaces, two equations are obtained at the two interfaces r = R and  $r = R_0$ :

$$\frac{\kappa_1}{\kappa_2} \frac{j'_n(\kappa_1 R)}{j_n(\kappa_1 R)} = \frac{B_n j'_n(\kappa_2 R) + C_n n'_n(\kappa_2 R)}{B_n j_n(\kappa_2 R) + C_n n_n(\kappa_2 R)},$$
(P4.29)

$$\frac{\kappa_{\rm e}}{\kappa_2} \frac{j'_n(\kappa_{\rm e}R_0) + D_n h'_n(\kappa_{\rm e}R_0)}{j_n(\kappa_{\rm e}R_0) + D_n h_n(\kappa_{\rm e}R_0)} = \frac{B_n j'_n(\kappa_2 R_0) + C_n n'_n(\kappa_2 R_0)}{B_n j_n(\kappa_2 R_0) + C_n n_n(\kappa_2 R_0)},$$
(P4.30)

where the superscript (1) to the spherical Hankel function has been dropped. By defining the left-hand side of (P4.29) as  $\chi_n, B_n$  can be solved in terms of  $C_n$  as

$$B_n = \mathcal{J}_n C_n, \tag{P4.31}$$

where

$$\mathcal{J}_n = \frac{-\chi_n n_n(\kappa_2 R) + n'_n(\kappa_2 R)}{\chi_n j_n(\kappa_2 R) - j'_n(\kappa_2 R)}.$$
(P4.32)

By substituting (P4.31) into (P4.30),  $D_n$  is solved to yield

$$D_n = \frac{-\kappa_{\rm e} j'_n(\kappa_{\rm e} R_0)/\kappa_2 + \mathcal{H}_n j_n(\kappa_{\rm e} R_0)}{\kappa_{\rm e} h'_n(\kappa_{\rm e} R_0)/\kappa_2 - \mathcal{H}_n h_n(\kappa_{\rm e} R_0)},\tag{P4.33}$$

where

$$\mathcal{H}_n = \frac{\mathcal{J}_n j'_n(\kappa_e R_0) + n'_n(\kappa_2 R_0)}{\mathcal{J}_n j_n(\kappa_e R_0) + n_n(\kappa_2 R_0)}.$$
(P4.34)

Exactly the same expressions hold for 2D if all the spherical Bessel functions are replaced with their normal counterpart, i.e.,  $j_n \to J_n$ ,  $n_n \to N_n$  and  $h_n^{(1)} \to H_n^{(1)}$ .

For 1D, a simple solution approach is the transfer matrix method. In Fig. P4.1 there are four regions. By starting from the leftmost region and going toward the right side, the right-going and left-going waves in respective regions can be written as  $\exp(i\kappa_e x) + \rho \exp(-i\kappa_e x)$ ,  $A \exp(i\kappa_1 x) + B \exp(-i\kappa_1 x)$ ,  $C \exp(i\kappa_2 x) + D \exp(-i\kappa_2 x)$ , and  $\tau \exp(i\kappa_e x)$ . The coefficients between the neighboring regions are related to one another by the continuity of  $\phi$  and  $\partial \phi / \partial x$  across each interface. These conditions may be written as  $2 \times 2$  matrices relating the wave amplitudes. Thus,

$$\begin{pmatrix} A\\B \end{pmatrix} = \mathbf{M}_0 \begin{pmatrix} 1\\\rho \end{pmatrix}, \tag{P4.35}$$

$$\begin{pmatrix} C\\ D \end{pmatrix} = \mathbf{M}_1 \begin{pmatrix} A\\ B \end{pmatrix}, \tag{P4.36}$$

$$\begin{pmatrix} \tau \\ 0 \end{pmatrix} = \mathbf{M}_2 \begin{pmatrix} C \\ D \end{pmatrix}, \tag{P4.37}$$

where

$$\mathbf{M}_{0} = \begin{bmatrix} \left(1 + \frac{\kappa_{e}}{\kappa_{1}}\right) \exp\left[-2\mathrm{i}(\kappa_{e} - \kappa_{1})p'R_{0}\right] & \left(1 - \frac{\kappa_{e}}{\kappa_{1}}\right) \exp\left[2\mathrm{i}(\kappa_{e} + \kappa_{1})p'R_{0}\right] \\ \left(1 - \frac{\kappa_{e}}{\kappa_{1}}\right) \exp\left[-2\mathrm{i}(\kappa_{e} + \kappa_{1})p'R_{0}\right] & \left(1 + \frac{\kappa_{e}}{\kappa_{1}}\right) \exp\left[2\mathrm{i}(\kappa_{e} - \kappa_{1})p'R_{0}\right] \end{bmatrix},$$
(P4.38)  
$$\mathbf{M}_{1} = \begin{bmatrix} \left(1 + \frac{\kappa_{1}}{\kappa_{2}}\right) & \left(1 - \frac{\kappa_{1}}{\kappa_{2}}\right) \\ \left(1 - \frac{\kappa_{1}}{\kappa_{2}}\right) & \left(1 + \frac{\kappa_{1}}{\kappa_{2}}\right) \end{bmatrix},$$
(P4.39)

and

$$\mathbf{M}_{2} = \begin{bmatrix} \left(1 + \frac{\kappa_{2}}{\kappa_{e}}\right) & \left(1 - \frac{\kappa_{2}}{\kappa_{e}}\right) \\ \times \exp[2i(\kappa_{2} - \kappa_{e})(1 - p')R_{0}] & \times \exp[-2i(\kappa_{2} + \kappa_{e})(1 - p')R_{0}] \\ \left(1 - \frac{\kappa_{2}}{\kappa_{e}}\right) & \left(1 + \frac{\kappa_{2}}{\kappa_{e}}\right) \\ \times \exp[2i(\kappa_{2} + \kappa_{e})(1 - p')R_{0}] & \times \exp[-2i(\kappa_{2} - \kappa_{e})(1 - p')R_{0}] \end{bmatrix},$$
(P4.40)

By multiplying the matrices together, one gets

$$\begin{pmatrix} \tau \\ 0 \end{pmatrix} = \mathbf{M}_2 \mathbf{M}_1 \mathbf{M}_0 \begin{pmatrix} 1 \\ \rho \end{pmatrix} = \mathbf{M} \begin{pmatrix} 1 \\ \rho \end{pmatrix}.$$
(P4.41)

Solution of (P4.41) gives  $\tau$  and  $\rho$  in terms of the matrix elements of M:

$$\rho = -\frac{(\mathbf{M})_{21}}{(\mathbf{M})_{22}},\tag{P4.42}$$

$$\tau = \frac{\det(\mathbf{M})}{(\mathbf{M})_{22}}.$$
 (P4.43)

Here det() means the determinant of the matrix.

From the solutions  $D_n(p')$  in 3D and 2D, and  $\tau(p') - 1$  in 1D, the forward scattering amplitudes can be obtained from (4.14). The consideration of different values of p' as constituting different structural units then yields the CPA condition as given by (4.24).

In the long-wavelength limit, expansion of the various functions in the  $n=0\ {\rm terms}$  gives

$$\chi_0 \simeq -\frac{\kappa_1}{\kappa_2} \frac{\kappa_1 R}{d},\tag{P4.44}$$

$$\mathcal{J}_0 \simeq \frac{A_d}{(\kappa_2 R)^d [1 - (\kappa_1/\kappa_2)^2]},$$
 (P4.45)

with

$$A_d = \begin{cases} 3 & \text{in 3D} \\ \frac{4}{\pi} & \text{in 2D,} \end{cases}$$
(P4.46)

$$\mathcal{H}_0 \cong -\frac{\kappa_2 R_0}{d} \left\{ 1 - p' \left[ 1 - \left(\frac{\kappa_1}{\kappa_2}\right)^2 \right] \right\},\tag{P4.47}$$

so that in 3D and 2D,

$$D_0 \simeq -\frac{\mathrm{i}}{A_d} (\kappa_\mathrm{e} R_0)^d \left\{ 1 - \left(\frac{\kappa_2}{\kappa_\mathrm{e}}\right)^2 + \left(\frac{\kappa_2}{\kappa_\mathrm{e}}\right)^2 p' \left[ 1 - \left(\frac{\kappa_1}{\kappa_2}\right)^2 \right] \right\}.$$
(P4.48)

In 1D, the expansion of the exponentials and the more tedious multiplication of matrices in the low frequency limit yield

$$\rho = \tau - 1 \cong -i(\kappa_e R_0) \left[ 1 - (1 - p') \left(\frac{\kappa_2}{\kappa_e}\right)^2 - p' \left(\frac{\kappa_1}{\kappa_e}\right)^2 \right].$$
(P4.49)

By identifying  $(\kappa_1/\kappa_e)^2 = \varepsilon_1/\bar{\varepsilon}$ ,  $(\kappa_2/\kappa_e)^2 = \varepsilon_2/\bar{\varepsilon}$ , and inserting the expressions of the forward scattering amplitudes into (4.24), the p' integral may be performed to give

$$p = \int_0^1 \mathcal{D}(p')p' dp', \qquad (P4.50)$$

and (4.25) results.

For the mean free path, it is noted that in the  $\kappa_e R_0 \to 0$  limit, the total scattering cross-section O(p') may be expressed [from (4.14) and (P3.14)] as

$$O(R_0, p') = \begin{cases} \frac{4\pi}{\kappa_e^2} |D_0|^2 & 3D\\ \frac{4}{\kappa_e} |D_0|^2 & 2D\\ |\rho|^2 + |\tau - 1|^2 & 1D. \end{cases}$$
(P4.51)

By using the optical theorem and (4.19), l is given by

$$l = C_d R_0^d \left\{ \int_0^1 \mathcal{D}(p') O(R_0, p') \, \mathrm{d}p' \right\}^{-1}, \qquad (P4.52)$$

where  $C_d = 4\pi/3, \pi, 2$ , in 3D, 2D, 1D, respectively. Equation (4.26) results from the substitution of (P4.48), (P4.49), and (P4.51) into (P4.52).

**Problem 4.5** Derive the 1D, 2D, and 3D electromagnetic CPA equations for the dispersion microgeometry in the long-wavelength limit, and give their respective mean free path expressions.

**Solution** In the solution to Problem 4.2, details about the electromagnetic boundary condition and the requirement of the vector character of the electromagnetic wave scattering have been discussed. In short, instead of the n = 0 term, the lowest-order scattering in the long-wavelength limit is given by the n = 1 term. Also, the boundary condition change means  $\kappa_1/\kappa_2$  in (P4.29) and  $\kappa_e/\kappa_2$  in (P4.30) should be replaced by  $(\kappa_1/\kappa_2)^3$  and  $(\kappa_e/\kappa_2)^3$ , respectively. In 1D the simple derivation of the CPA equation in the solution to Problem

4.2 shows it to be exact in the  $\kappa_{e}R_{0} \rightarrow 0$  limit. Therefore, both the 1D electromagnetic CPA equation and its relevant expression for the mean free path are the same as their counterparts in the scalar wave case. Only the 3D and 2D cases need to be detailed.

By expanding the various functions obtained in the solution to the previous problem, and making the necessary replacements consistently, it is found that

$$\chi_1 \cong \left(\frac{\kappa_1}{\kappa_2}\right)^3 \frac{1}{\kappa_1 R},\tag{P4.53}$$

$$\mathcal{J}_1 \cong A_d \frac{1}{(\kappa_2 R)^d} \frac{(\kappa_1/\kappa_2)^2 + (d-1)}{(\kappa_1/\kappa_2)^2 - 1},$$
(P4.54)

$$\mathcal{H}_1 \cong \frac{1}{\kappa_2 R_0} \frac{(\kappa_1/\kappa_2)^2 + (d-1) + (d-1)p'\left[(\kappa_1/\kappa_2)^2 - 1\right]}{(\kappa_1/\kappa_2)^2 + (d-1) - p'\left[(\kappa_1/\kappa_2)^2 - 1\right]}, \qquad (P4.55)$$

with  $A_d$  given by (P4.46), and

$$D_{1} \cong \frac{1}{A_{d}} (\kappa_{2}R_{0})^{d} \\ \times \frac{(\bar{\varepsilon} - \varepsilon_{2})[(d-1)\varepsilon_{2} + \varepsilon_{1}] + p'[(d-1)\varepsilon_{2} + \bar{\varepsilon}](\varepsilon_{2} - \varepsilon_{1})}{[(d-1)\varepsilon_{2} + \varepsilon_{1}][(d-1)\bar{\varepsilon} + \varepsilon_{2}] + (d-1)p'(\varepsilon_{2} - \varepsilon_{1})(\bar{\varepsilon} - \varepsilon_{2})}.$$
(P4.56)

The assumption about local compositional fluctuations and the use of (4.24) directly result in the CPA equation as given by (4.28). As for the mean free path, it is noted that to the leading order,

$$f = \begin{cases} -\frac{3\mathrm{i}}{\kappa_{\mathrm{e}}} D_{1} \cos \theta & 3\mathrm{D}, \\ 2\sqrt{\frac{2}{\pi\kappa_{\mathrm{e}}}} \exp\left(-\mathrm{i}\frac{\pi}{4}\right) D_{1} \cos \theta & 2\mathrm{D}, \end{cases}$$
(P4.57)

and

$$O = \begin{cases} 2\pi \int_{-1}^{1} |f|^2 d\cos\theta & 3D, \\ \int_{0}^{2\pi} |f|^2 d\theta & 2D. \end{cases}$$
(P4.58)

It follows from (P4.52) and (P4.58) that the mean free path expression has the form given by (4.29) if there is no fluctuation in p.

# **Diffusive Waves**

# 5.1 Beyond the Coherent Regime

In the last chapter it has been argued that a random medium can appear locally homogeneous to a propagating wave, with an effective material property that is a function of the material properties of the inhomogeneous components and their microstructure. Thus waves are locally coherent. However, when the transport distance is much larger than the mean free path l so that the description of wave properties is focused on a larger scale, wave incoherence dominates. This chapter is concerned with the wave transport behavior under this scenario.

As a function of increasing observation scale, there is a smooth transition in the transport characteristics from the coherent regime to the multiply scattered behavior, which is denoted the diffusive regime in anticipation of later developments. To better appreciate this transition, let us consider the situation where a plane wave has propagated the intermediate distance of two to three mean free paths into an inhomogeneous medium. With two to three scatterings, a certain fraction of the wave energy will have become incoherent, in the sense of departing from its original propagation direction (but still at the same frequency, since all scatterings considered here are elastic). However, there is still a finite portion of the wave energy which remains coherent. This is clear from the calculation of the scattering amplitude  $f(\theta)$ : Each scattering takes away only a fraction of the wave energy from the forward direction. Therefore, as the wave encounters successive scatterings, the component that remains coherent decreases geometrically, i.e., decays exponentially with the number of scatterings. The possibility of a certain part of the incoherent wave energy being returned to the coherent state can be discounted, because to be coherent means propagation not only in the original direction but also with the same phase relation. The combined requirements make the return probability essentially nil. Therefore, the transition from the coherent, CPA-dominated regime to the diffusive regime is governed by a smooth exponential decay of

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the coherent component, and the growth and saturation of the incoherent component, as the number of scattering increases.

The approach of this book is to treat wave transport only in the two limits and to approximate the transition behavior as a superposition of the coherent and the diffusive components. The most serious error inherent in such an approximation is that the incoherent component may not be fully diffusive in the transition regime. It remains an intriguing question as to how many scatterings are required before diffusive transport is firmly established. However, experimental evidence seems to indicate well-established diffusive behavior after only a few scatterings. The caveat here is that such evidence applies only to *configurationally averaged intensity*, which is expected to differ from the single-configuration behavior.

## 5.2 Pulse Intensity Evolution in a Random Medium

To address quantitatively the question of what happens beyond the coherent regime, let us approach the problem physically by following the *intensity* evolution of a pulse injected into a random medium at point  $\mathbf{r}'$  and t = 0. The focus on intensity is intended to enable us to track the wave transport behavior after the phase coherence is destroyed by random scatterings. By definition, the pulse intensity is given by

$$P(t, \mathbf{r}, \mathbf{r}') = |G(t, \mathbf{r}, \mathbf{r}')|^2, \qquad (5.1)$$

where  $|G|^2 = GG^*$ . We are interested in the behavior of P for large t and large distance  $|\mathbf{r} - \mathbf{r}'|$  of the measuring point  $\mathbf{r}$  from  $\mathbf{r}'$ . It is important to point out that although in that limit the wave energy is predominantly incoherent, the pulse front *is always coherent* because by definition it is the part of the wave train which reaches any given point the earliest in time. Since any scattering would slow down the wave by increasing its travel path, the pulse front thus consists of the energy component that has suffered no scattering. Although this coherent component is exponentially small in amplitude, its importance lies in the fact that it is through pulse front propagation that the causal nature of the wave becomes manifest, and the propagation speed of the coherent wave can be measured by tracking the front. However, pulse-front tracking is *not* the approach developed below. Instead, we intend to track the *dominant energy component* of the wave, which has an entirely different transport behavior.

For various reasons that will become clear later, the information on intensity transport is easier to obtain in the Fourier domain. Since in the frequency domain

$$G(t, \mathbf{r}, \mathbf{r}') = \frac{1}{2\pi} \int_{-\infty}^{\infty} G^+(\omega_1, \mathbf{r}, \mathbf{r}') \exp(-i\omega_1 t) d\omega_1, \qquad (5.2)$$

$$G^*(t, \mathbf{r}, \mathbf{r}') = \frac{1}{2\pi} \int_{-\infty}^{\infty} G^-(\omega_2, \mathbf{r}', \mathbf{r}) \exp(\mathrm{i}\omega_2 t) \mathrm{d}\omega_2, \qquad (5.3)$$

we have

$$|G(t, \mathbf{r}, \mathbf{r}')|^2 = \frac{1}{(2\pi)^2} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} G^+(\omega_1, \mathbf{r}, \mathbf{r}') G^-(\omega_2, \mathbf{r}', \mathbf{r}) \\ \times \exp(-\mathrm{i}\Delta\omega t) \mathrm{d}\omega_1 \mathrm{d}\omega_2.$$
(5.4)

Here  $\Delta \omega = \omega_1 - \omega_2$  is the frequency difference; (5.3) follows from the solution to Problem 2.3 [(P2.10) and (P2.12)], and  $G^+$ ,  $G^-$  denote the retarded and the advanced Green functions, respectively. In a uniform medium, they are seen to correspond to the outgoing (+) and the incoming (-) waves associated with a point source, as worked out in Chap. 2. The important point to note about (5.4) is that  $\Delta \omega$  is the conjugate variable to the travel time t for intensity transport. This is reasonable because  $\Delta \omega$  is the "beating frequency" relevant to the propagation of the wave packet *envelope*, which corresponds to the motion of the center of mass for a quantum particle (wave packet). Due to the nature of the conjugate variables, the  $t \to \infty$  limit corresponds to the  $\Delta \omega \to 0$ limit because for  $\Delta \omega > 1$  the rapid oscillations of the factor  $\exp(-i \Delta \omega t)$  in the integrand of (5.4) make the total integrated contribution negligible. The only significant contribution is thus from the region  $\Delta \omega \leq 1/t$ .

The spectral component of  $P(t, \mathbf{r}, \mathbf{r}')$  is given by

$$P(\bar{\omega}, \mathbf{r}, \mathbf{r}') = \int_{-\infty}^{\infty} P(t, \mathbf{r}, \mathbf{r}') \exp(\mathrm{i}\bar{\omega} t) \mathrm{d}t.$$
(5.5)

The substitution of (5.4) into (5.5) gives

$$P(\bar{\omega}, \mathbf{r}, \mathbf{r}') = \frac{1}{2\pi} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} G^{+}(\omega_{1}, \mathbf{r}, \mathbf{r}') G^{-}(\omega_{2}, \mathbf{r}', \mathbf{r}) \delta(\bar{\omega} - \Delta \omega) d\omega_{1} d\omega_{2}$$

$$= \frac{1}{2\pi} \int_{-\infty}^{\infty} G^{+}\left(\omega + \frac{\Delta \omega}{2}, \mathbf{r}, \mathbf{r}'\right) G^{-}\left(\omega - \frac{\Delta \omega}{2}, \mathbf{r}', \mathbf{r}\right) d\omega,$$
(5.6)

where the time integration has been performed to yield  $2\pi\delta(\bar{\omega} - \Delta\omega)$ , and the frequency integration variables have been changed from  $\omega_1$ ,  $\omega_2$  to  $\omega = (\omega_1 + \omega_2)/2$  and  $\Delta\omega = \omega_1 - \omega_2$ . The spectral frequency  $\bar{\omega}$  is seen to be identical to  $\Delta\omega$ . Since  $\Delta\omega$  is related to the frequency of the wave packet envelope and  $\omega$ to the wave frequency inside the wave packet as shown in Fig. 5.1, sometimes  $\Delta\omega$  is also called the external frequency, or the modulation frequency, and  $\omega$ is called the internal frequency, or the carrier frequency.

For the spatial spectral component of P, it is simpler to examine the configurationally averaged quantity

$$\left\langle P(\Delta\omega, \mathbf{r}, \mathbf{r}')\right\rangle_{c} = \frac{1}{2\pi} \int_{-\infty}^{\infty} \left\langle G^{+}\left(\omega + \frac{\Delta\omega}{2}, \mathbf{r}, \mathbf{r}'\right) G^{-}\left(\omega - \frac{\Delta\omega}{2}, \mathbf{r}', \mathbf{r}\right) \right\rangle_{c} \mathrm{d}\omega.$$
(5.7)



Fig. 5.1. An amplitude-modulated wave. The faster variation inside the envelope corresponds to frequency  $\omega$ . The frequency of the envelope corresponds to  $\Delta \omega$ .

In the solution to Problem 5.1, it is shown that

$$S(\Delta\omega, \mathbf{r} - \mathbf{r}'|\omega) = \left\langle G^{+} \left( \omega + \frac{\Delta\omega}{2}, \mathbf{r}, \mathbf{r}' \right) G^{-} \left( \omega - \frac{\Delta\omega}{2}, \mathbf{r}', \mathbf{r} \right) \right\rangle_{c}$$

$$= \frac{L^{3d}}{N^{3} (2\pi)^{3d}} \iiint \left\langle \left\langle \mathbf{k}_{+} | \mathbf{G}^{+}(\omega^{+}) | \mathbf{k}'_{+} \right\rangle \left\langle \mathbf{k}'_{-} | \mathbf{G}^{-}(\omega^{-}) | \mathbf{k}_{-} \right\rangle \right\rangle_{c}$$

$$\times \exp \left[ i\Delta \mathbf{k} \cdot (\mathbf{r} - \mathbf{r}') \right] d\mathbf{k} d\mathbf{k}' d\Delta \mathbf{k},$$

$$= \frac{1}{N^{3}} \sum_{\mathbf{k}} \sum_{\mathbf{k}'} \sum_{\Delta \mathbf{k}} \left\langle \left\langle \mathbf{k}_{+} | \mathbf{G}^{+} | \mathbf{k}'_{+} \right\rangle \left\langle \mathbf{k}'_{-} | \mathbf{G}^{-} | \mathbf{k}_{-} \right\rangle \right\rangle_{c}$$

$$\times \left\langle \mathbf{r} | \Delta \mathbf{k} \right\rangle \left\langle \Delta \mathbf{k} | \mathbf{r}' \right\rangle, \qquad (5.8)$$

where all the  ${\bf k}$  summations are understood to be within the first Brillouin zone in the lattice case, and

$$\mathbf{k}_{+} = \mathbf{k} + \frac{\Delta \mathbf{k}}{2},$$
  

$$\mathbf{k}_{-} = \mathbf{k} - \frac{\Delta \mathbf{k}}{2},$$
  

$$\mathbf{k}'_{+} = \mathbf{k}' + \frac{\Delta \mathbf{k}}{2},$$
  

$$\mathbf{k}'_{-} = \mathbf{k}' - \frac{\Delta \mathbf{k}}{2}.$$
(5.9)

The abbreviations  $\omega^+ = \omega + \Delta \omega/2$ ,  $\omega^- = \omega - \Delta \omega/2$  have been used. Equation (5.8) tells us that the averaged quantity depends only on  $\mathbf{r} - \mathbf{r}'$  and that the conjugate variable to the travel distance  $|\mathbf{r} - \mathbf{r}'|$  is the difference in the momenta,  $\Delta \mathbf{k}$ , of the two Green functions, which have a total of three distinct momenta after configurational averaging. Since  $\Delta \mathbf{k}$  and  $\mathbf{r} - \mathbf{r}'$  are conjugate variables, the  $|\mathbf{r} - \mathbf{r}'| \rightarrow \infty$  limit therefore corresponds to the  $|\Delta \mathbf{k}| \rightarrow 0$  limit. From (5.8) it is clear that the  $\Delta \mathbf{k}$  Fourier component of  $S(\Delta \omega, \mathbf{r} - \mathbf{r}'|\omega)$  is given by

$$S(\Delta\omega, \Delta \mathbf{k}|\omega) = \frac{1}{N^2} \sum_{\mathbf{k}} \sum_{\mathbf{k}'} \phi_{\mathbf{k},\mathbf{k}'}(\Delta\omega, \Delta \mathbf{k}|\omega), \qquad (5.10)$$
where

 $\phi$ 

$$_{\mathbf{k},\mathbf{k}'}(\Delta\omega,\Delta\mathbf{k}|\omega) = \left\langle \left\langle \mathbf{k}_{+} | \mathbf{G}^{+}(\omega^{+}) | \mathbf{k}'_{+} \right\rangle \left\langle \mathbf{k}'_{-} | \mathbf{G}^{-}(\omega^{-}) | \mathbf{k}_{-} \right\rangle \right\rangle_{c}.$$
(5.11)

The desired information is the behavior of  $S(\Delta \omega, \Delta \mathbf{k} | \omega)$  in the limit of  $\Delta \omega$ ,  $|\Delta \mathbf{k}| \rightarrow 0$ . It is important to be aware that in seeking the behavior of S, the approach to be followed later is analogous to that for obtaining the diffusive behavior from Brownian motion. Namely, instead of following the wave in a given medium for a long time and over many collisions, we look for *configura*tionally averaged behavior and assume the two to be the same. The fact that this is true is not trivial; it is a reflection of the ergodicity of the system; i.e., the infinite-time average of the transported intensity in an (infinite) sample is equivalent to its spatial average over many configurations. Therefore, random scatterings will not appear explicitly in the subsequent development. Instead, the effect of random scatterings is captured by configurational averaging. The limitation of this approach is that the derived behavior is only in the *expected* sense mathematically. For any given configuration, the expected behavior can deviate from the actual behavior. Such a deviation may be especially large at short propagation times when the number of scatterings is small, and the effect of time averaging is therefore minimal. However, for the long-time, averaged property this approach should be accurate.

In anticipation of the result that pulse intensity transport is diffusive in the long-time, large-travel-distance limit, the desired form of  $S(\Delta \omega, \Delta \mathbf{k} | \omega)$ may be obtained by assuming  $\langle P(t, \mathbf{r}, \mathbf{r}') \rangle_c$  to be diffusive so that

$$(-i\Delta\omega - D(\omega)\nabla^2)S(\Delta\omega, \mathbf{r} - \mathbf{r}'|\omega) = \text{constant} \cdot \delta(\mathbf{r} - \mathbf{r}'), \qquad (5.12)$$

where the factor  $-i\Delta\omega$  is recognized to originate from the first-order time derivative  $\partial/\partial t$  of the diffusion equation,  $D(\omega)$  is the wave diffusion constant, and  $\nabla^2$  is understood to operate on **r**. By writing  $S(\Delta\omega, \mathbf{r} - \mathbf{r}' | \omega)$  and  $\delta(\mathbf{r} - \mathbf{r}')$ as the sums of their spectral components, one immediately obtains

$$S(\Delta\omega, \Delta \mathbf{k}|\omega) \propto \frac{1}{-i\Delta\omega + D(\omega)|\Delta \mathbf{k}|^2}, \qquad \Delta\omega, \ |\Delta \mathbf{k}| \to 0.$$
 (5.13)

That is,  $S(\Delta \omega, \Delta \mathbf{k} | \omega)$  has a *diffusive pole*. The derivation of this form and the calculation of  $D(\omega)$  are detailed below.

# 5.3 The Bethe-Salpeter Equation and its Solution by Moments

The evaluation of the configurational average of two Green functions requires an equation that relates it to the effect of random scatterings. To that end, let us start with the operator formalism defined relative to the effect medium, where

$$\mathbf{G}^{+} = \mathbf{G}_{\mathrm{e}}^{+} + \mathbf{G}_{\mathrm{e}}^{+} \bar{\mathbf{T}}^{+} \mathbf{G}_{\mathrm{e}}^{+}, \qquad (5.14a)$$

$$\mathbf{G}^{-} = \mathbf{G}_{\mathrm{e}}^{-} + \mathbf{G}_{\mathrm{e}}^{-} \mathbf{\bar{T}}^{-} \mathbf{G}_{\mathrm{e}}^{-}, \qquad (5.14\mathrm{b})$$

and

$$\left\langle \bar{\mathbf{T}}^{+} \right\rangle_{c} = \left\langle \bar{\mathbf{T}}^{-} \right\rangle_{c} = 0$$
 (5.15)

to within the CPA accuracy. For intensity transport, what is required is the configurationally averaged outer product

$$\left\langle \mathbf{G}^{+} \otimes \mathbf{G}^{-} \right\rangle_{c} = \mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-} + \left\langle \mathbf{G}_{e}^{+} \bar{\mathbf{T}}^{+} \mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-} \bar{\mathbf{T}}^{-} \mathbf{G}_{e}^{-} \right\rangle_{c}, \qquad (5.16)$$

where terms involving single  $\bar{\mathbf{T}}^{\pm}$  are averaged to zero in accordance with the condition (5.15). Here the concept of the outer product (of two vectors) has been discussed earlier in Chap. 2 in connection with the Dirac notation. For two vectors, the outer product yields a matrix. For the operators in (5.16), which are matrices, the outer product gives an object with four indices, i.e., a fourth-rank tensor. We denote each index as representing a "channel," which can be either in real space or in momentum space. Configurational averaging means that the number of independent channel variables is generally one less than the number of channels. This is shown in the solution to Problem 5.1.

The second term of (5.16) can be written as

$$\mathbf{G}_{e}^{+} \left\langle \bar{\mathbf{T}}^{+} \mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-} \bar{\mathbf{T}}^{-} \right\rangle_{c} \mathbf{G}_{e}^{-} = \left( \mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-} \right) : \left\langle \bar{\mathbf{T}}^{+} \otimes \bar{\mathbf{T}}^{-} \right\rangle_{c} : \left( \mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-} \right), \tag{5.17}$$

$$= \left( \mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-} \right) : \mathbf{\Gamma} : \left( \mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-} \right).$$

## 5.3.1 Notations

Here an explanation of the notations is important for understanding the subsequent development.  $\Gamma$  is known as the "vertex" function. It has two input channels and two output channels. If  $\Gamma$  is expressed as  $\langle \bar{\mathbf{T}}^+ \otimes \bar{\mathbf{T}}^- \rangle_c$ , then the two input channels are on the outside left and right, whereas the two output channels are on the center left and the center right, facing each other across the outer product sign, i.e.,

Outputs  
Input 
$$\rightarrow \overline{T}^+ \stackrel{\frown}{\longrightarrow} \otimes \stackrel{\frown}{\longleftarrow} \overline{T}^- \leftarrow$$
 Input.

In the convention used later,  $(\bar{\mathbf{T}}^+ \otimes \bar{\mathbf{T}}^-) : (\mathbf{G}_e^+ \otimes \mathbf{G}_e^-)$  means  $\bar{\mathbf{T}}^+ \mathbf{G}_e^+ \otimes \mathbf{G}_e^- \bar{\mathbf{T}}^-$ , whereas  $(\mathbf{G}_e^+ \otimes \mathbf{G}_e^-) : (\bar{\mathbf{T}}^+ \otimes \bar{\mathbf{T}}^-)$  means  $\mathbf{G}_e^+ \bar{\mathbf{T}}^+ \otimes \bar{\mathbf{T}}^- \mathbf{G}_e^-$ ; One can see that graphically as

Input 
$$\rightarrow \overline{\mathbf{T}}^+ \cdot \mathbf{G}_e^+ \stackrel{\frown}{\longrightarrow} \otimes \stackrel{\frown}{\longrightarrow} \mathbf{G}_e^- \cdot \overline{\mathbf{T}}^- \leftarrow \text{Input}$$

being equivalent to

 $\begin{array}{l} + \mbox{ Channel input} \rightarrow \mathbf{\bar{T}}^+ \cdot \mathbf{G}_{\rm e}^+ \rightarrow + \mbox{ Channel output} \\ \otimes \quad \otimes \\ - \mbox{ Channel input} \rightarrow \mathbf{\bar{T}}^- \cdot \mathbf{G}_{\rm e}^- \rightarrow - \mbox{ Channel output} \end{array}$ 



Fig. 5.2. The schematic diagram for the vertex function  $\Gamma$  with its four channels, denoted by the two colons (*solid circles*).

Here the inner product between two operators is denoted by a dot, which is usually omitted.  $\Gamma$  can be represented graphically by a box with four leads, two in and two out, as shown in Fig. 5.2. The two top leads are the + channels, indicated by the upper dots of the two colons, and the two bottom leads are the – channels, indicated by the bottom dots of the two colons. In this alternative representation, the two channels to the left are the input ones, and the two to the right are the output ones. The purpose of the colons is to remind the readers that the channels indicated form inner products with the object(s) to its left or right in the manner just described. Here + means (1) the frequency is  $\omega + \Delta \omega/2$ , (2) the wave vector is  $\mathbf{k} + \Delta \mathbf{k}/2$ , and (3) the Green function is retarded. The – has similarly implied meanings. The directions of the arrows, representing the Green functions, are opposite for the + and – channels because  $\mathbf{G}_{\rm e}^+$  and  $\mathbf{G}_{\rm e}^-$  generally represent opposite directions of propagation.

In (5.17), we have used the following general equivalence:  $\mathbf{A}\overline{\mathbf{T}}^{+}\mathbf{B}\otimes\mathbf{B}'\overline{\mathbf{T}}^{-}\mathbf{A}' = (\mathbf{A}\otimes\mathbf{A}') : (\overline{\mathbf{T}}^{+}\otimes\overline{\mathbf{T}}^{-}) : (\mathbf{B}\otimes\mathbf{B}')$ . At the risk of belaboring the point, this equivalence can also be demonstrated directly through the summation of indices. Let  $(\mathbf{A})_{ij}\otimes(\mathbf{A}')_{kl} = (\mathbf{A}\otimes\mathbf{A}')_{iljk}$ , where *i*, *l* are the input channels and *j*, *k* are the output channels. Then

$$(\mathbf{A}\bar{\mathbf{T}}^{+}\mathbf{B})_{ij} = \sum_{l,k} (\mathbf{A})_{il}(\bar{\mathbf{T}}^{+})_{lk}(\mathbf{B})_{kj},$$
$$(\mathbf{B}'\bar{\mathbf{T}}^{-}\mathbf{A}')_{st} = \sum_{m,n} (\mathbf{B}')_{sm}(\bar{\mathbf{T}}^{-})_{mn}(\mathbf{A}')_{nt},$$

and therefore

$$(\mathbf{A}\bar{\mathbf{T}}^{+}\mathbf{B}\otimes\mathbf{B}'\bar{\mathbf{T}}^{-}\mathbf{A}')_{itjs} = \sum_{l,k,m,n} (\mathbf{A})_{il}(\bar{\mathbf{T}}^{+})_{lk}(\mathbf{B})_{kj}(\mathbf{B}')_{sm}(\bar{\mathbf{T}}^{-})_{mn}(\mathbf{A}')_{nt}$$
$$= \sum_{l,k,m,n} (\mathbf{A}\otimes\mathbf{A}')_{itln}(\bar{\mathbf{T}}^{+}\otimes\bar{\mathbf{T}}^{-})_{lnkm}(\mathbf{B}\otimes\mathbf{B}')_{kmjs}$$
$$= [(\mathbf{A}\otimes\mathbf{A}'):(\bar{\mathbf{T}}^{+}\otimes\bar{\mathbf{T}}^{-}):(\mathbf{B}\otimes\mathbf{B}')]_{itjs}.$$

# 5.3.2 Irreducible Vertex Function and the Bethe–Salpeter Equation

Let us introduce a new vertex function  ${\bf U}$  as

$$\Gamma = \mathbf{U} + \mathbf{U} : (\mathbf{G}_{\mathbf{e}}^+ \otimes \mathbf{G}_{\mathbf{e}}^-) : \Gamma, \qquad (5.18)$$

where the second term on the right-hand side means the + and – *output* channels of **U** form inner products with the respective + and – *input* channels of  $\mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-}$ , whose *output* channels form inner products with the respective + and – *input* channels of  $\Gamma$ . By substituting (5.18) into (5.16) and (5.17), one gets

$$\begin{split} \left\langle \mathbf{G}^{+}\otimes\mathbf{G}^{-}\right\rangle_{\mathrm{c}} &= \mathbf{G}_{\mathrm{e}}^{+}\otimes\mathbf{G}_{\mathrm{e}}^{-} + (\mathbf{G}_{\mathrm{e}}^{+}\otimes\mathbf{G}_{\mathrm{e}}^{-}):\mathbf{U}:(\mathbf{G}_{\mathrm{e}}^{+}\otimes\mathbf{G}_{\mathrm{e}}^{-})\\ &+ (\mathbf{G}_{\mathrm{e}}^{+}\otimes\mathbf{G}_{\mathrm{e}}^{-}):\mathbf{U}:(\mathbf{G}_{\mathrm{e}}^{+}\otimes\mathbf{G}_{\mathrm{e}}^{-}):\boldsymbol{\Gamma}:(\mathbf{G}_{\mathrm{e}}^{+}\otimes\mathbf{G}_{\mathrm{e}}^{-})\\ &= \mathbf{G}_{\mathrm{e}}^{+}\otimes\mathbf{G}_{\mathrm{e}}^{-} + (\mathbf{G}_{\mathrm{e}}^{+}\otimes\mathbf{G}_{\mathrm{e}}^{-}):\mathbf{U}:[\mathbf{G}_{\mathrm{e}}^{+}\otimes\mathbf{G}_{\mathrm{e}}^{-})\\ &+ (\mathbf{G}_{\mathrm{e}}^{+}\otimes\mathbf{G}_{\mathrm{e}}^{-}):\boldsymbol{\Gamma}:(\mathbf{G}_{\mathrm{e}}^{+}\otimes\mathbf{G}_{\mathrm{e}}^{-})]. \end{split}$$

Since the quantity in the square bracket is just  $\langle \mathbf{G}^+ \otimes \mathbf{G}^- \rangle_c$ , we have

$$\left\langle \mathbf{G}^{+} \otimes \mathbf{G}^{-} \right\rangle_{\mathrm{c}} = \mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-} + \left( \mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-} \right) : \mathbf{U} : \left\langle \mathbf{G}^{+} \otimes \mathbf{G}^{-} \right\rangle_{\mathrm{c}}.$$
 (5.19)

Equation (5.19), together with the definitions of **U** and  $\Gamma$ , is known as the Bethe-Salpeter equation. It is the analog of the Dyson equation for  $\langle \mathbf{G}^+ \rangle_c$ , and **U** is the two-Green-function analog of self-energy.

From (5.18),  $\Gamma$  may be expressed as a series of U:

$$\Gamma = \mathbf{U} + \mathbf{U} : \left(\mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-}\right) : \mathbf{U} + \mathbf{U} : \left(\mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-}\right) : \mathbf{U} : \left(\mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-}\right) : \mathbf{U} + \cdots$$
(5.20)

This series is represented graphically in Fig. 5.3. Since  $\Gamma$  can be represented as the sum of many terms, it is called the "reducible" vertex function. In contrast, **U** is called the "irreducible" vertex function because it cannot be further decomposed.

Leaving aside the calculation of U for the moment, let us explore the solution of (5.19) in the momentum representation, at the limit of small  $\Delta \omega$ 



Fig. 5.3. Diagrammatic representation of the reducible vertex function  $\Gamma$  in terms of repeated U's, the irreducible vertex function. The channels are denoted by *solid circles*.

and  $|\Delta \mathbf{k}|$ . In *k*-space,  $\langle \mathbf{G}^+ \otimes \mathbf{G}^- \rangle_c = \phi_{\mathbf{k},\mathbf{k}'}(\Delta \omega, \Delta \mathbf{k}|\omega)$ , and  $\mathbf{G}_e^+ \otimes \mathbf{G}_e^- = NG_e^+(\omega^+, \mathbf{k}^+) \ G_e^-(\omega^-, \mathbf{k}^-)\delta_{\mathbf{k},\mathbf{k}'}$ . Therefore, (5.19) may be written as

$$\phi_{\mathbf{k},\mathbf{k}'} = (G_{e}^{+}G_{e}^{-})_{\mathbf{k}} \left[ N\delta_{\mathbf{k},\mathbf{k}'} + \frac{1}{N}\sum_{\mathbf{k}_{1}} U_{\mathbf{k},\mathbf{k}_{1}}\phi_{\mathbf{k}_{1},\mathbf{k}'} \right].$$
(5.21)

Here **k** denotes the input channels  $(\mathbf{k}+\Delta\mathbf{k}/2 \text{ and } \mathbf{k}-\Delta\mathbf{k}/2)$ , and  $\mathbf{k}'$  the output channels. The inner product between  $(G_{\rm e}^+G_{\rm e}^-)_k$  and  $U_{\mathbf{k},\mathbf{k}_1}$  has already been performed in (5.21) as shown by the identical **k** index, a result of summation over a delta function. From (5.21) it is clear that the  $\mathbf{k}'$  variable may be summed to give an equation in terms of the quantity

$$\phi_{\mathbf{k}} = \frac{1}{N} \sum_{\mathbf{k}'} \phi_{\mathbf{k}, \mathbf{k}'}.$$
(5.22)

The new equation for  $\phi_{\mathbf{k}}(\Delta \omega, \Delta \mathbf{k} | \omega)$  is

$$\phi_{\mathbf{k}} = (G_{e}^{+}G_{e}^{-})_{\mathbf{k}} \left[ 1 + \frac{1}{N} \sum_{\mathbf{k}_{1}} U_{\mathbf{k}, \, \mathbf{k}_{1}} \phi_{\mathbf{k}_{1}} \right].$$
(5.23)

To proceed further, it is noted that since the product of two numbers a, b may be expressed as

$$ab = \frac{a-b}{b^{-1}-a^{-1}},\tag{5.24}$$

it follows that

$$G_{\rm e}^+G_{\rm e}^- = \frac{G_{\rm e}^+ - G_{\rm e}^-}{(G_{\rm e}^-)^{-1} - (G_{\rm e}^+)^{-1}}.$$
(5.25)

From the definition of  $G_{\rm e}^+(\omega, \mathbf{k})$  as an approximation to  $\langle G^+ \rangle_{\rm c}(\omega, k)$ , with  $\beta = 1$  [(3.51)],

$$\frac{N}{L^d} \left[ (G_{\mathbf{e}}^-)^{-1} - (G_{\mathbf{e}}^+)^{-1} \right] = -\frac{\mathrm{d}\kappa_0^2}{\mathrm{d}\omega} \Delta \omega + \nabla_k e(\mathbf{k}) \cdot \Delta \mathbf{k} + \Sigma^+(k_+, \omega^+) - \Sigma^-(k_-, \omega^-),$$
(5.26)

where  $\nabla_k$  means the gradient operator on the variables  $k_x$ ,  $k_y$ , and  $k_z$ . For simplicity, the definition of  $e(\mathbf{k})$  is generalized for the classical wave as  $e(\mathbf{k}) =$  $\mathbf{k} \cdot \mathbf{k}$ , so that  $\nabla_k e(\mathbf{k}) = 2\mathbf{k}$  in that case. On the other hand, in the limit of  $|\Delta \mathbf{k}|, \Delta \omega \to 0, G_e^+ - G_e^-$  may be approximated as

$$\lim_{|\Delta \mathbf{k}|, \, \Delta \omega \to 0} (G_{\rm e}^+ - G_{\rm e}^-) = (\Delta G_{\rm e})_{\mathbf{k}} = \frac{2i{\rm Im}\Sigma^+}{[\kappa_0^2 - e(\mathbf{k}) - {\rm Re}(\Sigma^+)]^2 + ({\rm Im}\Sigma^+)^2} \frac{N}{L^d}.$$
(5.27)

Equation (5.27) is obtained with the knowledge that the Im $\Sigma^+$  and Im $\Sigma^$ differ only by a sign. By defining  $\kappa_e^2 = \kappa_0^2 - \text{Re}(\Sigma^+)$  and evaluating  $\Sigma^+$  relative to the effective medium  $\kappa_e$ ,  $(\Delta G_e)_{\mathbf{k}}$  may be approximated by a delta function when  $|\text{Im}\Sigma^+|$  is small (as shown in the solution to Problem 2.1), given by

$$(\Delta G_{\rm e})_{\mathbf{k}} = 2\mathrm{i}\mathrm{Im}G_{\rm e}^+(\omega,\mathbf{k}) \simeq -2\pi\mathrm{i}\,\delta[\kappa_{\rm e}^2 - e(\mathbf{k})]\frac{N}{L^d}.$$
(5.28)

Here the negative sign comes from the fact that  $\text{Im}\Sigma^+ < 0$ , as required by the causal behavior of retarded Green function. Note that although  $\Sigma^+$  is evaluated relative to the effective medium, the overbar on  $\Sigma^+$  has been dropped for simplicity. The last equality of (5.28) can be seen alternatively from the fact that in the limit of  $|\Delta \mathbf{k}|$ ,  $\Delta \omega \to 0$ ,  $G_{\rm e}^+$ ,  $G_{\rm e}^-$  are complex conjugates of each other. The approximation expressed by (5.28) will be used extensively in the following calculations. The quantity  $(\Delta G_{\rm e})_{\mathbf{k}}$  is alternatively known as the spectral function because it peaks at the dispersion relation(s) of the effective medium excitation(s). Its use in extending the CPA to the intermediate-frequency regime has been presented in the last chapter.

With the conversion of  $G_e^+G_e^-$  to  $(\Delta G_e)_k/[(G_e^-)^{-1}-(G_e^+)^{-1}]$ , a new form of the Bethe-Salpeter equation is obtained:

$$\phi_{\mathbf{k}} = \frac{(\Delta G_{\mathrm{e}})_{\mathbf{k}}}{(\Delta G_{\mathrm{e}}^{-1})_{\mathbf{k}}} \left[ 1 + \frac{1}{N} \sum_{\mathbf{k}_{1}} U_{\mathbf{k}, \mathbf{k}_{1}} \phi_{\mathbf{k}_{1}} \right], \tag{5.29}$$

where  $(\Delta G_{\rm e}^{-1}) = (G_{\rm e}^{-})^{-1} - (G_{\rm e}^{+})^{-1}$ . It is now possible to use a simple model to see the emergence of the diffusive behavior, and its relation to an identity.

## 5.3.3 Diffusive Behavior from a Simplified Model

If we assume  $U_{\mathbf{k},\mathbf{k}_1} = U$  independent of the wavevectors [as in the case of Anderson model shown later, see (5.77)], then by recalling the definition

$$S(\Delta \omega, \Delta \mathbf{k} | \omega) = \frac{1}{N} \sum_{\mathbf{k}} \phi_{\mathbf{k}}$$

and summing both sides of (5.29) with respect to **k**, one immediately obtains

$$S = \left[\frac{\sum_{\mathbf{k}} (\Delta G_{\mathrm{e}})_{\mathbf{k}} / (\Delta G_{\mathrm{e}}^{-1})_{\mathbf{k}}}{1 - (U/N) \sum_{\mathbf{k}} (\Delta G_{\mathrm{e}})_{\mathbf{k}} / (\Delta G_{\mathrm{e}}^{-1})_{\mathbf{k}}}\right].$$
 (5.30)

This simple model is in fact reasonably realistic in many circumstances, as we will see later. Explicit expression for U may be obtained for the Anderson model, for example (see (5.105c)).

We wish to show that S displays a diffusive pole. Toward this goal, we recall that  $G_{\rm e}^+$  and  $G_{\rm e}^-$  differ in (1) the signs of the imaginary parts, (2) momenta being  $\mathbf{k} + \Delta \mathbf{k}/2$  and  $\mathbf{k} - \Delta \mathbf{k}/2$ , respectively, and (3) frequencies being  $\omega + \Delta \omega/2$ 

and  $\omega - \Delta \omega/2$ , respectively. If  $|\Delta \mathbf{k}|$ ,  $\Delta \omega \to 0$ , then the dominant form for  $(\Delta G_{\rm e})_{\mathbf{k}}$  is given by (5.27), whereas  $(\Delta G_{\rm e}^{-1})_{\mathbf{k}} = a^d (\Sigma^+ - \Sigma^-) = a^d \Delta \Sigma$  can be assumed to be independent of  $\mathbf{k}$  in the effective medium limit. Thus if we would like to expand  $\sum_{\mathbf{k}} (\Delta G_{\rm e})_{\mathbf{k}} / (\Delta G_{\rm e}^{-1})_{\mathbf{k}}$  to the leading orders of  $\Delta \omega$  and  $|\Delta \mathbf{k}|$ , the zeroth order term in the denominator of (5.30) is given by

$$1 - \frac{U}{L^d} \frac{\sum_{\mathbf{k}} (\Delta G_{\mathrm{e}})_{\mathbf{k}}}{\Delta \Sigma}.$$

This leading order term must vanish identically if S were to display a diffusive pole [see (5.13)]. In fact this proves to be the case – ensured by the so-called *Ward identity*:

$$a^d \Delta \Sigma = \frac{U}{N} \sum_{\mathbf{k}} (\Delta G_{\mathrm{e}})_{\mathbf{k}}.$$
 (5.31a)

The validity of (5.31a) will be demonstrated below within the CPA context. With the vanishing of the zeroth order term in the denominator of (5.30), the next leading order terms in the expansion must be  $C_1\Delta\omega + C_2|\Delta \mathbf{k}|$ . But it is easy to show that  $C_2 = 0$  because  $C_2 \propto \sum_{\mathbf{k}} (\nabla e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}}) \frac{\partial}{\partial e(\mathbf{k})} [(\Delta G_e)_{\mathbf{k}} / (\Delta G_e^{-1})_{\mathbf{k}}]$ . Here  $\Delta \hat{\mathbf{k}}$  denotes the unit vector in the direction of  $\Delta \mathbf{k}$ . Because the Green function's dependence on  $\mathbf{k}$  is through  $e(\mathbf{k})$ , which is an even function of  $\mathbf{k}$ , it follows that the derivative of the quantity in the square bracket with respect to  $e(\mathbf{k})$  must also be an even function of  $\mathbf{k}$ . But the gradient of  $e(\mathbf{k})$ must be an odd function of  $\mathbf{k}$ . Thus the summation of the product of an even function and an odd function of  $\mathbf{k}$  is identically zero due to antisymmetry. Thus the lowest order nonvanishing terms in the denominator of (5.30) must be

$$C_1 \Delta \omega + C_3 |\Delta \mathbf{k}|^2$$
.

This is the form of the diffusive pole we are looking for [see (5.13)]. By carrying out the necessary Taylor expansions, it is straightforward to obtain the actual formula for the diffusion coefficient in this simplified model. However, our task here is just to illustrate the inevitability of the diffusive pole and the connection between the diffusive behavior and the Ward identity. Expression for the diffusion coefficient will be obtained later in a more formal derivation.

#### 5.3.4 The Ward Identity in the Simplified Model

To complete the picture delineated by the simplified model, the validity of (5.31a) must be demonstrated. It is useful to note that  $(1/N) \sum_{\mathbf{k}} (\Delta G_{\mathrm{e}})_{\mathbf{k}}$  is exactly the Fourier transform of  $\Delta G(\mathbf{r} = \mathbf{r}')$  (or  $\mathbf{l} = \mathbf{l}'$ ) in real space. Hence (5.31a) is simply

$$U = \frac{a^d \Delta \Sigma}{\Delta G},\tag{5.31b}$$

where all quantities are scalars, which makes the demonstration especially simple. From the definition of U and  $\Gamma$ , (5.18) and (5.17), we obtain

$$U_i = U|i\rangle\langle i| = \frac{\langle \bar{t}_i^+ \bar{t}_i^- \rangle_c}{1 + G_e^+ G_e^- \langle \bar{t}_i^+ \bar{t}_i^- \rangle_c} |i\rangle\langle i|, \qquad (5.32)$$

where *i* denotes the site index (either **r** or **l**), averaged quantity is noted to be independent of the site index, and the overbar implies the quantity is evaluated with reference to the effective medium. Thus the task is to evaluate  $\langle \bar{t}_i^+ \bar{t}_i^- \rangle_c$ , with the CPA condition that [see (4.5)]

$$\langle \bar{t}_i^{\pm} \rangle_c = 0 = \left\langle \frac{a^d (\sigma_i^{\pm} - \Sigma^{\pm})}{1 - G_{\mathrm{e}}^{\pm} a^d (\sigma_i^{\pm} - \Sigma^{\pm})} \right\rangle_c$$
$$= \frac{1}{G_{\mathrm{e}}^{\pm}} \left( -1 + \frac{1}{G_{\mathrm{e}}^{\pm}} \left\langle \frac{1}{(G_{\mathrm{e}}^{\pm})^{-1} - a^d (\sigma_i^{\pm} - \Sigma^{\pm})} \right\rangle_c \right), \qquad (5.33)$$

where the last step is obtained through simple algebraic manipulation. By defining

$$L_i^{\pm} = \frac{1}{(G_{\rm e}^{\pm})^{-1} - a^d(\sigma_i^{\pm} - \Sigma^{\pm})}$$

so that  $\langle L_i^{\pm} \rangle_c = G_{\rm e}^{\pm}$  from (5.33), we obtain

$$\langle \bar{t}_i^+ \bar{t}_i^- \rangle_c = -\frac{1}{G_{\rm e}^+ G_{\rm e}^-} + \frac{1}{(G_{\rm e}^+ G_{\rm e}^-)^2} \langle L_i^+ L_i^- \rangle_c,$$
 (5.34)

so that the substitution of (5.34) into (5.32) yields

$$U = \frac{1}{G_{\rm e}^+ G_{\rm e}^-} - \frac{1}{\langle L_i^+ L_i^- \rangle_c}.$$
 (5.35)

Since  $\langle L_i^{\pm} \rangle_c = G_e^{\pm}$ , it follows that

$$\begin{aligned} G_{\rm e}^{+} - G_{\rm e}^{-} &= \langle L_{i}^{+} - L_{i}^{-} \rangle_{c} = \langle L_{i}^{+} [(L_{i}^{-})^{-1} - (L_{i}^{+})^{-1}] L_{i}^{-} \rangle_{c} \\ &= \langle L_{i}^{+} [(G_{\rm e}^{-})^{-1} - a^{d} (\sigma_{i}^{-} - \Sigma^{-}) - (G_{\rm e}^{+})^{-1} + a^{d} (\sigma_{i}^{+} - \Sigma^{+})] L_{i}^{-} \rangle_{c} \\ &= \langle L_{i}^{+} L_{i}^{-} \rangle_{c} \left( \frac{1}{G_{\rm e}^{-}} - \frac{1}{G_{\rm e}^{+}} - a^{d} \Delta \Sigma \right). \end{aligned}$$
(5.36)

Here we note that among the differences between the advanced and retarded quantities, only the difference in the signs of the imaginary parts is retained. Hence the scattering potentials in the + and – channels are taken to be the same, and they cancel. This is in accordance with the rule that the quantities in the Ward identity are evaluated at the  $\Delta \omega = |\Delta \mathbf{k}| = 0$  limit. Through simple algebraic manipulation, one obtains immediately from (5.36) that

$$\frac{1}{\langle L_i^+ L_i^- \rangle_c} = \frac{1}{\Delta G_{\rm e}} \left( \frac{\Delta G_{\rm e}}{G_{\rm e}^+ G_{\rm e}^-} - a^d \Delta \Sigma \right).$$
(5.37)

Substitution of (5.37) into (5.35) leads to the desired expression (5.31b). In what follows, a more formal derivation of the diffusive pole and the Ward identity will be presented. However, the basic physics is already foreshadowed by the simple model shown here.

# 5.3.5 Formal Derivation of the Diffusive Pole and the Diffusion Constant

We return to (5.29), where the multiplication of both sides by  $[(G_e^-)^{-1} - (G_e^+)^{-1}]$  and the substitution of (5.26) lead to

$$\frac{L^{d}}{N} \left( -\frac{\mathrm{d}\kappa_{0}^{2}}{\mathrm{d}\omega} \Delta \omega + \nabla_{k} e(\mathbf{k}) \cdot \Delta \mathbf{k} + \Sigma^{+} - \Sigma^{-} \right) \phi_{\mathbf{k}} \\
= (\Delta G_{\mathrm{e}})_{\mathbf{k}} \left[ 1 + \frac{1}{N} \sum_{\mathbf{k}'} U_{\mathbf{k}, \mathbf{k}'} \phi_{\mathbf{k}'} \right].$$
(5.38)

Equation (5.38) is the dynamical transport equation for wave intensity in the limit of  $\Delta \omega$ ,  $|\Delta \mathbf{k}| \rightarrow 0$ . It corresponds to the Boltzmann transport equation for classical particles. Just as in the classical case, one solution approach to (5.38) is by calculating the moments of  $\phi_{\mathbf{k}}$ . To do that, let us first sum both sides of (5.38) with respect to  $\mathbf{k}$ . This gives

$$-\frac{\mathrm{d}\kappa_0^2}{\mathrm{d}\omega}\Delta\omega S + 2\Delta\mathbf{k}\cdot\mathbf{J} + \frac{1}{N}\sum_{\mathbf{k}}\left(\Sigma^+ - \Sigma^-\right)\phi_{\mathbf{k}}$$
$$= \frac{2\mathrm{i}}{L^d}\sum_{\mathbf{k}}\mathrm{Im}G_{\mathrm{e}}^+(\omega,\mathbf{k}) + \frac{1}{L^dN}\sum_{\mathbf{k}}\sum_{\mathbf{k}'}\left(\Delta G_{\mathrm{e}}\right)_{\mathbf{k}}U_{\mathbf{k},\mathbf{k}'}\phi_{\mathbf{k}'}, \quad (5.39)$$

where S and J are the moments of  $\phi_{\mathbf{k}}$  defined as

$$S = \frac{1}{N} \sum_{\mathbf{k}} \phi_{\mathbf{k}} \,,$$

the energy density given by (5.10), and

$$\mathbf{J} = \frac{1}{2N} \sum_{\mathbf{k}} \nabla_k e(\mathbf{k}) \phi_{\mathbf{k}} \,, \tag{5.40}$$

which may be regarded as the energy current density. In (5.39) the first term on the right-hand side can be immediately evaluated to give

$$\frac{2\mathrm{i}}{L^d} \sum_{\mathbf{k}} \mathrm{Im} G_{\mathrm{e}}^+(\omega, \mathbf{k}) = -2\mathrm{i} \frac{\mathrm{d}\omega}{\mathrm{d}\kappa_{\mathrm{e}}^2} \pi \rho_{\mathrm{e}}(\omega) \frac{N}{L^d}.$$
(5.41)

Here  $\rho_{\rm e}(\omega)$  is the density of states of the effective medium. That leaves two terms in (5.39) that seem difficult to handle. However, by switching the summation labels **k** and **k'**, which effects no real change, we may invoke the Ward identity:

$$\Sigma^{+}(\omega^{+},k_{+}) - \Sigma^{-}(\omega^{-},k_{-}) = (\Delta\Sigma)_{\mathbf{k}} = \frac{1}{N} \sum_{\mathbf{k}'} (\Delta G_{\mathrm{e}})_{\mathbf{k}'} U_{\mathbf{k}',\mathbf{k}} \frac{N}{L^{d}}.$$
 (5.42)

Multiplying by  $\phi_{\mathbf{k}}$  and summation over k on both sides give

$$\frac{1}{N}\sum_{\mathbf{k}} \left(\Delta\Sigma\right)_{\mathbf{k}} \phi_{\mathbf{k}} = \frac{1}{L^d N} \sum_{\mathbf{k}'} \sum_{\mathbf{k}} \left(\Delta G_{\mathbf{e}}\right)_{\mathbf{k}'} U_{\mathbf{k}', \mathbf{k}} \phi_{\mathbf{k}}, \tag{5.43}$$

so the third term on the left hand side exactly cancels the second term on the right hand side, and (5.39) becomes

$$-\frac{\mathrm{d}\kappa_0^2}{\mathrm{d}\omega}\Delta\omega S + 2\Delta\mathbf{k}\cdot\mathbf{J} = -2\mathrm{i}\frac{\mathrm{d}\omega}{\mathrm{d}\kappa_\mathrm{e}^2}\pi\rho_\mathrm{e}(\omega)\frac{N}{L^d}.$$
 (5.44a)

The Ward identity is an important statement relating  $\Sigma$ , which is a quantity associated with the average of single Green function, to the vertex function **U**, which is associated with the average of two Green functions. The physical basis of the identity is energy conservation, and the formal demonstration of its validity is deferred to section 5.5.

Equation (5.44a) has two unknowns, S and  $\mathbf{J}$ . A physical interpretation of the equation may be obtained by interpreting the multiplication by  $-i\Delta\omega$ as the result of the  $\partial/\partial t$  operation, and the dot product with  $i\Delta \mathbf{k}$  as the result of the divergence operation ( $\nabla \cdot$ ). These interpretations originate from the fact that  $\Delta\omega$  and  $\Delta \mathbf{k}$  are the conjugate variables to travel time and travel distance, respectively, as pointed out earlier. Then (5.44a) has the form of a continuity equation with a source term, i.e.,

$$\frac{1}{2} \left( \frac{\mathrm{d}\kappa_0^2}{\mathrm{d}\omega} \right) \frac{\partial S}{\partial t} + \nabla \cdot \mathbf{J} = \frac{\mathrm{d}\omega}{\mathrm{d}\kappa_\mathrm{e}^2} \pi \rho_\mathrm{e}(\omega) \frac{N}{L^d}.$$
(5.44b)

#### 5.3.6 Higher Moment of the Transport Equation

In light of the above interpretation, a diffusive pole for S would be established if  $\mathbf{J}$  can be expressed as the gradient of S, i.e.,  $\mathbf{J} \sim |\Delta \mathbf{k}| S$ . In order to obtain an additional equation relating S and  $\mathbf{J}$ , one can take a higher moment of (5.38) by multiplying both sides by  $(1/2\nabla_k e(\mathbf{k}) \cdot \Delta \mathbf{k})$  and summing with respect to  $\mathbf{k}$ . However, since for  $|\Delta \mathbf{k}| \to 0$  this would result in a higher-order equation than (5.44a),  $G_e^+ - G_e^-$  has to be expanded consistently to the same order. That is,  $(\Delta G_e)_{\mathbf{k}}$  in (5.38) has to be replaced by

$$G_{\rm e}^{+} - G_{\rm e}^{-} = (\Delta G_{\rm e})_{\mathbf{k}} + \frac{\partial G_{\rm e}^{+}}{\partial e(\mathbf{k})} \left[ \nabla_{k} e(\mathbf{k}) \cdot \frac{\Delta \mathbf{k}}{2} \right] - \frac{\partial G_{\rm e}^{-}}{\partial e(\mathbf{k})} \left[ \nabla_{k} e(\mathbf{k}) \cdot \left( \frac{-\Delta \mathbf{k}}{2} \right) \right]$$
$$= (\Delta G_{\rm e})_{\mathbf{k}} + \left[ (G_{\rm e}^{+})_{\mathbf{k}}^{2} + (G_{\rm e}^{-})_{\mathbf{k}}^{2} \right] \left[ \frac{1}{2} \nabla_{k} e(\mathbf{k}) \cdot \Delta \mathbf{k} \right] \frac{L^{d}}{N}.$$
(5.45a)

The net result of the three operations [multiplication by  $(1/2\nabla_k e(\mathbf{k}) \cdot \Delta \mathbf{k})$ , expansion of  $G_{\rm e}^+ - G_{\rm e}^-$  to higher order, and summation with respect to  $\mathbf{k}$ ] is

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$$-\frac{\mathrm{d}\kappa_{0}^{2}}{\mathrm{d}\omega}\Delta\omega\Delta\mathbf{k}\cdot\mathbf{J} + \frac{2}{N}\sum_{\mathbf{k}}\left[\frac{1}{2}\nabla_{k}e(\mathbf{k})\cdot\Delta\mathbf{k}\right]^{2}\phi_{\mathbf{k}}$$

$$+\frac{1}{N}\sum_{\mathbf{k}}\mathrm{i}(\mathrm{Im}\Sigma^{+})\left[\nabla_{k}e(\mathbf{k})\cdot\Delta\mathbf{k}\right]\phi_{\mathbf{k}}$$

$$=\frac{1}{2L^{d}}\sum_{\mathbf{k}}\left[\nabla_{k}e(\mathbf{k})\cdot\Delta\mathbf{k}\right](\Delta G_{\mathrm{e}})_{\mathbf{k}}$$

$$+\frac{1}{2N^{2}}\sum_{\mathbf{k}}\sum_{\mathbf{k}'}\left(\Delta G_{\mathrm{e}})_{\mathbf{k}}\left[\nabla_{k}e(\mathbf{k})\cdot\Delta\mathbf{k}\right]\frac{N}{L^{d}}U_{\mathbf{k},\mathbf{k}'}\phi_{\mathbf{k}'}$$

$$+\frac{1}{4N}\sum_{\mathbf{k}}\left[\nabla_{k}e(\mathbf{k})\cdot\Delta\mathbf{k}\right]^{2}\left[\left(G_{\mathrm{e}}^{+}\right)_{\mathbf{k}}^{2}+\left(G_{\mathrm{e}}^{-}\right)_{\mathbf{k}}^{2}\right]$$

$$+\frac{L^{d}}{4N^{3}}\sum_{\mathbf{k}}\sum_{\mathbf{k}'}\left[\left(G_{\mathrm{e}}^{+}\right)_{\mathbf{k}}^{2}+\left(G_{\mathrm{e}}^{-}\right)_{\mathbf{k}}^{2}\right]\left[\nabla_{k}e(\mathbf{k})\cdot\Delta\mathbf{k}\right]^{2}\frac{N}{L^{d}}U_{\mathbf{k},\mathbf{k}'}\phi_{\mathbf{k}'}.$$
(5.45b)

The first term on the right-hand side of (5.45b) vanishes because  $(\Delta G_e)_{\mathbf{k}}$  is an even function of  $\mathbf{k}$  [ $e(\mathbf{k})$  is an even function of  $\mathbf{k}$  as can be seen from (2.37)], and  $[\nabla_k e(\mathbf{k}) \cdot \Delta \mathbf{k}]$  is an odd function of  $\mathbf{k}$  as can be seen from the solution to Problem 5.2, so the angular integration implicit in the  $\mathbf{k}$  summation forces it to be zero. Let us label the third term on the right-hand side as  $NR_0|\Delta \mathbf{k}|^2/L^d$ . If  $\Sigma^+$  is assumed to be  $\mathbf{k}$  independent, as in the case of CPA, then the third term on the left-hand side, denoted L3, may be written as

$$L3 = 2i(Im\Sigma^+)\Delta \mathbf{k} \cdot \mathbf{J}.$$

That leaves the second terms on the left-hand side and the second and the last terms on the right-hand side of (5.45b) to be treated, which we will label as L2, R2, and R4, respectively.

## 5.3.7 Expansion of $\phi_{\mathbf{k}}$

At this stage it is necessary to make a digression on the **k** dependence of  $\phi_{\mathbf{k}}$ From (5.23), the source term for  $\phi_{\mathbf{k}}$  is seen to be  $(G_{\mathrm{e}}^+G_{\mathrm{e}}^-)_{\mathbf{k}}$ , which can be expanded to obtain

$$G_{e}^{+}G_{e}^{-} = (G_{e}^{+})_{\mathbf{k}}(G_{e}^{-})_{\mathbf{k}}$$

$$+ \left[ (G_{e}^{-})_{\mathbf{k}}(G_{e}^{+})_{\mathbf{k}}^{2} - (G_{e}^{+})_{\mathbf{k}}(G_{e}^{-})_{\mathbf{k}}^{2} \right] \left[ \frac{1}{2} \nabla_{k} e(\mathbf{k}) \cdot \Delta \mathbf{k} \right] \frac{L^{d}}{N}$$

$$= \frac{(\Delta G_{e})_{\mathbf{k}}}{2i\mathrm{Im}\Sigma^{+}} \frac{N}{L^{d}} + \frac{(\Delta G_{e})_{\mathbf{k}}^{2}}{2i\mathrm{Im}\Sigma^{+}} \left[ \frac{1}{2} \nabla_{k} e(\mathbf{k}) \cdot \Delta \mathbf{k} \right], \qquad (5.46a)$$

where the identity

1

$$\frac{L^d}{N} (G_{\rm e}^+)_{\mathbf{k}} (G_{\rm e}^-)_{\mathbf{k}} = \frac{(\Delta G_{\rm e})_{\mathbf{k}}}{2\mathrm{i}\mathrm{Im}\Sigma^+}$$
(5.46b)

has been used. It follows that to the first approximation, if we take  $\Sigma^+$  to be only frequency dependent as in the CPA, then  $\phi_{\mathbf{k}}$  may be written as

$$\phi_{\mathbf{k}} \cong (\Delta G_{\mathbf{e}})_{\mathbf{k}} \left[ \phi^{(0)} + \frac{1}{2} (\nabla_k e(\mathbf{k}) \cdot \Delta \mathbf{k}) (\Delta G_{\mathbf{e}})_{\mathbf{k}} \frac{L^d}{N} \phi^{(1)} \right],$$
(5.47a)

where  $\phi^{(0)}$  and  $\phi^{(1)}$  are just functions of  $\Delta \omega$  and  $\omega$ . By summing both sides of (5.47a) with respect to  $\mathbf{k}$ ,  $\phi^{(0)}$  is identified as

$$\phi^{(0)} = \frac{\mathrm{i}}{2\pi} \frac{\mathrm{d}\kappa_{\mathrm{e}}^2}{\mathrm{d}\omega} \rho_{\mathrm{e}}^{-1}(\omega) S, \qquad (5.47\mathrm{b})$$

where the second term involving  $\nabla_k e(\mathbf{k}) \cdot \Delta \mathbf{k}$  vanishes for the same reason as stated earlier. By multiplying both sides of (5.47a) by  $(1/2)(\nabla_k e(\mathbf{k}) \cdot \Delta \mathbf{k})$ and summing with respect to  $\mathbf{k}$  again, the term involving  $\phi^{(0)}$  now vanishes; what remains is

$$\Delta \mathbf{k} \cdot \mathbf{J} = -|\Delta \mathbf{k}|^2 K_0 \phi^{(1)}, \qquad (5.48a)$$

with

$$K_0 = -\frac{L^d}{4N^2} \sum_{\mathbf{k}} \left[ \nabla_k e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}} \right]^2 (\Delta G_e)_{\mathbf{k}}^2, \qquad (5.48b)$$

where  $\Delta \hat{\mathbf{k}}$  is the unit vector in the direction of  $\Delta \mathbf{k}$ .  $K_0$  is noted to be a function of  $\omega$  only. For the purpose of later calculations, it would be useful to give a weak-scattering approximation to (5.48b). In that limit  $(\Delta G_e)^2_{\mathbf{k}}$  may be approximated by

$$(\Delta G_{\rm e})_{\mathbf{k}}^2 \cong -2(G_{\rm e}^+)_{\mathbf{k}}(G_{\rm e}^-)_{\mathbf{k}} = -\mathrm{i}\frac{(\Delta G_{\rm e})_{\mathbf{k}}}{-\mathrm{Im}\Sigma^+}\frac{N}{L^d}$$
$$\cong \frac{-2\pi}{-\mathrm{Im}\Sigma^+}\delta[\kappa_{\rm e}^2 - e(\mathbf{k})]\left(\frac{N}{L^d}\right)^2,$$

so that

$$K_0 = \frac{\pi}{2} \frac{1}{-\mathrm{Im}\Sigma^+} \sum_{\mathbf{k}} \left[ \nabla_k e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}} \right]^2 \delta[\kappa_{\mathrm{e}}^2 - e(\mathbf{k})] \frac{1}{L^d}.$$
 (5.48c)

The neglect of the  $(G_{\rm e}^+)_{\mathbf{k}}^2$  and  $(G_{\rm e}^-)_{\mathbf{k}}^2$  terms as being small compared to  $(G_{\rm e}^+)_{\mathbf{k}}(G_{\rm e}^-)_{\mathbf{k}}$  is justified if their respective magnitudes were compared when summed with respect to  $\mathbf{k}$ . Whereas the summation of  $(G_{\rm e}^+)_{\mathbf{k}}^2 = -\partial(G_{\rm e}^+)_{\mathbf{k}}/\partial\kappa_0^2$  gives something finite, the summation of  $(G_{\rm e}^+)_{\mathbf{k}}(G_{\rm e}^-)_{\mathbf{k}} \propto (\Delta G_{\rm e}^+)_{\mathbf{k}}/-\mathrm{Im}\Sigma^+$  gives  $\sim \rho(\omega)/-\mathrm{Im}\Sigma^+$ , which diverges as  $-\mathrm{Im}\Sigma^+ \to 0$ . Physically, in the limit of  $|\Delta \mathbf{k}| = \Delta \omega = 0$ ,  $(G_{\rm e}^+)_{\mathbf{k}}(G_{\rm e}^-)_{\mathbf{k}}$  is a real quantity, in contrast to  $(G_{\rm e}^+)_{\mathbf{k}}^2$  and  $(G_{\rm e}^-)_{\mathbf{k}}^2$ , which have finite imaginary parts, so that whereas  $(G_{\rm e}^+)_{\mathbf{k}}(G_{\rm e}^-)_{\mathbf{k}}$  represent evanescent modes.

The net result of the digression is that  $\phi_{\mathbf{k}}$  may be expanded as

$$\phi_{\mathbf{k}} \cong (\Delta G_{\mathbf{e}})_{\mathbf{k}} \left[ \frac{\mathrm{i}}{2\pi} \frac{\mathrm{d}\kappa_{\mathbf{e}}^2}{\mathrm{d}\omega} \rho_{\mathbf{e}}^{-1}(\omega) S - \frac{1}{2} |\Delta \mathbf{k}|^{-2} K_0^{-1} [\nabla_k e(\mathbf{k}) \cdot \Delta \mathbf{k}] (\Delta G_{\mathbf{e}})_{\mathbf{k}} \frac{L^d}{N} \Delta \mathbf{k} \cdot \mathbf{J} \right].$$
(5.49)

# 5.3.8 Solution for the Diffusive Pole and the Diffusion Constant

With this form of  $\phi_{\mathbf{k}}$  we return to (5.45b) and continue the solution of the Bethe-Salpeter equation. The substitution of (5.49) into the second term on the left side of (5.45b) gives

$$L2 = \frac{1}{N} \sum_{\mathbf{k}} \left[ \frac{1}{2} \nabla_k e(\mathbf{k}) \cdot \Delta \mathbf{k} \right]^2 2(\Delta G_e)_{\mathbf{k}} \frac{\mathrm{i}}{2\pi} \frac{\mathrm{d}\kappa_e^2}{\mathrm{d}\omega} \rho_e^{-1}(\omega) S$$

as a result. **J** does not appear here because of the presence of the odd function  $\nabla_k e(\mathbf{k}) \cdot (\Delta \mathbf{k})$ . The substitution of (5.49) into the last (fourth) term on the right-hand side of (5.45b) gives

$$R4 = \frac{L^d}{N^3} \sum_{\mathbf{k}} \sum_{\mathbf{k}'} \left[ \frac{1}{2} \nabla_k e(\mathbf{k}) \cdot \Delta \mathbf{k} \right]^2$$
$$\times \frac{N}{L^d} U_{\mathbf{k}, \mathbf{k}'} (\Delta G_e)_{\mathbf{k}'} \left[ (G_e^+)_{\mathbf{k}}^2 + (G_e^-)_{\mathbf{k}}^2 \right] \frac{i}{2\pi} \frac{d\kappa_e^2}{d\omega} \rho_e^{-1}(\omega) S.$$

From the Ward identity, (5.42), the summation over  $\mathbf{k}'$  directly gives  $\Sigma^+ - \Sigma^-$ . So a new form of the expression is

$$\mathbf{R4} = \frac{L^d}{N^2} \sum_{\mathbf{k}} \left[ \frac{1}{2} \nabla_k e(\mathbf{k}) \cdot \Delta \mathbf{k} \right]^2 (2i \mathrm{Im} \Sigma^+) \left[ (G_{\mathrm{e}}^+)_{\mathbf{k}}^2 + (G_{\mathrm{e}}^-)_{\mathbf{k}}^2 \right] \frac{\mathrm{i}}{2\pi} \frac{\mathrm{d}\kappa_{\mathrm{e}}^2}{\mathrm{d}\omega} \rho_{\mathrm{e}}^{-1}(\omega) S.$$

By noting that  $(\Delta G_e)_{\mathbf{k}}/\mathrm{iIm}\Sigma^+ = 2(G_e^+)_{\mathbf{k}}(G_e^-)_{\mathbf{k}}L^d/N$ , one can combine the two expressions L2 and R4 (with a minus sign), yielding

$$L2 - R4 = -\frac{L^d}{4N^2} \sum_{\mathbf{k}} \left[ \nabla_k e(\mathbf{k}) \cdot \Delta \mathbf{k} \right]^2 (\Delta G_e)_{\mathbf{k}}^2 \left( \frac{-\text{Im}\Sigma^+}{\pi \rho_e(\omega)} \frac{\mathrm{d}\kappa_e^2}{\mathrm{d}\omega} \right) S$$
$$= K_0 \left( \frac{-\text{Im}\Sigma^+}{\pi \rho_e(\omega)} \frac{\mathrm{d}\kappa_e^2}{\mathrm{d}\omega} \right) |\Delta \mathbf{k}|^2 S.$$

On the other hand, the substitution of (5.49) into the second term on the right side of (5.45b) gives

$$R2 = -\frac{1}{4} \frac{1}{N^2} \sum_{\mathbf{k}} \sum_{\mathbf{k}'} (\Delta G_e)_{\mathbf{k}} (\Delta G_e)_{\mathbf{k}'}^2 \left[ \nabla_k e(\mathbf{k}) \cdot \Delta \mathbf{k} \right] U_{\mathbf{k}, \mathbf{k}'}$$
$$\times \left[ \nabla_{k'} e(\mathbf{k}') \cdot \Delta \mathbf{k} \right] |\Delta \mathbf{k}|^{-2} K_0^{-1} \Delta \mathbf{k} \cdot \mathbf{J}$$
$$= -\frac{1}{4K_0} \left\{ \frac{L^d}{N^3} \sum_{\mathbf{k}} \sum_{\mathbf{k}'} (\Delta G_e)_{\mathbf{k}} (\Delta G_e)_{\mathbf{k}'}^2 \left[ \nabla_k e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}} \right] \right\}$$
$$\times \frac{N}{L^d} U_{\mathbf{k}, \mathbf{k}'} \left[ \nabla_{k'} e(\mathbf{k}') \cdot \Delta \hat{\mathbf{k}} \right] \right\} (\Delta \mathbf{k} \cdot \mathbf{J}). \tag{5.50}$$

The  $(\Delta G_{\rm e})_{\rm k}$ 's in (5.50) ensure that the magnitudes of  ${\bf k}'$  and  ${\bf k}$  are both determined by the constant-energy condition, i.e., they are both on the "energy shell." It follows that the real degrees of freedom lie in the angular part, i.e., the angle  $\theta$  between  ${\bf k}$  and  ${\bf k}'$ , and the angle of  ${\bf k}$  relative to the sample. If the sample is isotropic, only  $\theta$  matters in the summation. It can be seen from (5.50) that if  $U_{{\bf k},{\bf k}'}$  is independent of  $\theta$ , then the double summations can be performed independently, each giving zero. Therefore, this term is nonzero only when  $U_{{\bf k},{\bf k}'}$  is *anisotropic*. By defining

$$M_{0} = i \frac{L^{d}}{N^{3}} \sum_{\mathbf{k}} \sum_{\mathbf{k}'} (\Delta G_{e})_{\mathbf{k}} (\Delta G_{e})_{\mathbf{k}'}^{2} \left[ \nabla_{k} e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}} \right]$$
$$\times \frac{N}{L^{d}} U_{\mathbf{k}, \mathbf{k}'} \left[ \nabla_{k'} e(\mathbf{k}') \cdot \Delta \hat{\mathbf{k}} \right], \qquad (5.51a)$$

the second term on the right-hand side of (5.45b) has the form

$$\mathbf{R}2 = \frac{\mathbf{i}}{4K_0} M_0 \Delta \mathbf{k} \cdot \mathbf{J}.$$

In the weak-scattering limit,  $M_0$  may be alternatively expressed as

$$M_{0} = \frac{1}{N^{2}} \frac{1}{-\mathrm{Im}\Sigma^{+}} \sum_{\mathbf{k}} \sum_{\mathbf{k}'} (\Delta G_{\mathrm{e}})_{\mathbf{k}} (\Delta G_{\mathrm{e}})_{\mathbf{k}'} \left[ \nabla_{k} e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}} \right]$$
$$\times \frac{N}{L^{d}} U_{\mathbf{k}, \mathbf{k}'} \left[ \nabla_{k'} e(\mathbf{k}') \cdot \Delta \hat{\mathbf{k}} \right], \qquad (5.51b)$$

where  $(\Delta G_{\rm e})^2_{{\bf k}'}$  is again replaced by  $-i(\Delta G_{\rm e})_{{\bf k}'} N/(-\mathrm{Im}\Sigma^+)L^d$  as before. By combining terms, (5.45b) has the following form:

$$-\frac{\mathrm{d}\kappa_{0}^{2}}{\mathrm{d}\omega}\Delta\omega\Delta\mathbf{k}\cdot\mathbf{J} + \frac{1}{\pi\rho_{\mathrm{e}}(\omega)}\frac{\mathrm{d}\kappa_{\mathrm{e}}^{2}}{\mathrm{d}\omega}(-\mathrm{Im}\Sigma^{+})|\Delta\mathbf{k}|^{2}K_{0}S + 2\mathrm{i}\mathrm{Im}(\Sigma^{+})\Delta\mathbf{k}\cdot\mathbf{J}$$
$$= \frac{\mathrm{i}}{4K_{0}}M_{0}\Delta\mathbf{k}\cdot\mathbf{J} + \frac{N}{L^{d}}R_{0}|\Delta\mathbf{k}|^{2}, \qquad (5.52)$$

or

$$\Delta \mathbf{k} \cdot \mathbf{J} = \frac{-\mathrm{i}}{\pi \rho_{\mathrm{e}}(\omega)} \frac{K_0 (\mathrm{d}\kappa_{\mathrm{e}}^2/\mathrm{d}\omega)(-\mathrm{Im}\Sigma^+) |\Delta \mathbf{k}|^2 S - (N/L^d)\pi \rho_{\mathrm{e}}(\omega) R_0 |\Delta \mathbf{k}|^2}{(M_0/4K_0) - 2\mathrm{Im}(\Sigma^+)},$$
(5.53)

where the term  $-(d\kappa_0^2/d\omega)\Delta\omega \Delta \mathbf{k} \cdot \mathbf{J}$  has been dropped because it is small  $(\Delta\omega \to 0)$  compared to the other terms that are linear in  $\Delta \mathbf{k}$ . The  $R_0$  term (the third term on the right-hand side of (5.45b)) is not proportional to S, therefore it may be viewed as an additional source term when (5.53) is combined with (5.44a). However, compared with the source term on the right side of (5.44a), the  $R_0$  term is small because it is proportional to  $|\Delta \mathbf{k}|^2$ . Therefore, it can be dropped.

Because (5.53) states that  $|\mathbf{J}| \sim |\Delta \mathbf{k}| S$  (without the  $R_0$  term), its substitution into (5.44) immediately gives a diffusive pole for S, as anticipated:

$$S = \frac{2(\mathrm{d}\omega/\mathrm{d}\kappa_0^2)(\mathrm{d}\omega/\mathrm{d}\kappa_\mathrm{e}^2)\rho_\mathrm{e}(\omega)\pi}{-\mathrm{i}\Delta\omega + D(\omega)|\Delta\mathbf{k}|^2}\frac{N}{L^d},\tag{5.54}$$

with

$$D(\omega) = \frac{1}{\pi \rho_{\rm e}(\omega)} \frac{8K_0^2 [1 + d\text{Re}(\Sigma^+)/d\kappa_0^2](-\text{Im}\Sigma^+)}{M_0 - 8K_0 \text{Im}(\Sigma^+)},$$
(5.55)

where the definition  $\kappa_{\rm e}^2 = \kappa_0^2 - {\rm Re}(\Sigma^+)$  has been used. Earlier, (5.13) states that if the intensity transport behavior were diffusive in the  $\Delta \omega$ ,  $|\Delta \mathbf{k}| \rightarrow$ 0 limit, then S should have a diffusive pole. That assertion has now been substantiated. Since the derivation does not invoke the dispersion relation of the wave or whether it is on a lattice or in continuum, the diffusive behavior is thus a general property of multiply scattered wave intensity transport in the limit of long propagation time and long travel distance. Also, since  $U_{\mathbf{k},\mathbf{k}'}$ , is not specified or approximated, the fact that S has a diffusive pole is independent of the type of scattering. In fact, even if  $U_{\mathbf{k},\mathbf{k}'} = 0$  so that  $M_0 = 0$  (no vertex correction),  $D(\omega)$  would still be finite. As  $-\text{Im}\Sigma^+$  gives the mean free path in the effective medium, the effect of  $M_0$  may be viewed as providing a correction to the effective medium mean free path. This point will be made more explicit later in this chapter. The units of  $D(\omega)$  can also be checked to be correct: Since  $M_0 \sim [\text{length}]^{-d}$ ,  $K_0 \sim [\text{length}]^{2-d}$ ,  $\kappa_0^2$ ,  $\Sigma \sim [\text{length}]^{-2}$ , and  $\rho_e \sim [\text{time}][\text{length}]^{-d}$ , from (5.55)  $D(\omega)$  has the units of  $[\text{length}]^2/[\text{time}]$ , independent of the spatial dimensionality. The evaluation of  $D(\omega)$  is left to a later section, since it is necessary to first examine the vertex function and demonstrate formally the validity of the Ward identity. At this point it is instructive to reflect on the results obtained so far.

## 5.3.9 Configurational Averaging and Dissipation

In electronic systems, the diffusion constant is known to be proportional to the electrical conductivity via the Einstein relation. Since electrical conductivity is associated with dissipation, it may seem odd that although in our calculations so far there has been no consideration of dissipative mechanisms (as evidenced by the fact that the basis of our models is the wave equation, in which energy conservation is an inherent attribute), the final derived expression is proportional to the electrical conductivity. The answer to this puzzle is contained in the subtle effect of configuration averaging. By smoothing out fluctuations through configurational averaging, irreversibility is implicitly introduced. As irreversibility is implied by entropy generation and dissipation, it is not altogether surprising that the diffusion constant and the electrical conductivity may be related to quantities calculated through the route of configurational averaging. The validity of this approach is in fact the basis of the linear response theory.

# 5.4 The Vertex Function

The reducible vertex function is defined by

$$\Gamma = \left\langle \bar{\mathbf{T}}^+ \otimes \bar{\mathbf{T}}^- \right\rangle_{\rm c},\tag{5.56}$$

where the overbar indicates the quantity to be defined relative to a CPA effective medium (i.e., the embedding medium is the effective medium rather than the unperturbed uniform medium). In the solution to Problem 3.8, the  $\bar{\mathbf{T}}$  operator is shown to be decomposable into operators involving individual scatterers:

$$\bar{\mathbf{T}} = \sum_{i} \bar{\mathbf{R}}_{i},\tag{5.57}$$

where i is the site index in the lattice model and is the (structural unit) particle index in the continuum case. Furthermore, it is also shown that

$$\bar{\mathbf{R}}_{i} = \bar{\mathbf{t}}_{i} \left( \mathbf{I} + \mathbf{G}_{e} \sum_{n \neq i} \bar{\mathbf{R}}_{n} \right) = \left( \mathbf{I} + \sum_{n \neq i} \bar{\mathbf{R}}_{n} \mathbf{G}_{e} \right) \bar{\mathbf{t}}_{i}.$$
 (5.58)

#### 5.4.1 The Reducible Vertex and Diagrammatic Representations

Through iterative substitution of the right-hand side of (5.58) into the summation over  $\mathbf{\bar{R}}_n$  (with a change of index labeling from *i* to *n*), an operator equation is obtained:

$$\bar{\mathbf{T}} = \sum_{i} \bar{\mathbf{t}}_{i} + \sum_{i} \sum_{j \neq i} \bar{\mathbf{t}}_{i} \mathbf{G}_{e} \bar{\mathbf{t}}_{j} + \sum_{i} \sum_{j \neq i} \sum_{k \neq j} \bar{\mathbf{t}}_{i} \mathbf{G}_{e} \bar{\mathbf{t}}_{j} \mathbf{G}_{e} \bar{\mathbf{t}}_{k} + \cdots, \qquad (5.59)$$

where each succeeding index cannot coincide in value with the preceding index. So in the third term  $j \neq i$  and  $k \neq j$ , but k = i is allowed. In terms of individual scatterers, the structure of the vertex function is thus clear: It is the sum of configurational averages of infinite number of terms involving  $\langle \bar{\mathbf{t}}_i^+ \bar{\mathbf{t}}_j^- \rangle_c, \langle \bar{\mathbf{t}}_i^+ \mathbf{G}_e^+ \bar{\mathbf{t}}_j^+ \bar{\mathbf{t}}_k^- \rangle_c$ , etc. However, certain terms can be eliminated due to the CPA condition of  $\langle \bar{\mathbf{t}}_i^\pm \rangle_c = 0$ . For example, terms involving an odd number of  $\bar{\mathbf{t}}$ 's must vanish because at least one  $\bar{\mathbf{t}}$  would be independently averaged in the subsets of configurations where the other  $\bar{\mathbf{t}}$ 's are fixed. For terms involving an even number of  $\mathbf{t}$ 's, on the other hand, the indices must form pairs in order for them not to vanish. For example,

$$a_{123} = \left\langle \overline{\mathbf{t}}_1^+ \mathbf{G}_{\mathrm{e}}^+ \overline{\mathbf{t}}_2^+ \mathbf{G}_{\mathrm{e}}^+ \overline{\mathbf{t}}_3^+ \otimes \overline{\mathbf{t}}_3^- \mathbf{G}_{\mathrm{e}}^- \overline{\mathbf{t}}_2^- \mathbf{G}_{\mathrm{e}}^- \overline{\mathbf{t}}_1^- \right\rangle_{\mathrm{c}}$$
(5.60)

would not vanish, because  $\langle \bar{\mathbf{t}}_i^+ \bar{\mathbf{t}}_i^- \rangle_c \neq 0$ . But if  $\bar{\mathbf{t}}_1^-$  were replaced by  $\bar{\mathbf{t}}_4^-$ , the term would vanish because there is no corresponding  $\bar{\mathbf{t}}_4^+$ . However, even with



Fig. 5.4. Four diagrams corresponding to three scatterers. The symbols are explained in the text.

this rule there can still be many possibilities for the three-scatterers case. This is because the *order* in which the operators multiply matters, so for example

$$b_{123} = \left\langle \overline{\mathbf{t}}_1^+ \mathbf{G}_e^+ \overline{\mathbf{t}}_2^+ \mathbf{G}_e^+ \overline{\mathbf{t}}_3^+ \otimes \overline{\mathbf{t}}_2^- \mathbf{G}_e^- \overline{\mathbf{t}}_3^- \mathbf{G}_e^- \overline{\mathbf{t}}_1^- \right\rangle_c$$
(5.61)

is different from  $a_{123}$ , even though the same three scatterers are involved. This may be visualized by diagrams as shown in Fig. 5.4. Here a small cross denotes a scatterer. The same scatterers related to two (+ and -) Green functions are connected by a dashed "interaction" line. Together they represent  $\langle \bar{\mathbf{t}}_i^+ \bar{\mathbf{t}}_i^- \rangle_c$ . The Green functions are represented by solid lines, with the + channels on top and the - channels on the bottom. Due to configurational averaging, the input and output channels associated with each scattering vertex must have the property given by (5.9), i.e., in the k representation the total momentum of the input channels must equal to the momentum sum of the output channels. In Fig. 5.4a, which corresponds to  $a_{123}$ , the diagram is called reducible, as explained in relation to Fig. 5.3. In Fig. 5.4b, which corresponds to  $b_{123}$ , the diagram is also reducible. However, Figs. 5.4c, d are irreducible, and they correspond, respectively, to

$$c_{123} = \left\langle \overline{\mathbf{t}}_1^+ \mathbf{G}_e^+ \overline{\mathbf{t}}_2^+ \mathbf{G}_e^+ \overline{\mathbf{t}}_3^+ \otimes \overline{\mathbf{t}}_1^- \mathbf{G}_e^- \overline{\mathbf{t}}_2^- \mathbf{G}_e^- \overline{\mathbf{t}}_3^- \right\rangle_c$$

and

$$d_{123} = \left\langle \overline{\mathbf{t}}_1^+ \mathbf{G}_e^+ \overline{\mathbf{t}}_2^+ \mathbf{G}_e^+ \overline{\mathbf{t}}_1^+ \mathbf{G}_e^+ \overline{\mathbf{t}}_3^+ \otimes \overline{\mathbf{t}}_2^- \mathbf{G}_e^- \overline{\mathbf{t}}_3^- \right\rangle_c$$

## 5.4.2 CPA for $\Gamma$ : the Ladder Diagrams and the Irreducible Vertex

With the understanding of the vertex function  $\Gamma$ , let us rewrite (5.56) by using (5.57) as

$$\Gamma = \left\langle \sum_{i} \bar{\mathbf{R}}_{i}^{+} \otimes \sum_{j} \bar{\mathbf{R}}_{j}^{-} \right\rangle_{c} = \sum_{i} \sum_{j} \Gamma_{ij}, \qquad (5.62a)$$

where

$$\Gamma_{ij} = \left\langle \bar{\mathbf{t}}_{i}^{+} \left( \mathbf{I} + \mathbf{G}_{e}^{+} \sum_{n \neq i} \bar{\mathbf{R}}_{n}^{+} \right) \otimes \left( \mathbf{I} + \sum_{m \neq j} \bar{\mathbf{R}}_{m}^{-} \mathbf{G}_{e}^{-} \right) \bar{\mathbf{t}}_{j}^{-} \right\rangle_{c}$$
$$= \delta_{i,j} \left\langle \bar{\mathbf{t}}_{i}^{+} \otimes \bar{\mathbf{t}}_{i}^{-} \right\rangle_{c} + \left\langle \bar{\mathbf{t}}_{i}^{+} \mathbf{G}_{e}^{+} \left( \sum_{n \neq i} \bar{\mathbf{R}}_{n}^{+} \otimes \sum_{m \neq j} \bar{\mathbf{R}}_{m}^{-} \right) \mathbf{G}_{e}^{-} \bar{\mathbf{t}}_{j}^{-} \right\rangle_{c}, \quad (5.62b)$$

and terms with an odd number of  $\mathbf{\bar{t}}$  or  $\mathbf{\bar{R}}$  vanish for the reason already given. It should be noted here that the subscripts ij on  $\Gamma$  are the scatterer indices. They should not be confused with channel indices. The last term of (5.62b) may be approximated as

$$\left\langle \bar{\mathbf{t}}_{i}^{+} \mathbf{G}_{e}^{+} \left( \sum_{n \neq i} \bar{\mathbf{R}}_{n}^{+} \otimes \sum_{m \neq j} \bar{\mathbf{R}}_{m}^{-} \right) \mathbf{G}_{e}^{-} \bar{\mathbf{t}}_{j}^{-} \right\rangle_{c}$$

$$\cong \left\langle \bar{\mathbf{t}}_{i}^{+} \mathbf{G}_{e}^{+} \left\langle \sum_{n \neq i} \bar{\mathbf{R}}_{n}^{+} \otimes \sum_{m \neq j} \bar{\mathbf{R}}_{m}^{-} \right\rangle_{c} \mathbf{G}_{e}^{-} \bar{\mathbf{t}}_{j}^{-} \right\rangle_{c}$$

$$= \sum_{n \neq i} \sum_{m \neq j} \delta_{i,j} \left\langle \bar{\mathbf{t}}_{i}^{+} \otimes \bar{\mathbf{t}}_{i}^{-} \right\rangle_{c} : \left( \mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-} \right) : \Gamma_{mn}.$$

$$(5.63)$$

This approximation is in the same CPA spirit as discussed in the last chapter, because it reduces all calculations to one particle (or one site). As shall be seen later, the CPA approximation selects out a particular class of diagrams for the calculation of the vertex function.

Under the CPA approximation,  $\Gamma_{ij}$  satisfies the following equation:

$$\Gamma_{ij}^{(B)} = \delta_{i,j} \left\langle \bar{\mathbf{t}}_i^+ \otimes \bar{\mathbf{t}}_i^- \right\rangle_{c} + \delta_{i,j} \left\langle \bar{\mathbf{t}}_i^+ \otimes \bar{\mathbf{t}}_i^- \right\rangle_{c} : \sum_{n \neq i} \sum_{m \neq j} \left( \mathbf{G}_{e}^+ \otimes \mathbf{G}_{e}^- \right) : \Gamma_{mn}^{(B)}.$$
(5.64)

Here the superscript B is used to denote the quantities as evaluated under the CPA, which yields the classical Boltzmann diffusion constant as will be seen later. The order of the terms has been rearranged in (5.64) for simplicity in manipulation. By adding the excluded term (n = i, m = j) to both sides of (5.64), the constraints in the summations can be removed:

$$\left( \mathbf{I} + \left\langle \bar{\mathbf{t}}_{i}^{+} \mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-} \bar{\mathbf{t}}_{j}^{-} \right\rangle_{c} \right) : \mathbf{\Gamma}_{ij}^{(B)}$$

$$= \delta_{i,j} \left\langle \bar{\mathbf{t}}_{i}^{+} \otimes \bar{\mathbf{t}}_{i}^{-} \right\rangle_{c} : \left[ \mathbf{I} + \sum_{n} \sum_{m} \left( \mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-} \right) : \mathbf{\Gamma}_{nm}^{(B)} \right], \quad (5.65)$$

where **I** is short for  $\mathbf{I} \otimes \mathbf{I}$ . The term  $\langle \mathbf{\bar{t}}_i^+ \mathbf{G}_e^+ \otimes \mathbf{G}_e^- \mathbf{\bar{t}}_j^- \rangle_c$  in this equation is noted to have the input channels specified by  $\mathbf{\bar{t}}_i^+$  and  $\mathbf{\bar{t}}_j^-$ , respectively, i.e., particle

(site) *i* and *j*. The output channels are linked to  $\Gamma_{ij}^{(B)}$ , with *i* associated with the + channel and *j* the - channel. Since  $\langle \bar{\mathbf{t}}_i^+ \otimes \bar{\mathbf{t}}_j^- \rangle_c \propto \delta_{i,j}$ , however, all channels are reduced to the same particle (site).

The inverse of  $\mathbf{A} \otimes \mathbf{B}$  is  $\mathbf{A}^{-1} \otimes \mathbf{B}^{-1}$ , since  $(\mathbf{A} \otimes \mathbf{B}) : (\mathbf{A}^{-1} \otimes \mathbf{B}^{-1}) = \mathbf{I} \otimes \mathbf{I}$ . We wish to define

$$\frac{\mathbf{A}^{+} \otimes \mathbf{B}^{-}}{\mathbf{C}^{+} \otimes \mathbf{D}^{-}} \equiv (\mathbf{C}^{+})^{-1} \mathbf{A}^{+} \otimes \mathbf{B}^{-} (\mathbf{D}^{-})^{-1}.$$
 (5.66)

With this inverse notation, (5.65) may be written succinctly in terms of the irreducible vertex function  $\mathbf{U}_{i}^{(\mathrm{B})}$  as

$$\mathbf{\Gamma}_{ij}^{(\mathrm{B})} = \delta_{i,j} \mathbf{U}_{i}^{(\mathrm{B})} : \left[ \mathbf{I} + \sum_{n} \sum_{m} \left( \mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-} \right) : \mathbf{\Gamma}_{nm}^{(\mathrm{B})} \right],$$
(5.67)

where

$$\mathbf{U}_{i}^{(\mathrm{B})} = \frac{\langle \mathbf{\bar{t}}_{i}^{+} \otimes \mathbf{\bar{t}}_{i}^{-} \rangle_{\mathrm{c}}}{\mathbf{I} + \langle \mathbf{\bar{t}}_{i}^{+} \mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-} \mathbf{\bar{t}}_{i}^{-} \rangle_{\mathrm{c}}} = \mathbf{u}_{i}^{+} \otimes \mathbf{u}_{i}^{-}, \qquad (5.68)$$

with  $\mathbf{u}_i^+,\,\mathbf{u}_i^-$  associated with the + and - channels of particle (site) i respectively.

Equation (5.67) can be expressed as

$$\Gamma_{ij}^{(B)} = \delta_{i,j} \mathbf{U}_i^{(B)} + \delta_{i,j} \sum_n \sum_m \left( \mathbf{u}_i^+ \mathbf{G}_e^+ \otimes \mathbf{G}_e^- \mathbf{u}_i^- \right) : \Gamma_{nm}^{(B)}.$$
(5.69)

Through iterative substitutions, (5.69) becomes

$$\boldsymbol{\Gamma}_{ij}^{(\mathrm{B})} = \delta_{i,j} \left\{ \mathbf{u}_{i}^{+} \otimes \mathbf{u}_{i}^{-} + \mathbf{u}_{i}^{+} \left( \sum_{n} \mathbf{G}_{\mathrm{e}}^{+} \mathbf{u}_{n}^{+} \otimes \mathbf{u}_{n}^{-} \mathbf{G}_{\mathrm{e}}^{-} \right) \mathbf{u}_{i}^{-} + \mathbf{u}_{i}^{+} \left[ \sum_{n} \mathbf{G}_{\mathrm{e}}^{+} \mathbf{u}_{n}^{+} \left( \sum_{m} \mathbf{G}_{\mathrm{e}}^{+} \mathbf{u}_{m}^{+} \otimes \mathbf{u}_{m}^{-} \mathbf{G}_{\mathrm{e}}^{-} \right) \mathbf{u}_{n}^{-} \mathbf{G}_{\mathrm{e}}^{-} \right] \mathbf{u}_{i}^{-} + \cdots \right\},$$

$$(5.70)$$

where the  $\delta$  function in (5.69) reduces every double summation to a single summation at each stage of the iteration. The structure of the CPA approximation is thus clear: It consists of an infinite number of elementary diagrams like that shown in Fig. 5.5. Since every one of these diagrams looks like a



Fig. 5.5. Diagrammatic representation of  $\Gamma^{(B)}$ . The symbols are explained in the text. Because the diagrams look like ladders, they are called the ladder diagrams.

ladder lying on its side, they are called the "ladder diagrams." It should be noted that the diagrams here differ from those in Fig. 5.4 because the vertex for each scattering is different. Here each  $\mathbf{U}_i = \mathbf{u}_i^+ \otimes \mathbf{u}_i^-$  is represented by two open circles linked by a dashed line. It should be distinguished from the crosses shown in Fig. 5.4. Also, whereas in Fig. 5.4 one has to follow the rule that a succeeding particle (site) must differ from the preceding one, there is no such restriction here. The removal of this restriction is responsible for the denominator in the definition of  $\mathbf{U}_i$ . Without the restriction on the summation index means that the summation over i, j in (5.69) results in

$$\boldsymbol{\Gamma}^{(\mathrm{B})} = \mathbf{U}^{(\mathrm{B})} + \mathbf{U}^{(\mathrm{B})} : \left(\mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-}\right) : \boldsymbol{\Gamma}^{(\mathrm{B})}.$$
(5.71)

Comparison with (5.18) and (5.20) shows  $\mathbf{U}^{(\mathrm{B})}$  to be the irreducible CPA vertex function. As the effect of  $\mathbf{U}$  on the wave diffusion constant is embodied in the term  $M_0$ , (5.51), for  $\mathbf{U} = \mathbf{U}^{(\mathrm{B})}$  the  $M_0^{(\mathrm{B})}$  thus represents the total contribution of the ladder diagrams to the diffusion constant. The reducible vertex  $\boldsymbol{\Gamma}$ , on the other hand, is useful in the evaluation of scattered wave intensity. From (3.11) it is seen that in the presence of scattering disorder,

$$\phi(\omega, \mathbf{r}) = \phi_0(\omega, \mathbf{r}) + \int G_{\rm e}^+(\omega, \mathbf{r} - \mathbf{r}_1) \bar{T}^+(\omega, \mathbf{r}_1, \mathbf{r}_2) \phi_0(\omega, \mathbf{r}_2) \mathrm{d}\mathbf{r}_1 \mathrm{d}\mathbf{r}_2.$$

If  $\phi_0$  is the incident wave and the scattered wave amplitude  $\phi - \phi_0$  is denoted by  $\Psi_{\rm S}$ , then the *averaged scattering intensity* may be expressed as

$$\left\langle |\Psi_{\mathrm{S}}(\omega,\mathbf{r})|^2 \right\rangle_{\mathrm{c}} = \iiint G_{\mathrm{e}}^+(\mathbf{r}-\mathbf{r}_1) G_{\mathrm{e}}^-(\mathbf{r}-\mathbf{r}_2) \Gamma(\mathbf{r}_1,\mathbf{r}_2,\mathbf{r}_3,\mathbf{r}_4) \\ \times \phi_0(\mathbf{r}_3) \phi_0^*(\mathbf{r}_4) \,\mathrm{d}\mathbf{r}_1 \mathrm{d}\mathbf{r}_2 \mathrm{d}\mathbf{r}_3 \mathrm{d}\mathbf{r}_4,$$

where the  $\omega$  labeling is suppressed on the right-hand side.

#### 5.4.3 Irreducible Vertex: the Anderson Model

Let us now proceed to take a more detailed look at the irreducible vertex function  $\mathbf{U}^{(\mathrm{B})}$ . In the Anderson model, the vertex function is associated with a single site **l**, and  $\mathbf{U}_{\mathbf{l}}^{(\mathrm{B})}/a^d$  may be written as

$$\frac{1}{a^d} \mathbf{U}_{\mathbf{l}}^{(\mathrm{B})} = \frac{1}{a^d} U^{(\mathrm{B})}(\Delta \omega | \omega) \left( |\mathbf{l}\rangle_+ \langle \mathbf{l}|_+ \otimes |\mathbf{l}\rangle_- \langle \mathbf{l}|_- \right).$$
(5.72)

At  $\Delta \omega = 0$ , (5.68) gives

$$\frac{1}{a^d}U^{(B)} = \frac{C_1}{1+C_2},\tag{5.73a}$$

with

$$C_{1} = \frac{1}{a^{d}N} \sum_{\mathbf{l}} |\bar{t}_{1}^{+}|^{2}$$
$$= \frac{a^{d}}{2\sigma_{w}} \int_{-\sigma_{w}}^{\sigma_{w}} d\sigma \left| \frac{\sigma - \Sigma^{+}}{1 - (\sigma - \Sigma^{+})a^{d}G_{0}^{+}(\kappa_{0}^{2} - \Sigma^{+}, \mathbf{l} = \mathbf{l}')} \right|^{2}, \quad (5.73b)$$

and

$$C_{2} = \frac{1}{2\sigma_{w}} \int_{-\sigma_{w}}^{\sigma_{w}} \mathrm{d}\sigma \left| \frac{(\sigma - \Sigma^{+})a^{d}G_{0}^{+}(\kappa_{0}^{2} - \Sigma^{+}, \mathbf{l} = \mathbf{l}')}{1 - (\sigma - \Sigma^{+})a^{d}G_{0}^{+}(\kappa_{0}^{2} - \Sigma^{+}, \mathbf{l} = \mathbf{l}')} \right|^{2}.$$
 (5.73c)

Here  $\Sigma^+$  is obtained from the solution of the CPA equation given in Chap. 4. In real space, therefore,

$$\mathbf{U}^{(\mathrm{B})} = \sum_{\mathbf{l}} \mathbf{U}_{\mathbf{l}}^{(\mathrm{B})} = U^{(\mathrm{B})} \sum_{\mathbf{l}} |\mathbf{l}\rangle_{+} \langle \mathbf{l}|_{+} \otimes |\mathbf{l}\rangle_{-} \langle \mathbf{l}|_{-}.$$
 (5.74)

In the momentum representation, the input and output channels of U have the same structure as that of  $\Gamma$  because (5.18) shows

$$\boldsymbol{\Gamma}_{ij} = \mathbf{U}_{ij} + \sum_{n} \sum_{m} \left( \mathbf{u}_i^+ \mathbf{G}_e^+ \otimes \mathbf{G}_e^- \mathbf{u}_j^- \right) : \boldsymbol{\Gamma}_{nm},$$
(5.75)

where i, j, n, m are the particle (site) indices. Since  $\mathbf{G}_{\mathrm{e}}^{\pm}$  is diagonal in the **k** representation, the second term of (5.75) shows that the output momentum channels of  $\mathbf{U}_{ij}$  must coincide with the input momentum channels of  $\Gamma_{nm}$ . That means the same momentum-channel structure as that specified by (5.9). With this in mind,  $U_{\mathbf{k},\mathbf{k}'}^{(\mathrm{B})}$  for the Anderson model may be written in the momentum-channel representation by using (5.74) as the starting point. That is, by inserting identity operators, we have

$$\mathbf{U}^{(\mathrm{B})} = U^{(\mathrm{B})} \sum_{\mathbf{l}} |\mathbf{l}\rangle_{+} \langle \mathbf{l}|_{+} \otimes |\mathbf{l}\rangle_{-} \langle \mathbf{1}|_{-}$$

$$= \frac{U^{(\mathrm{B})}}{N^{2}} \sum_{\mathbf{l}} \sum_{\mathbf{k}_{+}, \mathbf{k}_{+}'} \sum_{\mathbf{k}_{-}, \mathbf{k}_{-}'} |\mathbf{k}_{+}\rangle \langle \mathbf{k}_{+}|\mathbf{l}\rangle \langle \mathbf{l}|\mathbf{k}_{+}'\rangle \langle \mathbf{k}_{+}'| \otimes |\mathbf{k}_{-}'\rangle \langle \mathbf{k}_{-}'|\mathbf{l}\rangle \langle \mathbf{l}|\mathbf{k}_{-}\rangle \langle \mathbf{k}_{-}|$$

$$= \frac{(2\pi)^{d}}{a^{d}} \frac{U^{(\mathrm{B})}}{N^{2}} \sum_{\mathbf{k}_{+}, \mathbf{k}_{+}'} \sum_{\mathbf{k}_{-}, \mathbf{k}_{-}'} \delta(\mathbf{k}_{+}' - \mathbf{k}_{-}' - \mathbf{k}_{+} + \mathbf{k}_{-}) |\mathbf{k}_{+}\rangle \langle \mathbf{k}_{+}'| \otimes |\mathbf{k}_{-}'\rangle \langle \mathbf{k}_{-}|$$

$$(5.76)$$

Equation (5.76) shows that in the momentum space, the irreducible vertex is independent of the scattering angle, i.e.,

$$U_{\mathbf{k},\mathbf{k}'}^{(B)} = U^{(B)}.$$
 (5.77)

## 5.4.4 Irreducible Vertex: Classical Waves

For scalar waves in continuum, the situation is more complicated because in real space  $\mathbf{u}_i^+$  and  $\mathbf{u}_i^-$  are full matrices, since each scatterer may be regarded as consisting of infinitely many sites, so  $\overline{\mathbf{t}}_i$  is no longer simple in real space as for the Anderson model. However, the calculation of  $\overline{\mathbf{t}}_i$  is simpler in the momentum representation because it may be accomplished through the solution of boundary-value problems, as seen from Chap. 3. Therefore, let us approach the evaluation of  $\mathbf{U}^{(B)}$  directly in the momentum-channel representation. In order to do that, one should first fix the input and output channels as  $\mathbf{k}_{\pm}$  and  $\mathbf{k}'_{\pm}$ , respectively, with  $|\mathbf{k}_{\pm}| = |\mathbf{k}'_{\pm}| = \kappa_{e} \pm \text{ term on the order}$ of  $|\Delta \mathbf{k}|/2$ . Since the output channels of the term  $\langle \mathbf{\bar{t}}_i^+ \mathbf{G}_e^+ \otimes \mathbf{G}_e^- \mathbf{\bar{t}}_i^- \rangle_c$  in the denominator of  $\mathbf{U}^{(B)}$  [see (5.68)] are linked directly to the input channels of  $\Gamma_{ij}^{(B)}$  [see (5.65)], they are thus fixed at  $\mathbf{k}_{\pm}$ . Moreover, the output channels of  $\mathbf{G}_{e}^{+}$  and  $\mathbf{G}_{e}^{-}$  in real space are *restricted to the same particle* due to the CPA condition. That means in momentum space they have to be convolved with the particle function H as described in the last chapter. The input channels of  $\langle \bar{\mathbf{t}}_i^+ \mathbf{G}_e^+ \otimes \mathbf{G}_e^- \bar{\mathbf{t}}_i^- \rangle_c$  are labeled as  $(\mathbf{k}_1)_{\pm}$ . For the  $\langle \bar{\mathbf{t}}_i^+ \otimes \bar{\mathbf{t}}_i^- \rangle_c$  term in the numerator, its output channels are fixed at  $\mathbf{k'}_{\pm}$ , whereas its input channels are labeled as  $(\mathbf{k}_2)_{\pm}$ .

The numerator of  $\mathbf{U}^{(\mathrm{B})}N/L^d$  with the specified output channels is given by

$$\frac{N}{L^{d}} \left\langle \mathbf{\bar{t}}_{i}^{+} \otimes \mathbf{\bar{t}}_{i}^{-} \right\rangle_{c} = \frac{1}{N^{2}} \sum_{\mathbf{k}_{2+}} \sum_{\mathbf{k}_{2-}} \delta_{\mathbf{k}_{2+}-\mathbf{k}_{2-}, \Delta \mathbf{k}} \\
\times \left( \sum_{m} n_{m} \left\langle \mathbf{k}_{2+} \right| \mathbf{\bar{t}}_{m}^{+} \left| \mathbf{k}_{+}^{\prime} \right\rangle \left\langle \mathbf{k}_{-}^{\prime} \right| \mathbf{\bar{t}}_{m}^{-} \left| \mathbf{k}_{2-} \right\rangle \right) \\
\times \left( \left| \mathbf{k}_{2+} \right\rangle \left\langle \mathbf{k}_{+}^{\prime} \right| \otimes \left| \mathbf{k}_{-}^{\prime} \right\rangle \left\langle \mathbf{k}_{2-} \right| \right).$$
(5.78)

Here the configuration averaging with the factor  $N/L^d$  yields the number density  $n_m$  for the *m*th scatterer species (obtained by fixing at a particular site and summing over all particles, with a probability 1/N that the center of a given particle would fall on that particular site), as well as a Kronecker delta relating  $\mathbf{k}_{2+}$  and  $\mathbf{k}_{2-}$ . With fixed  $\mathbf{k'}_{\pm}$ , the + channel operator may be regarded as a column vector and the – channel operator a row vector. The denominator of  $\mathbf{U}^{(\mathrm{B})}$  may be expressed as

$$\mathbf{I} + \left\langle \mathbf{\bar{t}}_{i}^{+} \mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-} \mathbf{\bar{t}}_{i}^{-} \right\rangle_{c} \\
= \frac{1}{N^{2}} \sum_{\mathbf{k}_{1+}} \sum_{\mathbf{k}_{1-}} \delta_{\mathbf{k}_{1+}-\mathbf{k}_{1-}, \Delta \mathbf{k}} \left\{ N \delta_{\mathbf{k}, \mathbf{k}_{1}} + \frac{1}{N} \sum_{\mathbf{k}''} \left( \sum_{m} n_{m} \left\langle \mathbf{k}_{1+} \right| \mathbf{\bar{t}}_{m}^{+} |\mathbf{k}''_{+}\rangle \left\langle \mathbf{k}''_{-} \right| \mathbf{\bar{t}}_{m}^{-} |\mathbf{k}_{1-}\rangle \right) \\
\times \frac{L^{d}}{N} G_{e}^{+}(\omega^{+}, \mathbf{k}''_{+}) G_{e}^{-}(\omega^{-}, \mathbf{k}''_{-}) H(\mathbf{k}'' - \mathbf{k}) \right\} \\
\times (|\mathbf{k}_{1+}\rangle \left\langle \mathbf{k}_{+}| \otimes |\mathbf{k}_{-}\rangle \left\langle \mathbf{k}_{1-}| \right\rangle. \tag{5.79}$$

Here one factor of  $L^d/N$  (in front of  $G_{\rm e}^+$ ) has been generated by configurational averaging.

To evaluate  $U_{\mathbf{k},\mathbf{k}'}^{(\mathrm{B})}$  it is simpler to invert the numerator and contract it with the denominator. By noting that the inverse of a column vector is a row vector and vice versa, and that  $\langle \mathbf{k}_{2+} | \mathbf{k}_{1+} \rangle = N \delta_{\mathbf{k}_{2+},\mathbf{k}_{1+}}$ , for example, the resulting expression is given by

$$\frac{L^{d}}{N} \left( \mathbf{U}_{\mathbf{k},\mathbf{k}'}^{(\mathrm{B})} \right)^{-1} \left( |\mathbf{k}'_{+}\rangle \left\langle \mathbf{k}_{+} | \otimes |\mathbf{k}_{-}\rangle \left\langle \mathbf{k}'_{-} | \right), \right.$$
(5.80a)

where

$$\frac{L^{d}}{N} \left( \mathbf{U}_{\mathbf{k},\mathbf{k}'}^{(\mathrm{B})} \right)^{-1} = \left( \sum_{m} n_{m} \langle \mathbf{k}_{+} | \, \bar{\mathbf{t}}_{m}^{+} | \mathbf{k}'_{+} \rangle \langle \mathbf{k}'_{-} | \, \bar{\mathbf{t}}_{m}^{-} | \mathbf{k}_{-} \rangle \right)^{-1} \\
+ \frac{1}{N} \sum_{\mathbf{k}''} \frac{1}{N} \sum_{\mathbf{k}_{1}} \frac{\sum_{m} n_{m} \langle \mathbf{k}_{1+} | \, \bar{\mathbf{t}}_{m}^{+} | \mathbf{k}''_{+} \rangle \langle \mathbf{k}''_{-} | \, \bar{\mathbf{t}}_{m}^{-} | \mathbf{k}_{1-} \rangle \\
\times \frac{N}{L^{d}} \left[ \frac{L^{2d}}{N^{2}} G_{\mathrm{e}}^{+} (\omega^{+}, \mathbf{k}''_{+}) G_{\mathrm{e}}^{-} (\omega^{-}, \mathbf{k}''_{-}) H(\mathbf{k}'' - \mathbf{k}) \right]. \quad (5.80b)$$

It should be noted that one of the Kronecker deltas (e.g.,  $\delta_{\mathbf{k}_1+-\mathbf{k}_1-,\Delta\mathbf{k}}$ ) is redundant because the inner product between  $|\mathbf{k}_{2+}\rangle$  and  $|\mathbf{k}_{1+}\rangle$  makes  $\mathbf{k}_2 = \mathbf{k}_1$ , so the two delta functions give the identical condition. One summation thus produces an extra factor of N, which cancels one  $N^{-1}$ . In the limit of  $\Delta\omega$ ,  $|\Delta\mathbf{k}| = 0$ , the first term is

$$\left(\sum_{m} n_{m} \left| \langle \mathbf{k} | \, \bar{\mathbf{t}}_{m}^{+} \, | \mathbf{k}' \rangle \right|^{2} \right)^{-1}.$$

The second term may be estimated as follows. The summation over  $\mathbf{k}_1$  consists of an integral over the magnitude of  $\mathbf{k}_1$  and an angular integral over its directions. Since it is assumed that  $|\mathbf{k}'| = \kappa_{\rm e}$ , and  $|\mathbf{k}''|$  is also equal to  $\kappa_{\rm e}$  as will be seen later, the magnitude dependence on  $\mathbf{k}_1$  in the numerator and denominator of the integrand should be nearly identical. That means  $(1/N)\Sigma_{\mathbf{k}_1}$  is just an angular average, given by

$$\frac{1}{S_d} \int \mathrm{d}\Omega_{\hat{\mathbf{k}}_1} \frac{\sum_m n_m \left| \langle \mathbf{k}_1 | \, \overline{\mathbf{t}}_m^+ \, | \mathbf{k}'' \rangle \right|^2}{\sum_m n_m \left| \langle \mathbf{k}_1 | \, \overline{\mathbf{t}}_m^+ \, | \mathbf{k}' \rangle \right|^2} = \frac{1}{S_d} F(\hat{\mathbf{k}}'' \cdot \hat{\mathbf{k}}'), \tag{5.81}$$

where  $\hat{\mathbf{k}}''$  and  $\hat{\mathbf{k}}'$  are, respectively, the unit vectors in the directions of  $\mathbf{k}''$  and  $\mathbf{k}'$ ,  $d\Omega_{\hat{\mathbf{k}}_1}$  denotes angular integration over the direction of  $\mathbf{k}_1$  (in 1D this would be just a summation over the two directions), and  $S_d = 4\pi$ ,  $2\pi$ , 2 in 3D, 2D, 1D, respectively. In the limit of  $|\Delta \mathbf{k}|$ ,  $\Delta \omega \to 0$ , the product  $L^{2d}G_{\rm e}^+G_{\rm e}^-/N^2$  may be converted, through the use of (5.25), (5.26), and (5.28), to the more useful form of

$$\frac{-2\pi \mathrm{i}\delta\left[\kappa_{\mathrm{e}}^2-(k'')^2\right]}{2\mathrm{i}\mathrm{Im}\Sigma^+}.$$

Since  $H(\mathbf{k}) = 1$  for  $|\Delta \mathbf{k}|$  less than  $\sim 1/R$ , where R is the radius of the particle, the summation over  $\mathbf{k}''$  may be performed with the assumption that  $|\mathbf{k}'' - \mathbf{k}|$  is less than 1/R. Together with the factors  $(1/N)N/L^d$ , the result is

$$\frac{1}{-\mathrm{Im}\Sigma^+} \frac{1}{S_d} \hat{F} \frac{\pi \kappa_{\mathrm{e}}^{d-2}}{2(2\pi)^d}$$

where both  $\Sigma^+$  and  $|\mathbf{k}''|$  in expression (5.81) are evaluated at  $|\mathbf{k}''| = \kappa_{\rm e}$ , and

$$\hat{F} = \int \mathrm{d}\Omega_{\hat{\mathbf{k}}''} F(\hat{\mathbf{k}}'' \cdot \hat{\mathbf{k}}').$$
(5.82)

Now (5.80a) can be inverted to give

$$\frac{N}{L^{d}}U_{\mathbf{k},\mathbf{k}'}^{(\mathrm{B})}\left(\left|\mathbf{k}_{+}\right\rangle\left\langle\mathbf{k}'_{+}\right|\otimes\left|\mathbf{k}'_{-}\right\rangle\left\langle\mathbf{k}_{-}\right|\right),$$

with

$$\frac{N}{L^{d}}U_{\mathbf{k},\mathbf{k}'}^{(\mathrm{B})} = \frac{\sum_{m} n_{m} \left|\langle \mathbf{k} | \,\bar{\mathbf{t}}_{m}^{+} \left| \mathbf{k}' \rangle \right|^{2}}{1 + \frac{\pi}{2} \frac{\kappa_{\mathrm{e}}^{d-2}}{(2\pi)^{d}} \frac{1}{S_{d}} \hat{F} \frac{\sum_{m} n_{m} \left|\langle \mathbf{k} | \,\bar{\mathbf{t}}_{m}^{+} \left| \mathbf{k}' \rangle \right|^{2}}{-\mathrm{Im}\Sigma^{+}}.$$
(5.83)

In the form of (5.83) the vertex function has similarities to (5.73). From the optical theorem,  $-\text{Im}\Sigma^+ \cong -\text{Im}\Sigma_m n_m \langle \mathbf{k} | \mathbf{\bar{t}}_m^+ | \mathbf{k} \rangle = \kappa_e \Sigma_m n_m \bar{O}_m$ , where

$$\bar{O}_m = \int \mathrm{d}\Omega_{\hat{\mathbf{k}}'} \left| \langle \mathbf{k} | \, \bar{\mathbf{t}}_m^+ \, | \mathbf{k}' \rangle \right|^2 \cdot \frac{\kappa_{\mathrm{e}}^{d-3} \pi}{2(2\pi)^d} \tag{5.84}$$

[see (P3.13), (P3.14)]. It follows that (5.83) may be re-expressed as

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$$\frac{N}{L^d}U_{\mathbf{k},\mathbf{k}'}^{(\mathrm{B})} \cong \frac{\sum_m n_m \left|\langle \mathbf{k} | \, \overline{\mathbf{t}}_m^+ \, | \mathbf{k}' \rangle\right|^2}{1 + Q(\mathbf{k},\mathbf{k}')} = \sum_m n_m \hat{O}_m(\mathbf{k},\mathbf{k}') \frac{2(2\pi)^d}{\pi \kappa_{\mathrm{e}}^{d-3}},\tag{5.85}$$

where  $\hat{O}_m(\mathbf{k}, \mathbf{k}')$  is a renormalized scattering cross-section, given by

$$\hat{O}_m(\mathbf{k}, \mathbf{k}') = \frac{\left|\langle \mathbf{k} | \, \bar{\mathbf{t}}_m^+ \, | \mathbf{k}' \rangle \right|^2}{1 + Q(\mathbf{k}, \mathbf{k}')} \frac{\kappa_{\rm e}^{d-3} \pi}{2(2\pi)^d},\tag{5.86a}$$

and Q is a dimensionless number given by

$$Q(\mathbf{k}, \mathbf{k}') = \frac{1}{S_d} \hat{F} \frac{\sum_m n_m \left| \langle \mathbf{k} | \, \bar{\mathbf{t}}_m^+ \, | \mathbf{k}' \rangle \right|^2}{\sum_m n_m \int \mathrm{d}\Omega_{\hat{\mathbf{k}}'} \left| \langle \mathbf{k} | \, \bar{\mathbf{t}}_m^+ \, | \mathbf{k}' \rangle \right|^2}.$$
 (5.86b)

In the limit of long wavelength where the scattering is *isotropic* and Rayleighlike, then  $\hat{F}/S_d \cong S_d$ , and the denominator of Q is  $\sim S_d \sum_m n_m |\langle \mathbf{k}' | \bar{\mathbf{t}}_m^+ | \mathbf{k} \rangle|^2$ , so that

$$Q \cong 1. \tag{5.87}$$

The meaning of the denominator in the vertex function may be traced to the removal of the restriction on *successive* scatterings from the same particle, which is a crucial step in the homogenization of the medium in the CPA sense. Therefore the denominator arises purely from the geometric effect that two different particles cannot overlap. In the generation of ladder diagrams, since it is implicitly assumed that the homogeneous medium outside the scattering particle already contains the same particle (successive scatterings from the same particle are allowed), the presence of the denominator is necessary to correct that.

# 5.5 The Ward Identity

The relationship between the irreducible vertex function and the self energy, known as the Ward identity (Ward 1950; Mahan 1981), is important to the diffusive solution of the Bethe-Salpeter equation as demonstrated in the case of the simplified model. To show its validity, it has to be stressed that the Ward identity holds only when both the irreducible vertex function and the self-energy are calculated to the *same* level of approximation. In this section the Ward identity is derived within the CPA framework using the operators. However, the line of logic follows closely that shown in (5.32)–(5.38) for the simplified case, appropriate for the Anderson model. Here the operator formalism is designed for the classical waves case.

It is convenient to define  $\tilde{\Sigma} = L^d \Sigma / N$ , and note that the CPA condition on  $\tilde{\Sigma}$  is given by either

$$\left\langle \mathbf{\bar{t}}_{i} \right\rangle_{c} = \left\langle \left( \mathbf{v}_{i} - \tilde{\Sigma} \right) \left[ \mathbf{I} - \tilde{\mathbf{G}}_{e} (\mathbf{v}_{i} - \tilde{\Sigma}) \right]^{-1} \right\rangle_{c} = 0,$$

or equivalently

$$\langle \mathbf{\bar{t}}_i \rangle_{\rm c} = \left\langle \left[ \mathbf{I} - (\mathbf{v}_i - \tilde{\Sigma}) \mathbf{\tilde{G}}_{\rm e} \right]^{-1} (\mathbf{v}_i - \tilde{\Sigma}) \right\rangle_{\rm c} = 0.$$

Here  $\tilde{\mathbf{G}}_{e}$  mean  $\mathbf{H}\mathbf{G}_{e}\mathbf{H}$ , where  $\mathbf{H}$  is the particle function as defined in Chap. 4 to restrict the action of  $\mathbf{G}_{e}$  to sites inside the scattering particle.  $\tilde{\mathbf{G}}_{e}$  is noted not to have an inverse in the general sense, because its matrix elements are zero outside the particle. However,  $(\tilde{\mathbf{G}}_{e})^{-1}$  may be defined on the sites inside the particle. In what follows this is taken to be the meaning of  $\tilde{\mathbf{G}}_{e}$ 's inverse. Also,  $\tilde{\mathbf{G}}_{e}^{\pm} = \mathbf{H}\mathbf{G}_{e}^{\pm}\mathbf{H}$ . For convenience, let us define

$$\mathbf{L}^{\pm} = \left[ (\tilde{\mathbf{G}}_{\mathrm{e}}^{\pm})^{-1} - (\mathbf{v}_{i}^{\pm} - \tilde{\Sigma}^{\pm}) \right]^{-1}.$$
 (5.88)

In this notation the CPA condition may be cast as

$$\left\langle \mathbf{L}^{\pm}\right\rangle_{\mathrm{c}} = \tilde{\mathbf{G}}_{\mathrm{e}}^{\pm},$$
 (5.89)

which is shown in the solution to Problem 5.3. In (5.88) the meaning of  $\mathbf{v}_i^{\pm}$  is that the values of  $\sigma_i^{\pm}$  are calculated at frequencies  $\omega^{\pm} = \omega \pm \Delta \omega/2$ . For the quantum case,  $\sigma_i$  is independent of the frequency, so  $\mathbf{v}_i^+ = \mathbf{v}_i^-$  always. However, this is not so for classical waves, where  $\sigma_i = (1 - \varepsilon_i) \omega^2 / v_0^2$  [see (2.15)]. While the validity of the Ward identity is not affected, it will be shown that the difference between the quantum case and the classical wave case has important consequences because of the extra term proportional to  $\Delta \omega$ , which can result in the classical scalar wave case.

The CPA irreducible vertex  $\mathbf{U}^{(B)}$  is given by (compare with (5.32))

$$\mathbf{U}^{(\mathrm{B})} = \frac{\left\langle \mathbf{\bar{t}}_{i}^{+} \otimes \mathbf{\bar{t}}_{i}^{-} \right\rangle_{\mathrm{c}}}{\mathbf{I} + \left\langle \mathbf{\bar{t}}_{i}^{+} \mathbf{\tilde{G}}_{\mathrm{e}}^{+} \otimes \mathbf{\tilde{G}}_{\mathrm{e}}^{-} \mathbf{\bar{t}}_{i}^{-} \right\rangle_{\mathrm{c}}}.$$
(5.90)

This is essentially (5.68) except  $\mathbf{G}_{e}^{+}$  is replaced with  $\tilde{\mathbf{G}}_{e}^{+}$  to stress that the realspace input and output channels are within the same particle. The numerator may be written as

$$\begin{split} \left\langle \mathbf{\tilde{t}}_{i}^{+} \otimes \mathbf{\tilde{t}}_{i}^{-} \right\rangle_{c} \\ &= \left\langle \left( \mathbf{v}_{i}^{+} - \tilde{\Sigma}^{+} \right) \left[ \mathbf{I} - \mathbf{\tilde{G}}_{e}^{+} (\mathbf{v}_{i}^{+} - \tilde{\Sigma}^{+}) \right]^{-1} \otimes \left[ \mathbf{I} - \left( \mathbf{v}_{i}^{-} - \tilde{\Sigma}^{-} \right) \mathbf{\tilde{G}}_{e}^{-} \right]^{-1} \left( \mathbf{v}_{i}^{-} - \tilde{\Sigma}^{-} \right) \right\rangle_{c} \\ &= \left\langle \left( \mathbf{v}_{i}^{+} - \tilde{\Sigma}^{+} \right) \left[ \left( \mathbf{\tilde{G}}_{e}^{+} \right)^{-1} - \left( \mathbf{v}_{i}^{+} - \tilde{\Sigma}^{+} \right) \right]^{-1} \left( \mathbf{\tilde{G}}_{e}^{+} \right)^{-1} \\ & \otimes \left( \mathbf{\tilde{G}}_{e}^{-} \right)^{-1} \left[ \left( \mathbf{\tilde{G}}_{e}^{-} \right)^{-1} - \left( \mathbf{v}_{i}^{-} - \tilde{\Sigma}^{-} \right) \right]^{-1} \left( \mathbf{v}_{i}^{-} - \tilde{\Sigma}^{-} \right) \right\rangle_{c} \\ &= \left\langle \left[ -\mathbf{I} + \left( \mathbf{\tilde{G}}_{e}^{+} \right)^{-1} \mathbf{L}^{+} \right] \left( \mathbf{\tilde{G}}_{e}^{+} \right)^{-1} \otimes \left( \mathbf{\tilde{G}}_{e}^{-} \right)^{-1} \left[ -\mathbf{I} + \mathbf{L}^{-} \left( \mathbf{\tilde{G}}_{e}^{-} \right)^{-1} \right] \right\rangle_{c} \\ &= \left[ - \left( \mathbf{\tilde{G}}_{e}^{+} \right)^{-1} \otimes \left( \mathbf{\tilde{G}}_{e}^{-} \right)^{-1} \right] + \left( \mathbf{\tilde{G}}_{e}^{+} \right)^{-1} \left\langle \mathbf{L}^{+} \left( \mathbf{\tilde{G}}_{e}^{+} \right)^{-1} \otimes \left( \mathbf{\tilde{G}}_{e}^{-} \right)^{-1} \mathbf{L}^{-} \right\rangle_{c} \left( \mathbf{\tilde{G}}_{e}^{-} \right)^{-1} . \end{split}$$
(5.91)

The last line is obtained from the previous line by the application of (5.89) twice, and by noting that  $(\tilde{\mathbf{G}}_{e}^{\pm})^{-1}$  can be taken out of the configurational averaging brackets since it is already an averaged quantity. Now the term  $\left\langle \bar{\mathbf{t}}_{i}^{+}\tilde{\mathbf{G}}_{e}^{+}\otimes \tilde{\mathbf{G}}_{e}^{-}\bar{\mathbf{t}}_{i}^{-}\right\rangle_{c}$  in (5.90) differs in treatment from  $\left\langle \bar{\mathbf{t}}_{i}^{+}\otimes \bar{\mathbf{t}}_{i}^{-}\right\rangle_{c}$  only slightly. From the last line of (5.91) it may be readily deduced that

$$\mathbf{I} + \left\langle \bar{\mathbf{t}}_{i}^{+} \tilde{\mathbf{G}}_{e}^{+} \otimes \tilde{\mathbf{G}}_{e}^{-} \bar{\mathbf{t}}_{i}^{-} \right\rangle_{c} = (\tilde{\mathbf{G}}_{e}^{+})^{-1} \left\langle \mathbf{L}^{+} \otimes \mathbf{L}^{-} \right\rangle_{c} (\tilde{\mathbf{G}}_{e}^{-})^{-1}.$$
(5.92a)

The inverse of (5.92a) is

$$\left[ (\tilde{\mathbf{G}}_{e}^{+})^{-1} \left\langle \mathbf{L}^{+} \otimes \mathbf{L}^{-} \right\rangle_{c} (\tilde{\mathbf{G}}_{e}^{-})^{-1} \right]^{-1} = (\mathcal{L}^{+})^{-1} \tilde{\mathbf{G}}_{e}^{+} \otimes \tilde{\mathbf{G}}_{e}^{-} (\mathcal{L}^{-})^{-1}, \quad (5.92b)$$

where  $\mathcal{L}^{\pm}$  is introduced to mean

$$\left\langle \mathbf{L}^{+} \otimes \mathbf{L}^{-} \right\rangle_{c} = \mathcal{L}^{+} \otimes \mathcal{L}^{-},$$
 (5.93a)

$$\left\langle \mathbf{L}^{+} \otimes \mathbf{L}^{-} \right\rangle_{c}^{-1} = (\mathcal{L}^{+})^{-1} \otimes (\mathcal{L}^{-})^{-1}.$$
 (5.93b)

The operations of taking the average and then the inverse are noted not to commute; i.e., their order cannot be interchanged. Therefore  $\langle \mathbf{L}^+ \otimes \mathbf{L}^- \rangle_c^{-1} \neq \langle (\mathbf{L}^+)^{-1} \otimes (\mathbf{L}^-)^{-1} \rangle_c$ . From the definition of inverse multiplication, (5.66), an expression for  $\mathbf{U}^{(B)}$  is obtained [see (5.35)]:

$$\mathbf{U}^{(\mathrm{B})} = \left[ -(\mathcal{L}^{+})^{-1} \tilde{\mathbf{G}}_{\mathrm{e}}^{+} (\tilde{\mathbf{G}}_{\mathrm{e}}^{+})^{-1} \otimes (\tilde{\mathbf{G}}_{\mathrm{e}}^{-})^{-1} \tilde{\mathbf{G}}_{\mathrm{e}}^{-} (\mathcal{L}^{-})^{-1} \right] \\ + \left[ (\mathcal{L}^{+})^{-1} \tilde{\mathbf{G}}_{\mathrm{e}}^{+} (\tilde{\mathbf{G}}_{\mathrm{e}}^{+})^{-1} \mathcal{L}^{+} (\tilde{\mathbf{G}}_{\mathrm{e}}^{+})^{-1} \otimes (\tilde{\mathbf{G}}_{\mathrm{e}}^{-})^{-1} \mathcal{L}^{-} (\tilde{\mathbf{G}}_{\mathrm{e}}^{-})^{-1} \tilde{\mathbf{G}}_{\mathrm{e}}^{-} (\mathcal{L}^{-})^{-1} \right] \\ = - \left\langle \mathbf{L}^{+} \otimes \mathbf{L}^{-} \right\rangle_{\mathrm{c}}^{-1} + (\tilde{\mathbf{G}}_{\mathrm{e}}^{+})^{-1} \otimes (\tilde{\mathbf{G}}_{e}^{-})^{-1}.$$
(5.94)

# 5.5.1 Cancellation of the Impurity Scattering Terms

To further simplify the term  $\langle \mathbf{L}^+ \otimes \mathbf{L}^- \rangle_c^{-1}$  with the aim of getting rid of the  $\mathbf{v}_i$ 's, it is necessary first to examine the expression

$$\begin{split} \left\langle \mathbf{L}^{+} - \mathbf{L}^{-} \right\rangle_{\mathrm{c}} &= \left\langle \mathbf{L}^{+} \otimes (\mathbf{L}^{-})^{-1} \mathbf{L}^{-} - \mathbf{L}^{+} (\mathbf{L}^{+})^{-1} \otimes \mathbf{L}^{-} \right\rangle_{\mathrm{c}} \\ &= \left\langle \mathbf{L}^{+} \otimes \left[ (\tilde{\mathbf{G}}_{\mathrm{e}}^{-})^{-1} - (\mathbf{v}_{i}^{-} - \tilde{\Sigma}^{-}) \right] \mathbf{L}^{-} \right\rangle_{\mathrm{c}} \\ &- \left\langle \mathbf{L}^{+} \left[ (\tilde{\mathbf{G}}_{\mathrm{e}}^{+})^{-1} - (\mathbf{v}_{i}^{+} - \tilde{\Sigma}^{+}) \right] \otimes \mathbf{L}^{-} \right\rangle_{\mathrm{c}}. \end{split}$$
(5.95)

In particular, let us pull out the two terms related to  $\mathbf{v}_i^{\pm}$  and look at them in the form of fixed real-space input and output channels:

$$\langle \mathbf{L}^{+}\mathbf{v}_{i}^{+} \otimes \mathbf{L}^{-} \rangle_{c} - \langle \mathbf{L}^{+} \otimes \mathbf{v}_{i}^{-}\mathbf{L}^{-} \rangle_{c}$$

$$= \langle (\mathbf{L}^{+})_{\mathbf{r}_{1}\mathbf{r}_{2}}(\mathbf{v}_{i}^{+})_{\mathbf{r}_{2}}(\mathbf{L}^{-})_{\mathbf{r}_{3}\mathbf{r}_{4}} - (\mathbf{L}^{+})_{\mathbf{r}_{1}\mathbf{r}_{2}}(\mathbf{v}_{i}^{-})_{\mathbf{r}_{3}}(\mathbf{L}^{-})_{\mathbf{r}_{3}\mathbf{r}_{4}} \rangle_{c},$$

$$(5.96)$$

where  $\mathbf{r}_1$ ,  $\mathbf{r}_2$ ,  $\mathbf{r}_3$ ,  $\mathbf{r}_4$ , are all within the same particle, with implied summations over  $\mathbf{r}_2$  in the first term and over  $\mathbf{r}_3$  in the second term. In addition, the fact that  $\mathbf{v}_i$  is diagonal in real space has been utilized. Here the two terms cancel if  $\mathbf{v}_i$  is a constant inside the particle. But even if  $\mathbf{v}_i$  is not constant, it is recalled from the definition of the pulse intensity, (5.4), that  $\mathbf{r}_1 = \mathbf{r}_4$  and  $\mathbf{r}_2$  $= \mathbf{r}_3$  because there are only two distinct spatial points, source position  $\mathbf{r}'$  and receiver position  $\mathbf{r}$ , in the problem (see the solution to Problem 5.1). Therefore, the two terms must cancel for the quantum case, and for the classical wave case they cancel only when  $\Delta \omega = 0$ , which is always the case for the derivation of the Ward identity. This is a crucial step. With  $\mathbf{v}_i^{\pm}$  eliminated from (5.95), the rest of (5.95) may be written as

$$\left\langle \mathbf{L}^{+} \otimes \mathbf{I} - \mathbf{I} \otimes \mathbf{L}^{-} \right\rangle_{c} = \tilde{\mathbf{G}}_{e}^{+} \otimes \mathbf{I} - \mathbf{I} \otimes \tilde{\mathbf{G}}_{e}^{-}$$
  
=  $\left\langle \mathbf{L}^{+} \otimes \mathbf{L}^{-} \right\rangle_{c} : \left[ \mathbf{I} \otimes (\tilde{\mathbf{G}}_{e}^{-})^{-1} - (\tilde{\mathbf{G}}_{e}^{+})^{-1} \otimes \mathbf{I} + \mathbf{I} \otimes \tilde{\Sigma}^{-} - \tilde{\Sigma}^{+} \otimes \mathbf{I} \right], \quad (5.97)$ 

where the first equality follows from the CPA relation, and the convention  $(\mathbf{L}^+ \otimes \mathbf{L}^-) : (\mathbf{I} \otimes (\tilde{\mathbf{G}}_e^-)^{-1}) = \mathbf{L}^+ \otimes (\tilde{\mathbf{G}}_e^-)^{-1} \mathbf{L}^-$  is followed; i.e., the quantity to the right always forms inner products with the output channels of the quantity to the left. From (5.97), one can use the inverse notation to get

$$-\left\langle \mathbf{L}^{+}\otimes\mathbf{L}^{-}\right\rangle _{c}^{-1}=\left[\mathbf{I}\otimes(\tilde{\mathbf{G}}_{e}^{-})^{-1}-(\tilde{\mathbf{G}}_{e}^{+})^{-1}\otimes\mathbf{I}\right.$$
$$+\mathbf{I}\otimes\tilde{\Sigma}^{-}-\tilde{\Sigma}^{+}\otimes\mathbf{I}\right]:\left(\mathbf{I}\otimes\tilde{\mathbf{G}}_{e}^{-}-\tilde{\mathbf{G}}_{e}^{+}\otimes\mathbf{I}\right)^{-1}.$$
(5.98)

This expression may be further simplified because

$$\mathbf{I} \otimes (\tilde{\mathbf{G}}_{e}^{-})^{-1} - (\tilde{\mathbf{G}}_{e}^{+})^{-1} \otimes \mathbf{I} \\
= (\tilde{\mathbf{G}}_{e}^{+})^{-1} \tilde{\mathbf{G}}_{e}^{+} \otimes \mathbf{I} (\tilde{\mathbf{G}}_{e}^{-})^{-1} - (\tilde{\mathbf{G}}_{e}^{+})^{-1} \mathbf{I} \otimes \tilde{\mathbf{G}}_{e}^{-} (\tilde{\mathbf{G}}_{e}^{-})^{-1} \\
= -\left[ (\tilde{\mathbf{G}}_{e}^{+})^{-1} \otimes (\tilde{\mathbf{G}}_{e}^{-})^{-1} \right] : \left( -\tilde{\mathbf{G}}_{e}^{+} \otimes \mathbf{I} + \mathbf{I} \otimes \tilde{\mathbf{G}}_{e}^{-} \right). \quad (5.99)$$

Substitution of (5.99) into (5.98) gives

$$-\left\langle \mathbf{L}^{+} \otimes \mathbf{L}^{-} \right\rangle_{c}^{-1} = -(\tilde{\mathbf{G}}_{e}^{+})^{-1} \otimes (\tilde{\mathbf{G}}_{e}^{-})^{-1} + (\mathbf{I} \otimes \tilde{\Sigma}^{-} - \tilde{\Sigma}^{+} \otimes \mathbf{I}) : \left( \mathbf{I} \otimes \tilde{\mathbf{G}}_{e}^{-} - \tilde{\mathbf{G}}_{e}^{+} \otimes \mathbf{I} \right)^{-1}.$$
 (5.100)

Substitution of (5.100) into (5.94) yields the Ward identity in the operator notation:

$$\mathbf{U}^{(\mathrm{B})} = \mathbf{u}^{+} \otimes \mathbf{u}^{-} = \left(\mathbf{I} \otimes \tilde{\Sigma}^{-} - \tilde{\Sigma}^{+} \otimes \mathbf{I}\right) : \left(\mathbf{I} \otimes \tilde{\mathbf{G}}_{\mathrm{e}}^{-} - \tilde{\mathbf{G}}_{\mathrm{e}}^{+} \otimes \mathbf{I}\right)^{-1}, \quad (5.101)$$

or

$$\mathbf{u}^{+} \otimes \tilde{\mathbf{G}}_{e}^{-} \mathbf{u}^{-} - \mathbf{u}^{+} \tilde{\mathbf{G}}_{e}^{+} \otimes \mathbf{u}^{-} = \mathbf{I} \otimes \tilde{\Sigma}^{-} - \tilde{\Sigma}^{+} \otimes \mathbf{I}.$$
(5.102)

By remembering that  $\tilde{\mathbf{G}}_{e}^{\pm} = \mathbf{H}\mathbf{G}_{e}^{\pm}\mathbf{H}$ , this last equation may be rewritten as

$$\mathbf{u}^{+}\mathbf{G}_{e}^{+}\mathbf{H}\otimes\mathbf{u}^{-}-\mathbf{u}^{+}\otimes\mathbf{H}\mathbf{G}_{e}^{-}\mathbf{u}^{-}=\tilde{\Sigma}^{+}\otimes\mathbf{I}-\mathbf{I}\otimes\tilde{\Sigma}^{-},$$
(5.103)

where  $\mathbf{u}^+ \mathbf{H}$  and  $\mathbf{H}\mathbf{u}^-$  are noted to be just  $\mathbf{u}^+$  and  $\mathbf{u}^-$  because they are already defined on the sites of the particle, so multiplying by the particle function  $\mathbf{H}$  does not alter them in any way. The usual way of displaying the Ward identity is in the **k**-representation, where  $\tilde{\Sigma}^{\pm}$  are diagonal. For input and output channels of (5.103) characterized by  $\mathbf{k}_{\pm}, \mathbf{k}'_{\pm}$  respectively, and with  $\mathbf{u}^+ \otimes \mathbf{u}^- = U^{(B)}_{\mathbf{k},\mathbf{k}'}(|\mathbf{k}_+\rangle \langle \mathbf{k}'_+| \otimes |\mathbf{k}'_-\rangle \langle \mathbf{k}_-|)$  the operator equation becomes

$$\begin{cases} \frac{1}{N} \sum_{\mathbf{k}'} U_{\mathbf{k},\mathbf{k}'}^{(\mathrm{B})} G_{\mathrm{e}}^{+}(\omega^{+},\mathbf{k}'_{+}) H(\mathbf{k}'_{+}-\mathbf{k}_{+}) - U_{\mathbf{k},\mathbf{k}'}^{(\mathrm{B})} G_{\mathrm{e}}^{-}(\omega^{-},\mathbf{k}'_{-}) H(\mathbf{k}'_{-}-\mathbf{k}_{-}) \end{cases} \\ \times (|\mathbf{k}_{+}\rangle \langle \mathbf{k}'_{+}| \otimes |\mathbf{k}'_{-}\rangle \langle \mathbf{k}_{-}|) = \left[ \tilde{\Sigma}^{+}(\omega^{+},\mathbf{k}_{+}) - \tilde{\Sigma}^{+}(\omega^{-},\mathbf{k}_{-}) \right] \\ \times (|\mathbf{k}_{+}\rangle \langle \mathbf{k}'_{+}| \otimes |\mathbf{k}'_{-}\rangle \langle \mathbf{k}_{-}|), \qquad (5.104) \end{cases}$$

where the **k** dependence of  $\tilde{\Sigma}$  is put in for accuracy, even though in the CPA it is **k** independent. By noting that  $H(\mathbf{k}) = 1$  for  $|\mathbf{k}| < |\mathbf{k}|_{\max} \approx 1/R$ , defining  $\Delta G_{\rm e}$  as  $G_{\rm e}^+ - G_{\rm e}^-$ , and returning  $\tilde{\Sigma}^{\pm}$  to  $\Sigma^{\pm}$ , (5.104) becomes

$$\frac{1}{N} \sum_{\mathbf{k}'} \frac{N}{L^d} U_{\mathbf{k},\mathbf{k}'}^{(B)} (\Delta G_e)_{\mathbf{k}'} = \Sigma^+(\omega^+,\mathbf{k}_+) - \Sigma^-(\omega^-,\mathbf{k}_-), \qquad (5.105a)$$

which is the usual form of the Ward identity, (5.42). The prime on the summation sign expresses the fact that the range of  $\mathbf{k}'$  covers only  $|\mathbf{k}' - \mathbf{k}| < |\mathbf{k}|_{\text{max}}$ . Since  $(\Delta G_{\text{e}})_{\mathbf{k}}$  is a peaked function, this restriction is essentially immaterial as long as the allowed  $|\mathbf{k}'|$  regime overlaps with  $\kappa_{\text{e}}$ .

## 5.5.2 Reduction to the Simplified Model

In the case of point scatterers such as encountered in the Anderson model and the long-wavelength limit of the CPA, the  $U_{\mathbf{k},\mathbf{k}'}^{(B)}$  is  $\mathbf{k}(\mathbf{k}')$  independent, and the summation over  $\mathbf{k}'$  in (5.105a) produces  $\Delta G_{\rm e}$  for a single site, i.e.,  $\Delta G_{\rm e}(\omega,\mathbf{l}=\mathbf{l}')$ . The Ward identity of the Anderson model is therefore given by an algebraic equation

$$\frac{N}{L^d} U^{(B)} \Delta G_{\mathbf{e}}(\omega, \mathbf{l} = \mathbf{l}') = \Sigma^+(\omega) - \Sigma^-(\omega).$$
 (5.105b)

This is precisely the case treated in the simplified model (5.31b). Since  $\Delta G_{\rm e}(\omega, \mathbf{l} = \mathbf{l}')$  is given by

$$\Delta G_{\rm e}(\omega, \mathbf{l} = \mathbf{l}') = \frac{1}{N} \sum_{\mathbf{k}} (\Delta G_{\rm e})_{\mathbf{k}} = -2 \,\mathrm{i} \frac{\mathrm{d}\omega}{\mathrm{d}\kappa_{\rm e}^2} \pi \rho_{\rm e}(\omega),$$

it follows that

$$U^{(B)} = \frac{-\mathrm{Im}\Sigma^{+}}{\pi\rho_{\mathrm{e}}(\omega)} \frac{\mathrm{d}\kappa_{\mathrm{e}}^{2}}{\mathrm{d}\omega} \frac{L^{d}}{N}.$$
 (5.105c)

An important observation here is that the *independence* of  $U^{(B)}$  from **k**, **k'** makes the solution of S possible [see (5.30)]. Note that the irreducible vertex is inherently a quantity in the second moment of the Green function, whereas the self energy is a quantity in the first moment of the Green function. The Ward identity thus expresses a relation between the two quantities in two different moments of Green functions.

## 5.5.3 Meaning of the Ward Identity

In (5.105a), if  $\Delta \omega$ ,  $|\Delta \mathbf{k}| = 0$ , then the right-hand side is ~ 2 iIm $\Sigma^+$  ~ 2 iIm  $\langle T \rangle_c$ . On the left-hand side,  $(\Delta G_e)_{\mathbf{k}'}$  may be approximated as  $-2\pi i$  times a delta function (5.28). The summation over  $\mathbf{k}'$  is therefore effectively an angular integration. If one looks only at the numerator of  $\mathbf{U}^{(B)}$ ,  $\langle \mathbf{t}_i^+ \otimes \mathbf{t}_i^- \rangle_c$  [see (5.83)], the angular integration yields the total scattering cross-section. The form of the Ward identity is therefore very close to the optical theorem. However, an important difference arises if the effective medium is allowed to be complex. In that case Im  $\langle \mathbf{t}_i \rangle_c = 0$ , as in the CPA for the Anderson model, and the optical theorem in the usual sense does not exist. Instead, energy conservation has to be expressed through the imaginary part of the self-energy, i.e.,  $-\text{Im}\Sigma^+$ , which is the quantity that appears in the Ward identity. Therefore the Ward identity may be regarded as a generalized optical theorem for a scattering medium.

# 5.6 Diffusion Constant Modification for Classical Waves

The derivation of the CPA Ward identity clearly demonstrates the necessity for  $\Sigma$  and  $\mathbf{U}$  to be evaluated to the same accuracy in order for the identity to hold. However, the condition of  $\mathbf{v}_i$  being frequency independent is not something that can be satisfied by classical waves in random media, regardless of the accuracy of  $\Sigma$  and  $\mathbf{U}$ . Therefore, the Ward identity in its form of (5.42) is strictly valid for classical scalar waves only in the  $\Delta \omega = 0$  limit ( $\omega^+ = \omega^-$ ). However, it should be noted that for the classical waves,  $\mathbf{v}_i^+$  does not cancel  $\mathbf{v}_i^-$  in (5.97), and in addition to  $\Sigma^+ - \Sigma^-$  on the right-hand side of (5.105a), there is an extra term of the form

$$-\frac{1}{N^2}\sum_{\mathbf{k}'}\sum_{\mathbf{k}''}\frac{N}{L^d}\left\langle\left\langle \mathbf{k} \mid \psi_{\mathbf{k}'}^+\right\rangle\left\langle\psi_{\mathbf{k}'}^+\left|\frac{\partial\mathbf{v}_i}{\partial\omega}\right|\psi_{\mathbf{k}''}^+\right\rangle\left\langle\psi_{\mathbf{k}''}^+\mid\mathbf{k}\right\rangle\right\rangle_{c}\Delta\omega,$$

where  $\psi_{\mathbf{k}'}^+$  is an eigenfunction with incident wave vector  $\mathbf{k}'$  for a single configuration (the superscript + denotes that the eigenfunction is calculated from the retarded Green function  $\mathbf{G}^+$ ). This term is evaluated at  $|\Delta \mathbf{k}| = 0$  and it does not affect the validity of the Ward identity, which is evaluated at  $\Delta \omega$ ,  $|\Delta \mathbf{k}| = 0$  in any case. In the expression earlier, everything inside the outer  $\langle \rangle_c$ , except the state  $|\mathbf{k}\rangle$ , is subject to averaging. This can be done as follows. Since  $\partial \mathbf{v}_i / \partial \omega$  is an operator that is *local* in character, one can fix  $\mathbf{v}_i$  and average over configurations that change everything else. In the weak-scattering limit, the resulting averaged  $\langle \mathbf{k} \mid \psi_{\mathbf{k}'}^+ \rangle$  may be approximated by the form  $\exp(i\mathbf{k} \cdot \mathbf{r}) + f(\theta) \exp(ik'r)/r^{(d-1)/2}$  away from  $\mathbf{v}_i$  where  $f(\theta)$  is the scattering amplitude for  $\mathbf{v}_i$ . This is reasonable because after configurational averaging, only the coherent part of  $\psi_{\mathbf{k}'}^+$  is expected to remain, and the earlier approximation treats the surrounding (within a mean free path) of  $\mathbf{v}_i$  as being homogeneous in the CPA sense. Since  $\exp(i\mathbf{k} \cdot \mathbf{r})$  has an intrinsic magnitude of N,  $\langle \langle \mathbf{k} \mid \psi_{\mathbf{k}'}^+ \rangle_c \approx N \delta_{\mathbf{k},\mathbf{k}'}$  where the scattering term is assumed to be small compared to N. Using similar reasoning on  $\langle \langle \psi_{\mathbf{k}''}^+ \mid \mathbf{k} \rangle \rangle$  and summing over  $\mathbf{k}'$  and  $\mathbf{k}''$  yield

$$-\frac{N}{L^d}\left\langle\left\langle\psi_{\mathbf{k}}^+ \left|\frac{\partial\mathbf{v}_i}{\partial\omega}\right|\psi_{\mathbf{k}}^+\right\rangle\right\rangle_{\mathrm{c}}\Delta\omega$$

as the extra term when  $\Delta \omega \neq 0$ . For classical waves, it is recalled that

$$\mathbf{v}_{i} = \frac{\omega^{2}}{v_{0}^{2}} \sum_{\mathbf{r}} \left[1 - \varepsilon_{i}(\mathbf{r})\right] (\delta r)^{d} \cdot |\mathbf{r}\rangle \langle \mathbf{r}|, \quad \mathbf{r} \in \text{particle } i.$$

Using this form of  $\mathbf{v}_i$  gives the following form for the extra term:

$$-\frac{2\omega}{v_0^2} \sum_m n_m \sum_{\mathbf{r}} (\delta r)^d \left\langle \psi_{\mathbf{k}}^+ \mid \mathbf{r} \right\rangle \left\langle \mathbf{r} \mid [1 - \varepsilon_m(\mathbf{r})] \mid \mathbf{r} \right\rangle \left\langle \mathbf{r} \mid \psi_{\mathbf{k}}^+ \right\rangle \Delta \omega$$
$$= (2\omega \,\Delta \omega / v_0^2) \delta, \tag{5.106a}$$

where m is the index for the scatterer type, and

$$\delta = \int \mathrm{d}\mathbf{r} \sum_{m} n_m |\psi_{\mathbf{k}}^{+(m)}(\mathbf{r})|^2 [\varepsilon_m(\mathbf{r}) - 1].$$
 (5.106b)

Here the integrand is nonzero only inside the scatterers, and the wave function may be obtained approximately by solving a scattering boundary value problem where a single scatterer of type m is embedded in an effective medium. This procedure has been demonstrated in the last chapter.

Since the additional term is left over after the vertex function cancels with  $\Sigma^+ - \Sigma^-$ , it combines with the coefficient of S in (5.44a) with a negative sign to give

$$-\frac{\mathrm{d}\kappa_0^2}{\mathrm{d}\omega}\left(1+\frac{2\omega}{v_0^2}\frac{\mathrm{d}\omega}{\mathrm{d}\kappa_0^2}\delta\right)\Delta\omega S = -\frac{2\omega}{v_0^2}(1+\delta)\Delta\omega S = -\frac{2\omega}{v_0^2}\Delta\omega S'.$$
 (5.107a)

Here  $S' = (1+\delta)S$ , and we have specialized to classical waves (since this effect appears only for classical waves) so that  $d\kappa_0^2/d\omega = 2\omega/v_0^2$ . The modification

of S thus leads directly to the modification of the wave diffusion constant (van Albada et al. 1991). This is easily seen if S is rewritten as  $S'/(1+\delta)$  in (5.52) and (5.53); then for classical waves the diffusion constant expression becomes

$$D(\omega) = \frac{1}{\pi \rho_e(\omega)} \frac{8K_0^2 (\mathrm{d}\kappa_e^2/\mathrm{d}\kappa_0^2)}{M_0 - 8K_0^2 \mathrm{Im}(\Sigma^+)} \frac{-\mathrm{Im}\Sigma^+}{1+\delta}.$$
 (5.107b)

Physically,  $\delta$  can be large when the incident wave frequency coincides with an internal resonance of the scatterer. When that happens,  $|\psi_{\mathbf{k}}^{+}|^{2}$  has large magnitude inside the scatterer, leading to a large  $\delta$ . Another interpretation of resonant scattering is that the incident wave can get trapped inside the scatterer, bouncing around many times before emerging again. This picture tells us in an intuitive way that wave diffusion has to slow down when resonant scattering is present. However, (5.106b) would not be accurate in that case because the strong multiple scattering *between* the particles is not taken into account. These issues will be further addressed in Chap. 9.

## 5.7 Evaluation of the Wave Diffusion Constant

The expressions for  $D(\omega)$ , given by (5.55) and (5.107b), are to be evaluated for the Anderson model and for classical scalar waves, both to the accuracy of the CPA. Here the scattering is assumed to be weak, so the weak-scattering versions of the  $K_0$  and  $M_0$  expression, (5.48c) and (5.51b), will be used.

## 5.7.1 Quantum Case: the Anderson Model

For the Anderson model, a great simplification is obtained in noting that  $U_{\mathbf{k},\mathbf{k}'}^{(\mathrm{B})}$ , given by (5.76), is isotropic. As a result,  $M_0^{(\mathrm{B})}$  is zero; i.e., the ladder diagrams sum to zero in the case of the Anderson model. That means

$$D^{(\mathrm{B})}(\omega) = \frac{K_0(\mathrm{d}\kappa_{\mathrm{e}}^2/\mathrm{d}\kappa_0^2)}{\rho_{\mathrm{e}}(\omega)\pi} = -\frac{\mathrm{d}\kappa_{\mathrm{e}}^2/\mathrm{d}\kappa_0^2}{\rho_{\mathrm{e}}(\omega)\pi} \frac{L^d}{4N^2} \sum_{\mathbf{k}} [\nabla_k e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}}]^2 (\Delta G)_{\mathbf{k}}^2.$$
(5.108a)

This expression for the diffusion constant is generally known as the Kubo-Greenwood formula (Kubo 1956; Greenwood 1958). By using the weak-scattering approximation for  $K_0$ , (5.48c), the expression becomes

$$D^{(\mathrm{B})}(\omega) = \frac{\frac{1}{2} \int_{\mathrm{IBZ}} \frac{\mathrm{d}\mathbf{k}}{(2\pi)^d} [\nabla_k e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}}]^2 \delta[\kappa_{\mathrm{e}}^2 - e(\mathbf{k})] (\mathrm{d}\kappa_{\mathrm{e}}^2/\mathrm{d}\kappa_0^2)}{(-\mathrm{Im}\Sigma^+) \int_{\mathrm{IBZ}} \frac{\mathrm{d}\mathbf{k}}{(2\pi)^d} \delta[\kappa_{\mathrm{e}}^2 - e(\mathbf{k})] (\mathrm{d}\kappa_{\mathrm{e}}^2/\mathrm{d}\omega)}.$$
 (5.108b)

The factor  $\beta$  has been taken to be 1 in this chapter. If  $\beta \neq 1$ , its effect is to modify  $\kappa_{\rm e}^2$  to give  $\kappa_{\rm e}^2/\beta$ . In (5.108),  $(d\kappa_{\rm e}^2/d\kappa_0^2)/(d\kappa_{\rm e}^2/d\omega)$  can be combined

to give  $d\omega/d\kappa_0^2 = \hbar/2m$ , which is the unit of quantum diffusion constant. The rest of the expression represents an average of  $|\nabla_k e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}}|^2$  on the energy shell (as required by the delta function), divided by  $-\text{Im}\Sigma^+$ . Since  $e(\mathbf{k})$  represents the dispersion relation,  $\nabla_k e(\mathbf{k})$  gives the group velocity. On the other hand,  $-\text{Im}\Sigma^+$  gives the rate of scattering. Therefore the diffusion constant has the form

$$D^{(B)}(\omega) = \frac{1}{d} v_t l, \qquad (5.109)$$

where

$$l = \frac{\bar{\kappa}_{\rm e}}{-{\rm Im}\Sigma^+} \tag{5.110}$$

is denoted the mean free path, and  $v_t$  is the transport velocity, given by

$$v_t = \frac{\hbar \bar{\kappa}_{\rm e}}{m},\tag{5.111}$$

with

$$\bar{\kappa}_{\rm e}^2 = \frac{d}{4a^2} \frac{\int_{\rm IBZ} \mathrm{d}\mathbf{k} [a\nabla_k e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}}]^2 \delta[\kappa_{\rm e}^2 - e(\mathbf{k})]}{\int_{\rm IBZ} \mathrm{d}\mathbf{k} \delta[\kappa_{\rm e}^2 - e(\mathbf{k})]}.$$
(5.112)

Here the integrations are over the first Brillouin zone, as denoted by IBZ. From the definition of  $\bar{\kappa}_{e}$ , (5.112), it is clear that  $v_t$  is just the average group velocity on the constant frequency surface. The strong-scattering version of (5.110) is given in Chap. 9 [(9.21)].

In (5.112), the integral in the numerator is noted to be independent of the direction of  $\Delta \hat{\mathbf{k}}$ . This can be seen from the fact that  $a\nabla_k e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}} = 2\sum_i (\sin k_i a) (\Delta \hat{\mathbf{k}})_i$ . When squared, the cross terms integrate to zero because  $\sin k_i a$  is an odd function of  $k_i$ , whereas  $\delta[\kappa_e^2 - e(\mathbf{k})]$  is an even function. For the terms that remain,  $4\sum_i (\sin^2 k_i a) (\Delta \hat{\mathbf{k}})_i^2$ , each  $\sin^2 k_i a$  term integrates to the same result independent of the direction i, so  $(\Delta \hat{\mathbf{k}})_i^2$  can be summed to give 1, the magnitude of  $\Delta \hat{\mathbf{k}}$ .

With the  $\bar{\kappa}_{e}^{2}$  given by (5.112), the mean free path expression, (5.110), is what has been plotted in Figs. 4.1–4.3. For electrons, the diffusion constant is directly proportional to the electrical conductivity  $\sigma_{E}$  through the Einstein relation:

$$\sigma_{\rm E} = \frac{e^2}{\hbar} \rho_{\rm e}(\omega) D^{\rm (B)}(\omega).$$
(5.113)

Here e is the electronic change and  $e^2/\hbar$  has the unit of conductance, given by 4,108 ohms<sup>-1</sup>. The unit of conductivity is thus  $(e^2/\hbar a^{d-2})$ . In Figs. 5.6– 5.8 the conductivity of the Anderson model is plotted as a function of the electron energy for d=1, 2, and 3, respectively, with  $a^2\sigma_{\omega} = 0.5$ . Since in actual materials only electrons at the Fermi surface can carry current, the electron energy may alternatively be regarded as the electron Fermi energy. The electronic spin degeneracy is not taken into account here.



**Fig. 5.6.** Conductivity of the 1D Anderson model for  $\sigma_w a^2 = 0.5$ ,  $\beta = 1$ , plotted as a function of the nondimensionalized Fermi energy  $u = (E_{\rm F} - \varepsilon_0)/(\beta\hbar^2/ma^2)$ . Here u = 0 means that the Fermi level is at the center of the band.



Fig. 5.7. Conductivity of the 2D Anderson model for  $\sigma_w a^2 = 0.5$ ,  $\beta = 1$ , plotted as a function of reduced Fermi energy  $u = (E_{\rm F} - \varepsilon_0)/(\beta\hbar^2/ma^2)$ . Here the dip in the band center is seen to reflect the effect of the small mean free path (see Fig. 4.2).



Fig. 5.8. Conductivity of the 3D Anderson model for  $\sigma_w a^2 = 0.5$ ,  $\beta = 1$ , plotted as a function of reduced Fermi energy  $u = (E_{\rm F} - \varepsilon_0)/(\beta\hbar^2/ma^2)$ . The structure in the vicinity of the band center is due to the interplay between l,  $v_t$ , and  $\rho_{\rm e}(\omega)$ .

# 5.7.2 Classical Wave Diffusion Constant

For classical scalar waves in continuum, the function  $e(\mathbf{k}) = k^2$ , so the constants  $K_0$  and  $M_0^{(B)}$  may be explicitly evaluated. Since the weak scattering versions of  $K_0$  and  $M_0$  are used, it is simpler to consider the combinations  $(-\mathrm{Im}\Sigma^+)K_0$  and  $(-\mathrm{Im}\Sigma^+)M_0$ . With these combinations the diffusion constant expression may be rewritten as

$$D(\omega) = \frac{[\pi \rho_{\rm e}(\omega)]^{-1} 8[(-\mathrm{Im}\Sigma^+) K_0]^2 (\mathrm{d}\kappa_{\rm e}^2/\mathrm{d}\kappa_0^2)(1+\delta)^{-1}}{(-\mathrm{Im}\Sigma^+) M_0 + 8[(-\mathrm{Im}\Sigma^+) K_0](-\mathrm{Im}\Sigma^+)}.$$
 (5.114)

Let us first calculate the denominator. For  $[(-\text{Im}\Sigma^+)K_0]$ , (5.48c) tells us that

$$(-\mathrm{Im}\Sigma^{+})K_{0} = \frac{\pi}{2} \int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^{d}} 4k^{2} \cos^{2}\theta_{1} \delta\left[\kappa_{\mathrm{e}}^{2} - k^{2}\right] = \frac{\pi}{(2\pi)^{d}} C^{d} \kappa_{\mathrm{e}}^{d}, \quad (5.115a)$$

where  $\theta_1$  is the angle between **k** and  $\Delta \mathbf{k}$ , and  $C_d = 4\pi/3$ ,  $\pi$ , and 2 in 3D, 2D, and 1D, respectively. For  $(-\text{Im}\Sigma^+)M_0^{(B)}$ , (5.51b) may be written as

$$(-\mathrm{Im}\Sigma^{+})M_{0}^{(\mathrm{B})} = \int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^{d}} \int \frac{\mathrm{d}\mathbf{k}'}{(2\pi)^{d}} (-4\pi^{2})\delta\left[\kappa_{\mathrm{e}}^{2} - k^{2}\right]$$
$$\times \delta\left[\kappa_{\mathrm{e}}^{2} - (k')^{2}\right] 4kk'\cos\theta_{1}\cos\theta_{2} U_{\mathbf{k},\mathbf{k}'}^{(\mathrm{B})}\frac{N}{L^{d}}, \quad (5.115b)$$

where  $\theta_2$  is the angle between  $\mathbf{k}'$  and  $\Delta \mathbf{k}$ . The magnitude part of the integrals can be performed immediately with the help of the delta functions. Also, from (5.85) the vertex function is given by

$$U_{\mathbf{k},\mathbf{k}'}^{(\mathrm{B})}\frac{N}{L^{d}} = \sum_{m} n_{m} \hat{O}_{m}(\mathbf{k},\mathbf{k}') \frac{2(2\pi)^{d}}{\pi\kappa_{\mathrm{e}}^{d-3}} = \sum_{m} n_{m} \hat{O}_{\kappa_{e}}^{(m)}(\theta) \frac{2(2\pi)^{d}}{\pi\kappa_{\mathrm{e}}^{d-3}}, \qquad (5.116)$$

where the magnitudes of  $\mathbf{k}$  and  $\mathbf{k}'$  are set at  $\kappa_{\rm e}$ , and a slight change in the notation is meant to bring out explicitly the dependencies on the scattering angle  $\theta$  between  $\mathbf{k}$  and  $\mathbf{k}'$ , as well as on the magnitude of the wave vector. The hat on O is a reminder that the quantity is evaluated with respect to the CPA effective medium and renormalized by 1 + Q (see (5.86)). Substitution of (5.116) into (5.115b) yields

$$(-\mathrm{Im}\Sigma^{+})M_{0}^{(\mathrm{B})} = -\frac{\kappa_{\mathrm{e}}^{d+1}}{(2\pi)^{d-2}}\frac{2}{\pi}\int\mathrm{d}\Omega_{\hat{\mathbf{k}}}\int\mathrm{d}\Omega_{\hat{\mathbf{k}}'} \times \cos\theta_{1}\cos\theta_{2}\sum_{m}n_{m}\hat{O}_{\kappa_{e}}^{(m)}(\theta), \qquad (5.117)$$

where  $d\Omega_{\hat{\mathbf{k}}}$  and  $d\Omega_{\hat{\mathbf{k}}'}$  denote angular integrations. In the integrand,  $\cos \theta_2$  can be expressed in terms of  $\theta$  and  $\theta_1$  as

$$\cos \theta_2 = \begin{cases} \cos \theta \cos \theta_1 + \sin \theta \sin \theta_1 \cos(\phi - \phi_1) \text{ 3D} \\ \cos \theta \cos \theta_1 - \sin \theta \sin \theta_1 & \text{2D} \\ \pm 1 & \text{1D.} \end{cases}$$
(5.118)

Substitution of (5.118) into (5.117) results in

$$(-\mathrm{Im}\Sigma^{+})M_{0}^{(\mathrm{B})} = -\frac{\kappa_{\mathrm{e}}^{d+1}}{(2\pi)^{d-2}}$$

$$= \sum_{\alpha,\beta} \int_{-1}^{2\pi} d\theta \cos\theta \sum_{m} n_{m} \hat{O}_{\kappa_{\mathrm{e}}}^{(m)}(\theta) = 3\mathrm{D}$$

$$\times C_d \frac{\tilde{-}}{\pi} \begin{cases} \int_0^{\infty} d\theta \cos \theta \sum_m n_m O_{\kappa_e}^{(m)}(\theta) & 2D\\ \sum_m n_m \left[ \hat{O}_{\kappa_e}^{(m)}(0) - \hat{O}_{\kappa_e}^{(m)}(\pi) \right] & . & 1D \end{cases}$$

For the additional  $8(-\mathrm{Im}\Sigma^+)[(-\mathrm{Im}\Sigma^+)K_0]$  term present in the denominator of the diffusion constant, we use the Ward identity in the  $\Delta\omega = 0$  limit to evaluate  $-\mathrm{Im}\Sigma^+$ :

$$-\mathrm{Im}\Sigma^{+} = \frac{\mathrm{i}}{2}(\Sigma^{+} - \Sigma^{-}) = \frac{\mathrm{i}}{2} \int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^{d}} \sum_{m} n_{m} \hat{O}_{\kappa_{\mathrm{e}}}^{(m)}(\theta)$$
$$\times \frac{2(2\pi)^{d}}{\pi \kappa_{\mathrm{e}}^{d-3}} \left\{ -2\pi \mathrm{i}\delta \left[\kappa_{\mathrm{e}}^{2} - k^{2}\right] \right\}$$
$$= \kappa_{\mathrm{e}} \begin{cases} 2\pi \int_{-1}^{1} \mathrm{d}(\cos\theta) \sum_{m} n_{m} \hat{O}_{\kappa_{\mathrm{e}}}^{(m)}(\theta) & 3\mathrm{D} \\ \int_{0}^{2\pi} \mathrm{d}\theta \sum_{m} n_{m} \hat{O}_{\kappa_{\mathrm{e}}}^{(m)}(\theta) & 2\mathrm{D} \\ \sum_{m} n_{m} \left[ \hat{O}_{\kappa_{\mathrm{e}}}^{(m)}(0) + \hat{O}_{\kappa_{\mathrm{e}}}^{(m)}(\pi) \right] & 1\mathrm{D}. \end{cases}$$
(5.120)

It is seen that the multiplication of  $-\mathrm{Im}\Sigma^+$  with  $8(-\mathrm{Im}\Sigma^+)K_0$  [(5.115a)] yields the exact prefactor (with a sign change) as that for  $(-\mathrm{Im}\Sigma^+)M_0^{(\mathrm{B})}$ , (5.119). Therefore the denominator of  $D^{(\mathrm{B})}(\omega)$  (5.114),  $(-\mathrm{Im}\Sigma^+)M_0^{(\mathrm{B})} + 8K_0(-\mathrm{Im}\Sigma^+)^2$ , may be combined to give  $(1-\cos\theta)\Sigma_m n_m \hat{O}_{\kappa_e}^{(m)}(\theta)$  as the angular integral's integrand. What remains for the diffusion constant calculation is the numerator of (5.114),  $8K_0^2(\mathrm{d}\kappa_e^2/\mathrm{d}\kappa_0^2)\pi^{-1}\rho_e^{-1}(\omega)(-\mathrm{Im}\Sigma^+)^2(1+\delta)^{-1}$ , which is reduced straightforwardly to

$$\frac{8\pi}{(2\pi)^d} \frac{C_d}{d} \kappa_{\rm e}^{d+2} \frac{v_0^2}{\omega} \frac{1}{1+\delta}.$$

By combining terms, a simple expression for  $D^{(B)}(\omega)$  is obtained:

$$D^{(B)}(\omega) = \frac{1}{d} \frac{v_0^2}{\omega} \frac{\kappa_e}{1+\delta} l^* = \frac{1}{d} \frac{v_0^2}{v_e(1+\delta)} l^* = \frac{1}{d} v_t l^*, \qquad (5.121)$$

where  $l^*$  is exactly the classical Boltzmann expression for the transport mean free path, given by
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$$l^{*} = \left[ \int d\Omega (1 - \cos \theta) \sum_{m} n_{m} \hat{O}_{\kappa_{e}}^{(m)}(\theta) \right]^{-1} = \begin{cases} \left[ 2\pi \int_{-1}^{1} d(\cos \theta) (1 - \cos \theta) \sum_{m} n_{m} \hat{O}_{\kappa_{e}}^{(m)}(\theta) \right]^{-1} & 3D \\ \left[ \int_{0}^{2\pi} d\theta (1 - \cos \theta) \sum_{m} n_{m} \hat{O}_{\kappa_{e}}^{(m)}(\theta) \right]^{-1} & 2D \quad (5.122) \\ \left[ 2\sum_{m} n_{m} \hat{O}_{\kappa_{e}}^{(m)}(\pi) \right]^{-1} & , & 1D \end{cases}$$

and

$$v_t = v_0^2 \frac{\kappa_{\rm e}}{\omega(1+\delta)} = v_0 \sqrt{1 - v_0^2 \frac{{\rm Re}\Sigma^+}{\omega^2}} \frac{1}{1+\delta}$$
(5.123)

is defined as the transport velocity. The last equality follows from the definition of  $\kappa_{\rm e} = \sqrt{\kappa_0^2 - {\rm Re}\Sigma^+}$ . Aside from the  $(1 + \delta)^{-1}$  factor,  $v_t$  is equal to the group velocity, given by  $(d\kappa_{\rm e}/d\omega)^{-1} = v_0^2\kappa_{\rm e}/\omega = v_0^2/(\omega/\kappa_{\rm e})$ , provided that the CPA is valid and the frequency dependence of  ${\rm Re}\Sigma^+$  is negligible compared to that of  $\kappa_0^2$ . The actual calculation and interpretation of  $v_t$  in a random medium are addressed in Chap. 9.

When the scattering is isotropic, the  $\cos \theta$  term integrates to zero, and  $l^* = l$ , the usual mean free path. In that case a comparison with the former definition of l, given by (4.9), shows there to be a difference in the definition of the scattering cross-sections. Here  $\hat{O}_{\kappa_e}^{(m)}(\theta)$  is defined with a 1 + Q denominator [see (5.86a)], which is absent in the former definition because the geometric repulsion effect (related to the successive scattering from the same particle) has been ignored.

An interesting observation is that at low frequencies the scalar wave scattering is always isotropic; therefore the contribution of the ladder diagrams is zero, i.e.,  $M_0^{(B)} = 0$ . On the other hand, for vector electromagnetic waves the long-wavelength scattering is always dipolar in its angular dependence, as noted in the last chapter. Therefore, in the present formalism  $M_0^{(B)} \neq 0$  for electromagnetic waves even in the long-wavelength limit. The only catch here is that the formalism developed so far is for scalar waves. Therefore whether  $M_0^{(B)}$  is truly nonzero for electromagnetic waves would have to be resolved by a fully vectorial calculation without any approximation that might destroy the polarization dependence of each scattering. Such a calculation is still lacking so far.

## 5.7.3 Anisotropic Scattering and the Transport Mean Free Path

When the scattering is anisotropic, the transport mean free path  $l^*$  is always larger than l because the  $(1 - \cos \theta)$  factor tends to emphasize the nonforward

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directions, and directional randomization would thus take more than one collision, hence  $l^* > l$ . Since anisotropic scalar-wave scattering implies that the wave can already resolve some features of the scatterer, the assumption about the k independence of  $\Sigma$ , as well as the effective medium description, may both be called into question. When that happens, the diffusive behavior still persists at distances larger than  $l^*$ , but  $\Sigma$  has to be calculated with the spectral function approach described in Chap. 4. However, a more serious correction to the diffusion constant is the contributions to **U** beyond the ladder diagrams. It turns out that there are indeed such contributions due to the so-called maximally crossed diagrams, which are manifest physically as the coherent backscattering effect, to be described in the next chapter. But before leaving the subject of diffusive waves it is important to point out an important application of diffusive waves for the optical measurement of dynamical properties in opaque materials.

## 5.8 Application: Diffusive Wave Spectroscopy

Most of the fluids we encounter daily are actually fluid suspensions. Coffee, milk, ink, mineral water, paint, etc. all contain a certain fraction of suspended solid particles which give the fluid some or most of its physical properties. In particular, when the concentration of the solid particles is high and their optical indices of refraction are sufficiently different from that of the solution liquid, the suspension would appear opaque due to the strong scattering by the particles, even though the solution liquid may be transparent in its pure form. As a result, traditional optical measurement of suspensions is usually difficult except in the extremely dilute case.

In the past decade, a new optical measurement technique has emerged which is capable of determining the dynamic properties of the suspended particles, such as their diffusion constant. The technique relies on the fact that the suspended particles are generally not stationary but move stochastically as Brownian particles due to the thermal agitation by the solution molecules. As a result, the light which is scattered by the suspended particles must also exhibit some stochastic time dependence, i.e., fluctuations. The desired dynamic information about the suspension may be retrieved by an analysis of the time correlation function of the scattered light. The diffusive character of the light (Genack 1990) enters precisely in such an analysis and is crucial for understanding the observed experimental results, hence the name of the technique - diffusive wave spectroscopy (DWS) (Pine et al. 1990a).

## 5.8.1 The Siegert Relation

To give an understanding of the theoretical basis of DWS, let us start with what can be measured experimentally, which is  $\langle I(\Delta t)I(0)\rangle_c / \langle I\rangle_c^2$ . Here *I* is the intensity of the scattered light at a given point,  $\Delta t$  is the time separation between two intensity measurements, and  $\langle \rangle_c$  denotes averaging, which is

generally a time average experimentally but can also be interpreted as an average over different scattering configurations of the sample. At  $\Delta t = 0$ , this quantity becomes  $\langle I^2 \rangle_c / \langle I \rangle_c^2$ , which is always greater than or equal to 1 and can be written as

$$\frac{\langle I^2 \rangle_c}{\langle I \rangle_c^2} = 1 + \text{normalized variance of intensity fluctuations}$$
$$= 1 + \frac{\langle I^2 \rangle_c - \langle I \rangle_c^2}{\langle I \rangle_c^2}.$$
(5.124)

In the solution to Problem 5.4, it is shown that for  $\Delta t \neq 0$ , a similar relation may be written as

$$\frac{\langle I(\Delta t)I(0)\rangle_{\rm c}}{\langle I\rangle_{\rm c}^2} = 1 + g_2(\Delta t), \qquad (5.125)$$

where

$$g_2(\Delta t) = \left| \frac{\langle \phi(\Delta t)\phi^*(0)\rangle_c}{\langle |\phi|^2\rangle_c} \right|^2 = |g_1(\Delta t)|^2,$$
(5.126)

and  $\phi$  is the wave amplitude (electric field amplitude of light). Equations (5.125) and (5.126) are known as the *Siegert relation*. This relation is widely used in optical measurements to relate the measured intensity to its theoretical interpretation, which generally begins with wave amplitudes. The crucial quantity here is therefore  $\langle \phi(\Delta t)\phi^*(0)\rangle_c$ . In order to arrive at a theoretical expression for  $\langle \phi(\Delta t)\phi^*(0)\rangle_c$ , the first question one may ask is: What is the difference between  $\phi(\Delta t)$  and  $\phi(0)$ ? Suppose the source of light is at a distance r away from the measuring point. In arriving at the measuring point the light would have encountered many scatterings in a zigzag path. Let us for the moment focus on one class of light trajectories that consist of precisely n scatterings and label the field amplitude associated with that class of trajectories  $\phi^{(n)}$ . The difference between  $\phi^{(n)}(0)$  and  $\phi^{(n)}(\Delta t)$  is clearly caused by the movement of the intervening scatterers during the time period  $\Delta t$ . By assuming all scatterings to be elastic, this fact can be expressed mathematically as

$$\left\langle \phi^{(n)}(\Delta t)\phi^{(n)*}(0)\right\rangle_{c} = \left\langle |\phi^{(n)}(0)|^{2}\right\rangle_{c} \left\langle \exp\left\{-i\sum_{i=1}^{n}\mathbf{q}_{i}\cdot\left[\mathbf{r}_{i}(\Delta t)-\mathbf{r}_{i}(0)\right]\right\}\right\rangle_{c} \right\rangle_{c} \left\langle 5.127\right\rangle_{c}$$

Here  $\mathbf{q}_i$  is the scattering wave vector and  $\mathbf{r}_i$  the position of the scatterer for the *i*th scattering. If  $\theta_i$  is the scattering angle, then

$$|\mathbf{q}_i| = 2\kappa_{\rm e} \sin \frac{\theta_i}{2}.\tag{5.128}$$

By decomposing  $g_1(\Delta t)$  as

$$g_1(\Delta t) = \sum_{n=1}^{\infty} g_1^{(n)}(\Delta t),$$
 (5.129a)

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from (5.126) and (5.127) we have

$$g_1^{(n)}(\Delta t) = P(n) \left\langle \prod_{i=1}^n \exp\{-\mathbf{q}_i \cdot [\mathbf{r}_i(\Delta t) - \mathbf{r}_i(0)]\} \right\rangle_{\rm c}, \qquad (5.129b)$$

where  $P(n) = \langle |\phi^{(n)}(0)|^2 \rangle_c / \langle |\phi|^2 \rangle_c$  is the probability that a trajectory of n scatterings ends at the measurement point. The averaging bracket  $\langle \rangle_c$  in (5.129b) has the double function of averaging over  $\mathbf{r}_i$  and over all possible directions of  $\mathbf{q}_i$ . Strictly speaking, for fixed incoming and outgoing wave vectors,  $\mathbf{r}_i$  and  $\mathbf{q}_i$  are correlated. However, for large n the correlation is weak. Here  $\mathbf{r}_i$  and  $\mathbf{q}_i$  will be treated as independent. This approximation allows the two averages to be done separately.

#### 5.8.2 Averaging Over the Diffusive Motion of Scatterers

To average over  $\mathbf{r}_i$ , the fact that the scatterers are undergoing Brownian (diffusive) motion means that the probability density of the quantity for  $|\mathbf{r}_i(\Delta t) - \mathbf{r}_i(0)|$  is given by the Gaussian distribution:

$$\frac{1}{(4\pi D_{\rm p}\Delta t)^{3/2}} \exp\left[-\left|\mathbf{r}_i(\Delta t) - \mathbf{r}_i(0)\right|^2 / 4D_{\rm p}\Delta t\right],\tag{5.130}$$

where d = 3 is assumed since experiments so far have all been done on 3D systems, and  $D_{\rm p}$  is the diffusion constant of the suspended particles. Averaging each term of (5.129b) with respect to the position of particle *i* means integrating the product of the Gaussian distribution (5.130) with each term, leading to

$$g_{1}^{(n)}(\Delta t) = P(n) \left\langle \prod_{i=1}^{n} \exp\left(-D_{\rm p}\Delta t q_{1}^{2}\right) \right\rangle_{\rm c}$$
$$= P(n) \left\langle \prod_{i=1}^{n} \exp\left(-4D_{\rm p}\kappa_{\rm e}^{2}\Delta t \sin^{2}\frac{\theta_{i}}{2}\right) \right\rangle_{\rm c}$$
(5.131)

through the well-known trick of completing the squares in the exponent (which can then be exactly integrated) so that the term left over is precisely that shown in (5.131). Since  $(D_{\rm p}\kappa_{\rm e}^2)^{-1}$  has the unit of time, it is denoted  $t_0$ . Physically,  $t_0$  is the time scale for a suspended particle to diffuse a distance on the order of a wavelength. It also determines the decay rate of the autocorrelation function  $g_1(\Delta t)$ .

## 5.8.3 Averaging Over the Scattering Angles

To the leading order of  $\Delta t/t_0$ , i.e., small  $\Delta t$ , the average over the scattering angle  $\theta_i$  in (5.131) may be approximated by

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$$\left\langle \prod_{i=1}^{n} \exp\left(-4D_{\rm p}\kappa_{\rm e}^2 \Delta t \sin^2\frac{\theta_i}{2}\right) \right\rangle_{\rm c} \approx \exp\left[-2n\frac{\Delta t}{t_0} \left\langle 1 - \cos\theta_i \right\rangle_{\rm c}\right], \quad (5.132)$$

which is equivalent to retaining the first term in the cumulant expansion. Since the average of  $1 - \cos \theta_i$  is precisely the factor that differentiates  $l^*$  from l, it follows from (5.122) that

$$\langle 1 - \cos \theta_i \rangle_{\rm c} = \frac{l}{l^*},\tag{5.133}$$

so that

$$g_1^{(n)}(\Delta t) \cong P(n) \exp\left[-\frac{2\Delta t}{t_0} \frac{l}{l^*}n\right] = P(s) \exp\left[-\frac{2\Delta t}{t_0} \frac{s}{l^*}\right], \quad (5.134)$$

where the discrete variable n is converted into the continuous path-length variable s = nl for ease of calculations later on.

## 5.8.4 Diffusive Transport of Waves

A crucial step in the evaluation of  $g_1(\Delta t)$  involves the calculation of P(s). Consistent with the fact that waves behave diffusively after many scatterings, it is reasonable to approximate P(s) by the solution to the diffusion equation, i.e.,

$$P(s) = \left(\frac{1}{4\pi D^{(\mathrm{B})}t}\right)^{3/2} \exp\left(-\frac{r^2}{4D^{(\mathrm{B})}t}\right)$$
$$= \left(\frac{3}{4\pi s l^*}\right)^{3/2} \exp\left(-\frac{3r^2}{4s l^*}\right).$$
(5.135)

Here t is the photon travel time (which is very small because  $v_t$  is so large), r is the separation between the source and the measurement point, and in the last equality  $D^{(B)}t$  is expressed as  $v_t t l^*/3 = s l^*/3$ . One should be aware that although the expression (5.135) is very similar to (5.130), physically (5.130) is for the diffusion of suspended particles, whereas (5.135) is for the diffusion of light, which are two very different phenomena. The diffusive property of light is embodied in  $l^*$  plus the form of P(s) used. From (5.129a) and (5.134), we have

$$g_{1}(\Delta t) = \sum_{n=1}^{\infty} g_{1}^{(n)}(\Delta t) = \int_{0}^{\infty} \mathrm{d}s \, P(s) \exp\left(-\frac{2\Delta t}{t_{0}} \frac{s}{l^{*}}\right)$$
$$= \int_{0}^{\infty} \mathrm{d}s \, P(s) \exp\left(\frac{3}{4\pi s l^{*}}\right)^{3/2} \exp\left[-\frac{3r^{2}}{4s l^{*}} - \frac{2\Delta t}{t_{0}} \frac{s}{l^{*}}\right]. \quad (5.136)$$

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The dominant behavior of  $g_1(\Delta t)$  may be readily inferred from (5.136) because the expression in the exponent has a peak, located at the point  $s = s_0$  where the derivative of the exponent vanishes, or

$$\frac{3r^2}{4l^*}\frac{1}{s_0^2} - \frac{2\Delta t}{t_0}\frac{1}{l^*} = 0,$$

$$s_0 = r_1 \sqrt{\frac{3t_0}{l^*}}$$
(5.137)

from which one obtains

$$S_0 = i \sqrt{8\Delta t}$$
 (0.137)  
is in the exponent, the position of the peak dominates the

Since the peak is in the exponent, the position of the peak dominates the behavior of  $g_1(\Delta t)$ . Therefore, substitution of  $s_0$  back into the exponent gives

$$g_1(\Delta t) \sim \exp\left[-\frac{r}{l^*}\sqrt{\frac{6\Delta t}{t_0}}\right],$$
 (5.138)

which implies

$$g_2(\Delta t) = \left|g_1(\Delta t)\right|^2 \sim \exp\left[-\frac{2r}{l^*}\sqrt{\frac{6\Delta t}{t_0}}\right].$$
(5.139)

Equation (5.137) and the resulting form of  $g_1(\Delta t)$ , (5.138), can be interpreted physically as follows.

The form of (5.138), the exponential decay with a square-root dependence on  $\Delta t$ , is noted to arise from the competition between the term that describes particle diffusion,  $\exp[-(2\Delta t/t_0)s/l^*]$ , which favors short paths because they dephase much less, and the term that describes light diffusion,  $\exp[-3r^2/4sl^*]$ , which favors long paths because there are so many more of them. The final result represents a compromise between the two. For  $\Delta t$  small, the dominant path length  $s_0$  is large [as seen from (5.137)] because at short times the suspended particles move very little, and only through the long paths can these little changes add up to anything that can cause the decay of the autocorrelation function. One has to remember that light travels essentially instantaneously compared with the movement of the particles; therefore even a small  $\Delta t$  on the scale of  $t_0$  means an almost infinite travel distance for light. Through this interpretation it is easy to see that one can measure the displacements of the suspended particles that are much smaller than the light wavelength, although in this case the measurement represents an average over many particles rather than that of a single particle, and therefore does not represent a violation of the wave resolution limit.

To compare theory with experimental measurement, the geometry of the experiment has to be taken into account. For this purpose it is much more convenient to retain the form of  $g_1(\Delta t)$  as

$$g_1(\Delta t) = \int_0^\infty \mathrm{d}s \, P(s) \exp\left(-\frac{2\Delta t}{t_0} \frac{s}{l^*}\right),\tag{5.140}$$

where P(s) has to be obtained from the solution to a boundary value problem of the (light) diffusion equation. Equation (5.140) states that  $g_1$  is the Laplace transform of P(s). It turns out that the relevant mathematics is more accessible in the Laplace-transformed domain. Therefore, let us now solve for  $g_1(\Delta t)$ directly with the geometry in which a coherent laser beam is uniformly incident on the front surface of a slab sample with thickness L, extending from z = 0 to L, and measurement is done either at the front surface (reflected light) or the back surface (transmitted light). Before the calculation, however, it is necessary to clarify two points.

## 5.8.5 Boundary Conditions for Diffusive Light Flux

The first point is that the measured quantity is the *flux* of diffusive light. In the solution to Problem 5.5 it is shown that the flux of diffusing photons across the sample boundary in the +z direction is given by

$$J_{+} = \frac{P v_t}{4} - \frac{D^{(B)}}{2} \left(\frac{\partial P}{\partial z}\right).$$
 (5.141a)

Similarly, the flux in the -z direction is given by

$$J_{-} = \frac{P v_t}{4} + \frac{D^{(B)}}{2} \left(\frac{\partial P}{\partial z}\right).$$
 (5.141b)

Here P is the *diffusive* photon density. It is a function of photon path length and detection location; i.e., P is a function of s and  $\mathbf{r}$ , and  $D^{(B)} = v_t l^*/3$  is the light diffusion constant. The boundary condition at z = L is that there should be no net flux of diffusing photons coming into the sample from the outside. The only  $J_{-}$  measured inside the sample has to be that reflected from the sample surface at z = L. If the reflection coefficient is  $\rho$ , then at z = L

$$J_{-}=\rho J_{+},$$

which means

$$P + \left(\frac{2l^*}{3}\frac{1+\rho}{1-\rho}\right)\frac{\partial P}{\partial z} = 0.$$
 (5.142)

If this expression is used to eliminate  $\partial P/\partial z$  in  $J_+$ , then the result is

$$J_{+} = \frac{P v_t}{2(1+\rho)} \tag{5.143a}$$

at z = L. Similarly,

$$J_{-} = \frac{P v_t}{2(1+\rho)} \tag{5.143b}$$

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at z = 0. Therefore, what can be measured is proportional to P. This discussion also clarifies the boundary conditions at z = 0, L, which may be summarized as

$$P + C \frac{\partial P}{\partial z} = 0 \quad \text{at } z = L,$$
 (5.144a)

$$P - C \frac{\partial P}{\partial z} = 0$$
 at  $z = 0$ , (5.144b)

where

$$C = \frac{2l^*}{3} \frac{1+\rho}{1-\rho}.$$
 (5.144c)

The second point is about the source of the diffusing light. Although experimentally the source is a steady laser beam, in the solution for P(s), which essentially counts the number of paths with a given path length s, we need a time-dependent solution of P where the source is a pulse of light. It is noted that the parameter s, the path length, is directly proportional to the photon travel time t as  $s = v_t t$ . On the scale of  $t_0$ , the photon travel time t is minuscule for a given s. We will denote the photon travel time by  $s/v_t$  so as to distinguish it from the particle diffusion time. Another aspect of the source is its location. Since the light entering the sample is coherent, it must take a few scatterings to make it diffusive. Therefore the diffusive light source should be located at some distance  $z = z_0 \approx l^*$  inside the sample.

#### 5.8.6 Solutions for the Slab Geometry

With the earlier boundary conditions and the initial condition, the (diffusion) equation for P is

$$D^{(B)}\frac{\partial^2}{\partial z^2}P(z,s) - v_t\frac{\partial P(z,s)}{\partial s} = \delta(z-z_0)\,\delta\left(\frac{s}{v_t}\right),\tag{5.145}$$

where the dependence on the x, y coordinates is absent because both the incident beam and the measurement points are assumed to cover whole surfaces, so the problem becomes mathematically one-dimensional. The solution of (5.145) consists of the addition of two parts,  $P = P^{(1)} + P^{(2)}$ . One part is that of the Green function in an infinite 1D medium, given by the well-known solution

$$P^{(1)}(z,t) = \frac{1}{(4\pi D^{(\mathrm{B})}s/v_t)^{1/2}} \exp\left[-(z-z_0)^2 v_t/4D^{(\mathrm{B})}s\right].$$
 (5.146)

The other part of the solution,  $P^{(2)}$ , is that of the homogeneous diffusion equation [i.e., right-hand side of (5.145) is zero], with coefficients that may be adjusted to satisfy the boundary conditions at z = 0, L given by (5.144a, b). Since in the end it is the Laplace-transformed solution that is needed, the problem should be reformulated directly in the Laplace-transformed domain, where  ${\cal P}^{(2)}$  is easy to obtain. If

$$\bar{P}^{(1,2)}(z,q) = \int_0^\infty P^{(1,2)}(z,s) \exp(-qs) \,\mathrm{d}s, \qquad (5.147a)$$

then

$$q = \frac{2\Delta t}{t_0 l^*} \tag{5.147b}$$

by comparison with (5.140). It is straightforward to get

$$\bar{P}^{(1)} = \exp\left[-\frac{1}{l^*}\sqrt{\frac{6\Delta t}{t_0}} |z - z_0|\right] / \frac{2}{3}\sqrt{\frac{6\Delta t}{t_0}}$$
(5.148)

by either performing the integral, (5.147a), or looking up a Laplace transform table.  $P^{(2)}$  is the solution of

$$D^{(B)} \frac{\partial^2}{\partial z^2} \bar{P}^{(2)} - v_t q \bar{P}^{(2)} = 0, \qquad (5.149)$$

or

$$\bar{P}^{(2)} = A \cosh\left(\sqrt{\frac{6\Delta t}{t_0}}\frac{z}{l^*}\right) + B \sinh\left(\sqrt{\frac{6\Delta t}{t_0}}\frac{z}{l^*}\right).$$
(5.150)

The transformed boundary conditions for  $\bar{P} = N_0(\bar{P}^{(1)} + \bar{P}^{(2)})$  are

$$\bar{P} + C \frac{\partial P}{\partial z} = 0$$
 at  $z = L$ ,  
 $\bar{P} - C \frac{\partial \bar{P}}{\partial z} = 0$  at  $z = 0$ .

The normalization constant  $N_0$  ensures that at  $\Delta t = 0$ ,  $\bar{P} = 1$ . The constants A and B are uniquely fixed by the two boundary conditions, and the results for reflection measurement are given by

$$g_{1}(\Delta t) = \bar{P}(\Delta t, z = 0)$$

$$= \frac{3L + [(1+\rho)/(1-\rho)] 4l^{*}}{3(L-z_{0}) + [(1+\rho)/(1-\rho)] 2l^{*}}$$

$$\times \left\{ \frac{\sinh\left(\sqrt{\frac{6\Delta t}{t_{0}}} \frac{L-z_{0}}{l^{*}}\right)}{\left(\frac{1+\rho}{1-\rho}\sqrt{\frac{8\Delta t}{3t_{0}}}\cosh\left(\sqrt{\frac{6\Delta t}{t_{0}}} \frac{L-z_{0}}{l^{*}}\right)\right)}{\left[1+\left(\frac{1+\rho}{1-\rho}\right)^{2} \frac{8\Delta t}{3t_{0}}\right]\sinh\left(\sqrt{\frac{6\Delta t}{t_{0}}} \frac{L}{l^{*}}\right)}{\left(1+\left(\frac{21+\rho}{1-\rho}\sqrt{\frac{8\Delta t}{3t_{0}}}\right)\cosh\left(\sqrt{\frac{6\Delta t}{t_{0}}} \frac{L}{l^{*}}\right)}\right)}\right\}.$$
(5.151)

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It is easily checked that  $g_1(0) = 1$ , as it should be. For a sample of infinite thickness,  $L = \infty$ , the expression simplifies to

$$g_1(\Delta t) = \frac{\exp\left[-\sqrt{\frac{6\Delta t}{t_0}}\frac{z_0}{l^*}\right]}{1 + \frac{1+\rho}{1-\rho}\sqrt{\frac{8\Delta t}{3t_0}}}.$$
 (5.152)

For small  $\Delta t/t_0$ , the time dependencies in the numerator and denominator means that  $g_1(\Delta t)$  behaves as  $\exp[-c_0\sqrt{6\Delta t/t_0}]$ , where

$$c_0 \cong \frac{z_0}{l^*} + \frac{2}{3} \frac{1+\rho}{1-\rho}.$$
(5.153)

The value of  $c_0$  is known to be on the order of 2. For example, if reflectively  $\rho \approx 0.4$  and  $z_0 \approx l^*$ , then  $c_0 \approx 2.5$ . The value of  $\rho \approx 0.4$  is not unreasonable for a ratio of indices of refraction of ~1.4 at the two sides of a sample interface. By knowing  $c_0$ ,  $t_0$  (and hence  $D_p$ ) is uniquely determined from the reflection measurements. In Fig. 5.9 it is shown that the experimentally measured  $g_2(\Delta t)$  follows very well the predicted behavior.

One can also measure the transmitted intensity of the diffusing light. The predicted form of  $g_1(\Delta t)$  is



Fig. 5.9. Measured autocorrelation function  $g_2(\Delta t)$  in the reflection geometry. It is seen that the data are well described by an exponential  $\sqrt{\Delta t}$  decay, as predicted. The slope directly gives the value of  $c_0$ . The figure is reproduced from Pine (1990b). In the figure,  $\Delta t$  is denoted by  $\tau$ ,  $t_0$  by  $\tau_0$ . The sample is an L = 5-mm - thick cellwith 5% 0.412-µm - diameter polystyrene spheres. The *solid line* is the theoretical fit to the data.

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$$g_{1}(\Delta t) = \bar{P}(\Delta t, z = L) = \frac{L/l^{*} + (4/3)(1+\rho)/(1-\rho)}{z_{0}/l^{*} + (2/3)(1+\rho)/(1-\rho)} \\ \times \begin{cases} \sinh\left(\sqrt{\frac{6\Delta t}{t_{0}}}\frac{z_{0}}{l^{*}}\right) \\ +\frac{1+\rho}{1-\rho}\sqrt{\frac{8\Delta t}{3t_{0}}}\cosh\left(\sqrt{\frac{6\Delta t}{t_{0}}}\frac{z_{0}}{l^{*}}\right) \\ \frac{\left[\left(\frac{1+\rho}{1-\rho}\right)^{2}\frac{8\Delta t}{3t_{0}} + 1\right]\sinh\left(\sqrt{\frac{6\Delta t}{t_{0}}}\frac{L}{l^{*}}\right)}{1-\rho} \\ +2\frac{1+\rho}{1-\rho}\sqrt{\frac{8\Delta t}{3t_{0}}}\cosh\left(\sqrt{\frac{6\Delta t}{t_{0}}}\frac{z_{0}}{l^{*}}\right)} \end{cases} \end{cases} \right\}.$$
(5.154)

In this case  $L/l^*$  is a sensitive parameter of the solution. Therefore  $l^*$  may be obtained by performing experiments on samples of different thicknesses. For light scattering in a solution with a 1% concentration of  $0.5-\mu$ m-diameter polystyrene spheres,  $l^*$  has been measured to be  $\cong 180 \,\mu$ m, implying a light diffusion constant of  $D^{(B)} \cong 2 \times 10^8 \text{ cm}^2/\text{s}$ . In strong-scattering media, values of  $D^{(B)}$  as low as  $10^5-10^6 \text{ cm}^2/\text{s}$  have been observed (Genack, 1990).

## **Problems and Solutions**

**Problem 5.1** Derive an expression for the spatial spectral decomposition of the configurationally averaged  $\langle G^+G^-\rangle_c$ .

Solution From definitions,

$$G^{+}\left(\omega + \frac{\Delta\omega}{2}, \mathbf{r}, \mathbf{r}'\right) = \frac{L^{2d}}{N^{2}(2\pi)^{2d}} \iint d\mathbf{k} d\mathbf{k}' \\ \times \exp[i(\mathbf{k} \cdot \mathbf{r}) - i(\mathbf{k}' \cdot \mathbf{r}')] \langle \mathbf{k} | G^{+}(\omega^{+}) | \mathbf{k}' \rangle, \quad (P5.1)$$
$$G^{-}\left(\omega - \frac{\Delta\omega}{2}, \mathbf{r}', \mathbf{r}\right) = \frac{L^{2d}}{N^{2}(2\pi)^{2d}} \iint d\mathbf{k}_{1} d\mathbf{k}'_{1} \\ \times \exp[i(\mathbf{k}'_{1} \cdot \mathbf{r}') - i(\mathbf{k}_{1} \cdot \mathbf{r})] \langle \mathbf{k}'_{1} | G^{-}(\omega^{-}) | \mathbf{k}_{1} \rangle, \quad (P5.2)$$

Let  $\mathbf{R}_{c} = (\mathbf{r} + \mathbf{r}')/2$  so that

$$\mathbf{r} = \mathbf{R}_{c} + \frac{(\mathbf{r} - \mathbf{r}')}{2},$$
$$\mathbf{r}' = \mathbf{R}_{c} - \frac{(\mathbf{r} - \mathbf{r}')}{2}.$$

As seen from Problem 3.1, configurational averaging may be written as

$$\langle G^+ G^- \rangle_{\mathbf{r}} = \frac{L^{4d}}{N^4 (2\pi)^{4d}} \cdot \int \int \int \int d\mathbf{k} \, d\mathbf{k}' \, d\mathbf{k}'_1 \, d\mathbf{k}_1 \times \frac{1}{L^d} \int d\mathbf{R}_c \, \langle G^+ G^- \rangle_{\mathbf{k}} \exp[i(\mathbf{k} - \mathbf{k}' + \mathbf{k}'_1 - \mathbf{k}_1) \cdot \mathbf{R}_c] \times \exp[i(\mathbf{k} + \mathbf{k}' - \mathbf{k}'_1 - \mathbf{k}_1) \cdot (\mathbf{r} - \mathbf{r}')/2],$$
(P5.3)

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where  $\langle \ \rangle_{\mathbf{r}}$  denotes the configurational average of the product of the lefthand sides of (P5.1) and (P5.2) in real space,  $\langle \ \rangle_{\mathbf{k}}$  denotes the average of the product of the two corresponding quantities in the integrals of (P5.1) and (P5.2) in  $\mathbf{k}$  space. It should be emphasized that the latter can depend on  $\mathbf{k}$ ,  $\mathbf{k}'$ , and  $\mathbf{k}_1$ ,  $\mathbf{k}'_1$ . Thus the subscript  $\mathbf{k}$  does not imply dependence only on  $\mathbf{k}$ . Integration over  $\mathbf{R}_c$  yields  $(2\pi)^d \delta(\mathbf{k} - \mathbf{k}' + \mathbf{k}'_1 - \mathbf{k}_1)$ . Integration with respect to  $\mathbf{k}_1$  then gives

$$\left\langle G^{+}G^{-}\right\rangle_{\mathbf{r}} = \frac{L^{3d}}{N^{3}(2\pi)^{3d}} \cdot \int \int \int d\mathbf{k} \, d\mathbf{k}' \, d\mathbf{k}'_{1} \\ \times \left\langle G^{+}G^{-}\right\rangle_{\mathbf{k}} \exp[\mathrm{i}(\mathbf{k}'-\mathbf{k}'_{1})\cdot(\mathbf{r}-\mathbf{r}')].$$
(P5.4)

One can introduce a new set of variables that satisfy the delta function constraint automatically. Denote  $\mathbf{k}' - \mathbf{k}'_1 = \mathbf{k} - \mathbf{k}_1 = \Delta \mathbf{k}$ , then

$$\mathbf{k} \rightarrow \mathbf{k} + \frac{\Delta \mathbf{k}}{2} = \mathbf{k}_{+}$$
$$\mathbf{k}' \rightarrow \mathbf{k}' + \frac{\Delta \mathbf{k}}{2} = \mathbf{k}'_{+}$$
$$\mathbf{k}_{1} \rightarrow \mathbf{k} - \frac{\Delta \mathbf{k}}{2} = \mathbf{k}_{-}$$
$$\mathbf{k}'_{1} \rightarrow \mathbf{k}' - \frac{\Delta \mathbf{k}}{2} = \mathbf{k}'_{-}, \qquad (P5.5)$$

and

$$\langle G^+ G^- \rangle_{\mathbf{r}} = \frac{L^{3d}}{N^3 (2\pi)^{3d}} \cdot \int \int \int \langle G^+ G^- \rangle_{\mathbf{k}} \times \exp[\mathrm{i}(\Delta \mathbf{k} \cdot (\mathbf{r} - \mathbf{r}'))] \mathrm{d}\mathbf{k} \, \mathrm{d}\mathbf{k}' \, \mathrm{d}\Delta \mathbf{k}.$$
 (P5.6)

It is noted that whereas only  $\Delta \mathbf{k}$  appears in the exponent,  $\langle G^+G^- \rangle_{\mathbf{k}}$  in general can depend on  $\mathbf{k}$ ,  $\mathbf{k}'$ , as well as  $\Delta \mathbf{k}$ . However,  $\langle G^+G^- \rangle_{\mathbf{r}}$  depends only on  $\mathbf{r} - \mathbf{r}'$  because for intensity transport,  $G^+$  and  $G^-$  have the same source and detection points [see (5.1)]. However, if the quantity to be averaged is  $G^+(\omega^+, \mathbf{r}_+, \mathbf{r}'_+)G^-(\omega^-, \mathbf{r}'_-, \mathbf{r}_-)$ , where  $\mathbf{r}'_+ \neq \mathbf{r}'_-$ ,  $\mathbf{r}_+ \neq \mathbf{r}_-$ , then substitution of these  $\mathbf{r}$ 's into (P5.1) and (P5.2) gives for the general case:

$$\left\langle G^{+}G^{-}\right\rangle_{\mathbf{r}} = \frac{L^{3d}}{N^{3}(2\pi)^{3d}} \cdot \int \int \int \left\langle G^{+}G^{-}\right\rangle_{\mathbf{k}} \exp[\mathrm{i}(\Delta\mathbf{k}\cdot\mathbf{R}_{m}) + \mathrm{i}\mathbf{k}\cdot(\mathbf{r}_{+}-\mathbf{r}_{-}) + \mathrm{i}\mathbf{k}'\cdot(\mathbf{r}'_{+}-\mathbf{r}'_{-})] \,\mathrm{d}\mathbf{k}\,\mathrm{d}\mathbf{k}'\,\mathrm{d}\Delta\mathbf{k}.$$
(P5.7)

Here  $\mathbf{R}_m = [(\mathbf{r}_+ - \mathbf{r'}_+) + (\mathbf{r}_- - \mathbf{r'}_-)]/2$  is the "average" vector that bisects the two vectors linking the input and output sites of the + and - channels.

**Problem 5.2** Show that  $\nabla_k e(\mathbf{k})$  is an odd function of  $\mathbf{k}$ .

**Solution** From the definition of  $e(\mathbf{k})$ , (2.37), we have

$$\nabla_k e(\mathbf{k}) = \frac{2}{a} \left( \sin k_x a \, \hat{\mathbf{i}} + \sin k_y a \, \hat{\mathbf{j}} + \sin k_z a \, \hat{\mathbf{k}} \right), \tag{P5.8}$$

where  $\hat{\mathbf{i}}, \hat{\mathbf{j}}, \hat{\mathbf{k}}$  are unit vectors along the x, y, z directions, respectively. Each component of the vector is an odd function of its argument. Therefore, if  $\mathbf{k} \to -\mathbf{k}$ , we have  $\nabla_k e(-\mathbf{k}) = -\nabla_k e(\mathbf{k})$ , i.e.,  $\nabla_k e(-\mathbf{k})$  is an odd function of  $\mathbf{k}$ .

**Problem 5.3** Show that the configurationally averaged value of  $\mathbf{L}^{\pm}$  is given by  $\tilde{\mathbf{G}}_{e}^{\pm}$ , where  $\mathbf{L}^{\pm}$  is defined by (5.88), and  $\tilde{\mathbf{G}}_{e}^{\pm}$  is  $\mathbf{G}_{e}^{\pm}$  restricted on the sites of one particle only.

Solution From the CPA condition,

$$\left\langle \left( \mathbf{v}_{i} - \tilde{\Sigma} \right) \left[ \mathbf{I} - \tilde{\mathbf{G}}_{e} \left( \mathbf{v}_{i} - \tilde{\Sigma} \right) \right]^{-1} \right\rangle_{c} = 0,$$

one can factor out  $\mathbf{\tilde{G}}_{\mathrm{e}}^{-1}$  to get

$$\left\langle \left( \mathbf{v}_{i} - \tilde{\Sigma} \right) \left[ \tilde{\mathbf{G}}_{e}^{-1} - \left( \mathbf{v}_{i} - \tilde{\Sigma} \right) \right]^{-1} \right\rangle_{c} = 0,$$
 (P5.9)

where the  $\pm$  superscripts have been suppressed. Equation (P5.9) may be alternatively written as

$$\left\langle -\mathbf{I} + \tilde{\mathbf{G}}_{\mathrm{e}}^{-1} \left[ \tilde{\mathbf{G}}_{\mathrm{e}}^{-1} - \left( \mathbf{v}_{i} - \tilde{\Sigma} \right) \right]^{-1} \right\rangle_{\mathrm{c}} = 0,$$
 (P5.10)

or

$$\left< \mathbf{L}^{\pm} \right>_{\mathrm{c}} = \mathbf{\tilde{G}}_{\mathrm{e}}^{\pm}$$

as asserted.

Problem 5.4 Derive the Siegert relation

$$\frac{\langle I(\Delta t)I(0)\rangle_{\rm c}}{\langle I\rangle_{\rm c}^2} = 1 + g_2(\Delta t),$$

where  $g_2(\Delta t) = |g_1(\Delta t)|^2$ .

Solution Expressed in terms of amplitudes,

$$\left\langle I(\Delta t)I(0)\right\rangle_{c} = \left\langle \phi(\Delta t)\,\phi^{*}(\Delta t)\,\phi(0)\,\phi^{*}(0)\right\rangle_{c}\,.\tag{P5.11}$$

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Each  $\phi(\Delta t)$  may be expressed as the sum of amplitudes resulting from different scattering paths. For example,

$$\phi(\Delta t) = \sum_{i} \phi_{i}(\Delta t), \qquad (P5.12)$$

where the index i denotes a particular scattering path. Therefore,

$$\left\langle I(\Delta t)I(0)\right\rangle_{\rm c} = \sum_{i,j,k,l} \left\langle \phi_i\left(\Delta t\right)\phi_j^*(\Delta t)\phi_k(0)\phi_l^*(0)\right\rangle_{\rm c}.$$
 (P5.13)

The sum is zero if one index is different from the other three, since that term may be averaged independently, and  $\langle \phi_i \rangle_c = 0$  because of random phase addition. Therefore, we are left with the possibility of i = j, k = l, or i = l, j = k, or i = k, j = l. The last case still gives zero because the phases do not cancel unless  $\phi$  is paired with  $\phi^*$ . Also, there is the term i = j = k = l. However, compared with the large number of other terms, the contribution of this term is negligible. Therefore,

$$\langle I(\Delta t)I(0)\rangle_{c} \cong \langle I(\Delta t)\rangle_{c} \langle I(0)\rangle_{c} + \langle \phi(\Delta t) \phi^{*}(0)\rangle_{c} \langle \phi^{*}(\Delta t) \phi(0)\rangle_{c} = \langle I\rangle_{c}^{2} + |\langle \phi(\Delta t) \phi^{*}(0)\rangle_{c}|^{2}.$$
 (P5.14)

Dividing by  $\langle I \rangle_{\rm c}^2$  on both sides and recalling that

$$g_1(\Delta t) = \frac{\langle \phi(\Delta t) \, \phi^*(0) \rangle_{\rm c}}{\langle I \rangle_{\rm c}}$$

we get the Siegert relation directly from (P5.14).

**Problem 5.5** Derive an expression for the diffusive particle flux across the surface defined by z = 0.

**Solution** Consider the flux in the -z direction across a surface element in the neighborhood of the origin. If the number density of diffusive particles at **r** is denoted by  $P(\mathbf{r})$ , then the contribution to the flux J from a small volume element centered at **r** is given by

$$P(\mathbf{r})\mathrm{d}\Omega\left(v_t\cos\theta\right)\exp(-r/l^*)/4\pi l^*.$$

Here r is the distance of the volume element to the surface element,  $\exp(-r/l^*)/l^*$  accounts for the probability of no scattering in traversing the distance r,  $v_t \cos \theta$  is the projected z component velocity, and  $d\Omega/4\pi$  gives the probability for a velocity in that particular direction,  $d\Omega$  being the solid angle differential  $d\phi \sin \theta d\theta$ , where  $\theta$  is defined relative to the surface normal. The total flux is therefore given by

$$J_{-} = \frac{v_{t}}{4\pi l^{*}} \int_{0}^{2\pi} \mathrm{d}\phi \int_{0}^{\pi/2} \mathrm{d}\theta \int_{0}^{\infty} \mathrm{d}r P(\mathbf{r}) \cos\theta \sin\theta \exp(-r/l^{*})$$

$$\cong \frac{v_{t}}{2l^{*}} \int_{0}^{\pi/2} \mathrm{d}\theta \int_{0}^{\infty} \mathrm{d}r \left[P(0) + \mathbf{r} \cdot (\nabla P(\mathbf{r}))_{\mathbf{r}=0}\right] \cos\theta \sin\theta \exp(-r/l^{*})$$

$$= \frac{v_{t}}{4} P(0) + \frac{v_{t}l^{*}}{6} \left. \frac{\partial P}{\partial z} \right|_{z=0}$$

$$= \frac{v_{t}}{4} P(0) + \frac{D^{(\mathrm{B})}}{2} \left. \frac{\partial P}{\partial z} \right|_{z=0}.$$
(P5.15)

In the equation above, it should be noted that  $\nabla P$  has only the z-component. For  $J_+$ , the only difference is in the second term, where  $\mathbf{r} \to -\mathbf{r}$  (due to the fact that the volume contributing to the flux has to be on the other side of the surface) results in a negative sign. Therefore

$$J_{+} = \frac{v_t P(0)}{4} - \frac{D^{(B)}}{2} \left. \frac{\partial P}{\partial z} \right|_{z=0}.$$
 (P5.16)

## The Coherent Backscattering Effect

## 6.1 Wave Diffusion versus Classical Diffusion

In the last chapter it was argued that wave intensity transport is diffusive in character after multiple scattering. The purpose of this and subsequent chapters is to point out the inadequacies of this description and the interesting and rich phenomena that can arise from the nonclassical character of wave transport.

The difference between classical diffusion and wave diffusion may be easily demonstrated. Take a slab of strong-scattering material and illuminate its front surface by a coherent laser beam. If only the total transmitted light is measured, then the measured value and its dependence on sample thickness should be well described by the solution to the diffusion equation with appropriate boundary conditions. However, whereas the diffusion equation predicts the transmitted intensity to have smooth spatial variation, in reality the transmitted intensity is not smooth but consists of interlaced bright and dark spots visible when the transmitted light is projected onto a screen, e.g., a sheet of paper. These intensity variations are called the speckle pattern, which is a result of wave interference in the absence of inelastic scattering. The essential point is, therefore, that the diffusive character and the interference effect are not mutually exclusive in the absence of inelastic scattering. In the case of the speckle pattern, its effect shows up after configurational averaging as increased magnitude of *intensity fluctuation* at a given spatial point, when compared to classical diffusion.

The coherent backscattering is another interference effect, sometimes also known as the weak localization effect, that manifests itself at the level of configurationally averaged *mean intensity*. It can therefore directly affect the diffusive behavior when its dynamic ramifications are taken into account. The purpose of this chapter is to introduce the effect in its static version. 184 6 The Coherent Backscattering Effect

## 6.2 Coherence in the Backscattering Direction

Figure 6.1 is a schematic illustration showing two ray paths in a scattering medium, composed of fixed, randomly placed scatterers. Both rays have the same frequency. Ray path A has amplitude  $A_i$  at some far-away source point  $\mathbf{R}_0$  and wave vector  $\mathbf{k}_i$ . After n scatterings, ray path A exits from the front surface with wave vector  $\mathbf{k}_f$ . The spatial locations of the n scatterers are labeled consecutively as  $\mathbf{r} = \mathbf{r}_1, \mathbf{r}_2, \mathbf{r}_3, \ldots, \mathbf{r}' = \mathbf{r}_n$ . Ray path B has amplitude  $B_i = A_i$  at the same source point  $\mathbf{R}_0$  and the same incident and exit wave vectors  $\mathbf{k}_i, \mathbf{k}_f$  as A. It also experiences n scatterings, but in exactly the reverse order to ray path A. Namely, ray path B starts with  $\mathbf{r}' = \mathbf{r}_n$  and ends with  $\mathbf{r} = \mathbf{r}_1$ . Since all scatterings are elastic, the complex amplitude  $A_0$  of ray path A at some far-away observation point  $\mathbf{r}_o$  may be related to  $A_i$  as

$$A_{0} = A_{i}G_{A} \exp[\mathbf{i}\mathbf{k}_{i} \cdot (\mathbf{r}_{1} - \mathbf{R}_{0}) + \mathbf{i}\mathbf{k}_{1,2} \cdot (\mathbf{r}_{2} - \mathbf{r}_{1}) + \cdots + \mathbf{i}\mathbf{k}_{n-1,n} \cdot (\mathbf{r}_{n} - \mathbf{r}_{n-1}) + \mathbf{i}\mathbf{k}_{f} \cdot (\mathbf{r}_{0} - \mathbf{r}_{n})].$$

$$(6.1)$$

Here  $\mathbf{k}_{n-1,n}$  is the wave vector of ray propagation between scatterers n-1 and n, and  $G_A$  is a factor that accounts for the scattering strength of each scatterer as well as other effects (such as geometric spreading) associated with the fact that we have only written out explicitly the phase factors of the Green functions, which are the mathematically accurate wave propagators between



**Fig. 6.1.** A schematic illustration of two interfering rays A and B. Both rays have the same incident wave vector  $\mathbf{k}_i$  and the exit wave vector  $\mathbf{k}_f$ . The scattering sites of ray path A are labeled consecutively as  $\mathbf{r} = \mathbf{r}_1$ ,  $\mathbf{r}_2 \dots$ ,  $\mathbf{r}_n = \mathbf{r}'$ . Ray path B has exactly the reverse scattering order. This choice is made purposely by following the time reversal rule in which ray path B may be regarded as ray path A propagating backward in time between  $\mathbf{r}_1$  and  $\mathbf{r}_n$ . If the system is time reversal invariant, A and B are perfectly coherent when  $\mathbf{k}_f = -\mathbf{k}_i$ . It should be noted that the backscattering direction is different from the direction of specular reflection. The two coincide only at normal incidence.

successive scatterers. Similarly, for ray path B, amplitude  $B_0$  at the same far-away observation point  $\mathbf{r}_0$  may be related to  $B_i = A_i$  as

$$B_0 = A_i G_B \exp[i\mathbf{k}_i \cdot (\mathbf{r}_n - \mathbf{R}_0) + i\mathbf{k}_{n,n-1} \cdot (\mathbf{r}_{n-1} - \mathbf{r}_n) + \cdots + i\mathbf{k}_{2,1} \cdot (\mathbf{r}_1 - \mathbf{r}_2) + i\mathbf{k}_f \cdot (\mathbf{r}_0 - \mathbf{r}_1)], \qquad (6.2)$$

where  $G_B$  is a factor similar to  $G_A$ . Since  $\mathbf{k}_{n,n-1} = -\mathbf{k}_{n-1,n}$ , it is clear that the sum of all the intervening phase factors, between the first term and the last term, is identical to that of (6.1). Also,  $G_A = G_B$  because identical scatters are involved. Therefore,

$$\frac{A_0}{B_0} = \exp[\mathrm{i}(\mathbf{k}_{\mathrm{i}} + \mathbf{k}_{\mathrm{f}}) \cdot (\mathbf{r} - \mathbf{r}')].$$
(6.3)

From Fig. 6.1, (6.3) has the simple physical interpretation that  $B_0$  has the extra phase  $-\mathbf{k}_i \cdot (\mathbf{r} - \mathbf{r}')$  due to the extra path length at incidence, whereas  $A_0$  has the extra phase  $\mathbf{k}_f \cdot (\mathbf{r} - \mathbf{r}')$  due to the extra path length at exit. The two cancels if  $\mathbf{k}_i = -\mathbf{k}_f$ , and the two outgoing rays are exactly coherent in the direction opposite to the incident direction.

## 6.2.1 Time Reversal Invariance

In actual scattering, the outgoing amplitude  $A_0$  should be the sum of an infinite number of scattering paths, of which the one shown in Fig. 6.1 is just a single example. For fixed  $\mathbf{r}_1 = \mathbf{r}$ ,  $\mathbf{r}_n = \mathbf{r}'$ , and  $\mathbf{k}_i$ ,  $\mathbf{k}_f$ ,  $A_0$  and  $B_0$  should be

$$A_0 = \exp[\mathrm{i}\mathbf{k}_{\mathrm{i}} \cdot (\mathbf{r} - \mathbf{R}_0) + \mathrm{i}\mathbf{k}_{\mathrm{f}} \cdot (\mathbf{r}_0 - \mathbf{r}')] \sum_{\{\alpha\}} A_\alpha, \qquad (6.4)$$

$$B_0 = \exp[\mathrm{i}\mathbf{k}_{\mathrm{i}} \cdot (\mathbf{r}' - \mathbf{R}_0) + \mathrm{i}\mathbf{k}_{\mathrm{f}} \cdot (\mathbf{r}_0 - \mathbf{r})] \sum_{\{\beta\}} B_\beta, \qquad (6.5)$$

where  $A_i$  is set equal to 1, and  $\alpha$ ,  $\beta$  are the indices for the scattering paths. Under a certain assumption, it is possible to show that there is a one-to-one correspondence between the terms of the two sums,  $\Sigma_{\alpha}A_{\alpha}$  and  $\Sigma_{\beta}B_{\beta}$ . This correspondence is already clear in the choice of ray paths A and B shown in Fig. 6.1. Namely, for every path  $\alpha$ , it is possible to find a *time-reversed* path  $\beta$  which has exactly the reverse order of scattering to  $\alpha$ . Also, for every path  $\beta$  one can find a corresponding path  $\alpha$  with the reverse order. Therefore, time reversal defines a one-to-one mapping between the paths in the set  $\{\alpha\}$  and those in the set  $\{\beta\}$ . As a result, *if the system is assumed to be invariant under time reversal*, then

$$\sum_{\{\alpha\}} A_{\alpha} = \sum_{\{\beta\}} B_{\beta},$$

and  $A_0/B_0$  is given exactly by (6.3). The assumption of time reversal invariance requires, for example, that there be no uniform velocity for the scatterers

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so that the time-reversed path sees the same scatterers' positions. But since the wave speed is usually fairly high, the uniform velocity has to be comparable to that in order to make any appreciable difference. A more serious requirement is the absence of magnetic field in electronic systems, since the vector potential associated with the magnetic field can give different phases to a scattering path and its time-reversed counterpart. When the time reversal invariance is broken, the coherence effect in the backscattering direction is not totally destroyed. It is only diminished in proportion to the degree of time reversal symmetry breaking.

It is interesting to note that the coherent backscattering effect was "discovered" independently in two different fields: as the basis of weak localization of electrons in dirty metals in solid-state physics (Langer and Neal 1966) and in optics as a reflection enhancement effect for electromagnetic wave propagation in a turbulent atmosphere (deWolf 1971). The word "discovered" was put in quotation marks to mean within the realm of recent literature. In any case, the parallel discoveries do reflect the common wave character of the coherent backscattering effect, regardless of whether the wave is quantum or classical in nature. This similarity, however, does not carry over to the ease of actual demonstration. In the quantum case, the difficulty in observing an electronic wave function means that there are so far only indirect demonstrations of the effect. But in the case of the electromagnetic wave there have been several experiments involving laser light reflection from disordered samples in which the coherent backscattering effect was directly observed and quantitatively measured (Tsang and Ishimaru 1984; van Albada and Lagendijk 1985; Wolf and Maret 1985). Below we specialize to the classical scalar wave case for a more explicit calculation of the effect in a fixed sample geometry.

## 6.3 Angular Profile of the Coherent Backscattering

Equation (6.2) implies that if a coherent beam is incident on a random medium, its backscattered intensity from the two rays A and B is given by

$$|A_{0} + B_{0}|^{2} = |A_{0}|^{2} + |B_{0}|^{2} + A_{0}B_{0}^{*} + A_{0}^{*}B_{0}$$
  
=  $|A_{0}|^{2} |1 + \exp[-i(\mathbf{k}_{i} + \mathbf{k}_{f}) \cdot (\mathbf{r} - \mathbf{r}')]|^{2}$   
=  $2 \{1 + \cos[(\mathbf{k}_{i} + \mathbf{k}_{f}) \cdot (\mathbf{r} - \mathbf{r}')]\} |A_{0}|^{2}.$  (6.6)

Inside the curly brackets, the constant term is due to incoherent scattering,  $|A_0|^2 + |B_0|^2$ , and the  $\cos[(\mathbf{k}_i + \mathbf{k}_f) \cdot (\mathbf{r} - \mathbf{r}')]$  term is due to the interference terms,  $A_0B_0^* + A_0^*B_0$ . It is seen that the coherence decreases as  $\mathbf{k}_f$  moves away from  $-\mathbf{k}_i$  and as  $|\mathbf{r} - \mathbf{r}'|$  becomes large. At its maximum, i.e.,  $\mathbf{k}_i + \mathbf{k}_f = 0$ , the coherence enhances the backscattering intensity by a factor of 2.

## 6.3.1 Single Configuration and Configurational Averaging

Equation (6.6) suggests a new way of looking at the reflection intensity. For three points,  $\mathbf{r}$ ,  $\mathbf{r}'$ , and  $\mathbf{r}''$ , we have pairs of rays  $A_0(\mathbf{r} \to \mathbf{r}')$  and  $B_0(\mathbf{r}' \to \mathbf{r})$ ,  $C_0(\mathbf{r} \to \mathbf{r}'')$  and  $D_0(\mathbf{r}'' \to \mathbf{r})$ , plus  $E_0(\mathbf{r}' \to \mathbf{r}'')$  and  $F_0(\mathbf{r}'' \to \mathbf{r}')$ . The net intensity in the direction of  $\mathbf{k}_{\rm f}$  is obtained by the addition of the six ray amplitudes, squared. That means

$$|A_{0} + B_{0} + C_{0} + D_{0} + E_{0} + F_{0}|^{2} = |A_{0} + B_{0}|^{2} + |C_{0} + D_{0}|^{2} + |E_{0} + F_{0}|^{2} + (A_{0} + B_{0}) (C_{0} + D_{0})^{*} + (A_{0} + B_{0})^{*} (C_{0} + D_{0}) + (A_{0} + B_{0})^{*} (E_{0} + F_{0}) + (A_{0} + B_{0}) (E_{0} + F_{0})^{*} + (C_{0} + D_{0})^{*} (E_{0} + F_{0}) + (C_{0} + D_{0}) (E_{0} + F_{0})^{*}.$$

There are two types of terms. The first type consists of square of the paired amplitudes, each giving a result similar to (6.6). Then there are the interference terms between the paired amplitudes, which would yield a reflected intensity profile that is particular to the configuration. In other words, the phase between any two paired amplitudes, e.g.,  $(C_0 + D_0)$  and  $(E_0 + F_0)$ , is expected to vary randomly from one configuration to the next because there is no reason for the scattering paths of two pairs to have any fixed relationship. Therefore, if configurational averaging is applied, only the first type of terms is expected to remain, whereas the second type would vanish. This configurational averaging effect has been demonstrated experimentally (Etemad et al. 1986). By using light scattering from fixed scatterer configurations, it was found that for each configuration, there were may intensity peaks occurring at arbitrary reflection angles. This is exactly the speckle pattern mentioned earlier (in reflection in this case). However, after configurational averaging only the backscattering peak emerged, and all the other peaks vanished. The reason is that while there may be a large peak at  $\theta \neq 0$  for a given configuration (where  $\theta$  denotes the angle of deviation from the backscattering direction), the intensity at  $\theta$  is likely to be small for most other configurations. At  $\theta = 0$ , however, the backscattering peak may be smaller than the other peaks for any given configuration, but it is always present in every configuration. Therefore, upon configurational averaging, whereas the other peaks diminish, the coherent backscattering peak persists. In other words, the speckle pattern gets smoothed over at the level of mean intensity, but the coherent backscattering peak does not. This fact implies that if it is the configurationally averaged result one desires (which may be realized physically if the scatterers are moving randomly, such as in a colloid due to the Brownian motion, and the measured quantity is the time-averaged intensity), then the sum of the amplitudes squared may be replaced by the sum of the *paired intensities*, where each pair sees an enhancement factor  $[1 + \cos(\mathbf{k}_i + \mathbf{k}_f) \cdot (\mathbf{r} - \mathbf{r}')]$ .

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## 6.3.2 Counting the Scattering Paths

The simplification of looking only at paired intensities enables us to calculate the coherent backscattering angular profile for light reflected from a slab of disordered scattering medium, extending from z = 0 to z = L. Since the coherent enhancement depends on  $\mathbf{r} - \mathbf{r}'$ , the intensity addition has to take into account the amount  $P(\mathbf{r}, \mathbf{r}')$  of the source intensity that a ray starting at  $\mathbf{r}$  would end up at  $\mathbf{r}'$ . The quantity  $P(\mathbf{r}, \mathbf{r}')$  may be approximated by the solution to the diffusion equation, just as in the case of DWS analyzed in the last chapter. The difference here is that  $P(\mathbf{r}, \mathbf{r}')$  is not concerned with a given path *length*. Rather, it is the sum of *all* paths that end at  $\mathbf{r}'$  and begin at  $\mathbf{r}$ , both points at or close to the surface of the sample. Since path length is proportional to time, the solution needed here is the time-dependent solution [(5.135)] *integrated over time*, i.e., the time-independent, *static* solutions to the diffusion equation. If the boundary conditions are ignored for the moment, then

$$P(\mathbf{r}, \mathbf{r}') = \int_0^\infty \frac{\exp\left[-(\mathbf{r} - \mathbf{r}')^2 v_t / 4D^{(B)}s\right]}{(4\pi D^{(B)}s/v_t)^{3/2}} \,\mathrm{d}s$$
$$= \frac{v_t}{4\pi D^{(B)} |\mathbf{r} - \mathbf{r}'|}.$$
(6.7)

 $P(\mathbf{r}, \mathbf{r}')$  is recognized to be the solution to the 3D Laplace equation (diffusion equation with the  $\partial/\partial t$  term dropped) with a point source.

#### 6.3.3 Backscattering from a Random Half Space

Because of the requirement that the source be a *diffusive* wave source, **r** cannot be located on the sample surface but must be on a plane  $z = z_0 \cong l^*$  inside the sample where the incident wave is likely to have experienced at least one scattering. In order to get a physical feeling first, let us simplify to the case of  $L \to \infty$ , with the boundary conditions at z = 0 [(5.144b)] given by

$$P - C\frac{\partial P}{\partial z} = 0,$$

where  $C = (2l^*/3)(1+\rho)/(1-\rho)$ ,  $\rho$  being the reflection coefficient at the sample surface. C is also called the extrapolation length because if the solution is linearly extrapolated in the negative z direction, then P vanishes at z = -C. An approximate solution can be obtained from (6.7) by the method of images:

$$P(\mathbf{r}, \mathbf{r}') = \frac{v_t}{4\pi D^{(B)}} \left\{ \frac{1}{\left[ (x - x')^2 + (y - y')^2 + (z_0 - z')^2 \right]^{1/2}} - \frac{1}{\left[ (x - x')^2 + (y - y')^2 + (z' + z_0 + 2C)^2 \right]^{1/2}} \right\}, \quad (6.8)$$

where it is seen that at z' = -C the condition of P = 0 is guaranteed by the subtraction of an "image" source term located at  $\mathbf{r} = (x, y, -2C - z_0)$ . Since  $\mathbf{r}'$  represents the position of the last scattering, it is reasonable to assume that  $z' \cong z_0$ . Then

$$P(\mathbf{r}, \mathbf{r}') = \frac{3}{4\pi l^*} \left\{ \frac{1}{r_{\parallel}} - \frac{1}{\left[r_{\parallel}^2 + 4(z_0 + C)^2\right]^{1/2}} \right\},\tag{6.9}$$

where  $r_{||} = [(x - x')^2 + (y - y')^2]^{1/2}$ . By adding the paired intensities for those paths that start at **r** and end at all possible **r**', the total scattered intensity in the direction of **k**<sub>f</sub> is obtained as

$$\frac{I(\mathbf{k}_{i}, \mathbf{k}_{f})}{I_{\text{inc.}}} \cong \frac{3}{(4\pi)^{2}l^{*}} \int d\mathbf{r}_{\parallel} \left\{ \frac{1}{r_{\parallel}} - \frac{1}{\left[r_{\parallel}^{2} + 4(z_{0} + C)^{2}\right]^{1/2}} \right\} \times \left\{ 1 + \cos[(\mathbf{k}_{i} + \mathbf{k}_{f}) \cdot (\mathbf{r} - \mathbf{r}')] \right\}.$$
(6.10)

The result is noted to be independent of the source position  $\mathbf{r}$ . Here the extra factor of  $1/4\pi$  gives the fraction of intensity scattered into the direction of  $\mathbf{k}_{\rm f}$  at  $\mathbf{r}'$ ,  $I_{\rm inc.}$  is the incident intensity, and  $\mathbf{r}_{\parallel} = (x - x', y - y')$ . In the solution to Problem 6.1 it is shown that the integral yields

$$\frac{I(\mathbf{k}_{i},\mathbf{k}_{f})}{I_{\text{inc.}}} \cong \frac{3(z_{0}+C)}{4\pi l^{*}} \left\{ 1 + \frac{1 - \exp[-2|\mathbf{k}_{i} + \mathbf{k}_{f}|(z_{0}+C)]}{2|\mathbf{k}_{i} + \mathbf{k}_{f}|(z_{0}+C)} \right\}.$$
 (6.11)

#### 6.3.4 Features of the Coherent Backscattering Peak

Several interesting features of (6.11) should be noted. First,  $z_0+C$  is just some constant times  $l^*$ . Because  $l^*|\mathbf{k}_i|$  is generally  $\gg 1$ , it follows that the factor  $2|\mathbf{k}_i + \mathbf{k}_f|(z_0 + C)$  can vary from 0 (at the backscattering direction  $\mathbf{k}_f = -\mathbf{k}_i$ ) to some value that is  $\gg 1$ . Thus (6.11) predicts an enhancement factor of 2 at the backscattering direction over the incoherent baseline, as anticipated.

Second, the coherent enhancement exists only over a fairly narrow angular cone whose width may be evaluated from (6.11). If the width is defined by the condition

$$2|\mathbf{k}_{i} + \mathbf{k}_{f}|(z_{0} + C) = 1, \qquad (6.12)$$

and the deviation of  $\mathbf{k}_{\rm f}$  from  $-\mathbf{k}_{\rm i}$  is denoted by an angle  $\theta$  so that  $\mathbf{k}_{\rm f} + \mathbf{k}_{\rm i} \approx 2\pi\theta/\lambda$ , where  $\lambda$  is the wavelength, then the condition (6.12) implies a backscattering angular width of

$$\Delta \theta \cong \frac{\lambda}{4\pi l^*} \left( 1 + \frac{2}{3} \; \frac{1+\rho}{1-\rho} \right)^{-1},\tag{6.13}$$

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where it is assumed that  $z_0 = l^*$ . The inverse relationship between  $\Delta \theta$  and  $l^*$  is not surprising, because once  $\mathbf{k}_{\mathrm{f}} \neq -\mathbf{k}_{\mathrm{i}}$ , the dephasing becomes a function of separation between  $\mathbf{r}$  and  $\mathbf{r'}$ , and the (incoherent) diffusive behavior is expected to be dominant for  $|\mathbf{r} - \mathbf{r'}|$  larger than  $l^*$ . Thus one can add amplitudes coherently only for  $|\mathbf{r}_{\parallel} - \mathbf{r'}_{\parallel}| \leq l^*$ . The path difference between two rays emerging at an angle  $\theta$  from the backscattering direction but with the separation  $|\mathbf{r}_{\parallel} - \mathbf{r'}_{\parallel}| \simeq l^*$  is roughly  $\theta l^*$ . Since constructive interference may be expected only when  $\theta l^* \leq \lambda$ , it follows that  $\Delta \theta l^* \cong \lambda$  should be the condition for the angular width  $\Delta \theta$  of the coherent enhancement cone, as given by (6.13). In general, for light scattering experiments  $\Delta \theta$  is on the order of a fraction of a degree. The small magnitude of  $\Delta \theta$  could be the reason that the definitive verification of the effect is only a relatively recent event.

The third feature to be noted from (6.11) is that at small  $\theta$ , the decrease from the peak maximum is linear in  $\theta$ , independent of the azimuthal angle. Therefore, the enhancement angular profile has a cusp, i.e., a derivative discontinuity, at  $\theta = 0$ .

## 6.3.5 Backscattering from a Finite Thickness Slab

To make the calculation more accurate means relaxing the condition that  $\mathbf{r}$  and  $\mathbf{r}'$  must be at  $z = z_0$ , and using a general solution  $P(\mathbf{r}, \mathbf{r}')$  for finite L. In the solution to Problem 6.2 it is shown that if the boundary conditions of (5.144) are approximated by the conditions that P vanishes at z = -C and z = L + C, then

$$P(\mathbf{r}, \mathbf{r}') = \frac{3}{\pi l^* (L+2C)} \sum_{n=1}^{\infty} \sin \frac{n\pi (z+C)}{L+2C} \sin \frac{n\pi (z'+C)}{L+2C} \times K_0 \left(\frac{n\pi}{L+2C} r_{\parallel}\right),$$
(6.14)

where  $K_0$  denotes the zeroth order modified Bessel function of the second kind. Since **r** and **r'** are the sites of the first scattering and the last scattering, respectively, the probability for this to happen is given by

$$\frac{1}{l^*} \exp\left(\frac{-z}{l^*}\right)$$

for  $\mathbf{r}$  being the first scattering position under normal incidence, and

$$\frac{1}{l^*} \exp\left(\frac{-z'}{l^*}\right)$$

for having no further scattering after  $\mathbf{r}'$ . Here we have assumed  $\mathbf{k}_i$ ,  $\mathbf{k}_f$  to be close to the normal to the front surface so that the distance projection factor  $\cos \theta \approx 1$ . By combining all the factors, including  $1/4\pi$  for the fraction scattered into the direction  $\mathbf{k}_f$ , the final result is

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$$\frac{I(\mathbf{k}_{i},\mathbf{k}_{f})}{I_{inc.}} = \frac{1}{4\pi(l^{*})^{2}} \int d\mathbf{r}_{\parallel} \int_{0}^{L} dz \int_{0}^{L} dz' \exp\left(\frac{-z}{l^{*}}\right) \times [1 + \cos(\mathbf{k}_{i} + \mathbf{k}_{f}) \cdot (\mathbf{r} - \mathbf{r}')] P(\mathbf{r},\mathbf{r}') \exp\left(\frac{-z'}{l^{*}}\right). \quad (6.15)$$

Here  $P(\mathbf{r}, \mathbf{r}')$  is understood to be given by (6.14).

## 6.4 Sample Size (Path Length) Dependence

The coherent backscattering angular profiles for  $L = 1.4l^*$ ,  $4l^*$ , and  $100l^*$ , calculated from (6.14) and (6.15), are shown in Fig. 6.2. It is seen that in the case of  $L = 100l^*$ , the calculated angular profile indeed has a sharp peak which decays quickly to the incoherent background, which is half the peak value. Here the incoherent background value 1 is noted to correspond to that for a semi-infinite sample. Figure 6.3 shows a comparison between a theoretically calculated profile and the experimental result for a semi-infinite sample of colloidal suspension (Akkermans et al. 1986), where the random movement of the scattering particles and the time average of the measured intensity assures that the measured result is configurationally averaged. Here the value  $l^* =$  $19\,\mu$ m was measured independently, so the fit has no adjustable parameter (apart from making the baseline agree at the value 1). The fact that the peak



Fig. 6.2. Calculated angular profiles of the coherent backscattering peak for three sample thicknesses  $L/l^* = 1.4$ , 4, and 100.  $\theta = 0$  is the backscattering direction. The wavelength  $\lambda$  is assumed to be that of the He–Ne laser light, 633 nm, and  $l^*$  is taken to be 11 µm. The coherent peak is seen to decrease in magnitude as the sample thickness decreases. The baseline of the profile, denoted as 1, represents the incoherent component of the scattering from an  $L/l^* = \infty$  sample. The incoherent baselines of the finite-thickness samples are exactly a factor of 2 lower than their respective coherent backscattering peaks.



Fig. 6.3. Comparison of the experimentally measured profile (*wiggly line*) and the theoretical prediction. The sample used in the experiment is a colloidal suspension with  $l^* = 19 \,\mu\text{m}$  and a wavelength  $\lambda = 515 \,\text{nm}$ . The theoretical profile shown in the inset is convolved with a known instrumental resolution function to obtain the profile shown by the smooth *solid line*. The incoherent baseline of the profile is normalized to 1. The figure is from Akkermans et al. (1986).

value in Fig. 6.3 is not 2 is due to the convolution of the theoretical result with a known instrumental resolution profile. The agreement is seen to be very good indeed.

For  $L = 1.4l^*$  and  $4l^*$ , the theoretical profiles for the coherent enhancement are seen to have lower peak values when they are normalized by the incoherent background for a semi-infinite sample. This is a consequence of the path length dependence, or sample size dependence, of the coherent backscattering effect. That is, as the sample becomes thinner, the long backscattering paths are cut off, so they do not contribute to the effect. To have a better feel for this path length dependence, let us refer to (6.7), where the static diffusion solution is expressed in terms of an integral over the solution for a given path length s. The integrand is a Gaussian where  $|\mathbf{r} - \mathbf{r}'|$  decays over the distance  $\sim \sqrt{l^*s}$ (which can be seen by expressing  $D^{(B)} = v_t l^*/3$ ). The decay length is the same for the image solution, (6.9), if it were also expressed as the superposition of solutions for different path lengths. By substituting the integral form of the point source solution, (6.7), into (6.10), there would be an integrand involving the product

$$\exp\left[-3(\mathbf{r}-\mathbf{r}')^2/4l^*s\right]\cos[(\mathbf{k}_{i}+\mathbf{k}_{f})\cdot(\mathbf{r}-\mathbf{r}')].$$

The path length dependence of the coherent enhancement effect is clear from this product. If  $(\mathbf{k}_i + \mathbf{k}_f)$  is expressed as  $2\pi\theta/\lambda$ , then by using the decay length of  $|\mathbf{r} - \mathbf{r}'|$ ,  $\sqrt{l^*s}$ , from the Gaussian factor, the condition for coherence is seen to be

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$$\frac{2\pi\theta}{\lambda}\sqrt{l^*s} \le \pi,\tag{6.16}$$

or

$$s \le \frac{\lambda^2}{4\theta^2 l^*}.\tag{6.17}$$

Equation (6.17) tells us that at any given angle  $\theta$  away from the backscattering direction, there is an upper limit to the path length which can contribute to the coherent backscattering effect. As  $\theta$  approaches zero this upper limit diverges. The angular profiles shown in Fig. 6.2 are entirely consistent with this interpretation. The thinner samples are seen to have lower peak values, but far away from the peak the profiles remain unchanged. The angles below which the profile of the finite-thickness samples deviate from the semi-infinite case may be estimated from (6.17). Since the number of scatterings for a diffusive ray to travel a net distance L is  $3(L/l^*)^2$ , where the factor 3 arises from the three directions, it follows that the path lengths being cut off are those with  $s \geq 3(L/l^*)^2 l^* = 3L^2/l^*$ . From (6.17) it follows that a sample of thickness L should have a backscattering profile which differs from the semi-infinite case at

$$\theta < \frac{\lambda}{2\sqrt{3}L}$$

For  $\lambda = 633 \,\mathrm{nm}$  (He–Ne laser) and  $L = 50 \,\mu\mathrm{m}$ , the difference occurs for  $\theta \leq 2.5 \,\mathrm{mrad}$  from the backscattering direction. In Fig. 6.4 the measured coherent backscattering profiles (Etemad et al. 1987) are plotted for four sample thicknesses. They correspond remarkably well with what are expected theoretically.



Fig. 6.4. Experimentally measured angular profiles of the coherent backscattering peak for four sample thicknesses (denoted by W in the figure) of L = 15, 45, 100, and 1,000 µm. The samples used are colloidal suspensions with  $l^* = 11 \,\mu\text{m}$ , and the light wavelength is 633 nm (He–Ne laser). The three cases of L = 15, 45 µm, and 1,000 µm should be compared with the theoretical profiles shown in Fig. 6.2 with similar  $L/l^*$  ratios. The figure is from Etemad et al. (1987).

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An important point to be noted is that the incoherent background for a thin sample is always decreased by the *same* factor as the coherent peak, so that the ratio 2 between the coherent peak and the incoherent background is preserved for any sample thickness. As noted before, the decrease in the coherent peak seen in Fig. 6.4 is relative to the incoherent background of the semi-infinite sample case.

## 6.4.1 Effect of Absorption

Another means of eliminating the long scattering paths is through absorption, which can be characterized by an inelastic scattering length  $l_{\rm in}$ . However, there is a subtle difference between the sample size limitation and the limitation through absorption because absorption is a probabilistic process, whereas sample size limitation is deterministic. Therefore, for a finite sample size L the paths with lengths shorter than L are not affected. But if the sample has absorption, there is a finite probability for all paths to be affected, regardless of their lengths (of course, paths longer than  $l_{\rm in}$  are expected to be affected more than those shorter than  $l_{\rm in}$ ). However, the difference between the two mechanisms is actually fairly small, and the two are roughly equivalent. If we regard paths with  $s > l_{\rm in}$  as cut off in the absorption case and those with  $s > 3L^2/l^*$  as cut off in the finite sample thickness case, then a conversion between the two mechanisms may be achieved by using the relation

$$L = \sqrt{l_{\rm in} l^*/3}.$$
 (6.18)

The effect of inelastic scattering is addressed with more detail in Chap. 10.

The sample size dependence of the coherent backscattering effect foreshadows many interesting developments, including the scaling theory of localization. Together with the requirement of time reversal invariance (and the consequent degradation of the effect when the requirement is not satisfied), these two aspects of the coherent backscattering effect play an important role in providing indirect evidence of weak localization in dirty electronic systems. At the same time, they also provide the means for manipulating the wave localization effect through external control. These points are addressed in Chapters 10 and 11.

## Problems and Solutions

**Problem 6.1** Show that the integral of (6.10) gives the result of (6.11).

Solution The integral to be performed may be written as

$$\frac{3}{(4\pi)^2 l^*} \left\{ \int_0^{2\pi} \mathrm{d}\phi \int_0^\infty \left( 1 - \frac{r}{\sqrt{r^2 + a^2}} \right) \mathrm{d}r + \int_0^\infty \int_0^{2\pi} \cos(\mathbf{q} \cdot \mathbf{r}) \right. \\ \left. \times \left[ 1 - \frac{r}{\sqrt{r^2 + a^2}} \right] d\phi dr \right\},$$

where  $\mathbf{q} = \mathbf{k}_{i} + \mathbf{k}_{f}$ ,  $a = 2(z_{0} + C)$ , the two-dimensional  $d\mathbf{r}_{\parallel}$  integral is written as  $d\phi r dr$ , where the subscript  $\parallel$  on r has been dropped, and  $\phi$  is the angle between  $\mathbf{q}$  and  $\mathbf{r}$ . For the first term inside the bracket, a change of variable

$$r = a \tan \theta$$

converts the r integration into a simple  $\theta$  integration:

First term = 
$$2\pi a \int_0^{2\pi} (\sec^2 \theta - \sec \theta \tan \theta) d\theta$$
  
=  $2\pi a \left[ \frac{\sin \theta - 1}{\cos \theta} \right]_0^{\pi/2}$   
=  $2\pi a.$  (P6.1)

For the second term inside the bracket, the  $d\phi$  integration yields  $2\pi J_0(qr)$ . As a result,

Second term = 
$$2\pi \int_0^\infty J_0(qr) dr - 2\pi \int_0^\infty J_0(qr) \frac{r}{\sqrt{r^2 + a^2}} dr.$$
 (P6.2)

Since the Laplace transform of  $J_0(t)$  is  $1/\sqrt{1+s^2}$ , it follows that the first term of (P6.2) can be obtained from the Laplace transform with s = 0, which yields  $2\pi/q$ . The second term of (P6.2) can be obtained from a table of Hankel transforms. It has the value

$$\frac{2\pi}{q}\exp(-aq)$$

Collecting terms, we get

$$\frac{3}{(4\pi)^2 l^*} 2\pi a \left[ 1 + \frac{1 - \exp(-aq)}{aq} \right],$$

which is exactly the result shown by (6.11). The value of the expression inside the square brackets approaches 2 as  $aq \to 0$  and approaches 1 as  $aq \to \infty$ .

Problem 6.2 Find the solution to the equation

$$\nabla_{\mathbf{r}}^2 P(\mathbf{r}, \mathbf{r}') = -\frac{v_t}{D^{(B)}} \delta(\mathbf{r} - \mathbf{r}'),$$

with the boundary conditions that  $P(\mathbf{r}, \mathbf{r}') = 0$  at z = -C, L + C.

**Solution** Without the boundary conditions,  $P(\mathbf{r}, \mathbf{r}')$  is given by (6.7). With the boundary conditions, it is advantageous to solve the problem in cylindrical coordinates, where the equation may be cast in the form

$$\nabla_{\mathbf{r}}^2 P(\mathbf{r}, \mathbf{r}') = -\frac{v_t}{D^{(\mathrm{B})}} \frac{1}{r_{\parallel}} \delta(r_{\parallel} - r'_{\parallel}) \delta(\phi - \phi') \delta(z - z').$$
(P6.3)

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The boundary conditions at z = -C, L + C suggest that  $\delta(z - z')$  should be expanded in terms of the complete orthonormal functions of the form

$$\sqrt{\frac{2}{L+2C}}\sin\frac{n\pi(z+C)}{L+2C}, \qquad n=1,2,3,....$$

In other words,

$$\delta(z - z') = \frac{2}{L + 2C} \sum_{n=1}^{\infty} \sin \frac{n\pi(z + C)}{L + 2C} \sin \frac{n\pi(z' + C)}{L + 2C}.$$
 (P6.4)

Equation (P6.4) is easily verified by integration with  $\sin m\pi(z'+C)/(L+2C)$ , for example. It is a statement of the completeness of the expansion basis functions. Similarly,

$$\delta(\phi - \phi') = \frac{1}{2\pi} \sum_{m=-\infty}^{\infty} \exp[im(\phi - \phi')].$$
 (P6.5)

By writing

$$P(\mathbf{r}, \mathbf{r}') = \frac{v_t}{D^{(B)}\pi(L+2C)} \sum_{m=-\infty}^{\infty} \sum_{n=1}^{\infty} \sin \frac{n\pi(z+C)}{L+2C}$$
$$\times \sin \frac{n\pi(z'+C)}{L+2C} \exp[im(\phi-\phi')]g_m(r_{\parallel}), \tag{P6.6}$$

and substituting it into (P6.3), with  $\delta(z-z')$  and  $\delta(\phi-\phi')$  on the right-hand side expanded in accordance with (P6.4) and (P6.5), the resulting equation for  $g_m$  is

$$\frac{\partial^2 g_m}{\partial r_{\parallel}^2} + \frac{1}{r_{\parallel}} \frac{\partial g_m}{\partial r_{\parallel}} - \left[\frac{m^2}{r_{\parallel}^2} + \frac{n^2 \pi^2}{(L+2C)^2}\right] g_m = -\frac{1}{r_{\parallel}} \delta(r_{\parallel} - r_{\parallel}').$$
(P6.7)

This is the equation for modified Bessel function of order m. One can verify that the solution is given by

$$g_m = I_m \left(\frac{n\pi}{L+2C}r_{\parallel}\right) K_m \left(\frac{n\pi}{L+2C}r'_{\parallel}\right) \quad \text{if } r'_{\parallel} > r_{\parallel}$$
$$= I_m \left(\frac{n\pi}{L+2C}r'_{\parallel}\right) K_m \left(\frac{n\pi}{L+2C}r_{\parallel}\right) \quad \text{if } r_{\parallel} > r'_{\parallel}. \quad (P6.8)$$

Here  $I_m$  and  $K_m$  are the modified Bessel functions of the first kind and the second kind, respectively. By letting  $r'_{\parallel} = 0$ , the problem becomes isotropic

with respect to  $\phi$ , so only the m = 0 term remains, and  $I_0(0) = 1$ . It follows that

$$P(\mathbf{r}, \mathbf{r}') = \frac{3}{\pi l^* (L+2C)} \sum_{n=1}^{\infty} \sin \frac{n\pi (z+C)}{L+2C} \times \sin \frac{n\pi (z'+C)}{L+2C} K_0 \left(\frac{n\pi}{L+2C} r_{\parallel}\right),$$

which is exactly (6.14).

# **Renormalized Diffusion**

# 7.1 Coherent Backscattering Effect in the Diagrammatic Representation

In the last chapter the coherent backscattering effect was seen as an enhancement of scattered intensity in the backward direction. The purpose of this chapter is to examine the *dynamic* implications of this effect in the context of diffusive transport behavior as described in Chap. 5. To quantify such implications, it is necessary to represent the coherent backscattering effect in terms of the Green functions. A first step in this direction is the association of the ray paths in Fig. 6.1 with products of Green functions and  $\bar{t}_i$ ' s. For example, for the ray path A,

$$A_{0} = A_{i} G_{0}^{+}(\omega, \mathbf{r}_{1}, \mathbf{R}_{0}) \bar{t}_{1}^{+} G_{e}^{+}(\omega, \mathbf{r}_{2}, \mathbf{r}_{1}) \bar{t}_{2}^{+} G_{e}^{+}(\omega, \mathbf{r}_{3}, \mathbf{r}_{2}) \cdots \times \bar{t}_{n}^{+} G_{0}^{+}(\omega, \mathbf{r}_{0}, \mathbf{r}_{n}).$$
(7.1)

This association is based on the physical picture in which each scatterer acts as a secondary wave source, so that the amplitude of reaching the next scatterer is given by the Green function. Since each scattering occurs in an environment of other scatterers,  $G_e^+$  is used so that the presence of the random environment is taken into account in an averaged sense. Similarly,  $\bar{t}_i^+$  gives the scattering amplitude for the *i*th scatter in an effective medium. For each Green function,  $G_e^+(\omega, \mathbf{r}_j, \mathbf{r}_i)$ ,  $\mathbf{r}_i$  is noted to be the source point and  $\mathbf{r}_j$  the detection point. Therefore, the line that represents the Green function in a diagram should have an arrow that points from  $\mathbf{r}_i$  toward  $\mathbf{r}_j$ .  $\mathbf{r}_0(\mathbf{R}_0)$  is defined to be the observation (source) point far from the sample. Comparison with (6.1) shows the factor  $G_A$  in that equation to be the product of  $\bar{t}_1^+ \cdots \bar{t}_n^+$  with the magnitude part of the Green functions.

Since we are interested in the behavior of intensity transport, it is necessary to examine the quantity  $A_0A_0^*$ , where

$$A_{0}^{*} = G_{0}^{-}(\omega, \mathbf{r}_{n}, \mathbf{r}_{0}) \bar{t}_{n}^{-} \cdots \times G_{e}^{-}(\omega, \mathbf{r}_{2}, \mathbf{r}_{3}) \bar{t}_{2}^{-} G_{e}^{-}(\omega, \mathbf{r}_{1}, \mathbf{r}_{2}) \bar{t}_{1}^{-} G_{0}^{-}(\omega, \mathbf{R}_{0}, \mathbf{r}_{1}) A_{i}^{*}.$$
 (7.2)



**Fig. 7.1.** Representation of  $A_0 A_0^*$  (or  $B_0 B_0^*$ ) as a ladder diagram.

Here we have used the fact that  $[G^+(\omega, \mathbf{r}_i, \mathbf{r}_j)]^* = G^-(\omega, \mathbf{r}_j, \mathbf{r}_i)$  and  $(\bar{t}_i^+)^* =$  $\bar{t}_i^-$ . The product  $A_0 A_0^*$  is thus the outer product of the right-hand sides of (7.1) and (7.2). In accordance with the rules of drawing diagrams from outer products as explained in Chap. 5, the diagram corresponding to  $A_0 A_0^*$ , for the part between the first and the nth scatterer, is given by Fig. 7.1, where we have made the slight change of replacing  $\bar{t}_i^+ \otimes \bar{t}_i^-$  by the CPA vertex  $U_i^{(B)}$ . The net effect of this replacement is recalled to be the removal of the restriction on the indices of the scatterers. In addition, the dynamics of the multiply scattered wave is introduced by letting  $A_0$  and  $A_0^*$  have frequencies  $\omega^+ = \omega + (\Delta \omega/2)$ and  $\omega^- = \omega - (\Delta \omega/2)$ , respectively. The Green functions are also labeled by their momenta, with the beginning and ending momenta  $\mathbf{k}_{i}^{+} = \mathbf{k}_{i} + (\Delta \mathbf{k}/2),$  $\mathbf{k}_{\mathrm{f}}^{+} = \mathbf{k}_{\mathrm{f}} + (\Delta \mathbf{k}/2) \text{ for } A_{0}, \text{ and } \mathbf{k}_{\mathrm{i}}^{-} = \mathbf{k}_{\mathrm{i}} - (\Delta \mathbf{k}/2), \mathbf{k}_{\mathrm{f}}^{-} = \mathbf{k}_{\mathrm{f}} - (\Delta \mathbf{k}/2) \text{ for } A_{0}^{*}.$ It should be noted that although the direction of the arrows on the  $A_0^*$  path is opposite to that of  $A_0$ , yet an advanced Green function is also time reversed compared with a retarded Green function. As a result, despite the arrows directions, the momenta of the + and - channels are the same except for the sign of  $\Delta \mathbf{k}$ , which is introduced here to delineate the dynamics of diffusive waves. Figure 7.1 is immediately recognized to be one of the ladder diagrams responsible for the wave diffusion constant evaluation given in Chap. 5. Since  $A_0 A_0^*$  and  $B_0 B_0^*$  constitute the incoherent scattering background as seen in the last chapter, we have thus established a correspondence between the incoherent scattering and the ladder diagrams.

For the coherent interference part of the scattering, it is necessary to look at terms such as  $A_0B_0^*$  and  $B_0A_0^*$ . From Fig. 6.1, it is straightforward to write down for the ray path B,

$$B_{0} = A_{i}G_{0}^{+}(\omega,\mathbf{r}_{n},\mathbf{R}_{0})\bar{t}_{n}^{+}G_{e}^{+}(\omega,\mathbf{r}_{n-1},\mathbf{r}_{n})\bar{t}_{n-1}^{+}G_{e}^{+}(\omega,\mathbf{r}_{n-2},\mathbf{r}_{n-1})$$

$$\times\cdots\bar{t}_{1}^{+}G_{0}^{+}(\omega,\mathbf{r}_{0},\mathbf{r}_{1}).$$
(7.3)

The product  $B_0A_0^*$  is shown in Fig. 7.2. The difference from Fig. 7.1 is clear. Because the order in which the scatterings occur in  $B_0$  is reversed from that in  $A_0^*$ , the dashed interaction lines that link the upper  $(G^+)$  line and the



**Fig. 7.2.** Representation of  $B_0 A_0^*$  (or  $A_0 B_0^*$ ) as a maximally crossed diagram.

lower  $(G^{-})$  line have to cross each other. Figure 7.2 represents one of the socalled maximally crossed diagrams. The evaluation of the maximally crossed diagrams and their dynamic implications is now the task at hand.

## 7.2 Evaluation of the Maximally Crossed Diagrams

To evaluate the maximally crossed (MC) diagrams, let us consider for the moment the simplest MC diagram shown in Fig. 7.3a for two scatterers. For the evaluation purpose, it makes no difference if we redraw it as in Fig. 7.3b, which can be obtained from 7.3a by regarding the dashed interaction lines as elastic rubber bands connected to two parallel broomsticks (the Green functions), and by turning the bottom broomstick  $180^{\circ}$  to untangle the two dashed lines. The only change, which is immaterial to the actual evaluation of the diagram, is that now the - input channel is on the right side and its output is on the left, reversed from the + channels. Because the order of the two scatterings is preserved and the momenta remain unchanged, diagram 7.3b is identical to the original 7.3a from which it is obtained. Figure 7.3b looks very much like a ladder diagram. The one difference from a proper ladder diagram is in the sign of the momenta in the - channel. In general, the sign of momenta is physically significant. For example, in the presence of a magnetic field two electrons traveling a loop in opposite directions can have different phases that depend on the vector potential (Schiff 1968). However, if the system is time reversal invariant, then reversing the momenta would yield the same physical state.

## 7.2.1 Momenta Transformations

Thus, by *assuming* the system to be time reversal invariant, the sign of momenta in the – channel may be reversed as shown in Fig. 7.3c, so that the – output and input channels are now also reversed back to the same order as the + channels. Now the diagram is almost like a proper ladder diagram.



Fig. 7.3. (a) An MC diagram with two scatterers. (b) Same as in (a), but with the bottom line turned 180°. This operation should have no effect on the evaluation of the diagram. (c) The sign of the momenta in the – channel is reversed from that of (b). This operation would leave the diagram the same only if the system is time-reversal invariant. (d) The momenta are relabeled in accordance with the rule  $\mathbf{k}_f + \mathbf{k}_i = \mathbf{Q}$ , and  $\mathbf{k}$ ,  $\mathbf{k}'$  as defined by (7.4). The result is a proper-looking ladder diagram.

To make it completely proper, it is noted that the difference in the + and – channel momenta, which is  $\Delta \mathbf{k}$  in the ladder diagrams, is now given by  $\mathbf{k}_i^+ - (-\mathbf{k}_f^-) = \mathbf{k}_i^+ + \mathbf{k}_f^- = [\mathbf{k}_i + (\Delta \mathbf{k}/2)] + [\mathbf{k}_f - (\Delta \mathbf{k}/2)] = \mathbf{k}_i + \mathbf{k}_f$ . Therefore, one can write

$$\mathbf{k}_{i}^{+} = \left(\frac{\mathbf{k}_{i} - \mathbf{k}_{f}}{2} + \frac{\Delta \mathbf{k}}{2}\right) + \frac{\mathbf{k}_{i} + \mathbf{k}_{f}}{2}, \tag{7.4a}$$

$$\mathbf{k}_{\rm f}^{-} = \left(\frac{\mathbf{k}_{\rm i} - \mathbf{k}_{\rm f}}{2} + \frac{\Delta \mathbf{k}}{2}\right) - \frac{\mathbf{k}_{\rm i} + \mathbf{k}_{\rm f}}{2},\tag{7.4b}$$

$$\mathbf{k}_{\rm f}^+ = \left(\frac{\mathbf{k}_{\rm f} - \mathbf{k}_{\rm i}}{2} + \frac{\Delta \mathbf{k}}{2}\right) + \frac{\mathbf{k}_{\rm i} + \mathbf{k}_{\rm f}}{2},\tag{7.4c}$$

$$\mathbf{k}_{i}^{-} = \left(\frac{\mathbf{k}_{f} - \mathbf{k}_{i}}{2} + \frac{\Delta \mathbf{k}}{2}\right) - \frac{\mathbf{k}_{i} + \mathbf{k}_{f}}{2}.$$
 (7.4d)

By relabeling

$$\begin{split} \frac{\mathbf{k}_i - \mathbf{k}_f}{2} + \frac{\Delta \mathbf{k}}{2} & \text{as } \mathbf{k}, \\ \frac{\mathbf{k}_f - \mathbf{k}_i}{2} + \frac{\Delta \mathbf{k}}{2} & \text{as } \mathbf{k}', \\ \mathbf{k}_i + \mathbf{k}_f & \text{as } \mathbf{Q}, \end{split}$$

the transformed diagram is now just the ladder diagram with the three  $\mathbf{k}$  variables transformed from  $\mathbf{k}_i$ ,  $\mathbf{k}_f$ ,  $\Delta \mathbf{k}$  to  $\mathbf{k}$ ,  $\mathbf{k}'$ , and  $\mathbf{Q}$ , respectively. The final transformed diagram is shown in Fig. 7.3d. The net outcome of this whole exercise may be summarized as follows. In a time reversal invariant system, an MC diagram can be converted into a ladder diagram with the provision that

$$\mathbf{Q} = \mathbf{k}_{\mathrm{i}} + \mathbf{k}_{\mathrm{f}} \tag{7.5}$$

plays the role of  $\Delta \mathbf{k}$  in the ladder diagrams. On the other hand, the initial and final momenta of the transformed ladder diagram are related by  $\mathbf{k} = -\mathbf{k}' + \Delta \mathbf{k}$ .

## 7.2.2 Summation of Maximally Crossed Diagrams

The MC diagrams modify the vertex function beyond its Boltzmann approximation. That is,

$$\mathbf{U} = \mathbf{U}^{(\mathrm{B})} + \mathbf{U}^{(\mathrm{MC})}.$$
(7.6)

We have seen from Chap. 5 and the discussion earlier that  $\mathbf{U}^{(B)}$  is the generator of ladder diagrams. Here  $\mathbf{U}^{(MC)}$  is defined to be the sum of all MC diagrams as shown in Fig. 7.4. By following the rule of transformation, every one of the diagrams can be made into a corresponding ladder diagram with the condition of (7.5). In operator form, that means

$$\begin{aligned} \mathbf{U}^{(\mathrm{M}\,\mathrm{C})} &= \mathbf{U}^{(\mathrm{B})} : (\mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-}) : \mathbf{U}^{(\mathrm{B})} + \mathbf{U}^{(\mathrm{B})} : (\mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-}) : \mathbf{U}^{(\mathrm{B})} : (\mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-}) : \mathbf{U}^{(\mathrm{B})} \\ &+ \mathbf{U}^{(\mathrm{B})} : (\mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-}) : \mathbf{U}^{(\mathrm{B})} : (\mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-}) : \mathbf{U}^{(\mathrm{B})} : (\mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-}) : \mathbf{U}^{(\mathrm{B})} \\ &= \mathbf{U}^{(\mathrm{B})} : \left[ \mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-} + (\mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-}) : \mathbf{U}^{(\mathrm{B})} : (\mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-}) + \cdots \right] : \mathbf{U}^{(\mathrm{B})} \\ &= \mathbf{U}^{(\mathrm{B})} : \left\langle \mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-} \right\rangle_{c}^{(\mathrm{B})} : \mathbf{U}^{(\mathrm{B})}. \end{aligned}$$
(7.7)

The last equality of (7.7) follows from (5.19). What (7.7) tells us is that by summing over all the MC diagrams, the task of evaluating  $U^{(MC)}$  actually becomes easy because one can rely on what is already known. Since we can make the correspondence

$$\langle \mathbf{G}_{\mathbf{e}}^{+} \otimes \mathbf{G}_{\mathbf{e}}^{-} \rangle_{\mathbf{c}}^{(\mathbf{B})} \Leftrightarrow \phi_{\mathbf{k},\mathbf{k}'}^{(\mathbf{B})}(\Delta \omega, \mathbf{Q}|\omega), \tag{7.8}$$

Fig. 7.4. The definition of  $\mathbf{U}^{(\mathrm{MC})}$  as the sum of all the maximally crossed diagrams.

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which follows from (5.11) with  $\Delta \mathbf{k}$  replaced by  $\mathbf{Q} = \mathbf{k}_i + \mathbf{k}_f$ , the MC vertex contribution, (7.7), may be written in the form

$$\left(\frac{N}{L^d}\right)^2 U^{(\mathrm{MC})} = \frac{1}{N^2} \sum_{\mathbf{k}} \sum_{\mathbf{k}'} \left[\frac{\mathrm{d}\,\kappa_{\mathrm{e}}^2}{\mathrm{d}\,\omega} \frac{1}{\pi\,\rho_{\mathrm{e}}(\omega)}\right]^2 (-\mathrm{Im}\Sigma^+)^2 \phi_{\mathbf{k},\mathbf{k}'}^{(\mathrm{B})},\qquad(7.9)$$

where the Ward identity, (5.105c), has been used to convert  $U^{(B)}$  into an expression that involves  $-\text{Im}\Sigma^+$ . If  $\Sigma^+$  is assumed to be **k**-independent (as in CPA), then for the evaluation of  $U^{(MC)}$  all those factors besides  $\phi_{\mathbf{k},\mathbf{k}'}^{(B)}$ , can be moved to the outside of the double-**k** summation. The summations over **k**, **k'** then yield  $S^{(B)}(\Delta\omega, \mathbf{Q}|\omega)$  by definition (5.10). From (5.54) it is known that

$$S^{(\mathrm{B})}(\Delta\omega, \mathbf{Q}|\omega) = \frac{N}{L^d} \frac{2\pi \rho_{\mathrm{e}}(\omega) \left(\mathrm{d}\,\omega/\mathrm{d}\,\kappa_{\mathrm{e}}^2\right) \left(\mathrm{d}\,\omega/\mathrm{d}\,\kappa_{0}^2\right)}{-\mathrm{i}\,\Delta\omega + D^{\,(\mathrm{B})}(\omega)|\mathbf{k}_{\mathrm{i}} + \mathbf{k}_{\mathrm{f}}|^2},\tag{7.10}$$

where we have imposed the MC diagram condition of  $\mathbf{Q} = \mathbf{k}_i + \mathbf{k}_f$ . Combining terms in (7.9) yields

$$\frac{N}{L^d} U_{\mathbf{k}_i, \mathbf{k}_f}^{(MC)} = \frac{2}{\rho_e(\omega)\pi} \frac{\mathrm{d} \kappa_e^2}{\mathrm{d} \kappa_0^2} \frac{(-\mathrm{Im}\Sigma^+)^2}{-\mathrm{i}\,\Delta\omega + D^{(\mathrm{B})}(\omega)|\mathbf{k}_i + \mathbf{k}_f|^2}.$$
(7.11)

Since  $U^{(MC)}$  peaks at  $\mathbf{k}_i = -\mathbf{k}_f$ , the imprint of the coherent backscattering effect is very much evident in the foregoing form.

## 7.3 Renormalized Diffusion Constant

The contribution of the maximally crossed vertex function to the diffusion constant is expressed through the integral  $M_0$ , (5.51a). By denoting the MC component of the contribution as  $M_0^{(MC)}$ , one obtains

$$M_0^{(\mathrm{MC})} = \mathrm{i} \frac{L^d}{N^3} \sum_{\mathbf{k}_{\mathrm{f}}} \sum_{\mathbf{k}_{\mathrm{f}}} (\Delta G_{\mathrm{e}})_{\mathbf{k}_{\mathrm{i}}} (\Delta G_{\mathrm{e}})_{\mathbf{k}_{\mathrm{f}}}^2 \times \left[ \nabla_{k_{\mathrm{i}}} e(\mathbf{k}_{\mathrm{i}}) \cdot \Delta \hat{\mathbf{k}} \right] \frac{N}{L^d} U_{\mathbf{k}_{\mathrm{i}}, \mathbf{k}_{\mathrm{f}}}^{(\mathrm{MC})} \left[ \nabla_{k_{\mathrm{f}}} e(\mathbf{k}_{\mathrm{f}}) \cdot \Delta \hat{\mathbf{k}} \right].$$
(7.12)

For the weak scattering limit, (7.12) may be reduced to the form

$$M_{0}^{(\mathrm{MC})} = \frac{1}{N^{2}} \frac{1}{-\mathrm{Im}\Sigma^{+}} \sum_{\mathbf{k}_{i}} \sum_{\mathbf{k}_{f}} (\Delta G_{e})_{\mathbf{k}_{i}} (\Delta G_{e})_{\mathbf{k}_{f}} \\ \times \left[ \nabla_{k_{i}} e(\mathbf{k}_{i}) \cdot \Delta \hat{\mathbf{k}} \right] \frac{N}{L^{d}} U_{\mathbf{k}_{i}, \mathbf{k}_{f}}^{(\mathrm{MC})} \left[ \nabla_{k_{f}} e(\mathbf{k}_{f}) \cdot \Delta \hat{\mathbf{k}} \right].$$
(7.13)

In contrast to the case of  $M_0^{(B)}$  for isotropic scattering,  $M_0^{(MC)}$  is clearly nonzero because the two summations are coupled through the  $|\mathbf{k}_i + \mathbf{k}_f|$  dependence of  $U^{(MC)}$ . Another way of saying the same thing is that the MC
diagram is inherently related to anisotropic scattering (in fact, it is peaked in the backward direction), so its total contribution to the diffusion constant does not vanish. In Chap.9 the integrals in (7.12) will be treated with more care. In this chapter the focus is to get a general idea about the dynamic implications of coherent backscattering. For this purpose an approximate evaluation of (7.13) suffices. To this end, the double integral in (7.13) is separated into two single integrals by observing that the diffusive pole of  $U_{\mathbf{k}_i, \mathbf{k}_f}^{(MC)}$  makes it peaked at  $\mathbf{k}_i + \mathbf{k}_f = 0$ . Therefore, the  $\mathbf{k}_i$  integral of  $(\Delta G_e)_{\mathbf{k}_i} (\Delta G_e)_{\mathbf{k}_f} [\nabla_{k_i} e(\mathbf{k}_i) \cdot \Delta \hat{\mathbf{k}}] [\nabla_{k_f} e(\mathbf{k}_f) \cdot \Delta \hat{\mathbf{k}}]$  may be evaluated approximately with the condition  $\mathbf{k}_i = -\mathbf{k}_f$ . That is, as

$$\frac{L^{d}}{N} \int \frac{\mathrm{d}\mathbf{k}_{\mathrm{i}}}{(2\pi)^{d}} \left(\Delta G_{\mathrm{e}}\right)_{\mathbf{k}_{\mathrm{i}}} \left(\Delta G_{\mathrm{e}}\right)_{-\mathbf{k}_{\mathrm{i}}} \left[\nabla_{k_{\mathrm{i}}} e(\mathbf{k}_{\mathrm{i}}) \cdot \Delta \hat{\mathbf{k}}\right]^{2} \\
= -\frac{L^{d}}{N} \int \frac{\mathrm{d}\mathbf{k}}{(2\pi)^{d}} \left(\Delta G_{\mathrm{e}}\right)_{\mathbf{k}}^{2} \left[\nabla_{k} e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}}\right]^{2} \\
= 4K_{0} \frac{N}{L^{d}},$$
(7.14)

where the minus sign is due to the fact that  $\nabla_k e(\mathbf{k})$  is an odd function of  $\mathbf{k}$ , whereas  $(\Delta G_e)_{\mathbf{k}}$  is an even function. The remaining  $\mathbf{k}_f$  integral in (7.13) may be converted to an  $\mathbf{Q}$  integral over the MC vertex function. That is,

$$M_0^{(MC)} = \frac{L^d}{N} \frac{1}{-\mathrm{Im}\Sigma^+} \int \frac{\mathrm{d}\,\mathbf{Q}}{(2\pi)^d} \, 4K_0 \frac{N}{L^d} \frac{N}{L^d} U_{\mathbf{Q}}^{(MC)}.$$
 (7.15)

Substitution of (7.11) into (7.15) gives

$$M_0^{(\mathrm{MC})} \cong \frac{N}{L^d} \frac{8K_0}{\rho_{\mathrm{e}}(\omega)\pi} \frac{\mathrm{d}\kappa_{\mathrm{e}}^2}{\mathrm{d}\kappa_0^2} \left(-\mathrm{Im}\Sigma^+\right) \frac{L^d}{N} \int \frac{\mathrm{d}\mathbf{Q}}{(2\pi)^d} \frac{1}{-\mathrm{i}\Delta\omega + D^{(\mathrm{B})}(\omega)|\mathbf{Q}|^2}$$
$$= 8D^{(\mathrm{B})}(\omega) \left(-\mathrm{Im}\Sigma^+\right) \int \frac{\mathrm{d}\mathbf{Q}}{(2\pi)^d} \frac{1}{-\mathrm{i}\Delta\omega + D^{(\mathrm{B})}(\omega)|\mathbf{Q}|^2}, \tag{7.16}$$

where (5.108a) is used for  $D^{(B)}(\omega)$ , which is also correct for the classical scalar wave if the scattering is isotropic, and  $e(\mathbf{k}) = k^2$ .

From (5.55) and (5.108a) it follows that

$$\frac{1}{D(\omega)} = \frac{1}{D^{(B)}(\omega)} + \frac{M_0^{(MC)} \pi \rho_e(\omega)}{8K_0^2 (d \kappa_e^2 / d \kappa_0^2) (-\text{Im}\Sigma^+)} \\
= \frac{1}{D^{(B)}(\omega)} \left[ 1 + \frac{(d \kappa_e^2 / d \kappa_0^2)}{\pi \rho_e(\omega)} \int \frac{d \mathbf{Q}}{(2\pi)^d} \frac{1}{-i \Delta \omega + D^{(B)}(\omega) |\mathbf{Q}|^2} \right], (7.17)$$

where  $D(\omega)$  represents the diffusion constant renormalized by the coherent backscattering effect. The  $(1 + \delta)$  factor for classical scalar waves is neglected in this chapter. A more detailed discussion of its effect is presented in Chap. 9. In order to get a better idea of the consequences of the correction term due to the MC diagrams, let us set  $\Delta \omega = 0$  so as to focus on the infinite travel-time

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limit. The correction term is clearly positive, which means  $D(\omega) < D^{(B)}(\omega)$ . By multiplying both sides by  $D(\omega)D^{(B)}(\omega)$  and assuming the correction term to be small so that  $D(\omega) \cong D^{(B)}(\omega)$ , (7.17) (with  $\Delta \omega = 0$ ) becomes

$$D(\omega) \cong D^{(\mathrm{B})}(\omega) - \frac{(\mathrm{d}\,\kappa_{\mathrm{e}}^2/\mathrm{d}\,\kappa_0^2)}{\pi\,\rho_{\mathrm{e}}(\omega)} \int \frac{\mathrm{d}\,\mathbf{Q}}{(2\pi)^d} \frac{1}{|\mathbf{Q}|^2}.$$
 (7.18)

Some striking implications of the correction term are shown below.

# 7.4 Sample Size and Spatial Dimensionality Dependencies of Wave Diffusion

The first step in the approximate evaluation of  $D(\omega)$  is to set limits on the **Q** integral in (7.18). The upper limit of the integral cannot be infinity because the diffusive behavior has a minimum length scale, given by the mean free path l. That means the upper limit of  $|\mathbf{Q}|$  should be  $\sim l^{-1}$ . On the other hand, the lower limit of  $|\mathbf{Q}|$  should be  $\sim L^{-1}$ , the inverse of the sample size.

# 7.4.1 Conductivity Correction

The resulting correction factor is more conveniently expressed in terms of the electrical conductivity in the electronic case:

$$\sigma_{\rm E}(\omega) = \sigma_{\rm E}^{\rm (B)}(\omega) - \delta \sigma_{\rm E}, \qquad (7.19)$$

with

$$\delta \sigma_{\rm E} = \frac{e^2}{2\hbar\pi^2} \begin{cases} \frac{1}{\pi} \left(\frac{1}{l} - \frac{1}{L}\right) & 3D\\ \ln \frac{L}{l} & 2D\\ 2(L-l) & 1D, \end{cases}$$
(7.20)

where we have used the Einstein relation, (5.113), and approximated  $d \kappa_e^2 / d \kappa_0^2 \approx$ 1. The conductivity here does not account for electron's spin degeneracy, which would increase both  $\sigma_E^{(B)}$  and  $\delta \sigma_E$  by a factor of 2.

# 7.4.2 Diffusion Constant Correction

For the classical wave case we have

$$D(\omega) \cong D^{(B)}(\omega) - \delta D(\omega), \qquad (7.21)$$

with

$$\delta D(\omega) = \frac{1}{\pi} v_t \kappa_e^{1-d} \begin{cases} \frac{1}{l} - \frac{1}{L} & 3D\\ \ln \frac{L}{l} & 2D\\ L-l & 1D. \end{cases}$$
(7.22)

What is remarkable about the correction to the conductivity and diffusion constant is that the sample size L matters, which means that the renormalized conductivity or diffusion constant is no longer an intensive physical quantity. The physical origin of this L dependence may be traced to the sample size dependence of the coherent backscattering effect. However, what is most significant is the fact that in 1D and 2D, the (negative) correction term diverges as  $L \to \infty$ . In other words, renormalized diffusion implies the absence of diffusion over infinite distances in 1D and 2D.

#### 7.4.3 The Emergence of Localization Length

Since the conductivity or the diffusion constant cannot be negative, the condition  $D(\omega) = 0$  defines a length, the localization length  $\xi$ , beyond which diffusive transport is no longer an accurate description. While (7.20) and (7.22) are obtained on the basis of  $D(\omega) \cong D^{(B)}(\omega)$  so that the correction term is small, we may nevertheless obtain a rough estimate of the 1D and 2D localization length by setting  $\delta \sigma_{\rm E} = \sigma_{\rm E}^{({\rm B})}(\omega)$ ,  $\delta D(\omega) = D^{({\rm B})}(\omega)$  at  $L = \xi$ . That means in the electronic case

$$\xi \cong l \exp(\rho_{\rm e}(\omega) v_t \, l \, \pi^2) \qquad \text{2D}$$
$$\cong (1 + \pi^2 \rho_{\rm e}(\omega) \, v_t) \, l \qquad \text{1D}, \tag{7.23}$$

and for the classical wave case

$$\xi \simeq l \exp\left(\frac{\pi}{2} \kappa_{e} l\right) \qquad 2D$$
$$\simeq (1+\pi) l \qquad 1D, \qquad (7.24)$$

where we have substituted  $v_t l/d$  for  $D^{(B)}$  ( $l^* = l$  for isotropic scattering). Although the numerical factors in (7.23) and (7.24) may not be accurate, they nevertheless tell us that in 1D, the localization length is directly proportional to the mean free path, whereas in 2D the localization length is an exponential function of l and therefore can be extremely large in the weak scattering limit.

The fact that the localization tendency always prevails in 1D and 2D does not invalidate what have been derived before about diffusive transport in these dimensions. The existence of a localization length, however, does imply that the diffusive behavior is restricted to a wave transport regime which is intermediate between l and  $\xi$ .

#### 7.4.4 Localization in 3D : the Ioffe – Regel Criterion

In 3D, the correction term does not diverge as  $L \to \infty$ . Therefore whether  $\sigma(\omega)$  or  $D(\omega)$  can vanish depends on the magnitude of the coefficient. In the

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electronic case, the condition for localization is obtained by setting  $\sigma_{\rm E}^{(\rm B)} = \delta \sigma_{\rm E}$ with  $L = \infty$ . That gives the localization condition

$$1 \ge \frac{2\pi^3}{3} \,\rho_{\rm e} \, v_t \, l^2. \tag{7.25}$$

Since  $\rho_{\rm e}$  can become vanishingly small near the band edge and  $v_t l^2$  generally does not diverge, the condition expressed by (7.25) is easily satisfied for the Anderson model near the band edges. We therefore expect localization of the electronic band-edge states in the presence of disorder. By using the spherical approximation for the Brillouin zone, i.e.,  $e(\mathbf{k}) = k^2$ , then

$$\rho_{\rm e} = \frac{m}{\hbar} \, \frac{\kappa_{\rm e}}{2\pi^2}$$

whereas  $v_t = \hbar \kappa_e/m$  [see (5.111) and (5.112)], whereupon the criterion for localization becomes

$$\kappa_{\rm e} \, l \le \sqrt{3/\pi} \cong 1. \tag{7.26}$$

In the classical wave case, the condition obtained by setting  $D^{(B)} = \delta D$  with  $L = \infty$  is exactly the same as (7.26). If the constant on the right-hand side of (7.26) is taken to be 1, i.e.,

$$\kappa_{\rm e} \, l \le 1,\tag{7.27}$$

then the inequality is known as the Ioffe–Regel criterion for localization (Ioffe and Regel, 1960; Mott, 1974). The original proposition of this criterion is based on the simple physical notion that if the mean free path becomes comparable to the effective wavelength, then a wave can no longer be plane wave-like, and wave localization is postulated to occur. A more formal derivation of this criterion is given in Chap. 9 for classical scalar waves.

# 7.5 Localization in One Dimension: the Herbert–Jones–Thouless Formula

The divergence of the MC diagrams' contribution in 1D means that the diffusive behavior is valid only within a limited transport range. In order to have a better idea about what happens beyond the diffusive regime, it is instructive to examine the 1D case with a different perspective from the one pursued so far.

# 7.5.1 Localization and Intensity Transmission

In 1D, it is easy to see heuristically why a wave cannot propagate indefinitely in the presence of disorder. Consider a 1D sample as composed of many segments, each of length  $\sim l$ . Since over that distance a collision is likely to occur so that the phase is randomized, the wave intensity transmission coefficient, 7.5 Localization in One Dimension: the Herbert–Jones–Thouless Formula 209

 $|\tau|^2 < 1$ , over a sample of length  $L \gg l$  may be written approximately as the product of intensity transmission coefficients, i.e.,

$$|\tau|^2 = \prod_i |\tau_i|^2, \tag{7.28}$$

where  $\tau_i$  is the amplitude transmission coefficient for the *i*th segment. The error made in this approximation is to be noted below. But for the moment let us assume its validity. By writing

$$|\tau|^2 = \exp(-2L/\xi), \tag{7.29}$$

the localization length  $\xi$  is obtained as

$$\xi = -2lN \bigg/ \sum_{i=1}^{N} \ln |\tau_i|^2, \tag{7.30}$$

where N = L/l. Equation (7.30) clearly demonstrates that  $\xi \propto l$  in 1D, in agreement with (7.24). It also shows that the wave intensity should decay exponentially inside a random medium, and  $\xi$  is that decay length.

### 7.5.2 Fluctuations in Intensity Transmission

However, the decay cannot be a smooth exponential because  $|\tau_i|^2$  is expected to be randomly varying, both as a function of frequency  $\omega$  and as a function of sample length L. The range of such variations can extend from nearly 0 to 1 because at special frequencies and sample lengths, (7.28) breaks down due to the fact that in the absence of inelastic scattering, there can be constructive or destructive interference effects over the whole sample, so that the transmission may reach the limiting value of 1 or nearly 0. For a fixed sample length L, the fact that special frequencies exist may be inferred from the wellknown resonant transmission phenomenon of periodic dielectric stacks. When a periodic stack becomes disordered, these frequencies would shift but do not disappear. As  $L \to \infty$ , the frequency width of each total transmission peak shrinks to zero, so it becomes difficult to achieve such resonant conditions. For a fixed frequency, the same effect can also appear as L is varied.

What are the statistics of such fluctuations? Since the total transmission  $|\tau|$  can vary by orders of magnitude as a function of L, it indicates that  $\ln |\tau|$  (rather than  $|\tau|$ ) is more likely to be Gaussian distributed. Physically, the rationale for  $\ln |\tau|$  to have Gaussian statistics is that wave scattering in a 1D system is *successive* in nature; i.e., the transmission probability of the whole sample is the product of transmission coefficients of the segments. That means if each scattering is independent, then  $\ln |\tau|$ 's are additive, independent random variables. In fact, if  $\ln |\tau|$ 's have Gaussian statistics, then for a fixed L, the fluctuation in  $1/\xi$ , which is determined by the mean of all the  $\ln |\tau|$ 's becomes accurate because the central limit theorem of the Gaussian statistics requires it to be so. In other words,  $1/\xi$  is a *self-averaging* quantity.

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# 7.5.3 Randomness is not Always a Sufficient Condition for 1D Localization

In the preceding discussion it is clear that the multiplicative nature of 1D wave transport is the basis of the 1D localization phenomenon. This multiplicative property can be expressed rigorously if transfer matrices were used in place of transmission coefficients. A transfer matrix is defined as the matrix which relates the two amplitudes of the left-going and right-going waves in one homogeneous segment of the sample (if the material property is continuously varying, the segment can be made infinitesimal in thickness so as to approximate homogeneity) to those of the next segment. For a piecewise homogeneous sample, the transfer matrix simply expresses the boundary conditions at the interface of the two segments. In terms of transfer matrices, Furstenberg (1963) proved that the multiplication of *random* transfer matrices always results in localization. However, care has to be exercised in equating random transfer matrices to random media because a random medium does not always imply random transfer matrices. A trivial example of this nonequivalence may be found in the case of a random binary system where component 1 has a fixed segment thickness, and there is a resonant transmission phenomenon at a set of discrete frequencies for one such segment of component 1 embedded in an otherwise homogeneous component 2. At those special frequencies, the segment of component 1 is invisible to the wave because resonant transmission implies an identity transfer matrix for the segment. It follows that if an identical second segment (of component 1) is put in, it would still be invisible at the same frequencies. One can easily deduce that a randomly placed collection of such segments would always be invisible to the wave at those particular frequencies because if transfer matrices corresponding to the segments of components 1 are identity matrices, then the overall transfer matrix must correspond to a uniform medium. Wave delocalization can thus occur at this discrete set of resonant transmission frequencies. Another possibility for nonlocalization in 1D is for the transfer matrix at *each interface* to be an identity matrix. This is the case, for example, for P-polarized electromagnetic wave propagation in a binary randomly layered medium at the Brewster angle (Sipe et al., 1988). At that particular angle (which has different values in the two components), there is total transmission through the stack, regardless of whether the medium is random or ordered.

# 7.5.4 Localization and Eigenfunctions in 1D: the Characteristic Function

The existence of a localization length directly implies a basic change in the character of the wave eigenfunction. Whereas a plane wave extends throughout the sample and has an intrinsic magnitude N, in the presence of disorder every eigenfunction in 1D and 2D exists in a finite spatial domain. Thus, their normalization does not involve N as for the plane wave case. Also, localization

causes transport properties of a system to become sample size dependent, because whereas for  $l < L \ll \xi$  the localization effect is not manifest so that the wave transport behavior is still diffusive, for  $L \gg \xi$  such is not the case because the wave cannot be transported beyond that scale. The localization length is thus the second length scale, aside from the mean free path, which marks a transition in the character of wave transport behavior. In 1D, it turns out that there is a simple approach to the accurate evaluation of the localization length, known as the Herbert–Jones–Thouless formula. This is presented below.

A special property of the 1D eigenfunctions is that they can be indexed sequentially in accordance with the number of nodes in each eigenfunction. For example, if the eigenfunction on a sample of length L (extending from x = 0 to x = L) is determined by the condition  $\phi(L) = 0$ , with  $\phi(0) = 0$ ,  $d\phi(0)/dx = 1$ being the initial conditions, then the lowest-frequency eigenfunction would have no node inside the interval, the next eigenfunction one node, etc. This is true regardless of the nature of disorder in the interval. With this knowledge, let us define a quantity,

$$\Omega_{\omega} = \lim_{L \to \infty} \frac{1}{L} \ln \phi_{\omega}(L), \qquad (7.31)$$

which is known as the characteristic function. The real part of  $\Omega_{\omega}$  yields the inverse of the localization length [since  $\phi \sim \exp(-L/\xi)$ ],  $\xi^{-1}$ , but with a positive sign. That is because in a disordered medium, if one of the two solutions to the wave equation is known to decay with the exponent  $\xi^{-1}$ , then the other solution must grow with exactly the same exponent. This is shown in the solution to Problem 7.1. For an arbitrary initial condition, the growing solution always dominates, which accounts for the expression  $\Omega_{\omega}$  giving  $\xi^{-1}$ as its real part.

For the imaginary part of  $\Omega_{\omega}$ , it is instructive to focus on what happens to  $\phi_{\omega}(L)$  as  $\omega$  varies across an eigenfrequency. Since  $\phi_{\omega}(L) = 0$  at an eigenfrequency, it is clear that  $\phi_{\omega}(L)$ , obtained by integrating from the initial conditions, must change sign as  $\omega$  varies across it so as to increase the number of nodes by 1. As a result,  $\ln \phi_{\omega}(L)$  jumps by an amount  $i\pi$  every time  $\omega$  hits an eigenfrequency. If the imaginary part of  $\Omega_{\omega}$  is specified to be a monotonic function of  $\omega$ , then  $\operatorname{Im}[\ln \phi_{\omega}(L)/\pi L]$  gives the *accumulated* density of states. In other words,

$$\frac{1}{\pi} \operatorname{Im} \left[ \frac{\mathrm{d}\Omega_{\,\omega}}{\mathrm{d}\omega} \right] = \bar{\rho}(\omega) = -\frac{1}{\pi} \frac{\mathrm{d}\kappa_0^2}{\mathrm{d}\omega} \operatorname{Im} \left\langle G^+(\omega, x = x') \right\rangle_{\mathrm{c}}.$$
 (7.32)

# 7.5.5 Relation Between the 1D Localization Length and Density of States

The causal nature of the wave equation demands that the imaginary part of  $d\Omega_{\omega}/d\omega$  be related to its real part through the Kramers–Kronig relation, as

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expressed in the solution to Problem 2.3 (P2.16):

$$\operatorname{Re}\left[\frac{\mathrm{d}\Omega_{\omega}}{\mathrm{d}\omega}\right] = \frac{\mathrm{d}\kappa_{0}^{2}(\omega)}{\mathrm{d}\omega}P\int\mathrm{d}\omega'\frac{\bar{\rho}(\omega')}{\kappa_{0}^{2}(\omega)-\kappa_{0}^{2}(\omega')}$$
$$= \operatorname{Re}\left\{\frac{\mathrm{d}\kappa_{0}^{2}(\omega)}{\mathrm{d}\omega}\left\langle G^{+}(\omega,x=x')\right\rangle_{c}\right\},$$
(7.33)

or

$$\operatorname{Re}\left[\frac{\mathrm{d}\Omega_{\omega}}{\mathrm{d}\kappa_{0}^{2}(\omega)}\right] = P \int \mathrm{d}\omega' \frac{\bar{\rho}(\omega')}{\kappa_{0}^{2}(\omega) - \kappa_{0}^{2}(\omega')}.$$
(7.34)

By integrating both sides with respect to  $\kappa_0^2(\omega)$ , a relation between  $\bar{\rho}(\omega)$  and  $\xi^{-1}(\omega)$  is obtained:

$$\operatorname{Re}\Omega_{\omega} = \xi^{-1}(\omega) = \operatorname{const.} + \int d\omega' \ln |\kappa_0^2(\omega) - \kappa_0^2(\omega')|\bar{\rho}(\omega').$$
(7.35)

This equation (the HJT formula) was first derived by Herbert and Jones (1971), and by Thouless (1972).

# 7.5.6 HJT Formula for the 1D Anderson Model

In the case of the Anderson model, one can write

$$a\xi^{-1}(u) = \text{const.} + \frac{\beta\hbar}{ma} \int_{-\infty}^{\infty} \mathrm{d}v \,\ln|u-v|\bar{\rho}(v), \qquad (7.36a)$$

where  $v, u = (\hbar \omega - \varepsilon_0)/(\beta \hbar^2/m a^2)$ . The constant of integration is uniquely determined by the condition that in a uniform lattice,  $\xi^{-1} = 0$ . Therefore,

const. = 
$$-\frac{2}{\pi} \int_0^1 \frac{\ln v}{\sqrt{1-v^2}} \, \mathrm{d}v = \ln 2.$$
 (7.36b)

It is noted that we have set u = 0 for the evaluation of the integral in (7.36b). This is possible because the integral has to be independent of u, a fact which may may be deduced by noting that if F(u) is the value of the integral, then dF(u)/du gives the real part of the 1D (diagonal) Green function, which is identically zero inside the band [see (2.65)]. Consequently F(u) is independent of u. In Fig. 7.5,  $\xi/a$  as calculated from (7.36) is plotted as a function of u for  $a^2\sigma_w = 0.5$ . In Fig. 7.6 the  $\sigma_w$  dependence of  $a\xi^{-1}$  is shown for u = 0 and u = 1. It is seen that at u = 0,  $\xi^{-1} \propto \sigma_w^2$  for small  $\sigma_w$ . At the band edge, however,  $\xi^{-1}$  is seen to vary as  $\sigma_w^{2/3}$  at small  $\sigma_w$ .

# 7.5.7 HJT Formula for 1D Classical Scalar Waves

For classical scalar waves in continuum, the constant of integration in (7.35) is given by the condition that as  $\omega \to 0$ , scattering becomes weak, and  $\xi^{-1} \to 0$ independent of  $\bar{\rho}(\omega)$ . That means



Fig. 7.5. Localization length for the Anderson model with  $a^2 \sigma_w = 0.5$ . This figure should be compared with Fig. 4.1 for the 1D mean free path. With the same amount of randomness,  $\xi$  is almost 4l [see (7.41)].



Fig. 7.6. Dependence of  $a \xi^{-1}$  as a function of randomness  $a^2 \sigma_w$  at two frequencies: one at the band center and one at the band edge. On the log–log scale, the straight line implies a power-law dependence, with the slope giving the exponent. At small  $\sigma_w a^2$  values, the band center case shows a  $\sigma_w^2$  dependence, whereas the band edge case shows a  $\sigma_w^{2/3}$  dependence.

$$\xi^{-1}(\omega) = \int_0^\infty \mathrm{d}u \,\ln \left| 1 - \frac{\omega^2}{u^2} \right| \,\bar{\rho}(u).$$
(7.37)

Here u has the dimension of frequency. In the solution to Problem 7.2 it is shown that for a uniform medium,  $\xi^{-1}(\omega) = 0$  identically, as it should be. For the purpose of numerical calculation, however, it is more convenient to integrate (7.33) directly:

$$\xi^{-1}(\omega) = \operatorname{Re} \int_0^\omega \mathrm{d}\omega' \, \frac{2\omega'}{v_0^2} \left\langle G^+(\omega', x = x') \right\rangle_{\rm c}.$$
 (7.38a)

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In the low-frequency regime where the effective medium description is valid, one can write (7.38a) as

$$\xi^{-1}(\omega) = \operatorname{Re} \int_{0}^{\omega} d\omega' \frac{2\omega'}{v_{e}^{2}} G_{e}^{+}(\omega', x = x'), \qquad (7.38b)$$

so that everything is consistently defined relative to the effective medium. From (2.33),

$$G_{\rm e}^+(\omega', x = x') = \frac{1}{2\,\mathrm{i}\,\kappa_{\rm e}^*},$$
 (7.39)

where  $\kappa_{\rm e}^* = \kappa_{\rm e} + i/2l$ . To be consistent, one has to write  $v_{\rm e}^2 = (\omega')^2/(\kappa_{\rm e}^*)^2$ . Therefore

$$\xi^{-1}(\omega) = \operatorname{Re} \int_{0}^{\omega} d\omega' \frac{2\omega'}{(\omega')^{2}} \frac{\kappa_{e}^{*}}{2i}$$
$$= \int_{0}^{\omega} d\omega' \frac{2\omega'}{4(\omega')^{2}l(\omega')}.$$
(7.40a)

In Chap. 3, it has been shown that at low frequencies l has the Rayleigh frequency dependence of  $\omega^{-2}$  [see (4.21)]. Therefore, for  $\omega \to 0$  the denominator of the integrand in (7.40a) is frequency independent. One can thus rewrite  $\xi^{-1}(\omega)$  in the  $\omega \to 0$  limit as

$$\lim_{\omega \to 0} \xi^{-1}(\omega) = \frac{1}{4\omega^2 l(\omega)} \int_0^\omega \mathrm{d}\omega' \, 2\omega', \qquad (7.40b)$$

which means

$$\xi^{-1}(\omega) = \frac{1}{4}l^{-1}(\omega).$$
(7.41)

This equation clearly demonstrates the proportionality relation between  $\xi$  and l in 1D. The factor 4 is valid at low frequencies. This factor turns out to be valid for the Anderson model as well (see Figs. 7.5 and 4.1). It is noted that in (7.24), a very rough estimate of the proportionality constant gives  $1 + \pi$ , which is close to what is obtained.

### 7.5.8 High Frequency and Low Frequency Limiting Behaviors

At high frequencies, the behavior of  $\xi^{-1}(\omega)$  is expected to depend on the details of each model. However, some general comments are possible based on the expression given by (7.37). For a 1D uniform medium,  $\bar{\rho}(\omega) = 1/\pi v_0$  is a constant. For an inhomogeneous medium it can be argued that heuristically,  $\bar{\rho}(\omega) \cong 1/\pi v(\omega)$ . The value of  $v(\omega)$  is known in the low- and high-frequency limits. At low frequencies, the effective-medium wave speed  $\bar{v}$  is given by

$$\frac{1}{\bar{v}^2} = \frac{p}{v_1^2} + \frac{1-p}{v_2^2}.$$

At the high-frequency geometric optics limit, the effective speed is given by the travel time average, i.e.,

$$\frac{1}{\bar{v}_H} = \frac{p}{v_1} + \frac{1-p}{v_2}.$$

Mathematically,  $\bar{v}_H > \bar{v}$  always, so from the expression  $\bar{\rho}(\omega) \cong 1/\pi v(\omega)$ the low-frequency density of states should always be higher than the highfrequency density of states. In (7.37), the  $\ln |1 - (\omega^2/u^2)|$  expression is positive for  $u \leq \omega/\sqrt{2}$  and negative for  $u > \omega/\sqrt{2}$ . Since the positive area of the expression exactly cancels the negative area [so for  $\bar{\rho}(\omega) = \text{const.}, \xi^{-1} = 0$ ], any increase in the low-frequency density of states gives the right-hand side of (7.37) a finite positive value. Moreover, as long as the low-frequency density of states is bounded,  $\xi^{-1}(\omega)$  cannot diverge. As a result, there are two possibilities when  $\omega \to \infty$ :  $\xi^{-1}(\omega)$  either approaches a constant or decreases to zero. It turns out that a more detailed investigation of the high-frequency behavior, using a very different theoretical technique, shows the first possibility applies when the interfaces in a medium are abrupt and sharp. The second possibility applies when the medium is smoothly varying (Sheng et al., 1986). Since  $\xi^{-1}(\omega) \sim \omega^2$  at low frequencies, it follows that in the second case  $\xi^{-1}(\omega)$  has a maximum.

A large amount of work has been done on 1D localization behavior, ranging from rigorous proof of the various localization properties to numerical simulations to experimental measurements, because the one-dimensionality allows the application of numerous theoretical techniques as well as detailed experiments. Once the problem deviates from 1D, however, the situation becomes much more complicated. The next chapter addresses a fruitful phenomenological approach to the localization problem in higher spatial dimensionalities.

# **Problems and Solutions**

**Problem 7.1** Show that for the two solutions to the 1D wave equation at a fixed frequency  $\omega$ , if one decays exponentially with an exponent  $\xi^{-1}$ , then the other one must grow exponentially with the same exponent.

Solution The wave equation,

$$\frac{\mathrm{d}^2}{\mathrm{d}x^2}\phi + \kappa^2(\omega, x)\phi = 0, \qquad (P7.1)$$

may be rewritten as a  $2 \times 2$  first-order system:

$$\frac{\mathrm{d}}{\mathrm{d}x} \begin{pmatrix} \phi \\ \phi' \end{pmatrix} = \begin{pmatrix} 0 & 1 \\ -\kappa^2 & 0 \end{pmatrix} \begin{pmatrix} \phi \\ \phi' \end{pmatrix}, \qquad (P7.2)$$

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where  $\phi' = d\phi/dx$ . The important feature of (P7.2) is that the 2×2 transfer matrix on the right-hand side has zero trace. By defining

$$\boldsymbol{\Phi} = \begin{pmatrix} \phi & \phi_1 \\ \phi' & \phi'_1 \end{pmatrix}, \tag{P7.3}$$

where  $\phi$  and  $\phi_1$  are the two eigensolutions of the system, (P7.2) can be put in the discrete matrix form as

$$\mathbf{\Phi}(x + \Delta x) = (\mathbf{I} + \mathbf{M} \Delta x) \mathbf{\Phi}(x), \qquad (P7.4)$$

where

$$\mathbf{M} = \begin{pmatrix} 0 & 1\\ -\kappa^2 & 0 \end{pmatrix}. \tag{P7.5}$$

By taking the determinant of both sides of (P7.4), one gets

$$det[\mathbf{\Phi}(x + \Delta x)] = det(\mathbf{I} + \mathbf{M} \Delta x) det[\mathbf{\Phi}(x)]$$
  

$$\cong [1 + (\mathrm{Tr} \ \mathbf{M})\Delta x] det[\mathbf{\Phi}(x)], \qquad (P7.6)$$

where only terms linear in  $\Delta x$  are retained. Here Tr **M** means the trace of the matrix **M**. As  $\Delta x \rightarrow 0$ , (P7.6) means

$$\frac{\mathrm{d}}{\mathrm{d}x} \det \mathbf{\Phi} = (\mathrm{Tr} \ \mathbf{M}) \det \mathbf{\Phi} = 0, \tag{P7.7}$$

since **M** has zero trace. Therefore, the determinant of the solution matrix is a constant, independent of x. It follows that if one eigensolution decays like  $\exp(-x/\xi)$ , then the other eigensolution must grow exponentially as  $\exp(x/\xi)$ , so that the determinant of the solution matrix can maintain its constancy as a function of x. In particular, the growth rate must accurately equal the decay rate.

**Problem 7.2** Show that (7.37) gives  $\xi^{-1} = 0$  for the 1D uniform continuum.

**Solution** In a uniform 1D medium,  $\bar{\rho}$  is a constant independent of the frequency. From the equation,

$$\xi^{-1} = \int_0^\infty \ln \left| 1 - \frac{\omega^2}{u^2} \right| \, \bar{\rho}(u) \, \mathrm{d}u,$$

it is clear that if  $\bar{\rho}(\omega)$  is constant [see (2.27)], then the Lyapunov exponent,  $\xi^{-1}$  is proportional to a simple integral,

$$\xi^{-1} \propto \int_0^\infty \ln \left| 1 - \frac{1}{v^2} \right| \,\mathrm{d}v,$$
 (P7.8)

where  $v = u/\omega$ . The integral may be rewritten as

where z = v - 1 from the previous line. From the last line of this equation it is clear that the first two terms cancel exactly because by setting z' = 1 - v the first integral becomes exactly the second one. The third and fourth integrals also cancel because their integrands are the negatives of each other. Therefore,  $\xi^{-1} = 0$  exactly in a 1D uniform medium, as it should be. Equation (7.37) tells us that a constant density of states is intimately related to this fact. Since  $\xi^{-1}(\omega)$  must be positive, (7.37) is also a constraint on the physically allowable form for  $\bar{\rho}(\omega)$ .

# 8.1 Distinguishing a Localized State from an Extended State

The basis of a phenomenological theory of localization is the simple observation that a localized state has a length scale beyond which it is not mobile. To build upon this observation, one can imagine the following thought experiment. Consider a sample whose wave transport property is to be determined in the absence of inelastic scattering. By dividing the sample into many smaller cubes of varying sizes, the experiment consists of introducing a single-frequency excitation source at the center of each sample and detecting for the response at the samples' surfaces (perimeter). By plotting the response as a function of sample size L, it is possible to distinguish a localized state, which is expected to exhibit a decreasing response as sample size increases, from an extended state, for which the integrated response around the sample surface should stay constant.

To quantify this approach, the above argument may be turned around and focused on the sensitivity of the wave function inside the sample to the boundary condition(s) at the surfaces (Edward and Thouless 1972; Licciardello and Thouless 1975; Thouless 1974). The rationale is that a localized eigenfunction should become insensitive to the boundary conditions as the sample size increases beyond its localization length, whereas an extended state would always be sensitive, regardless of the sample size. One way to characterize this sensitivity is to define a frequency shift  $\delta \omega$  that is the difference in the eigenvalues when the boundary condition is switched from the symmetric, periodic boundary condition to the antisymmetric boundary condition. If the sample is homogeneous, this change of boundary condition is equivalent to adding or subtracting an extra half a wavelength, resulting in a change in the wavevector by the amount

$$\delta k = \pi/L. \tag{8.1}$$

For the electronic case,  $\omega = \hbar k^2/2m$ , so  $\delta \omega = (\hbar k/m)\delta k = v\pi/L$ . For the classical scalar wave,  $\omega = v_0 k$ , so  $\delta \omega = v_0 \pi/L$  also. In both cases it is seen that  $\delta \omega$  corresponds to the inverse of a time scale, L/v, required for the wave to travel from one side of the sample to the other. If the sample is inhomogeneous and the eigenfunctions are extended, the calculation of  $\delta \omega$  is not trivial. Thouless has argued that the relevant time scale should correspond to the *diffusive* transport time to cover the length L, which means

$$\delta\omega \propto \frac{1}{L^2}.\tag{8.2}$$

 $\delta\omega$  may be interpreted physically as the frequency width arising from the (stochastic) time scale required for a change in the boundary condition to be communicated to the wave function, since it can be argued that in changing the boundary condition abruptly at one end of a sample, the wave function cannot fully adjust to it until the change signal is propagated to the other end. If the nature of signal propagation is diffusive, then the uncertainty principle,  $\delta\omega \sim (\delta t)^{-1}$ , leads to the *L* dependence of (8.2). To obtain a quantitative measure of the sensitivity to a boundary condition change, the frequency shift  $\delta\omega$  has to be compared with the average frequency separation  $\Delta\omega$  between two neighboring eigenvalues, and the dimensionless ratio  $\delta\omega/\Delta\omega$  is denoted as

$$\gamma = \frac{\delta\omega}{\Delta\omega}.\tag{8.3}$$

Since  $\Delta \omega \propto L^{-d}$ , for extended states the parameter  $\gamma$  is noted to increase as L in 3D and  $\gamma \propto L^{d-2}$  in general.

For a localized state  $\delta\omega$  must decrease exponentially when L increases beyond  $\xi$ , since it is sensitive to the boundary condition only to the degree that it can "feel" the boundary through its exponential tail. However,  $\Delta\omega$  should have the same variation with L regardless of whether the states are extended or localized. Therefore,  $\gamma$  always decreases exponentially as a function of Lfor localized states.

# 8.1.1 The Scaling Argument and Spatial Dimensionality

A fruitful way to examine the sample size variation of  $\gamma$  is through the consideration of joining smaller samples to form successively larger ones, i.e., by scaling the sample size. Let us first consider the d = 3 case. When  $\delta \omega \gg \Delta \omega$ , an eigenstate in the smaller sample can easily couple onto other eigenstates in the neighboring samples that overlap with its eigenfrequency, thus resulting in a new wave function that can extend throughout the larger sample. Since in the larger sample the ratio  $\delta \omega / \Delta \omega$  increases, the renormalization to even larger samples guarantees the formation of new extended states. On the

other hand, if  $\delta\omega \ll \Delta\omega$ , then the eigenstates in neighboring samples would find it difficult to overlap with each other through their finite level width. Because they cannot couple effectively, the states would stay localized in the smaller samples. That means in the larger sample  $\delta\omega$  is smaller by an exponential factor, since a localized state can feel the boundary only through its tail. Consequently, although  $\Delta\omega$  is smaller in the larger sample, the ratio  $\delta\omega/\Delta\omega$  still decreases, thus guaranteeing localization in the infinite sample size limit. The lesson of this scaling argument, due to Thouless (1979), is that a large  $\delta\omega/\Delta\omega$  starting ratio tends to renormalize to  $\gamma = \infty$ , and a small  $\delta\omega/\Delta\omega$  starting ratio tends to renormalize to  $\gamma = \infty$ , and a small  $\delta\omega/\Delta\omega$  starting ratio tends to renormalize to  $\gamma = \gamma_c$  at which one behavior crosses into another. At  $\gamma_c$  the value of  $\delta\omega/\Delta\omega$  is scale-invariant. This plausible assumption of a critical  $\gamma_c$  is an essential element for the scaling theory of localization.

For the case of d = 1,  $\gamma \propto L^{d-2}$  in the extended states regime implies that as the sample size increases,  $\gamma$  would continuously decrease, first as 1/Land then as  $\exp(-2L/\xi)$  in the localized states regime. Thus there can only be localized states in 1D random systems. The case of d = 2 is interesting because  $\delta \omega / \Delta \omega \sim \text{constant}$  in the "extended state" limit, so it can go either way. Thus, the scaling argument indicates that the localization phenomenon is sensitive to the spatial dimensionality of the sample, with d = 2 being the "marginal" dimension for localization. From (8.2) and (8.3) the origin of the marginal spatial dimension can be traced to the diffusive character of the multiply scattered wave.

#### 8.1.2 Dimensionless Conductance

The quantity  $\gamma$  may be heuristically interpreted as the dimensionless conductance in electronic systems. From the Einstein relation and the fact that conductance  $\Gamma_{\rm E} = \sigma_{\rm E} L^{d-2}$ , it follows that

$$\Gamma_{\rm E} = \frac{e^2}{\hbar} D\rho_{\rm e}(\omega) L^{d-2}.$$

By using the Thouless argument that  $\delta \nu L^2 = D$ , where  $\delta \nu = \delta \omega / 2\pi$ , we have

$$\Gamma_{\rm E} = \frac{e^2}{\hbar} \frac{\delta\omega}{2\pi} L^2 \frac{1}{\delta\omega L^d} L^{d-2}$$

$$= \frac{e^2}{h} \gamma,$$
(8.4)

where  $\gamma = \delta \omega / \Delta \omega$ . From (8.2) and (8.3), it is seen that

$$\gamma\left(L\right) \propto L^{d-2} \tag{8.5}$$

in the extended regime, just what is expected for the conductance of an electric conductor with cubic (square) geometry. In the localized regime, it follows from the arguments presented above that  $\gamma$  should have the behavior

$$\gamma\left(L\right) \propto \exp(-2L/\xi) \tag{8.6}$$

for  $L \gg \xi$ .

In the context of a critical  $\gamma_c$  an interesting phenomenon, first pointed out by Mott, is that the electrically conducting materials all have 3D conductivities larger than a certain minimum value (Mott et al. 1975), given approximately by  $0.026 - 0.1e^2/(\hbar a)$ , denoted Mott's minimum metallic conductivity, where a is the atomic unit cell size. In terms of conductance, that translates into the condition of  $\gamma > \gamma_c \simeq 0.16 - 0.6$  for conducting behavior, since a is the smallest physical length scale in the problem. Due to the electron's spin degeneracy, the actual  $\gamma_c$  to be compared with the nondegenerate case is 0.008 - 0.3. While there is certainly the additional physics of electron–electron interaction underlying this phenomenon, Mott's observation nevertheless has an important influence on the development of the scaling theory of Anderson localization.

# 8.2 The Scaling Hypothesis and Its Consequences

One fact that has not been mentioned so far is that  $\gamma$  for a finite sample is a highly configuration-dependent quantity. Moreover, the measured  $\gamma$  in an electronic system is dependent on how the leads are attached to it. Suppose the latter problem is solved by defining a uniform way of measuring  $\gamma$ , e.g., by attaching leads of width L to two opposite sides of the sample; then there is still the generic problem of how to define  $\gamma$  for a finite sample. Since we want to associate a single value of  $\gamma$  to a given sample size, that value must be the configurational average of the  $\gamma$  values. The question is: What is the correct average to take? It turns out that averaging  $\gamma$  and averaging  $\ln \gamma$  can give quite different answers. In order to define the proper average for  $\gamma$ , it is necessary first to know the distribution of  $\gamma$ . Ideally, one would like to have a Gaussian distribution for  $\gamma$  so that the central limit theorem applies to its mean. Evidence from numerical simulations in 2D indicated that, as with the 1D transmission coefficient, whereas  $\gamma$  indeed displays Gaussian statistics when it is large in value (delocalized limit), in the limit of small  $\gamma$ ,  $\ln \gamma$ is the random variable that displays Gaussian statistics (Zhang and Sheng 1991a). Therefore an attractive choice is to define  $\langle \gamma \rangle_c$  or  $\langle \ln \gamma \rangle_c$  as the quantity to be associated with a given sample size in the extended or the localized regime, respectively. What is meant by  $\gamma$  in the localized regime is then defined by  $\exp \langle \ln \gamma \rangle_{c}$ . Of course, a Gaussian distribution has two parameters, the mean and the variance. The basis of the one-parameter scaling theory, to be discussed below, is that the mean alone is sufficient to describe most of the important features of the localization phenomena in different spatial dimensionalities. For simplicity, the configurational averaging brackets  $\langle \rangle_c$  will be dropped in the following so that  $\gamma$  and  $\ln \gamma$  should be understood as their configurationally averaged values in their respective regimes.

#### 8.2.1 Defining the Scaling Function

Let us summarize what has been discussed as follows. In terms of  $\gamma(L)$ , it is known that

$$\ln \gamma(L) \propto \begin{cases} (d-2) \ln L & \text{extended states} \\ -\exp(\ln L) & \text{localized states.} \end{cases}$$
(8.7)

Note that as L increases, whereas  $\ln \gamma$  increases for an extended state in 3D, it decreases for a localized state. In other words, the derivative  $d \left[ \ln \gamma \left( L \right) \right] / (d \ln L)$  is positive for 3D extended states but is negative for localized states. When combined with the fact that in 3D there is a critical point at  $\gamma_c$ , it is clear that if the transition from the localized to the extended state is continuous, then  $d \left[ \ln \gamma \left( L \right) \right] / (d \ln L) = 0$  at  $\gamma = \gamma_c$ . In other words,  $\ln \gamma_c$  is a *fixed point* in the scaling variation of  $\ln \gamma$  at which sample size variation has no effect on its value. The *scaling function* is defined by the rate of change of  $\ln \gamma$  as a function of  $\ln L$ :

$$\beta = \frac{\mathrm{d}\ln\gamma}{\mathrm{d}\ln L}.\tag{8.8}$$

(Here the  $\beta$  notation is used to be consistent with the literature notation; it should be distinguished from the effective-mass  $\beta$  factor introduced in Chap. 2). From (8.7) it follows that

$$\beta \propto \begin{cases} d-2 & \text{large } \gamma \\ \ln \gamma & \text{small } \gamma \\ 0 & \text{at } \gamma = \gamma_{c} \end{cases}$$
(8.9)

### 8.2.2 Scaling and the Effect of Randomness

In terms of the scaling function  $\beta$ , the scaling hypothesis states that  $\beta$  is a function of  $\gamma$  only. This hypothesis, first proposed by Abrahams et al. (1979), is consistent with the right-hand side of (8.9), and it is automatically satisfied if  $\gamma$  can be expressed as a function of L only, so that if  $\beta = f[\gamma(L), L]$ , then L may be solved in terms of  $\gamma$ . Since  $\gamma$  must also depend on the amount of randomness in the system, the scaling hypothesis thus implies that the effect of increasing or decreasing the randomness may be completely compensated by varying L. One way this can happen is that  $\gamma$  is a function of only one variable which combines both randomness and L, and the domain of values for that one variable is  $\xi/L$  in the localized domain and  $\zeta/L$  in the extended domain. The physical significance of  $\zeta$ , denoted the correlation length, will be clarified later. Both  $\xi$  and  $\zeta$  are functions of randomness, but if  $\gamma$  is a function only of the ratio  $\xi/L$  or  $\zeta/L$ , then varying L completely compensates for the effect of randomness.

In (8.9), the large- $\gamma$  limiting value of  $\beta = d-2$  reflects the spatial dimensionality dependence of conductance for a conventional conductor. In this limit conductivity remains an intensive quantity. The small- $\gamma$  limiting behavior, on the other hand, reflects the universal behavior of localization. It differs from the conventional notion for an insulator because here the conductance refers to the zero temperature conductance, which is zero for an infinite sample.

#### 8.2.3 Behavior of the Scaling Function

With the scaling hypothesis, a schematic picture of  $\beta$  (ln  $\gamma$ ) is shown in Fig. 8.1, where the *continuity* of the  $\beta$  function is assumed. Some interesting behavior can be immediately deduced simply by observing that  $\beta$  represents the "rate of change" for ln  $\gamma$  as a function of the sample size. For d = 3, if one starts at a point ln  $\gamma > \ln \gamma_c$  so that  $\beta$  is positive, then the definition of  $\beta = (d \ln \gamma) / (d \ln L)$  implies that as L increases, ln  $\gamma$  increases also, which would make  $\beta$  even larger. The arrow on the curve thus indicates the direction of variation of ln  $\gamma$  as L increases. If  $\beta$  (ln  $\gamma$ ) is negative, on the other hand, then as L increases ln  $\gamma$  should decrease, which makes  $\beta$  even more negative. Therefore  $\gamma_c$  is an unstable fixed point where the direction of variation of  $\beta$  (ln  $\gamma$ ) points away from  $\gamma_c$ . It is seen that  $\beta$  (ln  $\gamma$ ) embodies the scaling property of  $\gamma$  discussed earlier.

For the d = 2 case, (8.9) states that  $\beta = 0$  to the leading order in the large- $\gamma$  limit. The sign of the next-order term is therefore important in deciding whether there is a fixed point in 2D. In drawing Fig. 8.1, which shows no mobility edge for 2D, use has been made of the information given by (7.19) and (7.20) for the electronic systems, which show that in 2D,  $\sigma_{\rm E} = \Gamma_{\rm E}$ ,

$$\gamma = \gamma^{(B)} - \delta\gamma(L), \qquad (8.10)$$



Fig. 8.1. A schematic picture showing the behavior of the scaling function. The picture is constructed from the known asymptotic behavior of  $\beta (\ln \gamma)$  in different spatial dimensionalities, plus the assumption of continuity. The arrows indicate the directions of variation as the sample size L increases.

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with

$$\delta\gamma(L) = \frac{1}{\pi} \ln\left(\frac{L}{l}\right). \tag{8.11}$$

The constant here,  $1/\pi$ , is for a  $\gamma$  that has no spin degeneracy. It follows from (8.11) that

$$\frac{\mathrm{d}\gamma}{\mathrm{d}\left(\ln L\right)} = -\frac{1}{\pi}.\tag{8.12}$$

By dividing both sides by  $\gamma$ , we get

$$\frac{\mathrm{d}\,(\ln\gamma)}{\mathrm{d}\,(\ln L)} \cong -\frac{1}{\pi\gamma} \tag{8.13}$$

in 2D. That means the direction of variation for  $\beta$  with increasing L is as indicated in Fig. 8.1, so that all states are localized in 2D, no matter how weak the randomness. The same conclusion can be reached for the 1D case.

# 8.2.4 When System Size is Less than the Localization Length

The significant feature of the 1D or 2D scaling function  $\beta (\ln \gamma)$  is the switchover from the linear behavior at negative  $\ln \gamma$  to the constant behavior at large  $\ln \gamma$ . This switch-over is a manifestation of the effect of the localization length; for  $L < \xi$  the measured behavior should approximate that of a conductor, but as L increases beyond the localization length the insulating behavior prevails. This understanding may be used to estimate the variation of the localization length with randomness. For example, from the integration of (8.12) one gets,

$$\gamma = \gamma_0 - \frac{2}{\pi} \ln\left(\frac{L}{l}\right),\tag{8.14}$$

as the 2D behavior of  $\gamma$  in the regime  $L < \xi$ , where  $\gamma_0$  denotes the value of  $\gamma$  at L = l, the mean free path, and may be regarded as a measure of randomness. The value of  $\gamma = \gamma_s$ , at the switch-over is defined by setting  $L = \xi$ :

$$\gamma_{\rm s} = \gamma_0 - \frac{2}{\pi} \ln\left(\frac{\xi}{l}\right),$$
$$\frac{\xi}{l} = \exp\left[\frac{\pi}{2}(\gamma_0 - \gamma_{\rm s})\right]. \tag{8.15}$$

As  $\gamma_0$  becomes larger than  $\gamma_s$  in the weak-scattering regime, the localization length grows exponentially. This result reproduces what was deduced in the last chapter, but here the continuity of the  $\beta$  function into the strongscattering regime shows that the whole phenomenon can be viewed in a unified framework through the variation of the sample size L.

There is an interesting issue about taking the limit of  $\lambda \to \infty$ ,  $L \to \infty$  for the classical wave in 1D and 2D. It is known that as  $\lambda \to \infty$  the classical

or

wave scattering strength approaches zero as  $\lambda^{-(d+1)}$ . Therefore, by fixing L and letting  $\lambda \to \infty$ , the conducting limit is always attained. However, if  $\lambda$  is fixed and  $L \to \infty$ , then the renormalization arrow tells us that the limiting behavior is the localized state. The noncommuting limits are seen to be caused by the existence of *three length scales* in the problem:  $R, \lambda$ , and L. As  $R/\lambda$  determines  $\xi$ , taking the  $\lambda \to \infty$  limit first means  $\xi \to \infty$ , so  $\xi/L \to \infty$  always. However, if the limit  $L \to \infty$  is taken first, then  $\xi/L \to 0$ .

# 8.2.5 Wave Transport Near the 3D Mobility Edge

In 3D, the most intriguing feature of the  $\beta$  function is the existence of a critical  $\gamma_{\rm c}$ . Near  $\gamma_{\rm c}$ ,

$$\beta(\gamma) \cong s \left[ \ln \gamma - \ln \gamma_{\rm c} \right] = s \ln \frac{\gamma}{\gamma_{\rm c}},\tag{8.16}$$

where s is the slope of the  $\beta$  function when it crosses zero at  $\ln \gamma_c$ . From the definition of  $\beta (\ln \gamma)$ , one gets

$$\frac{\mathrm{d}\ln\gamma}{\ln\left(\gamma/\gamma_{\mathrm{c}}\right)} = s \,\mathrm{d}\left(\ln L\right). \tag{8.17}$$

Integration of (8.17) yields

$$\frac{\ln\left(\gamma/\gamma_{\rm c}\right)}{\ln\left(\gamma_0/\gamma_{\rm c}\right)} = \left(\frac{L}{l}\right)^s,\tag{8.18}$$

where  $\gamma_0$  is the value of  $\gamma$  at L = l, determined solely by the amount of randomness. Let us define

$$\varepsilon = \ln \gamma_0 - \ln \gamma_c = \ln \frac{\gamma_0}{\gamma_c} = \ln \left( \frac{\gamma_0 - \gamma_c}{\gamma_c} + 1 \right) \cong \frac{\gamma_0 - \gamma_c}{\gamma_c}, \tag{8.19}$$

as the distance to the mobility edge in terms of the randomness. Then (8.18) and (8.19) imply

$$\ln\left(\frac{\gamma}{\gamma_{\rm c}}\right) = \varepsilon \left(\frac{L}{l}\right)^s. \tag{8.20}$$

When  $\varepsilon$  is large enough such that

$$\beta \cong s \,\ln\frac{\gamma}{\gamma_{\rm c}} \cong 1,\tag{8.21}$$

the same switch-over in the behavior of  $\beta$  occurs in the positive  $\beta$  region. From (8.20) and (8.21), this condition is satisfied at a length scale  $\zeta$  such that  $\ln (\gamma/\gamma_c) \approx 1/s$ , or

$$\frac{\zeta}{l} = (s\varepsilon)^{-1/s}.\tag{8.22}$$

Therefore, in 3D there is a length scale,  $\zeta$ , denoted the correlation length, which defines the property of a strongly scattered wave field in the delocalized

regime.  $\zeta$  is seen to diverge at the mobility edge where  $\varepsilon \to 0$ . From the arrow attached to the  $\beta$  curve in the positive  $\beta$  region, one can deduce that for  $L > \zeta$  so that  $\beta \cong 1$ , the classical behavior prevails. For  $L < \zeta$ , on the other hand, the dimensionless conductance follows the behavior given by (8.18), or

$$\gamma = \gamma_{\rm c} \left( 1 + \frac{\gamma_0 - \gamma_{\rm c}}{\gamma_{\rm c}} \right)^{(L/l)^s} \cong \gamma_{\rm c} + \left( \frac{L}{l} \right)^s (\gamma_0 - \gamma_{\rm c}) \,.$$
(8.23)

As  $\gamma_0$  approaches  $\gamma_c$ , it is seen that  $\gamma$  becomes *independent of* L. In other words, the conductivity, which is given by  $\gamma/L$ , is no longer intensive in nature, even though the wave is still delocalized. Since  $D \propto \gamma/L$ , it follows from (8.23) that

$$\frac{\gamma}{L} \cong \frac{\gamma_{\rm c}}{L} + \frac{1}{L} \left(\frac{L}{l}\right)^s (\gamma_0 - \gamma_{\rm c}) \\ \cong \frac{\gamma_0}{L}$$
(8.24)

when  $\gamma_0 \cong \gamma_c$ . Observing that  $\gamma_0/l \propto D^{(B)}$  from the definition of  $\gamma_0$  we obtain from (8.24)

$$D = D^{(B)} \frac{l}{L}, \qquad (8.25a)$$

for  $l < L < \zeta$ . Continuity and the requirement that D be L independent for  $L > \zeta$  means

$$D = D^{(B)} \frac{l}{\zeta} \tag{8.25b}$$

in the  $L \to \infty$  limit. For weak scattering,  $\varepsilon$  is large and from (8.22)  $\zeta$  is expected to be on the order of l so that  $D \cong D^{(B)}$ . For strong scattering, however,  $\zeta \gg l$ , and the renormalized D can be significantly smaller than its Boltzmann value. A natural formula for interpolation between the two limits is

$$D = D^{(B)}l\left(\frac{1}{\zeta} + \frac{1}{L}\right).$$
(8.26)

# 8.2.6 Anomalous Diffusion and the Minimum Metallic Conductance in 3D

It is interesting to observe that near the mobility edge, i.e., when  $L\ll\zeta$ , the diffusion is anomalous in its time dependence. That is, from the diffusion relation

$$L^2 = Dt.$$

the L dependence of D as given by (8.25a) directly implies

$$L^{2} = D^{(B)} \frac{l}{L} t,$$

$$L \propto t^{1/3}.$$
(8.27)

Therefore, near the mobility edge diffusion has anomalous time dependence.

Equation (8.25a) also enables us to estimate the value of  $\gamma_c$ , a dimensionless constant, in terms of the Ioffe-Regel criterion. In 3D, we have the relation

$$\gamma = 2\pi\rho_e(\omega)DL. \tag{8.28}$$

For classical waves, the 3D density of states at a frequency  $\omega$  is given by

$$\rho_e(\omega) = \frac{4\pi\kappa_e^2\Delta\kappa_e}{(2\pi)^3\Delta\omega} = \frac{4\pi\kappa_e^2}{(2\pi)^3\upsilon},\tag{8.29}$$

in the approximation where  $e(\mathbf{k})$  is replaced by  $k^2$ , and  $\Delta \omega / \Delta \kappa_e$  is identified as v. By writing D as

$$D = \frac{1}{3}\upsilon l \frac{l}{L},\tag{8.30}$$

for D close to  $\gamma_c$ , substitution of (8.29) and (8.30) into (8.28) leads to

$$\gamma \cong \frac{(\kappa_e l)^2}{3\pi}, \qquad \gamma \cong \gamma_c.$$
 (8.31)

In particular,

or

$$\gamma_{\rm c} = \frac{1}{3\pi} \simeq 0.106 \tag{8.32}$$

when  $\kappa_{\rm e} l$  is set equal to 1 at  $\gamma_{\rm c}$  as dictated by the Ioffe-Regel criterion. This value of  $\gamma_{\rm c}$  is noted to be on the same order as what has been deduced from Mott's minimum metallic conductivity criterion.

# 8.2.7 Scaling Behavior in the 3D Localized Regime

When  $\beta$  is negative in the 3D case, i.e., in the localized regime, (8.20) may be written as

$$\ln\left(\frac{\gamma}{\gamma_{\rm c}}\right) = -|\varepsilon| \left(\frac{L}{l}\right)^s = -\left(\frac{L}{l|\varepsilon|^{-1/s}}\right)^s.$$
(8.33)

Here, however, one has to remember that for  $\gamma$  to be close to  $\gamma_c$  so that (8.33) is valid, L cannot be very large since increasing L would move  $\ln \gamma$  away from  $\ln \gamma_c$  as indicated by the arrow on the  $\beta$  curve shown in Fig. 8.1. If we let L become large,  $\ln \gamma$  has to be proportional to -L as specified by (8.7). In that case (8.33) has to become

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$$\ln\left(\frac{\gamma}{\gamma_c}\right) = -\frac{L}{l|\varepsilon|^{-1/s}} = -\frac{L}{\xi},\tag{8.34}$$

so that the localization length is seen to depend on  $|\varepsilon|$  as

$$\xi \propto |\varepsilon|^{-1/s}.\tag{8.35}$$

One way to manipulate the randomness as seen by a wave is to change its frequency. For example, in the Rayleigh scattering regime the strength of scattering by the random inhomogeneities increases with frequency. Therefore  $\gamma_0$  can also be regarded as a function of frequency. In that case if  $\omega_c$  is defined by the condition  $\gamma_c = \gamma_0 (\omega_c)$ , then

$$\varepsilon \cong \frac{[\gamma_0(\omega) - \gamma_0(\omega_c)]}{\gamma_0(\omega_c)} \propto (\omega - \omega_c), \qquad (8.36)$$

so that

$$\xi \propto |\omega - \omega_{\rm c}|^{-1/s} \tag{8.37}$$

near the mobility edge. Similarly, from (8.22) it may be concluded that

$$\zeta \propto |\omega - \omega_{\rm c}|^{-1/s} \,. \tag{8.38}$$

Therefore, both the correlation length and the localization length diverge at the mobility edge with the exponent 1/s.

The earlier discussion shows that  $\gamma$  is a function of  $L / l|\varepsilon|^{-1/s}$ . From (8.22) and (8.34) it is seen that  $l|\varepsilon|^{-1/s}$  is proportional to either  $\zeta$  or  $\xi$ , depending on whether the regime is localized or delocalized. Therefore  $\gamma$  is always a function of only one variable,  $L/\xi$  or  $L/\zeta$ , where  $\xi$  and  $\zeta$  are functions of randomness. This feature embodies the (previously) anticipated scaling property of  $\gamma$ .

# 8.3 Numerical Evaluation of the Scaling Function

The scaling hypothesis is a powerful assumption from which many localization properties may be deduced. One way to check its validity is via numerical simulations, which have gained acceptance over the past few decades as a means of verifying theoretical hypothesis and predicting the consequences of theories where analytical results are impossible or not yet available. Since numerical simulations can deal only with finite samples, the variation of  $\gamma$  with a finite sample size L is a near-perfect case for numerical simulation. In fact, it is shown below that in 1D,  $\beta$  (ln  $\gamma$ ) may be written down explicitly, and numerical simulations in 2D and 3D can be carried out by a scheme, commonly known as the finite-size scaling approach (MacKinnon and Kramer 1981, 1983), which is a generalization of the 1D problem to systems of finite cross-sections. This approach is numerically efficient and can provide the necessary accuracy for

checking the validity of the scaling hypothesis, as well as giving the actual functional form of  $\beta (\ln \gamma)$ .

In the finite-size scaling approach, one considers a strip of length N and cross-sectional dimension M, where  $N \gg M$  are both in units of the grid size. In the Anderson model, the grid coincides with the lattice. The condition that  $N \gg M$  is crucial here because in that limit a unique localization length  $\xi_{\rm M}(W)$ , independent of N, may be associated with a given cross-sectional dimension and randomness. Here W is used as a general parameter to denote the amount of randomness.

# 8.3.1 The Finite-Size Scaling Function

The central dimensionless quantity considered in the finite-size scaling approach is  $\Lambda = \xi_{\rm M}(W)/M$ , i.e., the localization length of the strip in units of its cross-sectional dimension. Just as in the scaling hypothesis, here the main finite-size scaling assumption, which is to be verified numerically, is that as M varies, the behavior of  $\Lambda = \xi_{\rm M}(W)/M$  depends only on  $\Lambda$ , not on W or M separately. That is, a finite-size scaling function is defined as

$$\frac{\mathrm{d}\left(\ln\Lambda\right)}{\mathrm{d}\left(\ln M\right)} = \chi\left(\ln\Lambda\right). \tag{8.39}$$

Equation (8.39) can be integrated to get

$$h(\Lambda) = \int \frac{\mathrm{d}\ln\Lambda}{\chi\left(\ln\Lambda\right)} = \ln M - \ln\left(\mathrm{constant}\right) = -\ln\frac{\mathrm{constant}}{M},\qquad(8.40)$$

where the "constant" means independence from M, so it can depend only on W. By denoting f as the inverse function of h, one obtains

$$\Lambda = f\left[\frac{\text{some function }(W)}{M}\right].$$
(8.41)

### 8.3.2 Limiting Behaviors

When  $\Lambda$  is small, i.e.,  $M \gg \xi_{\mathrm{M}}(W)$ , the system is essentially in the 2D or 3D limit, and  $\xi_{\mathrm{M}}(W)$  should be the true localization length in 2D or 3D,  $\xi(W)$ , where  $\xi(W)$  is independent of M. That gives

$$\frac{\mathrm{d}(\ln\Lambda)}{\mathrm{d}(\ln M)} = \frac{\mathrm{d}\left\{\ln\left[\xi(W)/M\right]\right\}}{\mathrm{d}(\ln M)} = -1 = \chi\left(\ln\Lambda\right). \tag{8.42}$$

Integration of (8.42) gives back

$$\Lambda = \frac{\xi(W)}{M},\tag{8.43}$$

as the small  $\Lambda$  limiting behavior. That is, when  $\Lambda \ll 1$  we have f(x) = x if  $\xi(W)$  is identified as the *W*-dependent constant in (8.41).

For  $\Lambda$  large,  $\xi_{\rm M}(W) \gg M$ , the problem is effectively 1D; and the wave amplitude is expected to be evenly distributed across the cross-section of the strip so that the *effective* 1D disorder W' as seen by the wave is averaged over the width of the strip. By assuming Gaussian statistics, that means

$$W^{\prime 2} = \frac{W^2}{M^{d-1}}.$$
(8.44)

Since the 1D localization length in the Anderson model is proportional to  $\sigma_w^{-2}$  (for  $\sigma_w$  not too large and for energy not close to be band edge), and  $\sigma_w \propto W$ , it means

$$\xi_{\rm M} \sim (W')^{-2} = W^{-2} M^{d-1},$$
 (8.45)

so that

$$\Lambda = \frac{\xi_{\rm M}}{M} \propto W^{-2} M^{d-2}.$$

For classical waves,  $\xi_{\rm M}$  is proportional to the mean free path in 1D, which in turn is proportional to  $(\kappa_{\rm e} R)^{-2}$  (see Chap. 3). That means  $W \sim$  frequency  $\omega$  for classical waves. In any case,

$$\chi(\ln\Lambda) = d - 2,\tag{8.46}$$

and the integration of (8.39) yields

$$\Lambda = \left[\frac{\xi_{\infty}(W)}{M}\right]^{2-d},\tag{8.47}$$

as the large  $\Lambda$  limiting behavior, i.e.,  $f(x) = x^{2-d}$  when  $\Lambda \gg 1$ . Here  $\xi_{\infty}(W)$ is a length that depends only on the randomness. In 2D,  $\xi_{\infty}(W) = \xi(W)$ . In 3D,  $\xi_{\infty}(W) = \zeta(W)$ , the correlation length, in the extended regime and  $\xi_{\infty}(W) = \xi(W)$  in the localized regime. For the 1D case, where M = 1 (the smallest discrete unit of length) and  $\Lambda > 1$  always,  $\chi(\ln \Lambda) = -1$  exactly for all values of  $\Lambda$ . From (8.43) and (8.47), it is seen that in 3D,  $\Lambda$  behaves as  $M^{-1}$  when  $\Lambda$  is small but as M when  $\Lambda$  is large. Thus in 3D the finite size scaling function can have a negative branch (small  $\Lambda$ , see (8.42)), and a positive branch (large  $\Lambda$ , see (8.46)). In 2D, however,  $\Lambda$  behaves as  $M^{-1}$  when  $\Lambda$  is small and as a constant in the large  $\Lambda$  limit.

### 8.3.3 Relation to the Scaling Functions in 1D, 2D, and 3D

By combining the behaviors of  $\Lambda$  at the two limits, it can be concluded that  $\Lambda = f(x)$ , with  $x = \xi(W)/M$  in the localized regime and  $x = \zeta(W)/M$  in the delocalized regime.  $\Lambda$  is clearly a quantity very similar to  $\gamma$ . The scaling

variable  $x = \xi_{\infty}(W)/M$  is a combination of both the randomness W and size M, and the domain of positive x values is completely covered by varying M, no matter what the value of W [and consequently  $\xi_{\infty}(W)$ ]. Therefore,  $\Lambda$  may be regarded as a function of only one variable, which ensures the validity of the scaling relation (8.39). It is also seen that  $\Lambda$  behaves the same as  $\gamma$  when both are large.  $\chi(\ln \Lambda)$  is therefore similar to  $\beta(\ln \gamma)$ . To relate the two so that  $\beta(\ln \gamma)$  may be explicitly calculated, we let M = L and write

$$\beta = \frac{\mathrm{d}(\ln \gamma)}{\mathrm{d}(\ln L)} = \frac{\mathrm{d}(\ln \gamma)}{\mathrm{d}(\ln \Lambda)} \frac{\mathrm{d}(\ln \Lambda)}{\mathrm{d}(\ln L)} = -\beta_1 (\ln \gamma) \chi(\ln \Lambda).$$
(8.48)

Since in the 1D case the value of  $\chi(\ln \Lambda)$  is identically -1,  $\beta_1(\ln \gamma)$  denotes the  $\beta$  function in 1D. In 2D and 3D the  $\beta$  function may be obtained from  $\beta_1(\ln \gamma)$  plus the numerical simulation of  $\chi(\ln \Lambda)$ .

To calculate  $\beta_1(\ln \gamma)$ , it is necessary to have an expression for  $\gamma$ . For a conductor,  $\gamma$  should be  $\infty$  in the absence of inelastic scattering. However, for an insulator  $\gamma$  is proportional to  $\exp(-2L/\xi) = \exp(-2/\Lambda)$ , remembering L = M. Based on an analogy with the tunnel junction, Landauer (1970) proposed that  $\gamma$  may be expressed in terms of the 1D transmission coefficient,  $|\tau|^2 = \exp(-2/\Lambda)$ , as

$$\gamma = \frac{|\tau|^2}{1 - |\tau|^2}.\tag{8.49}$$

This definition of  $\gamma$  is seen to satisfy the conditions for both the conducting limit and the insulating limit, since  $|\tau|^2 = 1$  for a conductor so that  $\gamma \to \infty$ , and  $|\tau|^2 \to 0$  as  $L \to \infty$  for an insulator, so  $\gamma \sim |\tau|^2$ . A more thorough discussion of the Landauer formula, including its generalization to 2D and 3D, is given in Chap. 11. Here it is easy to see that  $\gamma$  as given by (8.49) is a function of  $\Lambda$ . Therefore  $\beta_1(\ln \gamma)$  may be easily calculated to be:

$$\beta_{1} = -\frac{\mathrm{d}\ln\gamma}{\mathrm{d}\ln\Lambda} = -\frac{\Lambda}{\gamma}\frac{\mathrm{d}\gamma}{\mathrm{d}\Lambda} = -\frac{\Lambda}{\gamma}\frac{\mathrm{d}\gamma}{\mathrm{d}\tau}\frac{\mathrm{d}\tau}{\mathrm{d}\Lambda}$$
$$= (1+\gamma)\ln\left(\frac{\gamma}{1+\gamma}\right). \tag{8.50}$$

Since  $\gamma = \exp(\ln \gamma)$ ,  $\beta_1$  can thus be plotted as a function of  $\ln \gamma$  as shown in Fig. 8.2. For  $\gamma$  large,

$$\beta_1 \simeq -\gamma \ln\left(1+\frac{1}{\gamma}\right) \to -1,$$
(8.51)

and for  $\gamma$  small,  $\beta_1 \sim \ln \gamma$ , which are exactly the expected limits.

Figures 8.3a,b give the numerically calculated values of  $\Lambda$  as a function of  $\xi_{\infty}(W)/M$  for the 2D and 3D Anderson models. Data for many different values of W are included. A numerical approach for the evaluation of  $\xi_{\rm M}$ , called the recursive Green function method, is detailed in the solution to Problem 8.1. Here the value of  $\xi_{\infty}(W)$  is chosen to fit all the data onto a single curve. The fitted values of  $\xi_{\infty}(W)$  are shown in the inset as a function of W, expressed in



**Fig. 8.2.** The 1D scaling function  $\beta_1(\ln \gamma)$ .



Fig. 8.3. (a) Numerically calculated  $\xi_M/M$  (shown as  $\lambda_M/M$ ) plotted as a function of  $\xi/M$  (shown as  $\lambda_{\infty}/M$ ) for the 2D Anderson model at the band center. The value of  $\xi(W)$ , determined by fitting all data with different W's onto a single curve, is shown in the inset. Here  $W = 2\sigma_w a^2$ . Continued.



Fig. 8.3. (b) Same as (a) for the 3D Anderson model at the band center. There are now two segments. The top segment corresponds to delocalized states and is obtained for  $\sigma_w^2 a > 8$ . The fitted value of  $\xi_{\infty}$  in this case is the correlation length  $\zeta(W)$ . The lower segment corresponds to localized states with  $\sigma_w^2 a < 8$ . The fitted values of  $\xi_{\infty}$  in this case corresponds to  $\xi(\omega)$ . Both  $\zeta$  and  $\xi$  are shown as a function of W in the inset. Both figures are reproduced from MacKinnon and Kramer (1981).

units of  $V = \hbar^2/2ma^2(W/V = 2\sigma_w a^2)$ . The data show convincingly that the scaling hypothesis is valid at least to the first approximation (numerically it is not possible to verify the accuracy of the scaling hypothesis to all orders). For the 2D case  $\Lambda$  is seen (in Fig. 8.3a) to be a monotonically increasing function of  $x = \xi_{\infty}/M$ . Therefore,  $\chi = d(\ln \Lambda)/d(\ln M)$  is always negative, since  $d\ln M = -d\ln x$ . As a result, from (8.48)  $\beta(\ln \gamma)$  for the 2D case always has the same sign as  $\beta_1(\ln \gamma)$ . For the 3D case, the function  $\Lambda$  is composed of two different segments with opposite trends as a function of increasing x.  $\chi$  can therefore be both positive and negative as deduced before. In this case the upper segment corresponds to  $\sigma_w a^2 \leq 8$ , whereas the lower segment is for  $\sigma_w a^2 \geq 8$ . The fitted values of  $\xi(W)$  and  $\zeta(W)$  are both seen to increase dramatically at  $\sigma_w a^2 \cong 8$  (see inset). It follows that  $\sigma_w a^2 \cong 8$  defines the mobility edge for the 3D Anderson model. By evaluating  $\chi$  numerically for the 2D and 3D cases, the resulting 2D and 3D scaling functions may be obtained from (8.48) and (8.50). They are plotted in Fig. 8.4. Similarity with the



Fig. 8.4. The numerically calculated scaling function from the data shown in Figs. 8.3a, b. Note the similarity with Fig. 8.1. Here  $\gamma$  is shown as g. The figure is reproduced from MacKinnon and Kramer (1981).

expected behavior, shown in Fig. 8.1, is noted. In 3D,  $\gamma_c \cong \exp(-3.4) \cong 0.03$ , which is on the same order as but smaller than that predicted by the Ioffe-Regel criterion. The slope *s* at which the 3D  $\beta$  function crosses zero may also be evaluated. It is approximately 1 - 1.5. From (8.37), that directly gives the exponent for the divergence of the localization length at the mobility edge.

# 8.4 Universality and Limitations of the Scaling Theory Results

While the finite-size scaling calculations have been carried out by using the Anderson model, the scaling function should be the same for classical scalar waves because the basis of the scaling argument – that the effect of randomness (or frequency) is manifest only through a length scale, and that the

dimensionless scaling parameter  $\ln \gamma$  is dependent only on the ratio x between that length scale and the sample size – is wave-type as well as lattice-type independent. In other words,  $\ln \gamma(x)$  may be regarded as the composition of two functions. The outer function depends only on the sample's spatial dimensionality, but the inner variable x is a model-dependent function of W,  $\omega$ , wave type, lattice type, etc. The scaling theory conclusions are based on the universality of the outer function.

# 8.4.1 Localization in 3D : Lattice vs. Continuum

While the scaling theory should hold for both quantum and classical waves, it is important to point out that in 3D, it is entirely possible for classical scalar waves not to have the  $\beta < 0$  branch of the scaling function, i.e., no localized states, unless some other condition(s) is satisfied, such as the ratio between the dielectric constants of the scatterers and the matrix exceeding a critical value (see Chap. 9 for more discussion on this point). This is because, whereas in a lattice model wave localization is aided by the presence of the band edges, where there is a deficit in the density of states so that wave localization is facilitated by the resulting confinement effect (see Chap. 1), for waves in the continuum the absence of an upper band edge means that the coherent backscattering effect must work by itself, thus making localization more difficult. One is reminded in this regard that classical waves can never be localized at the lower band edge, i.e.,  $\omega \to 0$ , because the effect of randomness is always seen by the wave as  $\omega^2[\varepsilon(\mathbf{r})-1]/v_0^2$ , so the scattering vanishes in the zero-frequency limit. This point should also be obvious from the frequency dependence of Rayleigh scattering.

Although the localization of classical scalar waves in 3D is potentially more difficult, but if localization *does occur*, it is expected to have the same  $\beta(\ln \gamma)$  function as shown in Fig. 8.3b; that is, universality is not affected.

# 8.4.2 Implicit Assumptions of the Scaling Theory

Before leaving this chapter, some implicit assumptions of the scaling theory should be made explicit so that we are aware of the theory's limitations. The three most important assumptions of the scaling theory are homogeneity, isotropy, and the absence of inelastic scattering. Homogeneity means that the density of the scatterers is constant beyond a finite sampling scale. Most of the disordered media are expected to be homogeneous. An example of an *inhomogeneous* system is one in which the wave scatterers are distributed as a fractal; i.e., the scatterer density decays as a power law function of the sampling scale. In such a system a wave can delocalize even in 2D, because whereas the effect of the coherent backscattering varies as  $\ln L$  in the weak scattering limit [as derived in the last chapter; see (7.19) and (7.20)], the density of the scatterers decays even faster (as a power law), thus making it possible for the wave to delocalize (Zhang and Sheng 1991b).

Isotropy is another important assumption. In the Anderson model, there are two ways to introduce anisotropy. One is to make each scatterer anisotropic, e.g., by making the effective mass anisotropic so that the scattering and consequently the mean free paths become anisotropic as well. This model proves to have the same scaling behavior as the isotropic case because one can scale the sample size in different directions by its respective anisotropic mean free paths. After scaling the model returns to being isotropic (Li et al. 1989). Therefore this type of anisotropy does not invalidate the scaling property. Another way to introduce anisotropy is to make the diagonal terms have both an isotropic and an anisotropic component. For example, the anisotropic component can be a randomly layered system where the randomness is correlated over infinite distances in the directions parallel to the layers. Therefore, while each scatterer is isotropic, the anisotropy resides in its (infinite-range) spatial correlation. In fact, such models are relatively realistic for most of the Earth's subsurface, which is known to be predominantly layered. The coherent backscattering effect then has two components, one isotropic and the other one anisotropic. It can be shown that scaling breaks down in this case, and novel localization behavior emerges (Xue et al. 1989).

Another implicit assumption of the scaling theory is the absence of inelastic scattering. The effect of inelastic scattering will be discussed in Chap. 10, in relation to the physical manifestations of localization.

#### **Problem and Solution**

**Problem 8.1** Give a formulation of the numerical approach for the calculation of  $\xi_{\rm M}$ .

**Solution** In order to illustrate the numerical scheme, let us restrict ourselves to the 2D Anderson model defined by (2.51), (2.53), and (2.54). Consider a strip of width M and length  $N \gg M$ . Both N and M are in units of the lattice (discretization) constant. This strip is the disordered material whose  $\xi_{\rm M}$  we wish to measure. In order to do so, the strip is connected on both the left and the right to two perfect leads of the same width but infinite length. The perfect leads have no disorder, i.e.,  $\sigma(\mathbf{l}) = 0$ . In analogy with a 1D system where  $\xi$ is calculated as  $-2Na/\langle \ln |\tau|^2 \rangle_c$ ,  $\xi_M$  will be calculated similarly. However, the calculation of  $\tau$  is more complicated here because there are more than one plane wave for a given frequency. To see that, we recall that in a 2D perfect lead,  $e(\mathbf{k}) = 2(2 - \cos k_x a - \cos k_z a)/a^2$  so that at a fixed frequency,  $\omega = \hbar e(\mathbf{k})/2m$ , there can be multiple  $(k_x, k_z)$  combinations that satisfy the dispersion relation. Here x is defined as the width direction and z the length direction of the strip. We shall impose the periodic boundary condition in the width direction, so that there is only a discrete set of allowed  $k_x$  values, given by

$$k_x^{(i)} = \frac{2\pi}{aM}i, \qquad i = 1, \cdots, M.$$
 (P8.1)

 $k_z^{(i)}$  is then defined as the value of  $k_z$  that satisfies  $e(\mathbf{k})|_{k_x^{(i)}} = 2m\omega/\hbar$ . However, for fixed  $\omega$  and  $k_x^{(i)}$  there is not always a real solution for  $k_z^{(i)}$ . These cases correspond to  $k_z$  being imaginary, i.e., an evanescent mode. It follows that the number of propagating modes,  $M_0$ , is less than M in general. For the propagating modes on both the left and right sides, a transmission coefficient  $\tau_{ij}$  may be defined, where *i* corresponds to the left-hand side  $k_z^{(i)}$  and *j* corresponds to the right-hand side  $k_z^{(j)}$ . The total transmission  $|\tau|^2$  is given by

$$|\tau|^{2} = \sum_{i}^{M_{0}} \sum_{j}^{M_{0}} \tau_{ij} \tau_{ij}^{*}$$
  
= Tr(\(\tau\tau\tau\)), (P8.2)

where  $\dagger$  means Hermitian conjugate. In getting to the last line, the *j* summation is seen to be just the matrix multiplication, and Tr (the operation of taking trace of a matrix) is the summation over *i*. This  $|\tau|^2$  is used to calculate  $\xi_{\rm M}$  as  $-2aN/\langle \ln |\tau|^2 \rangle_c$ .

To calculate  $\tau_{ij}$ , Lee (1981) pioneered a recursive Green function method that is accurate and numerically stable. This method is presented below.

First,  $\tau_{ij}$  has to be related to the Green function. On physical ground,

$$\tau_{ij} \sim G_{ij}^+(0, N+1),$$
 (P8.3)

where  $G_{ij}^+(0, N+1)$  stands for the Green function that gives the amplitude of the wave in momentum channel j on the right side, at z = N+1, that started as a plane wave in momentum channel i on the left side. The Green function of (P8.3) is a bit unusual in that the indices i,j indicate the momentum channels of the perfect lead, whereas 0, N+1 indicate the real space positions of the incoming and the outgoing waves. The question now is, what is the proportionality factor? In Chap. 2, the 1D lattice Green function is shown to exhibit the form (2.65b)

$$G_0^+(\omega, l-l') = -\frac{\mathrm{i}}{\partial\beta e(k)/\partial k|_{k_0}} \exp(\mathrm{i}k_0|l-l'|),$$

where  $k_0$  is the solution to the equation  $\kappa_0^2 = \beta e(k_0), e(k)$  being the 1D dispersion function  $2(1 - \cos ka)/a^2$ . From this 1D expression, we see that since the  $\tau_{ij}$  is defined relative to the incoming wave flux in the z direction,  $G_{ij}^+$  first has to be multiplied by  $i\bar{k}_z(i)$  so that the incident plane wave has unit amplitude (apart from a phase factor), and then it has to be multiplied by  $\sqrt{\bar{k}_z(j)/\bar{k}_z(i)}$  in order to normalize the outgoing z flux relative to the incoming z flux. Here

$$\bar{k}_z(i) = \left. \frac{\partial \beta e(\mathbf{k})}{\partial k_z} \right|_{k_x^{(i)}, k_z^{(i)}}$$

is proportional to the group velocity in the z direction, and  $e(\mathbf{k})$  here is the 2D expression  $2(2 - \cos k_x a - \cos k_z a)/a^2$ . The net result is

$$|\tau_{ij}| = a\sqrt{\bar{k}_z(i)\bar{k}_z(j)} |G^+_{ij}(0,N+1)|.$$
(P8.4)

It is noted that  $G_{ij}^+(0, N+1)$  is a 2D Green function, which is dimensionless in the notations of this book. The task now is to calculate the Green function. To do so, we show first how to calculate  $G_{ij}^{L}(0,0)$ ,  $G_{ij}^{L}(1,1)$ , ...,  $G_{ij}^{L}(N,N)$ , which are the half-space (the left half) Green functions built up from the perfect-lead Green function from the left side, from which  $G_{ij}(0, N+1)$  may be obtained (superscript + is suppressed here and in the following).

First consider  $G_{ij}^{\rm L}(0,0)$ . This is the value of the Green function at the end of the perfect lead on the left side. It should be emphasized that  $G_{ij}^{\rm L}(0,0) \neq G_{ij}(0,0)$ , which is the true Green function of the full system at (0,0).  $G_{ij}^{\rm L}(0,0)$ has two properties. First, it is diagonal in the i, j indices, i.e.,

$$G_{ij}^{\rm L}(0,0) = \delta_{i,j} g_i^{-1}, \qquad (P8.5)$$

because *i* and *j* are the momentum eigenstates of the perfect lead, so they must be orthogonal, and the task is to solve for  $g_i^{-1}$ . Second, a new ending to the perfect lead may be generated by joining onto it a strip with  $\sigma(\mathbf{l}) = 0$ , where  $\mathbf{l} = (0, 1, \dots, M - 1)$  in the *x* direction. The new Green function,  $[G_{ij}^{\mathrm{L}}(0, 0)]'$ , is expressible by the perturbation series

$$\begin{bmatrix} G_{ij}^{\rm L}(0,0) \end{bmatrix}' = G_{ij}^{0} + G_{ij}^{0} v_0 G_{ij}^{\rm L}(0,0) v_0 G_{ij}^{0} + G_{ij}^{0} v_0 G_{ij}^{\rm L}(0,0) v_0 G_{ij}^{0} v_0 G_{ij}^{\rm L}(0,0) v_0 G_{ij}^{0} + \cdots, \quad (\text{P8.6})$$

which should be *identical* to  $G_{ij}^{L}(0,0)$  since the addition of a perfectly ordered strip to a perfect lead still results in a perfect lead. Here  $v_0 = -(\beta/a^2)a^d = -\beta$  is the constant off-diagonal matrix element [see (2.51)], and

$$G_{ij}^{0} = \frac{\delta_{i,j}}{\kappa_0^2 a^2 - 2\beta (1 - \cos k_x^{(i)} a)}$$
(P8.7)

is the Green function for the (isolated) ordered 1D strip that has been added. The meaning of this series is fairly straightforward: It involves  $G_{ij}^0$  plus all the interaction paths with the rest of the perfect lead, which is completely given by  $G_{ij}^{\rm L}(0,0)$ . The interaction in the z direction involves just the multiplication of  $G_{ij}^0$  of the isolated strip by  $v_0$  in linking it to the perfect lead, and the same in the other direction. The perturbation series may be summed to get

$$G_{ij}^{L}(0,0) = G_{ij}^{0} \left[ 1 - v_0^2 G_{ij}^{L}(0,0) G_{ij}^0 \right]^{-1}$$
$$= \left[ \left( G_{ij}^0 \right)^{-1} - v_0^2 G_{ij}^{L}(0,0) \right]^{-1}.$$
 (P8.8)

Since everything is diagonal in i, j, (P8.8) represents a simple algebraic equation to get  $g_i$  and hence  $G_{ij}^{L}(0,0)$ :

$$g_i^{-1} = \frac{1}{m_0 a^2 - 2v_0 \cos k_x^{(i)} a - v_0^2 g_i^{-1}},$$
 (P8.9)

where  $m_0 a^2 = \kappa_0^2 a^2 - 2\beta$ . Solution of (P8.9) yields

$$G_{ij}^{\rm L}(0,0) = g_i^{-1} \delta_{i,j}$$
  
=  $\frac{2\delta_{i,j}}{m_0 a^2 - 2v_0 \cos k_x^{(i)} a + \sqrt{(m_0 a^2 - 2v_0 \cos k_x^{(i)} a)^2 - 4v_0^2}}$ . (P8.10)

The + sign in the denominator is chosen because the square root is imaginary, and we want the retarded Green function, which must have a negative imaginary part.

The technique of generating (P8.8) may now be used to calculate  $G_{ij}^{L}(1,1)$ . In the operator notation, adding a *disordered* strip can be done in the same way as through the perturbation series, (P8.6). It follows that

$$G_{ab}^{\rm L}(1,1) = \left\{ \left[ \mathbf{G}^{0}(1,1) \right]^{-1} - v_0^2 \mathbf{G}^{\rm L}(0,0) \right\}_{ab}^{-1}, \qquad (P8.11)$$

where a, b are the real space coordinates along the x axis, and  $\mathbf{G}^{0}(1, 1)$  is the Green function of the first strip of the disordered medium. In real space, it is given by

$$\begin{bmatrix} \mathbf{G}^{0}(1,1) \end{bmatrix}^{-1} = \begin{pmatrix} [m_{0} - \sigma_{1}(1)]a^{2} & v_{0} & 0 & \cdots \\ v_{0} & [m_{0} - \sigma_{1}(2)]a^{2} & v_{0} \\ 0 & v_{0} & [m_{0} - \sigma_{1}(3)]a^{2} & \cdots \\ \vdots & \vdots & \vdots \end{pmatrix},$$
(P8.12)

where  $\sigma_1(m)$  means the random perturbation at site ma of strip 1. To be consistent,  $\mathbf{G}^{\mathrm{L}}(0,0)$  also has to be expressed in real space; i.e.,  $G_{ab}^{\mathrm{L}}(0,0)$  may be obtained from  $G_{ij}^{\mathrm{L}}(0,0) = \delta_{i,j}g_i^{-1}$  as

$$G_{ab}^{\rm L}(0,0) = \frac{1}{M} \sum_{i=1}^{M} g_i^{-1} \exp\left[\mathrm{i}k_x^{(i)}(x_a - x_b)\right],\tag{P8.13}$$

where  $x_a$ ,  $x_b$  is an integral multiple of a, ranging from a to Ma. From (P8.11)  $G_{ab}^{L}(n,n)$  can be obtained recursively as

$$\mathbf{G}^{\mathrm{L}}(n,n) = \left\{ [\mathbf{G}^{0}(n,n)]^{-1} - v_{0}^{2}\mathbf{G}^{\mathrm{L}}(n-1,n-1) \right\}^{-1}.$$
 (P8.14)

Here  $\mathbf{G}^{0}(n, n)$  is the Green function of the *n*th strip of the disordered medium. Numerically, the price one has to pay is a matrix inversion at each step.

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This can continue until  $G_{ab}^{L}(N, N)$  is reached. The strip N is linked to the perfect lead to the right and the strip N-1 to the left. That means

$$\mathbf{G}(N,N) = \left\{ [\mathbf{G}^{0}(N,N)]^{-1} - v_{0}^{2}\mathbf{G}^{\mathrm{L}}(N-1,N-1) - v_{0}^{2}\mathbf{G}^{\mathrm{R}}(N+1,N+1) \right\}^{-1}, \qquad (P8.15)$$

where the superscript L is dropped on  $\mathbf{G}(N, N)$  since it is now a Green function of the full system.  $\mathbf{G}^{\mathrm{R}}(N+1, N+1)$  is the Green function of the perfect lead on the right side. In real space, it is given by

$$G_{ab}^{\rm R}(N+1,N+1) = \frac{1}{M} \sum_{j=1}^{M} g_j^{-1} \exp\left[\mathrm{i}k_x^{(j)}(x_a - x_b)\right].$$
 (P8.16)

Having obtained the diagonal terms of the Green functions, we proceed to calculate  $\mathbf{G}(0, N + 1)$ . To that effect, it is noted that

$$\mathbf{G}^{\mathrm{L}}(0,1) = \mathbf{G}^{\mathrm{L}}(0,0)v_{0}\mathbf{G}^{0}(1,1) + \mathbf{G}^{\mathrm{L}}(0,0)v_{0}\mathbf{G}^{0}(1,1)v_{0}\mathbf{G}^{\mathrm{L}}(0,0)v_{0}\mathbf{G}^{0}(1,1) + \cdots = \mathbf{G}^{\mathrm{L}}(0,0)v_{0}\mathbf{G}^{\mathrm{L}}(1,1),$$
(P8.17)

and  $\mathbf{G}^{\mathrm{L}}(0,2) = \mathbf{G}^{\mathrm{L}}(0,1)\upsilon_0\mathbf{G}^{\mathrm{L}}(2,2)$ . It follows straightforwardly that

$$\mathbf{G}(0, N+1) = \mathbf{G}^{\mathrm{L}}(0, N)v_0\mathbf{G}(N+1, N+1)$$
(P8.18)

may be obtained recursively. Once  $\mathbf{G}(0, N+1)$  is evaluated,  $|\tau|^2$  is given by

$$|\tau|^2 = a^2 \sum_{i}^{M_0} \sum_{j}^{M_0} \bar{k}_z(i) \bar{k}_z(j) \left| G_{ij}^+(0, N+1) \right|^2, \qquad (P8.19)$$

where

$$G_{ij}^{+}(0, N+1) = \exp\left[\mathrm{i}k_{z}^{(j)}(N+1)\right] \sum_{x_{a}}^{Ma} \sum_{x_{b}}^{Ma} \times G_{ab}^{+}(0, N+1) \exp\left[\mathrm{i}(k_{x}^{(i)}x_{a}-k_{x}^{(j)}x_{b})\right].$$
 (P8.20)

The values of  $x_a, x_b$  range from a to Ma. By configurationally averaging  $\ln |\tau|^2$ ,  $\xi_{\rm M}$  can be easily obtained as  $-2Na/\left\langle \ln |\tau^2| \right\rangle_c$ .
# 9.1 The Self-Consistent Theory of Localization

The purposes of this chapter are to substantiate the scaling theory results by continuing the Green function development discussed in Chap. 7, and to fill in the details about the localized states and the approach to localization. It may be recalled that aside from the description of 1D localization, the estimates of 2D and 1D localization lengths in Chap. 7 [(7.23) and (7.24)] have been based on the weak scattering approximation. Besides being inconsistent, the mathematical framework cannot demonstrate how the localized states come into being. Therefore, further development obviously calls for the extension of the Green function formalism into the strong scattering regime, where localization is possible.

## 9.1.1 Transition to the Strong Scattering Regime

The starting point of our consideration is (7.17), without the approximate evaluation of  $M_0^{(M C)}$ :

$$\frac{1}{D(\omega)} = \frac{1}{D^{(\mathrm{B})}(\omega)} + \frac{\pi \rho_{\mathrm{e}}(\omega)}{8K_0^2 (\mathrm{d}\kappa_{\mathrm{e}}^2/\mathrm{d}\kappa_0^2)(-\mathrm{Im}\Sigma^+)} M_0^{(\mathrm{MC})}.$$

Multiplying both sides by  $D(\omega)D^{(B)}(\omega)$  yields

$$D^{(\mathrm{B})}(\omega) = D(\omega) + D(\omega)D^{(\mathrm{B})}(\omega) \frac{\pi \rho_{\mathrm{e}}(\omega)}{8K_0^2(-\mathrm{Im}\Sigma^+)(\mathrm{d}\kappa_{\mathrm{e}}^2/\mathrm{d}\kappa_0^2)} M_0^{(\mathrm{MC})}.$$

After rearrangement, this becomes

$$D(\omega) = D^{(B)}(\omega) \left[ 1 - \frac{\pi \rho_{e}(\omega)}{8K_{0}^{2}(-\mathrm{Im}\Sigma^{+})(\mathrm{d}\kappa_{e}^{2}/\mathrm{d}\kappa_{0}^{2})} D(\omega)M_{0}^{(\mathrm{MC})} \right]$$
$$= D^{(B)}(\omega)(1-C).$$
(9.1)

From (7.11) and (7.12), it is straightforward to write the dimensionless correction term C as

$$C = \frac{\pi \rho_{\rm e}(\omega)}{8K_0^2(-\mathrm{Im}\Sigma^+)(\mathrm{d}\kappa_{\rm e}^2/\mathrm{d}\kappa_0^2)} D(\omega)M_0^{\rm (M\,C)}$$
  
$$= \frac{\mathrm{i}}{4K_0^2} \frac{L^{3d}}{N^3} \iint \frac{\mathrm{d}\mathbf{k}_{\rm i}}{(2\pi)^d} \frac{\mathrm{d}\mathbf{k}_{\rm f}}{(2\pi)^d}$$
$$\times (\Delta G_{\rm e})_{\mathbf{k}_{\rm i}} (\Delta G_{\rm e})_{\mathbf{k}_{\rm f}}^2 \left[ \nabla_{k_{\rm i}} e(\mathbf{k}_{\rm i}) \cdot \Delta \hat{\mathbf{k}} \right]$$
$$\times \left[ \nabla_{k_{\rm f}} e(\mathbf{k}_{\rm f}) \cdot \Delta \hat{\mathbf{k}} \right] \frac{-\mathrm{Im}\Sigma^+}{-\frac{\mathrm{i}\Delta\omega}{D(\omega)} + \frac{D^{\rm (B)}(\omega)}{D(\omega)} |\mathbf{k}_{\rm i} + \mathbf{k}_{\rm f}|^2}. \tag{9.2}$$

A crucial step in extending the Green function formalism from the weak scattering regime to the strong regime is to replace  $D^{(B)}(\omega)$  in the integrand of (9.2) by the renormalized diffusion constant,  $D(\omega)$ . That is, instead of using the irreducible vertex  $U^{(B)}$  in the evaluation of  $U^{(MC)}$  as given by (7.7), one uses the more accurate form  $U^{(B)} + U^{(MC)}$ . This has the effect of making the theory self-consistent as first proposed by Vollhardt and Wölfle (1980). Mathematically, this replacement is equivalent to extending the theory to the strong scattering regime. The correction term C after this replacement becomes

$$C = i \left(\frac{-\mathrm{Im}\Sigma^{+}}{4K_{0}^{2}}\right) \frac{L^{3d}}{N^{3}} \int \int \frac{\mathrm{d}\mathbf{k}_{\mathrm{i}}}{(2\pi)^{d}} \frac{\mathrm{d}\mathbf{k}_{\mathrm{f}}}{(2\pi)^{d}} \\ \times \left(\Delta G_{\mathrm{e}}\right)_{\mathbf{k}_{\mathrm{i}}} \left[\nabla_{k_{\mathrm{i}}} e(\mathbf{k}_{\mathrm{i}}) \cdot \Delta \hat{\mathbf{k}}\right] \\ \times \frac{1}{-\frac{\mathrm{i}\Delta\omega}{D(\omega)} + |\mathbf{k}_{\mathrm{i}} + \mathbf{k}_{\mathrm{f}}|^{2}} (\Delta G_{\mathrm{e}})_{\mathbf{k}_{\mathrm{f}}}^{2} \left[\nabla_{k_{\mathrm{f}}} e(\mathbf{k}_{\mathrm{f}}) \cdot \Delta \hat{\mathbf{k}}\right].$$
(9.3)

The term  $-i\Delta\omega/D(\omega)$  has two distinct limits when  $\Delta\omega \to 0$ . Since in the electronic case  $D(\omega)$  may be related to the conductivity by the Einstein relation, we can use the conductivity behavior to guide us in getting the limits. In the conducting regime,  $D(\omega)$  is always finite so

$$\lim_{\Delta\omega\to 0} \frac{-\mathrm{i}\Delta\omega}{D(\omega)} = 0.$$

# 9.1.2 Response of the Localized State at Finite Modulation Frequency

On the other hand, in the insulating (localized) regime the dc diffusivity  $D(\omega) = 0$  by definition, so the limiting behavior is not immediately clear. However, in the solution to Problem 9.1 it is shown that at finite modulation frequency  $\Delta \omega$  the conductivity may be regarded as complex, in which the imaginary part represents the manifestation of displacement current. Therefore even if the real part of the conductivity is zero, there can still be a 90° out-of-phase (imaginary) component of the conductivity whose physical origin lies in the polarizability of the localized state, in exact analogy to the polarizability of a bound electron in an atom. In that case

$$D(\omega) = -i\Delta\omega \text{ (length scale)}^2. \tag{9.4a}$$

Physically, the only length scale in the problem is the size of the localized state as characterized by the localization length  $\xi$ . Therefore

$$\frac{-\mathrm{i}\Delta\omega}{D(\omega)} = \nu^2 \propto \xi^{-2} \tag{9.4b}$$

in the localized regime. The (dimensionless) constant of proportionality,  $\nu\xi$ , evaluated in the 1D classical wave case is of the order of 2 (~1.84). For the Anderson model the comparison of numerical results with those calculated from the self-consistent theory gives a constant that is much larger (~5–10). This difference could be due to the difference in the dispersion relations. In the following this proportionality constant in various cases will be fixed by comparison to simulation results whenever available.

#### 9.1.3 Qualitative Evaluation of the Correction Term

From (9.1) and (9.4b), the condition for determining the mobility edge is C = 1 with  $\nu^2 = 0$  (i.e., divergent localization length). Since  $D(\omega)$  cannot be negative, the same C = 1 condition gives an equation for  $\nu^2$  in the localized regime. In fact, a quick glimpse of the consequences of the self-consistent theory may be obtained by evaluating the multiple integral of (9.3) in the manner prescribed in Chap. 7. That is, by setting  $\mathbf{k}_i = -\mathbf{k}_f$  so that  $\nabla_{k_i} e(\mathbf{k}_i) = -\nabla_{k_f} e(\mathbf{k}_f)$ , and approximating  $(\Delta G_e)_{\mathbf{k}}^2 \cong [(\Delta G_e)_{\mathbf{k}}/(-\mathrm{iIm}\Sigma^+)](N/L^d)$ , (9.3) becomes

$$C = \frac{1}{K_0} \int \frac{\mathrm{d}\mathbf{Q}}{(2\pi)^d} \frac{1}{\nu^2 + Q^2}.$$
 (9.5a)

It may be argued that there is a physical upper cutoff for the magnitude of Q, defined by the width of the coherent backscattering peak,  $\alpha/l$ , where  $\alpha$  is some adjustable constant. Alternatively, one can argue that the diffusive behavior exists only for length scale > l, and as a result the maximum value of Q is defined by  $\alpha/l$ . In any case, by using such a cutoff, we get

$$C = \frac{S_d}{(2\pi)^d K_0} \begin{cases} \frac{\alpha}{l} - \nu \tan^{-1} \frac{\alpha}{\nu l} & 3D\\ \frac{1}{2} \ln \left( 1 + \frac{\alpha^2}{\nu^2 l^2} \right) & 2D\\ \frac{1}{\nu} \tan^{-1} \frac{\alpha}{\nu l} & 1D, \end{cases}$$
(9.5b)

where  $S_d = 4\pi$ ,  $2\pi$ , 2 in 3D, 2D, 1D, respectively. It is seen that in 3D,  $\nu = 0$ and C = 1 give the mobility edge equation

$$l = \frac{\alpha}{2\pi^2 K_0},\tag{9.6}$$

which yields a critical randomness for every frequency. In 2D and 1D,  $\nu$  has to be finite in order for C not to diverge, and the equations for the localization length are

$$\frac{1}{\nu} = \frac{l}{\alpha} [\exp(4\pi K_0) - 1]^{1/2}, \quad 2D$$
(9.7a)

$$\pi K_0 = \frac{1}{\nu} \tan^{-1} \frac{\alpha}{\nu l}.$$
 1D (9.7b)

Although the approximate evaluation of C in the above manner can give us all the qualitative localization behavior, it is nevertheless desirable to have a look at the consequences of the theory with as little approximation as feasible, so that the true nature of the self-consistent theory may be manifest. Below we treat the Anderson model and the classical scalar wave case separately.

# 9.2 Localization Behavior of the Anderson Model

The C factor involves a 2d-fold integral for the MC diagram contributions. To locate the mobility edge and to calculate the localization length in d = 3 and 2, this integral is difficult to evaluate due to the nonspherical (noncircular) Brillouin zone. What one can do is to define a wave vector  $\bar{k}$  that is the average over the constant energy surface. In accordance with the rules worked out in the solution to Problem 9.2, that means in 3D

$$a\,\bar{k}(E) = \frac{16\int_0^{\pi} \int_0^{\pi} \int_0^{\pi} \mathrm{d}x\,\mathrm{d}y\,\mathrm{d}z\,\sqrt{x^2 + y^2 + z^2}\,\sqrt{\sin^2 x + \sin^2 y + \sin^2 z}}{\frac{\times\delta\left[E + 2\cos x + 2\cos y + 2\cos z\right]}{S_0(E)}},$$
(9.8a)

where

$$S_0(E) = 16 \int_0^{\pi} \int_0^{\pi} \int_0^{\pi} dx \, dy \, dz \, \sqrt{\sin^2 x + \sin^2 y + \sin^2 z}$$

$$\times \delta \left[ E + 2\cos x + 2\cos y + 2\cos z \right],$$
(9.8b)

is the constant E surface area in the **k**-space, in units of  $1/a^2$ . Figure 9.1 gives a plot of  $a \bar{k}(E)$  vs. E, for -6 < E < 6. Here E = 2u, where u is the dimensionless energy/frequency parameter defined in Chap. 2.

Another factor that should be considered in the strong scattering limit is the width of  $(\Delta G_{\rm e})_{\rm k}$ . Since  $-{\rm Im}\Sigma^+$  may be appreciable, the delta-function approximation is no longer adequate, and one should use the accurate representation

$$\frac{L^d}{N} (\Delta G_{\rm e})_{\bf k} = \frac{2\,\mathrm{i}\,\mathrm{Im}\Sigma^+}{[\kappa_0^2 - \mathrm{Re}\Sigma^+ - e({\bf k})]^2 + (\mathrm{Im}\Sigma^+)^2}.$$
(9.9)



Fig. 9.1. Average wave vector  $\bar{k}$  over the constant energy surface for the 3D Anderson model plotted as a function of E = 2u, where u is the reduced frequency variable as defined in Fig. 2.5. Here  $\bar{k}$  is defined relative to the lower band edge.

# 9.2.1 Evaluation of the Maximally Crossed Diagrams' Contribution

With these considerations in mind, we can approach the evaluation of the maximally crossed (MC) diagrams' contribution by converting the **k** integrals into dEdS integrals as shown in Problem 9.2, where dS denotes surface integration over the constant-frequency surface. To see it explicitly, let us denote the double **k**-integral factor in C (see 9.3) as (MC); then

$$(MC) = i \frac{L^{3d}}{N^3} \frac{1}{(2\pi)^6} \int d\mathbf{k}_i \int d\mathbf{k}_f \left[ \nabla_{k_i} e(\mathbf{k}_i) \cdot \Delta \hat{\mathbf{k}} \right] \\ \times \left[ \nabla_{k_f} e(\mathbf{k}_f) \cdot \Delta \hat{\mathbf{k}} \right] (\Delta G_e)_{\mathbf{k}_i} (\Delta G_e)_{\mathbf{k}_f}^2 \frac{1}{\nu^2 + |\mathbf{k}_i + \mathbf{k}_f|^2} \\ \cong -\frac{8}{(2\pi)^6} \int_{E_l}^{E_u} dE_x \int_{E_l}^{E_u} dE_y \int dS_x \int dS_y f(E_x) f^2(E_y) \\ \times \frac{\cos \theta_x \cos \theta_y}{\nu^2 a^2 + x^2 + y^2 + 2xy \cos \theta_{xy}}.$$
(9.10)

Here in the last line the direction of  $\nabla_k e(\mathbf{k})$  is approximated as being the same as  $\mathbf{k}$ . This is certainly true at the lower band edge, where  $e(\mathbf{k}) = k^2$  so that  $\nabla_k e(\mathbf{k}) = 2\mathbf{k}$ . Also, in (9.10) the parameters are nondimensionalized, with  $y_0 = a^2 \Sigma^+$  denoting the dimensionless self-energy, and

$$f(E) = \frac{-\mathrm{Im}\,y_0}{E^2 + (\mathrm{Im}\,y_0)^2}$$

denoting the dimensionless form of  $(\Delta G_{\rm e})_{\bf k}$  to be used extensively below. In addition,

$$E_{l} = 2u - \operatorname{Re} y_{0} - 2d,$$
  

$$E_{u} = 2u - \operatorname{Re} y_{0} + 2d,$$
  

$$u = \frac{\kappa_{0}^{2}a^{2} - 2d}{2}, \qquad (\beta = 1)$$
(9.11)

and

$$x = a|\mathbf{k}_{i}|, \qquad y = a|\mathbf{k}_{f}|.$$

Here  $\theta_x$ ,  $\theta_y$  are the angles between  $\Delta \hat{\mathbf{k}}$  and  $\mathbf{k}_i$ ,  $\mathbf{k}_f$ , respectively.  $\theta_{xy}$  is the angle between  $\mathbf{k}_i$  and  $\mathbf{k}_f$ , and  $dS_x$ ,  $dS_y$  denote the constant  $e(\mathbf{k}_i)$ ,  $e(\mathbf{k}_f)$  surface differentials. In (9.10),  $x = a|\mathbf{k}_i|$  is related to  $E_x$  via the dispersion relation

$$E_x = \kappa_0^2 a^2 - \operatorname{Re} y_0 - a^2 e(\mathbf{k}_i)$$

and similarly for y to  $E_y$ . The dS integrals are thus coupled to the dE integrals. To facilitate the evaluation of (MC), let us approximate x, y by

$$x = a k_{i} (E'_{x}),$$
  
$$y = a \bar{k}_{f} (E'_{y}),$$

where  $\bar{k}_i$ ,  $\bar{k}_f$  are defined in accordance with (9.8a), and

$$E'_{x,y} = 2u - \operatorname{Re} y_0 - E_{x,y}$$

The approximation has the effect of decoupling the dS integrals from the dE integrations, i.e.,

$$(MC) \approx -\frac{8}{(2\pi)^6} \int_{E_l}^{E_u} dE_x \int_{E_l}^{E_u} dE_y f(E_x) f^2(E_y) \frac{S_0(E'_x)}{4\pi} \frac{S_0(E'_y)}{4\pi} \times \left\{ \int_0^{2\pi} d\phi_x \int_0^{2\pi} d\phi_y \int_{-1}^1 d\cos\theta_x \int_{-1}^1 d\cos\theta_y \times \frac{\cos\theta_x \cos\theta_y}{\nu^2 a^2 + x^2 + y^2 + 2xy \cos\theta_{xy}} \right\}.$$
(9.12a)

The angular integration in the curly brackets can be explicitly evaluated as illustrated in Problem 9.3, resulting in

$$\{\} = \frac{8\pi^2}{3} \frac{1}{xy} \left[ 1 - \frac{x^2 + y^2 + \nu^2 a^2}{4xy} \ln \frac{(x+y)^2 + \nu^2 a^2}{(x-y)^2 + \nu^2 a^2} \right].$$
(9.12b)

Substitution of (9.12b) into the previous expression yields

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(MC) = 
$$\frac{4}{3} \frac{1}{(2\pi)^6} A_3 \left[ u, y_0(\sigma_w, u), a^2 \nu^2 \right],$$
 (9.13a)

$$A_{3}\left[u, y_{0}(\sigma_{w}, u), a^{2}\nu^{2}\right] = \int_{E_{l}}^{E_{u}} dE_{x} \int_{E_{l}}^{E_{u}} dE_{y} f(E_{x}) f^{2}(E_{y}) S_{0}(E_{x}') S_{0}(E_{y}') (xy)^{-1} \\ \times \left[\frac{x^{2} + y^{2} + \nu^{2}a^{2}}{4xy} \ln \frac{(x+y)^{2} + \nu^{2}a^{2}}{(x-y)^{2} + \nu^{2}a^{2}} - 1\right].$$

$$(9.13b)$$

The subscript 3 is used to indicate that the function  $A_3$  is for 3D. The logarithmic divergence present in the integrand of (9.13b) is integrable; i.e., it is nondivergent upon integration. This fact is crucial for setting apart the 3D localization behavior from those of 2D, and 1D, as will be seen later.

An important point about the (MC) contribution is that it does not obey the particle-hole symmetry about the center of the band. This is traced to the form of  $U^{(\text{MC})} \propto 1/[\nu^2 + |\mathbf{k}_i + \mathbf{k}_f|^2]$ , where  $\mathbf{k}_i$  and  $\mathbf{k}_f$  are not the arguments of any trigonometric function. That means in our present calculation if the particle-hole symmetry were to be restored, one heuristic way would be to use (9.8a) only for  $-3 \le u \le 0$  ( $-6 \le E \le 0$ ). For  $3 \ge u \ge 0$  the magnitude of  $\mathbf{k}$ should be defined relative to the upper band edge. It has to be emphasized, however, that the problem of asymmetry with respect to u = 0 is not the result of any approximation used to evaluate  $M_0^{(\text{MC})}$ , but is rather in the form of  $U^{(\text{MC})}$ , which is only approximate due to the small Q expansion.

## 9.2.2 Mobility Edge in 3D Anderson Model

Apart from the factor (MC) in (9.2), for the  $2K_0$  factor in C accurate numerical evaluation may be carried out by expressing it in the following form:

$$2K_{0} = \frac{2}{(2\pi)^{3}} \int d\mathbf{k} \left[ \nabla_{k} e(\mathbf{k}) \cdot \Delta \, \hat{\mathbf{k}} \right]^{2} \\ \times \left\{ \frac{-\operatorname{Im}\Sigma^{+}}{[\kappa_{0}^{2} - \operatorname{Re}\Sigma^{+} - e(\mathbf{k})]^{2} + (\operatorname{Im}\Sigma^{+})^{2}} \right\}^{2} \\ = \frac{32}{(2\pi)^{3}a} B_{3} \left[ u, \, y_{0}(\sigma_{w}, \, u) \right],$$
(9.14)

where

$$B_3[u, y_0(\sigma_w, u)] = \int_{E_l}^{E_u} dE \ W_3(E') f^2(E), \qquad (9.15)$$

and

$$W_3(E') = \int_0^{\pi} \int_0^{\pi} \text{Re}\sqrt{1 - \left(\frac{E'}{2} + \cos x + \cos y\right)^2} dxdy,$$
(9.16)

with  $E' = 2u - \operatorname{Re} y_0 - E$ . Derivation of (9.15) and (9.16) is shown in the solution to Problem 9.2. The mobility edge equation can now be expressed as

$$1 = C = \frac{-\operatorname{Im}\Sigma^{+}}{(2K_{0})^{2}} (MC)$$
  
=  $\frac{1}{768} \left[-\operatorname{Im} y_{0}(\sigma_{w}, u)\right] \frac{A_{3} \left[u, y_{0}(\sigma_{w}, u), a^{2}\nu^{2} = 0\right]}{\left\{B_{3} \left[u, y_{0}(\sigma_{w}, u)\right]\right\}^{2}}.$  (9.17)

Equation (9.17) represents a condition on  $\sigma_w$  for a given u, and vice versa. It can be solved numerically by first tabulating the functions x, y, and  $W_3(E)$ and then carrying out the integrations in (9.15) and (9.13b). The results are plotted in Fig. 9.2 as the solid curve. At u = 0, the mobility edge is found to be at  $(\sigma_w a^2)_c = 6$ , which compares reasonably with the numerical simulation result of  $(\sigma_w a^2)_c = 8 - 8.25$  (Zdetsis et al. 1985). The comparison with numerical simulation over the whole mobility-edge curve shows general qualitative agreement. The fact that there is no adjustable parameter in the calculation makes it apparent that there is some overestimation of the localization effect in the self-consistent theory. This could be due to the approximate form of  $U^{(MC)}$  (owing to the small  $|\mathbf{k}_i + \mathbf{k}_f|$  expansion). Higher-order terms in  $|\mathbf{k}_i + \mathbf{k}_f|$ may provide a natural cutoff for the integral and make it smaller.

From Fig. 9.2 it is seen that as  $\sigma_w a^2$  increases from zero, the first occurrence of localization is near the band edge. Physically, this is due to the small group velocity of the band-edge states as well as the wave confinement effect caused by the deficit of states. As  $\sigma_w a^2$  continues to increase, however, the band-edge states are seen to delocalize. The reason for this behavior is that with the increase of randomness the CPA band edge moves toward larger



Fig. 9.2. Mobility-edge trajectory (solid line) for the 3D Anderson model calculated from the self-consistent theory. The u < 0 half is symmetric to the u > 0 half. The dashed line is the CPA band edge, defined by the condition  $-\text{Im}\Sigma^+ = 0$ . Numerical simulation results (Zdetsis et al. 1985) are shown as open circles.

frequencies, thus making the original band-edge states more like those inside the band. These states will eventually localize, however, with further increase in randomness. The rather flat top of the mobility edge curve means that as the randomness is increased toward  $(\sigma_w a^2)_c$  and crosses it, the mobility edge closes in very rapidly from the band edges, localizing all the states in the band.

# 9.2.3 Localization Length and Correlation Length in the 3D Anderson Model

When C < 1, (9.3) may be written alternatively as

$$D(\omega) = D^{(B)}(\omega) \frac{l}{\zeta} = D^{(B)}(\omega) (1 - C), \qquad (9.18)$$

(with  $\nu^2 = 0$  in the  $A_3$  factor) in accordance with the scaling concept of the correlation length  $\zeta$ , which determines the scale within which the classical diffusion behavior is significantly modified. It follows that

$$\frac{\zeta}{l} = \frac{1}{1-C}.\tag{9.19}$$

In the localized regime, the equation for the localization length is

$$1 = C = \frac{1}{768} \left[ -\text{Im} \, y_0(\sigma_w, \, u) \right] \frac{A_3 \left[ u, \, y_0(\sigma_w, \, u), \, \nu^2 a^2 \right]}{\left\{ B_3 \left[ u, \, y_0(\sigma_w, \, u) \right] \right\}^2}.$$
 (9.20)

At u = 0, the solution of (9.20) for  $\xi/a$  is shown in Fig. 9.3 as a function of  $\sigma_w a^2$ . The value of  $\zeta/l$  on the other side of the mobility edge is also plotted on the same graph, as well as the value of l/a, expressed as

$$\frac{l}{a} = \sqrt{6} \left\{ \frac{B_3 \left[ u, \, y_0(\sigma_w, \, u) \right]}{C_3 \left[ u, \, y_0(\sigma_w, \, u) \right]} \right\}^{1/2} \, \frac{1}{(-\mathrm{Im} \, y_0)^{1/2}},\tag{9.21}$$

where

$$C_3[u, y_0(\sigma_w, u)] = \int_{E_l}^{E_u} dE \, V_3 \left(2u - \operatorname{Re} y_0 - E\right) f(E), \qquad (9.22)$$

$$V_3(E) = \int_0^{\pi} \int_0^{\pi} \operatorname{Re} \left[ 1 - \left( \frac{E}{2} + \cos x + \cos y \right)^2 \right]^{-1/2} \mathrm{d}x \mathrm{d}y. \quad (9.23)$$

Equation (9.21) is the strong-scattering version of the definition for l. In Fig. 9.3, the functional form of  $\xi$  is found to accurately reflect the relation

$$\xi^{-1} \propto [\sigma_w a^2 - (\sigma_w a^2)_{\rm c}],$$



Fig. 9.3. Localization length  $\xi$ , correlation length  $\zeta$ , and mean free path l (all in unit of a) for the 3D Anderson model, plotted as a function of randomness  $\sigma_w a^2$ . Here u = 0 (band center).  $\xi$  and  $\zeta$  are both calculated from the self-consistent theory.

which means that the exponent of divergence is 1 in the self-consistent theory of localization. The value of  $\bar{\kappa}_e l$  can also be calculated, given by

$$\bar{\kappa}_{\rm e}l = \frac{\bar{\kappa}_{\rm e}^2}{(-\,{\rm Im}\Sigma^+)} = 6 \frac{B_3 \left[u, \, y_0(\sigma_w, \, u)\right]}{C_3 \left[u, \, y_0(\sigma_w, \, u)\right]}.\tag{9.24}$$

The value of  $\bar{\kappa}_{el}$  along the mobility edge is found to be

$$(\bar{\kappa}_{\rm e}l)_{\rm c} = 0.8 - 1.$$

The narrow range of variation for this dimensionless parameter may be regarded as support for the Ioffe–Regel criterion in the Anderson model (Zdetsis et al. 1985).

# 9.2.4 2D Anderson Model

In 2D, a similar approach may be applied to the calculation of C. To make the dispersion relation isotopic, one uses the relation

$$a\,\bar{k}(E') = \frac{8\int_0^{\pi} \int_0^{\pi} \mathrm{d}x\mathrm{d}y\,\sqrt{x^2 + y^2}\,\sqrt{\sin^2 x + \sin^2 y}\,\delta\left[E + 2\cos x + 2\cos y\right]}{L_0(E)},\tag{9.25}$$

where

$$L_0(E) = 8 \int_0^{\pi} \int_0^{\pi} \mathrm{d}x \mathrm{d}y \sqrt{\sin^2 x + \sin^2 y} \,\delta \left[E + 2\cos x + 2\cos y\right], \quad (9.26)$$

with  $-4 \leq E \leq 4$ . Let us first consider the multiple integral factor in C that may be written as

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$$(\mathrm{MC}) = \frac{\mathrm{i}}{(2\pi)^4} \int \mathrm{d} \,\mathbf{k}_{\mathrm{i}} \int \mathrm{d} \,\mathbf{k}_{\mathrm{f}} \left[ \nabla_{k_{\mathrm{i}}} e(\mathbf{k}_{\mathrm{i}}) \cdot \Delta \,\hat{\mathbf{k}} \right] \left[ \nabla_{k_{\mathrm{f}}} e(\mathbf{k}_{\mathrm{f}}) \cdot \Delta \,\hat{\mathbf{k}} \right] \times (\Delta G_{\mathrm{e}})_{\mathbf{k}_{\mathrm{i}}} (\Delta G_{\mathrm{e}})_{\mathbf{k}_{\mathrm{f}}}^2 \frac{1}{\nu^2 + |\mathbf{k}_{\mathrm{i}} + \mathbf{k}_{\mathrm{f}}|^2} \frac{L^{3d}}{N^3} \cong -\frac{8a^2}{(2\pi)^4} \int_{E_l}^{E_u} \mathrm{d} E_x \int_{E_l}^{E_u} \mathrm{d} E_y \frac{L_0(E'_x)}{2\pi} \frac{L_0(E'_y)}{2\pi} f(E_x) f^2(E_y) \times \left\{ \int_0^{2\pi} \mathrm{d} \,\theta_x \int_0^{2\pi} \mathrm{d} \,\theta_y \frac{\cos\theta_x \cos\theta_y}{a^2\nu^2 + x^2 + y^2 + 2xy \cos\theta_{xy}} \right\}.$$
(9.27)

Just as in the case of 3D, here the direction of  $\nabla e(\mathbf{k})$  is approximated as being the same as  $\mathbf{k}$ , and  $x = a\bar{k}_i(E'_x)$ ,  $y = a\bar{k}_f(E'_y)$  are evaluated in accordance with (9.25), with the magnitude of  $\mathbf{k}_i$  defined from the lower band edge for  $u \leq 0$ , and from the upper band edge for u > 0.  $E'_{x,y} = 2u - \operatorname{Re} y_0 - E_{x,y}$ ,  $\theta_{x,y}$  is the angle between  $\Delta \hat{\mathbf{k}}$  and  $\mathbf{k}_{i,f}$  and  $\theta_{xy} = \theta_y - \theta_x$  so that  $\cos \theta_y = \cos \theta_x \cos \theta_{xy} - \sin \theta_x \sin \theta_{xy}$ . The angular integral in the curly brackets may be performed to yield

$$\{\} = \int_{0}^{2\pi} \mathrm{d}\,\theta_{x}\cos^{2}\theta_{x} \int_{0}^{2\pi} \mathrm{d}\,\theta_{xy} \frac{\cos\theta_{xy}}{a^{2}\nu^{2} + x^{2} + y^{2} + 2xy\cos\theta_{xy}} \\ = \frac{\pi}{2xy} \int_{0}^{2\pi} \mathrm{d}\,\theta_{xy} \left[ 1 - \frac{a^{2}\nu^{2} + x^{2} + y^{2}}{a^{2}\nu^{2} + x^{2} + y^{2} + 2xy\cos\theta_{xy}} \right] \\ = \frac{\pi^{2}}{xy} \left[ 1 - \frac{a^{2}\nu^{2} + x^{2} + y^{2}}{\sqrt{(a^{2}\nu^{2} + x^{2} + y^{2})^{2} - 4x^{2}y^{2}}} \right],$$
(9.28)

where the  $\sin \theta_x \sin \theta_{xy}$  term is noted to integrate to zero. Substitution of (9.28) into (9.27) yields

(MC) = 
$$\frac{2a^2}{(2\pi)^4} A_2 \left[ u, y_0(\sigma_w, u), a^2 \nu^2 \right],$$
 (9.29a)

$$A_{2}\left[u, y_{0}(\sigma_{w}, u), a^{2}\nu^{2}\right] = \int_{E_{l}}^{E_{u}} \mathrm{d}E_{x} \int_{E_{l}}^{E_{u}} \mathrm{d}E_{y} f(E_{x}) f^{2}(E_{y}) L_{0}(E'_{x}) L_{0}(E'_{y})(xy)^{-1} \\ \times \left[\frac{a^{2}\nu^{2} + x^{2} + y^{2}}{\sqrt{(a^{2}\nu^{2} + x^{2} + y^{2})^{2} - 4x^{2}y^{2}}} - 1\right], \qquad (9.29b)$$

$$x = a \bar{k}_i (2u - \operatorname{Re} y_0 - E_x),$$
 (9.30a)

$$y = a \,\bar{k}_{\rm f} (2u - {\rm Re} \, y_0 - E_y),$$
 (9.30b)

and  $E_u = 2u - \text{Re } y_0 + 4$ ,  $E_l = 2u - \text{Re } y_0 - 4$ . From (9.29b) it is evident that  $A_2$  diverges when  $\nu \to 0$ ; i.e., the contribution of the MC diagrams dictates

that in order to avoid divergence,  $\nu \neq 0$  in 2D, so waves are localized for all u and  $\sigma_w$ . For the factor  $2K_0$ , we have

$$2K_{0} = \frac{2}{(2\pi)^{2}} \int d\mathbf{k} \left[ \nabla_{k} e(\mathbf{k}) \cdot \Delta \hat{k} \right]^{2} \left\{ \frac{-\mathrm{Im}\Sigma^{+}}{[\kappa_{0}^{2} - \mathrm{Re}\Sigma^{+} - e(\mathbf{k})]^{2} + (\mathrm{Im}\Sigma^{+})^{2}} \right\}^{2}$$
$$= \frac{2}{(2\pi)^{2}} \int dE' \int d\mathbf{k} \left[ \nabla_{k} e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}} \right]^{2} \delta \left[ E' + 4 - a^{2} e(\mathbf{k}) \right]$$
$$\times \left\{ \frac{-\mathrm{Im}\Sigma^{+}}{[\kappa_{0}^{2} - \mathrm{Re}\Sigma^{+} - e(\mathbf{k})]^{2} + (\mathrm{Im}\Sigma^{+})^{2}} \right\}^{2}$$
$$= \frac{16}{(2\pi)^{2}} B_{2} \left[ u, y_{0}(\sigma_{w}, u) \right], \tag{9.31}$$

where  $y_0 = a^2 \Sigma^+$ ,

$$B_2[u, y_0(\sigma_w, u)] = \int_{E_l}^{E_u} \mathrm{d}E \ W_2(E') f^2(E), \qquad (9.32)$$

$$W_2(E') = \int_0^{\pi} \text{Re}\sqrt{1 - \left(\frac{E'}{2} + \cos x\right)^2} \, \mathrm{d}x, \qquad (9.33)$$

with  $E' = 2u - \operatorname{Re} y_0 - E$ .

# 9.2.5 Localization Length in the 2D Anderson Model

Combining terms yields the 2D localization-length equation as

$$1 = C = \frac{1}{128} \left[ -\text{Im} \, y_0(\sigma_w, \, u) \right] \frac{A_2 \left[ u, \, y_0(\sigma_w, \, u), \, \nu^2 a^2 \right]}{\left\{ B_2 \left[ u, \, y_0(\sigma_w, \, u) \right] \right\}^2}.$$
 (9.34)

Since  $A_2$  diverges as  $\ln \nu$ , one can write

$$A_2 = c_1 \ln \nu + c_2.$$

Then (9.34) implies

$$\frac{a}{\xi} \propto \exp\left[\frac{(B_2)^2}{-\text{Im}y_0} \frac{128}{c_1} - \frac{c_2}{c_1}\right].$$

With numerical evaluation of  $c_1$  and  $c_2$ , the solution of (9.34) is plotted as a function of  $\sigma_w a^2$  in Fig. 9.4a. At small  $\sigma_w a^2$ ,  $\xi/a$  is transcendentally large. In the solution to Problem 9.4 it is shown that the 2D localization length should behave as

$$\ln\frac{\xi}{a} \propto \frac{1}{(\sigma_w a^2)^2} \tag{9.35}$$



Fig. 9.4. (a) Log of localization length (solid line) as a function of  $(\sigma_w a^2)^{-2}$  for the 2D Anderson model, calculated from the self-consistent theory at u = 0 (band center). (b) 2D localization length (solid line) as a function of u, calculated from the self-consistent theory at  $\sigma_w a^2 = 2.5$ . Numerical simulation results (Zdetsis et al. 1985) are shown as open circles.

in the limit of small randomness. Accordingly, the plot is done in the form of  $\ln(\xi/a)$  vs.  $(\sigma_w a^2)^{-2}$ . It is seen that at small  $\sigma_w a^2$ ,  $\ln(\xi/a)$  is indeed linear in  $(\sigma_w a^2)^{-2}$ , whereas at large randomness  $\ln(\xi/a)$  varies with a smaller power of  $1/(\sigma_w a^2)$ . By comparing the value of  $1/\nu$  with simulation results on the localization length, we obtain the value of the proportionality factor  $\nu\xi$  to be ~8 (see (9.4b)). For a fixed  $\sigma_w a^2 = 2.5$ , the localization length solved from (9.34) is plotted in Fig. 9.4b. The circles are the data from numerical simulation. The fact that the localization length is smaller near the center of the band is the result of the large density of states at those frequencies (density of states is divergent at the band center in the ordered case). As a result, the average  $\bar{\kappa}_e$  is small, so that the group velocity  $\hbar \bar{\kappa}_e/m$  is small at the band center. It is seen that there is overall qualitative agreement with the simulation results.

## 9.2.6 1D Anderson Model

In 1D, no approximation is needed for the evaluation of C. The factor  $2K_0$  may be expressed as

$$2K_{0} = 2a \int_{-\pi}^{\pi} \frac{d(ka)}{2\pi} 4\sin^{2} ka \left[\frac{-\operatorname{Im} y_{0}}{(2u - \operatorname{Re} y_{0} + 2\cos ka)^{2} + (\operatorname{Im} y_{0})^{2}}\right]^{2}$$

$$= \frac{4a}{\pi} \int_{-\pi}^{\pi} dx \sin^{2} x \left[\frac{-\operatorname{Im} y_{0}}{(2u - \operatorname{Re} y_{0} + 2\cos x)^{2} + (\operatorname{Im} y_{0})^{2}}\right]^{2}$$

$$= \frac{4a}{2\pi} \int_{-2}^{2} d(2\cos x) \sqrt{1 - \cos^{2} x}$$

$$\times \left[\frac{-\operatorname{Im} y_{0}}{(2u - \operatorname{Re} y_{0} + 2\cos x)^{2} + (\operatorname{Im} y_{0})^{2}}\right]^{2}$$

$$= \frac{2a}{\pi} B_{1} [u, y_{0}(\sigma_{w}, u)], \qquad (9.36a)$$

where

$$B_1[u, y_0(\sigma_w, u)] = \int_{E_l}^{E_u} dE \operatorname{Re} \sqrt{1 - \left(\frac{E}{2} - u + \frac{\operatorname{Re} y_0}{2}\right)^2 f^2(E)}, \quad (9.36b)$$

 $E_u = 2u - \operatorname{Re} y_0 + 2$ ,  $E_l = 2u - \operatorname{Re} y_0 - 2$ , and  $u = (\kappa_0^2 a^2 - 2)/2$ . The maximally crossed diagrams' contribution, on the other hand, is given by

(MC) = 
$$[32a^4/(2\pi)^2] A_1 [u, y_0(\sigma_w, u), a^2\nu^2],$$

where

$$A_{1}\left[u, y_{0}(\sigma_{w}, u), a^{2}\nu^{2}\right] = \int_{-\pi}^{\pi} \mathrm{d}x \int_{-\pi}^{\pi} \mathrm{d}y \\ \times \left[\frac{-\mathrm{Im}\,y_{0}}{(2u - \mathrm{Re}\,y_{0} + 2\cos x)^{2} + (\mathrm{Im}\,y_{0})^{2}}\right] \\ \times \left[\frac{-\mathrm{Im}\,y_{0}}{(2u - \mathrm{Re}\,y_{0} + 2\cos y)^{2} + (\mathrm{Im}\,y_{0})^{2}}\right]^{2} \frac{\sin x \sin y}{a^{2}\nu^{2} + \{(x + y)^{2}\}}.$$
(9.37)

Here  $\{(x+y)^2\}$  means a higher-order evaluation of the term using (5.30). In the present case it has the following precise meaning:

$$\left\{ (x+y)^2 \right\} = \frac{1}{2} \frac{\int_{-\pi}^{\pi} \left[ g_+(\theta) g_-(\theta) - g_+ \left(\theta + \frac{x+y}{2}\right) g_- \left(\theta - \frac{x+y}{2}\right) \right] \mathrm{d}\theta}{\int_{-\pi}^{\pi} \sin^2\theta g_+^2(\theta) \,\mathrm{d}\theta},$$
(9.38a)

$$g_{\pm}(\theta) = \frac{1}{2u - \operatorname{Re} y_0 + 2\cos\theta \mp \mathrm{i}\operatorname{Im} y_0}.$$
(9.38b)

# 9.2.7 Localization Length in the 1D Anderson Model

The 1D localization length equation has the following form:

$$1 = C = 2 \left[ -\text{Im} \, y_0(\sigma_w, \, u) \right] \frac{A_1 \left[ u, \, y_0(\sigma_w, \, u), \, \nu^2 a^2 \right]}{\left\{ B_1 \left[ u, \, y_0(\sigma_w, \, u) \right] \right\}^2}.$$
(9.39)

The divergence of  $A_1$  as  $\nu \to 0$  means there is no solution to (9.39) unless  $\nu^2$  is finite. This fact directly implies the *localization of all 1D waves*. The proportionality constant (with a value ~ 5) between  $\nu$  and  $\xi^{-1}$  [see (9.4b)] is set by requiring the calculated  $\xi$  equal to the exact result at u = 0 and  $\sigma_w a^2 = 0.5$  ( $\xi \cong 105a$ ) (Economou 1983). The numerically evaluated 1D localization length at u = 0 is plotted as a function of  $\sigma_w a^2$  in Fig. 9.5a. In anticipation of  $\xi/a$  varying as some inverse power of  $\sigma_w a^2$ , the plot is done in the log-log scale. It is seen that at small  $\sigma_w a^2$ ,  $a\xi^{-1}$  varies as ( $\sigma_w a^2$ )<sup>2</sup>, in complete agreement with what has been calculated in Chap. 7. In fact, comparison with Fig. 7.6 shows good agreement between the two calculations performed via very different routes. In Fig. 9.5b a plot is made of the localization length as a function of u for  $\sigma_w a^2 = 0.5$ . Comparison with Fig. 7.5 again shows reasonable agreement except near the band edge.

#### 9.2.8 Input Parameters to Localization Calculations

From the foregoing discussion, it is clear that all the localization calculations involve two elements. One element is the  $\langle G G^* \rangle_c$  calculation, which provides the framework for deriving the functional form of the equations. The other equally important element is the CPA calculation of  $\langle G \rangle_c$ , which gives  $\text{Re}\Sigma^+$ and  $\text{Im}\Sigma^+$  as inputs to the equations. It is through  $\Sigma^+$  that the localization parameters are linked to the randomness of the model. Both elements are essential to the accurate determination of the mobility edge and the localization length. In Sect. 9.3 it will be seen that in the context of the two elements described above, the difficulty with determining the classical wave localization parameters lies mostly in the accurate evaluation of  $\Sigma^+$  in the intermediate frequency regime.

#### 9.3 Classical Scalar Wave Localization

For classical scalar wave in continuum,  $e(\mathbf{k}) = k^2$ , and it becomes possible to express the correction term C in terms of just one dimensionless parameter. To proceed, let us define a complex  $\kappa_{e}^{*} = \kappa + i/2l$  such that

$$G_{\rm e}^{+} = \frac{1}{(\kappa_{\rm e}^{*})^2 - k^2} \frac{N}{L^d} = \frac{1}{\kappa_0^2 - k^2 - \Sigma^+} \frac{N}{L^d}.$$
 (9.40)

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**Fig. 9.5.** (a) Log of localization length (*solid line*) as a function of  $\ln(\sigma_w a^2)$  for the 1D Anderson model, calculated from the self-consistent theory at u = 0 (*band center*). Numerical simulation data (Economou et al. 1984) are shown as *open circles*. (b) Localization length (*solid line*) as a function of u, calculated from the self-consistent theory at  $\sigma_w a^2 = 0.5$ . This figure should be compared with Fig. 7.5.

That means

$$\operatorname{Re}(\kappa_{e}^{*})^{2} = \kappa^{2} - \frac{1}{4l^{2}} = \kappa_{0}^{2} - \operatorname{Re}\Sigma^{+},$$
 (9.41a)

$$\operatorname{Im}(\kappa_{\rm e}^*)^2 = \frac{\kappa}{l} = -\operatorname{Im}\Sigma^+.$$
(9.41b)

Note that when  $\kappa l \gg 1$ ,  $\kappa^2 \cong \kappa_{\rm e}^2 = \kappa_0^2 - {\rm Re}\Sigma^+$  as defined before. However, the possibility that  $\kappa l \cong 1$  has now to be taken into account. As a result,  $(\Delta G_{\rm e})_{\bf k}$  should be written as

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$$(\Delta G_{\rm e})_{\bf k} = -2i \frac{-\mathrm{Im}\Sigma^{+}}{(\kappa_{0}^{2} - \mathrm{Re}\Sigma^{+} - k^{2})^{2} + (\mathrm{Im}\Sigma^{+})^{2}} \frac{N}{L^{d}}$$
$$= -2i \frac{(\kappa/l) \cdot l^{4}}{\left(\kappa^{2}l^{2} - \frac{1}{4} - k^{2}l^{2}\right)^{2} + \kappa^{2}l^{2}} \frac{N}{L^{d}}$$
$$= -2i l^{2} \frac{\kappa l}{\left(\kappa^{2}l^{2} - \frac{1}{4} - k^{2}l^{2}\right)^{2} + \kappa^{2}l^{2}} \frac{N}{L^{d}}.$$
(9.42)

By treating  $\kappa l = \mu$  as a parameter of the problem,  $(\Delta G_{\rm e})_{\bf k}$  may be expressed as

$$(\Delta G_{\rm e})_{\mathbf{k}} = -2\,\mathrm{i}\,l^2\,\frac{\mu}{\left(\mu^2 - \frac{1}{4} - k^2l^2\right)^2 + \mu^2}\,\frac{N}{L^d}.\tag{9.43}$$

From (9.41a), we have the condition that  $\mu > 0.5$ , since otherwise the wave would be evanescent. Equation (9.43) will be used below for the evaluation of C as given by (9.3), in the case of classical waves.

# 9.3.1 One Parameter Criterion for the Classical Wave 3D Mobility Edge

To determine the 3D mobility edge, it is important to first note that for the classical scalar wave case, the  $D(\omega)$  expression as given by (5.107b) is different from that of the quantum case by a factor  $(1 + \delta)^{-1}$ , which is due to the noncancellation of  $\mathbf{v}^+$  and  $\mathbf{v}^-$  in the derivation of the Ward identity when  $\Delta \omega \neq 0$ . However, since this factor is just a multiplier for the whole  $D(\omega)$  expression, it *does not alter the location of the mobility edge*, although it does affect the diffusion constant through the transport velocity  $v_t$ . The latter effect will be examined in Sect. 9.4. For the present case, let us consider the factor  $-\text{Im}\Sigma^+/(2K_0)^2$  in C:

$$\frac{-\mathrm{Im}\Sigma^{+}}{4K_{0}^{2}} = \frac{-\mathrm{Im}\Sigma^{+}}{4} \left\{ \frac{4}{(2\pi)^{3}} \int \mathrm{d}\mathbf{k} \, k^{2} \cos^{2}\theta \right. \\ \left. \times \left[ \frac{-\mathrm{Im}\Sigma^{+}}{(\kappa_{0}^{2} - \mathrm{Re}\Sigma^{+} - k^{2})^{2} + (\mathrm{Im}\Sigma^{+})^{2}} \right]^{2} \right\}^{-2}.$$
(9.44)

The expression in the curly brackets,  $K_0$ , may be written as

$$K_0 = \frac{1}{l} \frac{2}{3\pi^2} \int_0^\infty \mathrm{d}x \, x^4 f_\mu^2(x), \qquad (9.45)$$

where x = kl and

$$f_{\mu} = \frac{\mu}{\left(\mu^2 - \frac{1}{4} - x^2\right)^2 + \mu^2}.$$
(9.46)

It follows that

$$\frac{-\mathrm{Im}\Sigma^+}{4K_0^2} = \mu \left[\frac{4}{3\pi^2} \int_0^\infty \mathrm{d}x \ x^4 f_\mu^2(x)\right]^{-2}.$$
(9.47)

For the double- $\mathbf{k}$  integral in (9.2), it is straightforward to write it as

$$(MC) = i \iint \frac{d \mathbf{k}_{i}}{(2\pi)^{3}} \frac{d \mathbf{k}_{f}}{(2\pi)^{3}} \frac{L^{d}}{N} (\Delta G_{e})_{\mathbf{k}_{i}} \frac{L^{2d}}{N^{2}} (\Delta G_{e})_{\mathbf{k}_{f}}^{2} \left[ \nabla_{\mathbf{k}_{i}} e(\mathbf{k}_{i}) \cdot \Delta \hat{\mathbf{k}} \right] \times \frac{1}{\nu^{2} + |\mathbf{k}_{i} + \mathbf{k}_{f}|^{2}} \left[ \nabla_{\mathbf{k}_{f}} e(\mathbf{k}_{f}) \cdot \Delta \hat{\mathbf{k}} \right] = -\frac{32}{(2\pi)^{4}} \int_{0}^{\infty} \int_{0}^{\infty} dx dy x^{2} y^{2} f_{\mu}(x) f_{\mu}^{2}(y) xy \int_{-1}^{1} \int_{-1}^{1} d \cos \theta_{i} d \cos \theta_{f} \times \frac{\cos \theta_{i} \cos \theta_{f}}{l^{2}\nu^{2} + x^{2} + y^{2} + 2xy \cos \theta_{if}} = \frac{4}{3\pi^{4}} \int_{0}^{\infty} \int_{0}^{\infty} dx dy x^{2} y^{2} f_{\mu}(x) f_{\mu}^{2}(y) \times \left[ \frac{l^{2}\nu^{2} + x^{2} + y^{2}}{4xy} \ln \frac{(x + y)^{2} + \nu^{2} l^{2}}{(x - y)^{2} + \nu^{2} l^{2}} - 1 \right].$$
(9.48)

Here  $x = |\mathbf{k}_i|l$ ,  $y = |\mathbf{k}_f|l$ , and the angular integration is done exactly the same way as shown in Problem 9.3.

From (9.47) and (9.48), the mobility-edge condition is an equation of only one variable,  $\mu$ ,

$$1 = C = \frac{3}{4} \mu \frac{A_3^{(c)}(\mu, \nu^2 = 0)}{\left[B_3^{(c)}(\mu)\right]^2},$$
(9.49)

where the superscript (c) and subscript 3 are used to denote classical wave and d = 3, respectively, and

$$A_{3}^{(c)}(\mu, \nu^{2}l^{2}) = \int_{0}^{\infty} \int_{0}^{\infty} dx dy \, x^{2}y^{2} f_{\mu}(x) f_{\mu}^{2}(y) \\ \times \left[ \frac{l^{2}\nu^{2} + x^{2} + y^{2}}{4xy} \ln \frac{(x+y)^{2} + \nu^{2}l^{2}}{(x-y)^{2} + \nu^{2}l^{2}} - 1 \right]. \quad (9.50)$$

$$B_3^{(c)}(\mu) = \int_0^\infty \mathrm{d}x \, x^4 f_\mu^2(x). \tag{9.51}$$

Both  $A_3^{(c)}$  and  $B_3^{(c)}$  are convergent integrals regardless of the value of  $\nu^2$ . Equation (9.49) can be solved numerically, giving

$$\mu_{\rm c} = (\kappa l)_{\rm c} = 0.985 \tag{9.52}$$

as the critical value for the mobility edge. Equation (9.52) may be regarded as a formal expression of the Ioffe–Regel criterion (van Tiggelen 1992) in the classical scalar wave case. The window for classical wave localization is therefore  $0.5 \leq \mu \leq 0.985$ . It is indeed remarkable that for this case the two elements of the localization calculation can be so clearly separated – the  $\langle GG* \rangle_c$  calculation determines the value of  $\mu$ , and the  $\langle G \rangle_c$  calculation links  $\mu$  to the parameters of the model.

# 9.3.2 Input Parameter Values and Wave Scattering Regimes

From (9.41a) and (9.41b),  $\kappa$ , l may be expressed as

$$l = \left\{ \frac{(\kappa_0^2 - \text{Re}\Sigma^+) + \sqrt{(\kappa_0^2 - \text{Re}\Sigma^+)^2 + (\text{Im}\Sigma^+)^2}}{2(\text{Im}\Sigma^+)^2} \right\}^{1/2}, \quad (9.53)$$

and  $\kappa = (-\mathrm{Im}\Sigma^+)l$ , so that  $\mu = (-\mathrm{Im}\Sigma^+)l^2$ . Here the question is, what should be the value of  $\kappa_0$ ? In the long-wavelength limit where the CPA is valid,  $\kappa_0$  should be taken to be  $\kappa_e = \bar{\varepsilon} \omega / v_0$ , and  $\Sigma^+$  is calculated relative to the effective medium according to the rules prescribed in Chap. 3 (e.g.,  $\text{Re}\Sigma^+ = 0$ ). In this way the parameters of the system, such as the dielectric constant ratio,  $R/\lambda$ , and the amount of disorder, may be linked to the localization criterion. However, localization is very unlikely in the long-wavelength limit because the Rayleigh frequency dependence tells us that the scattering must be weak in the long-wavelength limit, in contrast to the quantum case. At high frequencies geometric optics becomes a good description for classical wave propagation. In that regime the effect of scattering, as expressed by the mean free path, is expected to saturate at a scale comparable to the particle size or interparticle separation, whereas  $\kappa$ , being proportional to  $\omega$ , diverges. Therefore  $\mu \to \infty$  again and localization is unlikely in the high-frequency limit either. It follows that the localization of classical waves is most likely in the intermediate-frequency regime, where the wavelength is comparable to the size of the scatterers.

In the intermediate-frequency regime, the CPA may not have a solution, and one needs a way to determine the effective  $\kappa_e$  for the medium to be used for the calculation of l and  $\kappa$ . In Chap. 4 a spectral function approach to extend the CPA has been presented. In Sect. 9.4, on the transport velocity of classical waves, it is proposed that in the delocalized regime a  $\kappa_m$  may be determined on the basis of wave energy uniformity over the scale of microstructural correlation length. While the basic principle here is very different from that of the CPA, e.g., it does not involve phase coherence at all but rather with energy, at low frequencies it turns out  $\kappa_m = \kappa_e$ , the CPA value, because the two requirements are mutually compatible. The advantage of  $\kappa_m$  is that it has an unique solution at every frequency, and it plays a natural role in determining the wave transport velocity,  $v_t$ . Therefore,  $\kappa_m$  is the natural "effective medium" wave vector to be used in the calculation of  $\kappa$  and l. Accordingly,

$$l = \frac{1}{-\mathrm{Im}\Sigma_m^+} \frac{1}{\sqrt{2}} \left\{ \kappa_m^2 - \mathrm{Re}\Sigma_m^+ + \sqrt{(\kappa_m^2 - \mathrm{Re}\Sigma_m^+)^2 + (\mathrm{Im}\Sigma_m^+)^2} \right\}^{1/2},$$
(9.54a)
$$\kappa_m = (-\mathrm{Im}\Sigma_m^+) \, l,$$
(9.54b)

where

$$\Sigma_m^+ \cong \sum_i n_i \left\langle \kappa_m | \bar{\mathbf{t}}^+ | \kappa_m \right\rangle, \tag{9.55}$$

is calculated by solving a scattering problem where the scattering structural unit is embedded in a medium characterized by  $v_m = \omega/\kappa_m$ .  $\Sigma_m^+$  is approximated by the sum of the forward scattering components of the **t** matrix for the various scattering units, weighted by their respective number densities.

#### 9.3.3 Scattering and Localization of Classical Waves in 3D

For a system of randomly dispersed spherical scatterers of dielectric constant  $\varepsilon_1$  and volume fraction p in a matrix of dielectric constant  $\varepsilon_2$ , the dispersion microstructure applies, and one can write, from (4.14),

$$\Sigma_m^+ \cong n \left\langle \kappa_m | \bar{\mathbf{t}}^+ | \kappa_m \right\rangle = \frac{4\pi i}{\kappa_m} n \sum_{j=0}^\infty (2j+1) D_j,$$

where  $D_j$  is given by (P4.33). Here *n* is the number density of the coated sphere structural unit. With this  $\Sigma_m^+$  and the value of  $\kappa_m$  determined in the manner specified in Sect. 9.4, the value of  $\mu = \kappa l$  may be calculated. In Fig. 9.6 the results for p = 0.1 are plotted as a function of the reduce frequency variable  $\kappa_0 R$  ( $\varepsilon_2 = 1$  and  $\kappa_0 = \omega/v_0$ ). It is seen that for  $\sqrt{\varepsilon_1/\varepsilon_2} = 5$ , there is no frequency regime which satisfies the Ioffe–Regel criterion. However, the value of  $\kappa l$  is indeed lowest at the intermediate-frequency range. Although the results shown in Fig. 9.6 do not constitute a proof that localization is impossible for classical waves (the fact that the medium surrounding the scatterer is homogenized in our calculation certainly underestimates the actual scattering involved), they do confirm the difficulty of classical wave localization in 3D. In particular, for localization to occur the dielectric constant of the scatterers must be larger than  $\sqrt{\varepsilon_1/\varepsilon_2} = 5$ . This turns out to be indeed the case (Wiersma et al. 1997).

In Fig. 9.6, every dip in the  $\kappa l$  value is noted to be associated with a sphere resonance. Physically, it is a well-known fact that for almost every kind of resonance, the real part of the response function goes through a sign change at the resonance frequency, with the higher-frequency side being negative. In analogy with optics, a negative dielectric constant means the index of refraction is purely imaginary, thus implying a nonpropagating, exponentially decaying wave. Not only is the wave being excluded from the inside of the scatterers in this frequency region, but the fact that the scattering cross-section can be larger than the physical cross-section of each scatterer means that the wave could be attenuated in the interstitial space as well. However, one cannot carry this line of reason to the conclusion that wave must localize at a high volume fraction of spheres. This is because if the sphere concentration is too high, the effective medium in which the coated sphere is embedded can approach the



Fig. 9.6. Calculated value of  $\mu = \kappa l$  as a function of reduced frequency  $\kappa_0 R$  for the classical scalar wave in a strong-scattering medium, as determined by the approach described in the text. The random medium is a dispersion of dielectric spheres with  $\sqrt{\varepsilon_1} = 5$  and p = 0.1. The Ioffe–Regel criterion,  $\mu_c = 0.985$ , is indicated by the dashed line. No localized state is seen in this case, despite the fact that the value p = 0.1 is optimized to yield the lowest  $\mu$  values in the intermediate-frequency range.

property of the scatterer itself, thus making the effective scattering weaker rather than stronger. Another way of saying the same thing is that the multiple scattering between the spheres can renormalize the resonance property of the single scatterer, and at a high enough scatterer concentration the whole medium is inverted, with the scatterer material resembling the matrix and the interstitial spaces resembling the scatterers. Therefore, if classical scalar wave localization were to occur, it must happen not only within narrow frequency ranges in the intermediate-frequency regime but also at intermediate concentrations of the scatterers.

An obvious reason for the difficulty of classical wave localization in 3D lies in the absence of a minimum length scale for classical waves in continuum. The existence of such a scale in the Anderson model is responsible for the existence of an upper band edge. Near the band edge the slow group velocity implies that the lattice wave is literally "waiting" to be localized by the addition of randomness, as seen from the fact that as randomness increases from zero, the mobility edge of the Anderson model always nucleates from the CPA band edges.

#### 9.3.4 Periodicity in Scattering Structures and 3D Localization

One can deduce from such reasoning that one way to facilitate the localization of 3D classical waves is to start with a periodic system and then add randomness to it (John 1987; Yablonovitch 1987). Since periodicity would induce band edges, in the simple way of thinking one might expect 3D localization must occur, in direct analogy to the lattice models. However, complications

arise because the scatterers in continuum are not point particles; therefore as frequency increases there can be additional frequency bands associated with the various scatterer resonances. The presence of additional bands means that while a gap always exists for a wave of a fixed propagation direction, such a gap need not be isotropic with respect to the propagation directions. In particular, the upper band edge of the lower-frequency band in *one direction* can overlap with the lower band edge of the higher-frequency band in another direction. If randomness is introduced into such a periodic system, randomization of the wave propagation direction would mean that there is always some escape direction(s) for the random wave. Therefore, for the periodic scheme to work, it is not enough just to have periodicity. Instead, it is necessary that there be a "true" frequency gap in the sense that what described above does not happen. This proves to be possible only for certain periodic structures in which  $\sqrt{\varepsilon_1/\varepsilon_2}$ has to exceed a certain minimum threshold. For electromagnetic waves, it has been found that whereas it is possible for the diamond structure to have a true gap, it is not that easy for many other periodic structures to realize a true gap (Ho et al. 1990; Leung and Liu 1990). It turns out that the scalar wave is a poor approximation to the vector wave in the selection of periodic structures possessing true frequency gaps. In this particular case, the vector character of the electromagnetic wave plays an important role.

Suppose one finds a localized classical-wave state near the band edge of a "true gap" system. It may be argued that such a localized state is different from a localized state of a purely random system. Aside from the wave confinement effect (see Chap. 1) due to the deficit of wave density of states near the band edge, there can also be a difference due to their respective sensitivities to the environment beyond their localization length. For the "true gap + randomness" system, the localization length can be extremely small, e.g., comparable to the interparticle separation for the states well inside the gap region. Suppose the material around such a localized state, beyond a few  $\xi$ , is removed. From the Thouless argument presented in Chap. 8, there should be very little effect on the state because a localized state "feels" its environment only through its exponential tail. However, in the present case the existence of the localized state is dependent on the existence of a frequency gap, which in turn is dependent on the existence of long-range order. As material is removed so that only a small cube remains, the frequency gap would be drastically affected. As true gap becomes not so much of a gap any more, the "localized" state is also expected to be not so localized any more, much more so than is implied by the Thouless argument. In other words, a "localized" state in the "true gap + randomness" system may have sensitivity to its environment beyond its localization length.

Classical wave localization has been an active research area in recent years. Strong indications of localization have been found (Drake and Genack 1989). Moreover, light localization was observed in a system consisting of semiconducting powders in which  $\sqrt{\varepsilon_1/\varepsilon_2} \sim 3.5$  (Wiersma et al. 1997). Clearly the scalar wave calculations can offer only a qualitative guide to the prediction of vector wave localization.

# 9.3.5 Localization of 2D Classical Scalar Wave

In 2D, similar calculations as those in 3D yield

$$\frac{-\mathrm{Im}\Sigma^{+}}{4K_{0}^{2}} = \frac{\mu}{l^{2}} \left[\frac{2}{\pi} \int_{0}^{\infty} \mathrm{d}x \, x^{3} f_{\mu}^{2}(x)\right]^{-2}, \qquad (9.56)$$

whereas the double-**k** integral in (9.2) may be expressed as  $2A_2^{(c)}(\mu, \nu^2 l^2)/\pi^2$ , with

$$A_{2}^{(c)}(\mu,\nu^{2}l^{2}) = \int_{0}^{\infty} \int_{0}^{\infty} dx dy \, xy \, f_{\mu}(x) f_{\mu}^{2}(y) \\ \times \left[ \frac{l^{2}\nu^{2} + x^{2} + y^{2}}{\sqrt{(l^{2}\nu^{2} + x^{2} + y^{2})^{2} - 4x^{2}y^{2}}} - 1 \right].$$
(9.57)

The 2D localization length equation for classical scalar wave is therefore

$$1 = C = \frac{\mu}{2} \frac{A_2^{(c)}(\mu, \nu^2 l^2)}{\left[B_2^{(c)}(\mu)\right]^2},$$
(9.58)

where

$$B_2^{(c)}(\mu) = \int_0^\infty \mathrm{d}x \, x^3 f_\mu^2(x). \tag{9.59}$$

To obtain  $\xi/R$  as a function of reduced frequency  $\kappa_0 R$ , it is noted that the argument of wave energy uniformity no longer applies here because the states are localized. It is therefore necessary to use the CPA solution, where it is well-defined. In the same spirit as in the 3D case,  $\Sigma^+$  may be approximated by [see (4.14)]

$$\Sigma^{+} \cong 4in \sum_{j=-\infty}^{\infty} D_j, \qquad (9.60)$$

where n is the density per unit area of the scatterers,  $D_j$  is given by (P4.33) with the spherical Bessel functions replaced by the regular Bessel functions, and the value of  $\kappa_e$  is taken to be  $k_{\max}$  where the spectral function  $-\text{Im }G_e(\omega, \mathbf{k})$  [see (4.51)] has its maximum. For a given  $\mu$ , the solution of (9.58) gives  $\nu l$ , which is multiplied by 0.46R/l to get  $R/\xi$ . The factor 0.46 is obtained from the 1D calculation, shown below.

Figure 9.7 shows the results of the 2D localization length calculations for a system of randomly dispersed disks. The value of p is chosen to be 0.35, and  $\sqrt{\varepsilon_1} = 5$ ,  $\varepsilon_2 = 1$  (matrix). In accordance with the solution to Problem 9.4, it is expected that at small  $\omega$ 

$$\ln\frac{\xi}{R} \propto l \propto \frac{1}{\omega^3} \tag{9.61}$$

from Rayleigh scattering. Therefore,  $\ln(\xi/R)$  is plotted as a function of  $(\kappa_0 R)^{-3}$ . Indeed, at small  $\omega$  a good straight-line correlation is obtained.



Fig. 9.7. 2D localization length as a function of reduced frequency  $\kappa_0 R$  for a classical scalar wave in a medium consisting of randomly dispersed disks with p = 0.35 and  $\sqrt{\varepsilon_1} = 5$ , calculated in the framework of self-consistent theory with the self-energy input from the single-site CPA. At low frequencies it is seen that  $\xi \propto \omega^{-3}$ .

# 9.3.6 Localization of 1D Classical Scalar Wave

In 1D, the equation for the localization length is

$$1 = C = \frac{1}{4} \mu \frac{A_1^{(c)}(\mu, \nu^2 l^2)}{\left[B_1^{(c)}(\mu)\right]^2},$$
(9.62)

where

$$A_1^{(c)}(\mu,\nu^2 l^2) = \int_0^\infty \int_0^\infty dx dy \, xy \, f_\mu(x) f_\mu^2(y) \\ \times \left[ \frac{1}{\nu^2 l^2 + (x-y)^2} - \frac{1}{\nu^2 l^2 + (x+y)^2} \right], \qquad (9.63)$$

and

$$B_1^{(c)}(\mu) = \int_0^\infty \mathrm{d}x \, x^2 f_\mu^2(x). \tag{9.64}$$

Equation (9.62) can be solved for  $\nu l$  in terms of  $\mu$ . The results show  $\nu l \approx 0.46$  for all values of  $\mu$ . That implies  $\xi$  is directly proportional to l. Since in the  $\omega \to 0$  limit  $l/\xi = 0.25$  as shown in Chap. 7 [(7.41)], it follows that  $1/\nu \approx 0.5\xi$  in the scalar wave case. This may be interpreted as an indication that  $1/\nu$  gives the decay length for wave intensity rather than for wave amplitude.

With the proportionality constant thus fixed, the 1D localization length calculation can be carried out in exactly the same manner as in the 2D case. For a binary system in which the combined thickness of two components is fixed but the relative fraction fluctuates over a flat distribution, one has



Fig. 9.8. 1D localization length as a function of reduced frequency  $\kappa_0 R$  for classical scalar wave in a binary random medium defined in the text, with  $\sqrt{\varepsilon_1} = 5$ ,  $\varepsilon_2 = 1$ , and p = 0.35, calculated in the framework of self-consistent theory with the self-energy input from the single-site CPA. At low frequencies  $\xi \propto \omega^{-2}$ , and at high frequencies  $\xi \sim \text{constant}$ . Numerical simulation results are shown as *open circles*.

$$\Sigma^+ \cong 2ik_{\max}n \int D(p') [\tau(p', k_{\max}) - 1] dp',$$
 (9.65)

where n is the lineal density of the  $\varepsilon_1$  component. Here  $\tau$  is given by (P4.43). The localization length has been calculated for the case of p = 0.35, D(p') = 5/3 for  $0 \le p' \le 0.6$  and zero otherwise, and  $\sqrt{\varepsilon_1} = 5$ ,  $\varepsilon_2 = 1$ . The results are plotted in Fig. 9.8 as a function of  $\kappa_0 R$ , where 2R is the mean width of the  $\varepsilon_1$  component, and  $\kappa_0$  is defined in the  $\varepsilon_2$  component. The results of numerical simulations are also shown in the figure. Good agreement is obtained at low frequencies. At high frequencies the jitter in the calculated results reflects the instability of the CPA solution in that frequency regime. Surprisingly, however, the numerical value of  $\xi$  never departs far from the simulated results.

# 9.4 Transport Velocity of Classical Scalar Waves

In the beginning of Sect. 9.3.1 it was mentioned that the factor  $\delta$ , which is particular to the classical waves, has no effect on the mobility edge and localization length calculations. However,  $\delta$  does have an effect on the transport velocity and the overall magnitude of the diffusion constant.

From statistical mechanics, the diffusion constant of a randomly scattered particle can be written as

$$D = \frac{1}{3}vl$$

in 3D, where v is the ballistic particle velocity between collisions and l is the mean free path. For wave diffusion, the relevant velocity is  $v_t$ , which is shown in Chap. 5 to be [(5.123)]

$$v_{\rm t} = \frac{v_0^2}{(\omega/\kappa_{\rm e})(1+\delta)}$$

for the scalar wave. It is clear that  $v_t$  is not the phase velocity of the effective medium, which is given by  $\omega/\kappa_e$ , nor is it the group velocity due to the  $1 + \delta$ factor. Here  $\delta$  may be viewed as a delay additional to the group velocity, and which can be especially large near a scatterer resonance. To get an interpretation of  $v_t$ , we recall that in Chap. 5, it has been shown that the first moment of the Bethe–Salpeter equation may be interpreted as a continuity equation, i.e.,

$$\frac{1}{2} \left( \frac{\mathrm{d}\kappa_0^2}{\mathrm{d}\omega} \right) \frac{\partial S}{\partial t} + \nabla \cdot \mathbf{J} = \text{source term.}$$

Since from (5.53) one can write

$$\Delta \mathbf{\hat{k}} \cdot \mathbf{J} = -\frac{\mathrm{i}}{2} D(\omega) |\Delta \mathbf{k}| S, \qquad (9.66)$$

it follows that  $v_t$ , which is part of  $D(\omega)$ , is proportional to the ratio  $|\mathbf{J}|/S$ , i.e., as the ratio of the local energy flux to the local energy density.

#### 9.4.1 Transport Velocity and Scattering Resonances

It follows from above that the resonances of a continuum scatterer are expected to have a significant effect on  $v_t$  because S is large near a resonance (since the density of states peaks at the resonant frequency), and  $v_t$  should therefore exhibit dips near resonant frequencies. However, in the regime of strong resonant scattering the interparticle scattering is also strong. As a result, the "effective medium" surrounding a given scatterer may approach the property of the scatterer itself, and the resonances will thus become "leaky" and broadened in frequency (resonances would disappear if the medium surrounding a scatterer has the same property of the scatterer). This point has been demonstrated in some detail in Chap. 4 through the spectral function approach. From this intuitive point of view,  $v_t$  is expected to exhibit pronounced dips as a function of frequency only at low scatterer concentrations. At high scatterer concentrations the variation with frequency should be reduced.

#### 9.4.2 Effective Medium Based on Energy Homogenization

To capture the effect of resonances and interparticle interaction, it is proposed that the choice of the medium property be based on the principle that the wave energy density should be uniform on the scale of the microstructural correlation length or larger. That is, consider the case of the dispersed microstructure where the structural unit may be approximated by a coated sphere. Let  $\kappa_m = \omega/v_m$  characterize the embedding medium. The expectation value of the energy may be written as

wave energy in a structural unit 
$$\propto \langle \psi | \nabla^2 + \kappa_m^2 | \psi \rangle$$
  
+  $\langle \psi | \kappa^2(\mathbf{r}) - \kappa_m^2 | \psi \rangle$  (9.67)  
=  $\langle \psi | \nabla^2 + \kappa^2(\mathbf{r}) | \psi \rangle$ ,

where  $\psi$  is the scattering wave function for a plane wave incident on a coated sphere. Its solution is given in Chap. 4 (Problem 4.4). The condition for the *detennination* of  $\kappa_m$  is that it should homogenize  $\kappa(\mathbf{r})$ , i.e.,

$$\left\langle \psi \left| \nabla^{2} + \kappa_{m}^{2} \right| \psi \right\rangle = \left\langle \psi \left| \nabla^{2} + \kappa(\mathbf{r}) \right| \psi \right\rangle,$$
$$\left\langle \psi \left| \kappa^{2}(\mathbf{r}) - \kappa_{m}^{2} \right| \psi \right\rangle = 0.$$
(9.68)

That immediately implies

or

$$\delta_m = \frac{3p}{4\pi R_0^3} \int d\mathbf{r} |\psi(\mathbf{r})|^2 \left[\varepsilon(\mathbf{r}) - \varepsilon_m\right] = 0, \qquad (9.69)$$

where  $\varepsilon_m = (\kappa_m/\kappa_0)^2$ , and  $\delta_m$  is the  $\delta$  parameter defined relative to the medium  $\kappa_m$ . Physically, (9.69) means that relative to the homogenized medium, there is no additional delay. In (9.69),  $\varepsilon_m$  is the parameter to be determined. It should be noted that  $\psi(\mathbf{r})$ , being a scattering wave function, implicitly depends on  $\varepsilon_m$ . When (9.69) is satisfied we have

$$v_{\rm t} = \frac{v_m^2}{\omega} \sqrt{\kappa_m^2 - {\rm Re}\Sigma_m^+},\tag{9.70}$$

where  $\Sigma_m^+$  is calculated with the embedding medium characterized by  $\kappa_m$ . Equations (9.69) and (9.70) together define a mean field approach to the calculation of the transport velocity. The interparticle interaction is implicitly taken into account by the requirement of energy density homogeneity. This condition is reasonable because over the scale of the structural correlation length or larger, the medium is geometrically homogeneous. For a delocalized wave, its energy density should be homogeneous over the same spatial scale. On the other hand, this condition should not hold when the states are localized because each localized state has a center, and the decay of the localized states can exhibit large fluctuations even when the scale of observation is larger than the structural correlations length. As a result, (9.69) and (9.70) make sense only in the 3D delocalized regime.

Based on the above scheme, the calculation of  $v_t$  for  $\sqrt{\varepsilon_1} = 2.91$  (the index of refraction for alumina at the microwave frequency) has been carried out for three values of p=0.15, 0.3, and 0.45. The results are plotted as a function of  $\kappa_0 R$  in Fig. 9.9. It is seen that whereas at the low volume fractions there



Fig. 9.9. Calculated transport velocity  $v_t$  plotted as a function of the reduced frequency  $\kappa_0 R$  for classical scalar wave in a 3D medium, consisting of a random dispersion of identical spheres with  $\sqrt{\varepsilon_1} = 2.91$  and p = 0.15, 0.3, and 0.45.

are pronounced dips in  $v_{\rm t}$ , as p increases the dips become smeared out, as expected.

The energy homogenization scheme offers an alternative to the spectral function approach (Chap. 4), where the emphasis is on local coherence. The energy homogenization scheme does not involve coherence at all. This basic idea has been further developed by Busch and Soukoulis (1995, 1996).

# 9.5 The Scaling Function Evaluation

Up to now all the localization calculations in this chapter have been for infinite-sized samples. However, Chap. 8 has shown that an examination of wave localization in finite-sized samples can afford some insight into the crossover behavior from an extended state to a localized state. The key quantity to be examined in this regard is the parameter  $\gamma = \delta \omega / \Delta \omega$ . As shown in Chap. 8,  $\gamma$  may be expressed as

$$\gamma = 2\pi \rho_{\rm e}(\omega) D(\omega) L^{d-2} \tag{9.71}$$

for delocalized states. In order to use the Green function approach for the calculation of a  $\gamma(L)$  that can interpolate between the extended and the localized regimes, two aspects of (9.71) need to be modified.

## 9.5.1 Finite Size Modification of the Diffusion Constant

The first aspect concerns the expression for  $D(\omega)$ 

$$D(\omega) = D^{(B)}(\omega)(1 - C).$$

It is important to recognize that C depends on the sample size L, which was already demonstrated in Chap. 7. In order to make that dependence more precise, it is preferred to have an evaluation of C in which one focuses not on its numerical accuracy but rather on its correct variation with L. Since the end result, the function  $\beta$ , depends only on the functional form of  $\gamma(L)$ , this approach is well suited for our purpose. To this end, let us rewrite (9.3) in the form of (9.5a):

$$C = \frac{1}{K_0} \int \frac{\mathrm{d}\mathbf{Q}}{(2\pi)^d} \frac{1}{\nu^2 + Q^2}.$$
(9.72)

Since  $D^{(B)}(\omega) = K_0 (d\kappa_e^2/d\kappa_0^2)/\pi \rho_e(\omega)$ , an equivalent form of (9.72) is

$$C = \frac{\mathrm{d}\kappa_{\rm e}^2/\mathrm{d}\kappa_0^2}{\pi\rho_{\rm e}(\omega)D^{\rm (B)}(\omega)} \frac{1}{L^d} \sum_{\mathbf{Q}} \frac{1}{\nu^2 + Q^2}.$$
 (9.73)

The sample size dependence of C comes in as the lower limit to the magnitude of Q, 1/L, to be summed over. This is physically reasonable from the discussion in Chap. 6 because the sample size restricts the scattering path length, which may be expressed as a constraint on the diffusion-mode wave vector Qbe larger than 1/L. If in addition we use the fact that in the localized regime the equation for the determination of  $\nu$  is C = 1, with lower limit Q = 0, then (1 - C) is essentially the difference between two integrals which differ only in the lower integration limits, i.e.,

$$\sum_{Q=0}^{1/l} - \sum_{Q=1/L}^{1/l} = \sum_{Q=0}^{1/L}$$

so

$$D(L) = D^{(B)}(\omega)(1-C) = \frac{d\kappa_{e}^{2}/d\kappa_{0}^{2}}{\pi\rho_{e}(\omega)} \frac{1}{L^{d}} \sum_{\mathbf{Q}}^{1/L} \frac{1}{\nu^{2}+Q^{2}}$$

$$= \frac{d\kappa_{e}^{2}/d\kappa_{0}^{2}}{\pi\rho_{e}(\omega)} \begin{cases} \frac{1}{2\pi^{2}} \frac{1}{L} \left[ 1 - \nu L \tan^{-1} \left( \frac{1}{L\nu} \right) \right] & 3D \\ \frac{1}{4\pi} \ln \left[ 1 + \frac{1}{(\nu L)^{2}} \right] & 2D \\ \frac{1}{\pi\nu} \tan^{-1} \left( \frac{1}{L\nu} \right) & 1D. \end{cases}$$
(9.74)

#### 9.5.2 Finite Size Effect for Density of States

The second aspect of (9.71) that requires modification concerns the amount of mobile flux for a finite sample. For delocalized states, a pulse injected at origin would result in a measurable diffusive flux that is proportional to

the density of states  $\rho_{e}(\omega)$ . However, for localized states the measured total flux is expected to decrease as L increases. To account for this phenomenon quantitatively, it is necessary to recall (5.44), the first moment of the Bethe-Salpeter equation. As interpreted in Chap. 5, the  $\nabla \cdot \mathbf{J}$  term in that equation represents the outgoing flux due to a source term (the right-hand side of the equation), i.e.,

$$\nabla \cdot \mathbf{J} = \frac{\mathrm{d}\omega}{\mathrm{d}\kappa_{\mathrm{e}}^2} \pi \rho_{\mathrm{e}}(\omega) \frac{N}{L^d} + \frac{1}{2} \frac{\mathrm{d}\kappa_0^2}{\mathrm{d}\omega} \mathrm{i}\Delta\omega S,$$

or

$$i\Delta \mathbf{k} \cdot \mathbf{J} = \frac{d\omega}{d\kappa_{e}^{2}} \pi \rho_{e}(\omega) \frac{N}{L^{d}} \left[ 1 + \frac{i\Delta\omega}{-i\Delta\omega + D(\omega)|\Delta \mathbf{k}|^{2}} \right], \quad (9.75)$$

where the expression for S, (5.54), has been used. By expressing  $\nu^2 = -i\Delta\omega/D(\omega)$  in the  $\Delta\omega \to 0$  limit [see (9.4b)], it is seen that

$$i\Delta \mathbf{k} \cdot \mathbf{J} = \frac{\mathrm{d}\omega}{\mathrm{d}\kappa_{\mathrm{e}}^2} \frac{N}{L^d} \pi \rho_{\mathrm{e}}(\omega) \frac{|\Delta \mathbf{k}|^2}{\nu^2 + |\Delta \mathbf{k}|^2}.$$
(9.76)

Equation (9.76) tells us that for a unit source, the amount of disturbance away from the source is given by

$$\chi(\Delta \mathbf{k}) = \frac{|\Delta \mathbf{k}|^2}{\nu^2 + |\Delta \mathbf{k}|^2} \frac{N}{L^d}.$$
(9.77)

Not surprisingly, for  $\nu = 0$  (delocalized states), this factor is 1 (up to the normalization factors  $N/L^d$ ).

Let us now consider a finite-sized electronic sample on which a potential U(z) is applied along one direction, given by

$$U(z) = L - z. (9.78)$$

The applied potential will cause a change in the electronic density  $\delta \rho$  such that the net diffusion current would exactly cancel the electric current. To calculate  $\delta \rho$ , one first Fourier transforms (9.77) to get

$$\chi(z) = \frac{L^d}{N} \int_{-\infty}^{\infty} \exp(i\Delta kz) \chi(\Delta \mathbf{k}) \frac{\mathrm{d}(\Delta k)}{2\pi} = \delta(z) - \frac{\nu}{2} \exp(-\nu|z|).$$
(9.79)

The first term on the right-hand side is the local term of the response function. It would be the only term present if the states are delocalized. The second term is nonlocal in character. It is particular to the localized states. The presence of the nonlocal term means that

$$\delta \rho = \rho_{\rm e}(\omega) \int_0^L \chi(z - z') U(z') \mathrm{d}z'.$$
(9.80)

For delocalized states,  $\rho_{\rm e}(\omega)$  in (9.71) gives the density of states participating in electrical transport. In the more general case this may be measured by the quantity

$$-\left[\frac{\mathrm{d}(\boldsymbol{\delta}\boldsymbol{\rho})}{\mathrm{d}z}\right]_{z=L},$$

since it gives the number of states participating in the diffusive counter current. When  $\chi(z) = \delta(z)$ , we get

$$-\frac{\mathrm{d}(\delta\rho)}{\mathrm{d}z}\bigg|_{z=L}$$

$$= -\rho_{\mathrm{e}}(\omega)\lim_{\Delta z \to 0} \frac{\int_{0}^{L} \delta(L + \Delta z - z')U(z')\mathrm{d}z' - \int_{0}^{L} \delta(L - \Delta z - z')U(z')\mathrm{d}z'}{2\Delta z}$$

$$= \frac{1}{2}\rho_{\mathrm{e}}(\omega). \tag{9.81}$$

The factor 1/2 comes from the fact that for the current, there is an abrupt drop to zero outside the sample, so at the boundary what is measured is the mean between zero and one. For delocalized states, therefore,  $-2 \left[ d(\delta \rho)/dz \right]$ at z = L is equivalent to  $\rho_e(\omega)$ . In the case where  $\chi(z)$  is given by (9.79), direct integration plus some care at z = L gives

$$-2\frac{\mathrm{d}(\boldsymbol{\delta}\rho)}{\mathrm{d}z}\Big|_{z=L} = (\nu L+1)\exp(-\nu L)\rho_{\mathrm{e}}(\omega).$$
(9.82)

In the general case, therefore,  $\rho_{\rm e}(\omega)$  should be replaced by  $(\nu L+1) \exp(-\nu L)$  $\rho_{\rm e}(\omega)$  (Vollhardt and Wölfle 1982). It is seen that when  $\nu = 0$ , (9.82) recovers  $\rho_{\rm e}(\omega)$  exactly.

#### 9.5.3 Scaling Function and its Limiting Behaviors

By substituting (9.82) and (9.74) into (9.71), a generalized expression for  $\gamma$  is obtained which includes the finite size effect:

$$\gamma = 2(\nu L + 1) \exp(-\nu L) \pi \rho_{e}(\omega) D(L) L^{d-2}$$

$$= (\nu L + 1) \exp(-\nu L) \begin{cases} \frac{1}{\pi^{2}} \left(1 - \nu L \tan^{-1} \frac{1}{L\nu}\right) & 3D\\ \frac{1}{2\pi} \ln \left[1 + \frac{1}{\nu^{2}L^{2}}\right] & 2D\\ \frac{2}{\pi\nu L} \tan^{-1} \frac{1}{L\nu} & 1D, \end{cases}$$
(9.83)

where  $d\kappa_e^2/d\kappa_0^2$  is set equal to 1; i.e., the starting medium is taken to be the effective medium. Since  $1/\nu$  is proportional to  $\xi$ , it is seen that  $\gamma$  is a function of  $\xi/L$  only, thus guaranteeing the scaling behavior as pointed out in the last

chapter. By writing  $x = \nu L$ , the scaling function in the localized regime is given by

$$\beta = \frac{\mathrm{d} \, \ln \gamma}{\mathrm{d} \, \ln x} = x \frac{\mathrm{d} \ln \gamma}{\mathrm{d} x}$$

$$= \begin{cases} -\frac{x^2}{1+x} + \frac{x^2}{1+x \tan^{-1} x^{-1}} \left[ \frac{1}{1+x^2} - \frac{1}{x} \tan^{-1} \frac{1}{x} \right] & 3\mathrm{D} \\ -\frac{x^2}{1+x} - \frac{2}{1+x^2} \left[ \ln \left( 1 + \frac{1}{x^2} \right) \right]^{-1} & 2\mathrm{D} \\ -\frac{1+x+x^2}{1+x} - \frac{x}{\tan^{-1} x^{-1}} \frac{1}{1+x^2} & 1\mathrm{D} \end{cases}$$
(9.84)

Note that in order to write  $\beta$  in terms of  $\ln \gamma$ , x has to be solved in terms of  $\gamma$  (or  $\ln \gamma$ ) through (9.83).

The limiting behaviors of  $\beta$  may be obtained by letting  $x \to 0$  and  $x \to \infty$ . For 2D and 1D, the  $x \to 0$  limit is given by

$$\lim_{x \to 0} \beta = \begin{cases} \frac{1}{\ln x} = -\frac{1}{\pi \gamma} & 2D\\ -1 & -\frac{2}{\pi}x = -1 & -\frac{2}{\pi \gamma}, & 1D \end{cases}$$
(9.85)

where we have used (9.83). The  $x \to \infty$  limit is given by

$$\lim_{x \to \infty} \beta = -x = \ln \gamma + \text{const.}$$
(9.86)

to the leading order for d=1, 2, and 3. These are precisely the desired leading order asymptotic behaviors as spelled out in Chap. 8. In 3D, the  $x \to 0$  limit yields

$$\lim_{x \to 0} \beta = -\frac{\pi}{2}x = (\pi^2 \gamma - 1). \tag{9.87}$$

The fixed point at which  $\beta = 0$  therefore occurs at

$$\gamma_{\rm c} = \frac{1}{\pi^2}.\tag{9.88}$$

This value is very close to the  $\gamma_c$  value of  $1/3\pi$  estimated in Chap. 8. There is another branch of the scaling function in 3D associated with the delocalized states. For that branch,  $\nu = 0$ , so that

$$D(L) = D^{(B)} \left[ 1 - \frac{1}{\pi \rho_{e}(\omega)} \frac{1}{D^{(B)}} \frac{1}{2\pi^{2}} \left( \frac{\alpha}{l} - \frac{1}{L} \right) \right],$$
(9.89)

where  $\alpha/l$  is the upper cutoff of the Q integral. It follows that

$$\gamma = 2\pi\rho_{\rm e}(\omega)D(L)L = \frac{1}{\pi^2} + 2\pi\rho_{\rm e}D^{\rm (B)}l\frac{L}{\zeta}.$$
(9.90)



**Fig. 9.10.** Calculated scaling function  $\beta(\ln \gamma)$  for d = 1, 2, and 3.

By labeling  $x = L/\zeta$ , we have  $x(d\gamma/dx) = \gamma - 1/\pi^2$  or

$$\beta = \frac{x}{\gamma} \frac{\mathrm{d}\gamma}{\mathrm{d}x} = 1 - \frac{1}{\pi^2 \gamma} \tag{9.91}$$

as the delocalized branch of  $\beta$  in 3D. When  $\gamma \to \infty$ ,  $\beta \to 1$  as expected. The scaling function  $\beta(\ln \gamma)$  is plotted in Fig. 9.10 for d = 1, 2, 3. In the case of 1D, the present scaling function is not identical to that expressed by (8.50). However, since the asymptotic behaviors are similar, the difference is not that significant.

The derivation of the scaling function  $\beta(\ln \gamma)$  indicates that the mechanism of coherent backscattering, with the added self-consistency assumption, is compatible with the scaling hypothesis as enunciated in the last chapter.

# **Problems and Solutions**

**Problem 9.1** Derive from the Maxwell equations an expression for the complex conductivity.

Solution Only one of the Maxwell equations is needed for this purpose:

$$\nabla \times \mathbf{H} = \frac{4\pi}{c} \mathbf{j} + \frac{1}{c} \frac{\partial \mathbf{D}}{\partial t}, \tag{P9.1}$$

where

$$\mathbf{D} = \varepsilon \mathbf{E},$$
$$\mathbf{j} = \sigma_{\mathrm{E}} \mathbf{E},$$

c is the speed of light, **E** is the electric field, **j** is the current density, and **H** is the magnetic field. For a harmonic  $[\exp(-i\Delta\omega t)]$  time variation, (P9.1) becomes

$$\nabla \times \mathbf{H} = \frac{4\pi}{c} \left( \sigma_{\rm E} - \frac{\mathrm{i}\Delta\omega}{4\pi} \varepsilon \right) \mathbf{E}$$
$$= \frac{4\pi}{c} \sigma_{\rm E}^* \mathbf{E}. \tag{P9.2}$$

Consistent with the notations used in this book, here  $\Delta \omega$  is used to denote the modulation frequency.  $\sigma_{\rm E}^* \mathbf{E}$  is a complex current density whose imaginary part (the component that is 90° out of phase with the electric field) represents the dielectric polarization (displacement) current, which arises from bound (localized) electrons.

**Problem 9.2** Express the **k** integral of  $K_0$  in terms of a frequency/energy integration.

**Solution** The essential point in the conversion of the **k** integration to a frequency/energy integration is the division of d**k** into  $dS dk_{\perp}$ , where S is the equal-energy surface and  $k_{\perp}$  is the wavevector normal to that surface. Since

$$\frac{\mathrm{d}k_{\perp}}{\mathrm{d}e} = \frac{1}{\nabla_k e(\mathbf{k})},\tag{P9.3}$$

it follows that

$$d\mathbf{k} = dS \frac{dk_{\perp}}{de} de = dS de \frac{1}{|\nabla_k e(\mathbf{k})|},$$
(P9.4)

and

$$S_{0}(e = \kappa_{0}^{2}) = \int d\mathbf{k} |\nabla_{k} e(\mathbf{k})| \delta \left[\kappa_{0}^{2} - e(\mathbf{k})\right]$$
$$= \int dS \, de \, \delta \left[\kappa_{0}^{2} - e(\mathbf{k})\right]$$
$$= \int_{e = \kappa_{0}^{2}} dS$$
(P9.5a)

is the constant energy surface area for the ordered lattice. In particular, for the 3D simple cubic lattice the **k**-space surface area (in units of  $a^{-2}$ ) is

$$S_0(E) = 16 \int_0^{\pi} \int_0^{\pi} \int_0^{\pi} dx dy dz \sqrt{\sin^2 x + \sin^2 y + \sin^2 z} \times \delta \left[ E + 2\cos x + 2\cos y + 2\cos z \right],$$
(P9.5b)

where  $-6 \le E \le 6$ .

~

With this preliminary, it is easy to write

$$2K_{0} = \frac{2}{(2\pi)^{3}} \int d\mathbf{k} \left[ \nabla_{k} e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}} \right]^{2} \left\{ \frac{-\mathrm{Im}\Sigma^{+}}{\left[\kappa_{0}^{2} - \mathrm{Re}\Sigma^{+} - e(\mathbf{k})\right]^{2} + (\mathrm{Im}\Sigma^{+})^{2}} \right\}^{2}$$

$$= \frac{1}{(2\pi)^{3}} \frac{2}{3} \int dE' \int d\mathbf{k} |\nabla_{k} e(\mathbf{k})|^{2} \delta \left[E' - e(\mathbf{k})\right]$$

$$\times \left\{ \frac{-\mathrm{Im}\Sigma^{+}}{\left[\kappa_{0}^{2} - \mathrm{Re}\Sigma^{+} - e(\mathbf{k})\right]^{2} + (\mathrm{Im}\Sigma^{+})^{2}} \right\}^{2}$$

$$= \frac{1}{(2\pi)^{3}} \frac{2}{3} \int_{0}^{12/a^{2}} dE' \int_{E'} dS |\nabla_{k} e(\mathbf{k})|$$

$$\times \left\{ \frac{-\mathrm{Im}\Sigma^{+}}{\left[\kappa_{0}^{2} - \mathrm{Re}\Sigma^{+} - E'\right]^{2} + (\mathrm{Im}\Sigma^{+})^{2}} \right\}^{2}.$$
(P9.6)

In (P9.6) the integral of  $(\nabla_k e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}})^2$  is written as the integral of  $|\nabla_k e(\mathbf{k})|^2/3$ because  $\Delta \hat{\mathbf{k}}$  is an independent unit vector ( $\Delta \mathbf{k}$  is recalled to be the vector that is the difference between the + channels and the – channels in the  $\langle GG* \rangle_{\rm c}$ evaluation (see Chap. 5) and therefore may be taken as the z axis). Since the integral is symmetrical with respect to the three directions, it follows that the integration of  $|\nabla_k e(\mathbf{k})|^2 = 4 (\sin^2 k_x a + \sin^2 k_y a + \sin^2 k_z a)$  would give three times that of  $|\nabla_k e(\mathbf{k}) \cdot \Delta \hat{\mathbf{k}}|^2 = 4 \ (\sin^2 k_z)$ . To evaluate the dS integral in (P9.6), it is noted that from (P9.5),

$$\begin{split} \int_{E'} \mathrm{d}S \ |\nabla_k e(\mathbf{k})| &= \int \mathrm{d}\mathbf{k} |\nabla_k e(\mathbf{k})|^2 \delta \left[ E' - e(\mathbf{k}) \right] \\ &= \frac{1}{a^3} \int_{-\pi}^{\pi} \int_{-\pi}^{\pi} \int_{-\pi}^{\pi} \mathrm{d}x \, \mathrm{d}y \, \mathrm{d}z \\ &\times 4(\sin^2 x + \sin^2 y + \sin^2 z) \ \delta \left[ a^2 E' - a^2 e \right] \\ &= \frac{96}{a^3} \int_0^{\pi} \int_0^{\pi} \int_0^{\pi} \mathrm{d}x \, \mathrm{d}y \, \mathrm{d}z \ \sin^2 z \\ &\times \delta \left[ E - 6 + 2\cos x + 2\cos y + 2\cos z \right] \\ &= \frac{96}{a^3} \frac{1}{2} \int_0^{\pi} \int_0^{\pi} \int_0^{\pi} \mathrm{d}x \, \mathrm{d}y \frac{\mathrm{d}(2\cos z)}{|-\sin z|} \\ &\times \sin^2 z \ \delta \left[ E - 6 + 2\cos x + 2\cos y + 2\cos z \right] \\ &= \frac{48}{a^3} \int_0^{\pi} \int_0^{\pi} \mathrm{Re} \sqrt{1 - \left( \frac{E - 6}{2} + \cos x + \cos y \right)^2} \mathrm{d}x \ \mathrm{d}y.(\mathrm{P9.7}) \end{split}$$

Combining (P9.6) and (P9.7) yields

$$2K_0 = \frac{32}{(2\pi)^3 a} \int_{E_l}^{E_u} dE \ W_3(2u - \operatorname{Re} y_0 - E) \left\{ \frac{-\operatorname{Im}\Sigma^+}{E^2 + (\operatorname{Im}y_0)^2} \right\}^2, \quad (P9.8)$$

where  $(\kappa_0^2 - \text{Re}\Sigma^+ - E/a^2)a^2$  is relabeled as  $E, y_0 = a^2\Sigma^+$ , and  $u = (\kappa_0^2a^2 - 6)/2\beta$ .  $E_l = 2u - \text{Re}y_0 - 6$ ,  $E_u = 2u - \text{Re}y_0 + 6$ , and

$$W_3(E') = \int_0^{\pi} \int_0^{\pi} \operatorname{Re} \sqrt{1 - \left(\frac{E'}{2} + \cos x + \cos y\right)^2} \, \mathrm{d}x \, \mathrm{d}y.$$
(P9.9)

**Problem 9.3** Carry out the angular integration in (9.12).

**Solution** There are three unit vectors in the integration:  $\Delta \hat{\mathbf{k}}$ ,  $\hat{\mathbf{k}}_i$ , and  $\hat{\mathbf{k}}_f$ . Since it does not matter which one of the three vectors is held fixed in the angular integration, let us choose  $\hat{\mathbf{k}}_i$  as being the z axis. Then, from the addition theorem of spherical harmonics,

$$\cos\theta_{\rm f} = \cos\theta_{\rm i}\cos\theta_{\rm if} + \sin\theta_{\rm i}\sin\theta_{\rm if}\cos\phi_{\rm if},\tag{P9.10}$$

where  $\cos\theta_{\rm f} = \hat{\mathbf{k}}_{\rm f} \cdot \Delta \hat{\mathbf{k}}$ ,  $\cos\theta_{\rm i} = \hat{\mathbf{k}}_{\rm i} \cdot \Delta \hat{\mathbf{k}}$ , and  $\cos\theta_{\rm if} = \hat{\mathbf{k}}_{\rm i} \cdot \hat{\mathbf{k}}_{\rm f}$ . It follows that the expression in the curly brackets of the (9.12a) can be written as

$$\begin{split} &\int_{0}^{2\pi} \mathrm{d}\phi_{\mathrm{i}} \int_{0}^{2\pi} \mathrm{d}\phi_{\mathrm{f}} \int_{-1}^{1} \mathrm{d}\cos\theta_{\mathrm{i}} \int_{-1}^{1} \mathrm{d}\cos\theta_{\mathrm{f}} \frac{\cos\theta_{\mathrm{i}}\cos\theta_{\mathrm{f}}}{a^{2}\nu^{2} + x^{2} + y^{2} + 2xy\cos\theta_{\mathrm{if}}} \\ &= \int \mathrm{d}\Omega_{\Delta\hat{\mathbf{k}}} \int \mathrm{d}\Omega_{\hat{\mathbf{k}}_{\mathrm{f}}} \frac{\left(\Delta\hat{\mathbf{k}}\cdot\hat{\mathbf{k}}_{\mathrm{i}}\right)\left(\Delta\hat{\mathbf{k}}\cdot\hat{\mathbf{k}}_{\mathrm{f}}\right)}{a^{2}\nu^{2} + x^{2} + y^{2} + 2xy\left(\hat{\mathbf{k}}_{\mathrm{i}}\cdot\hat{\mathbf{k}}_{\mathrm{f}}\right)} \\ &= \int \mathrm{d}\Omega_{\Delta\hat{\mathbf{k}}} \int_{0}^{2\pi} \mathrm{d}\phi_{\mathrm{if}} \int_{-1}^{1} \mathrm{d}\cos\theta_{\mathrm{if}} \\ &\times \frac{\cos\theta_{\mathrm{i}}(\cos\theta_{\mathrm{i}}\cos\theta_{\mathrm{if}} + \sin\theta_{\mathrm{i}}\sin\theta_{\mathrm{if}}\cos\phi_{\mathrm{if}})}{a^{2}\nu^{2} + x^{2} + y^{2} + 2xy\cos\theta_{\mathrm{if}}} \\ &= 2\pi \int_{-1}^{1} \mathrm{d}(\cos\theta_{\mathrm{i}})\cos^{2}\theta_{\mathrm{i}} \left[2\pi \int_{-1}^{1} \mathrm{d}u \frac{u}{(a^{2}\nu^{2} + x^{2} + y^{2}) + 2xyu}\right] \\ &= 4\pi^{2}\frac{2}{3} \left[1 - \frac{a^{2}\nu^{2} + x^{2} + y^{2}}{4xy}\ln\left(\frac{(x+y)^{2} + a^{2}\nu^{2}}{(x-y)^{2} + a^{2}\nu^{2}}\right]\frac{1}{xy}. \end{split}$$
(P9.11)

**Problem 9.4** Determine the functional dependence of the 2D localization length in the weak scattering limit of the Anderson model.

**Solution** To determine the functional dependence of  $\xi$  in 2D, it is simpler to use the approximate evaluation of the maximally crossed diagrams' contribution, and the resulting equation (9.7a):

$$\frac{1}{\nu} = \frac{l}{\alpha} \left[ \exp(4\pi K_0) - 1 \right]^{1/2}, \qquad (P9.12)$$

where  $\xi \propto \nu^{-1}, \alpha/l$  is the (upper) cutoff for the Q wave vector, and  $K_0 \propto D^{(B)}\rho_e \propto l$ . We know  $l \propto (\sigma_w a^2)^{-2}$  in the weak scattering limit because the
mean free path is inversely proportional to the scattering cross-section, which is the angular integral of |scattering amplitude|<sup>2</sup>. The scattering amplitude must be proportional to  $\sigma_w a^2$  to the first order because the scattering amplitudes  $\rightarrow 0$  as  $\sigma_w \rightarrow 0$ , and, moreover, the real part of the scattering amplitude (the t matrix element) changes sign as  $\sigma_w$  changes sign. Together they imply linear dependence. It follows that  $\ln \xi \propto 1/(\sigma_w a^2)^2$  (to the leading order) in the weak scattering limit.

# Localization Phenomena in Electronic Systems

# 10.1 Finite Temperatures and the Effect of Inelastic Scattering

In previous chapters the term "scattering" has meant elastic scattering in which the wave frequency is not altered. However, the absence of inelastic scattering is at best an idealization. To be more realistic means to account for inelastic scattering in the interpretation of observed localization phenomenon.

The effect of inelastic scattering is generally expressed through a length scale, the inelastic scattering length  $l_{\rm in}$ . The value of  $l_{\rm in}$  is usually very different for the quantum case and the classical wave case. For classical waves,  $l_{\rm in}$  may be macroscopic in size and is generally insensitive to temperature. Just as in the observation of the coherent backscattering effect, classical wave localization effects may thus be observed with macroscopic samples, relatively free from the hindrance of inelastic scattering. This is not the case, however, with the electronic case. The electronic  $l_{\rm in}$  generally varies with temperature as  $1/T^p$ , where p is some constant on the order of 1–2, and at finite temperatures  $l_{\rm in}$  is invariably microscopic in scale. Therefore, the manifestation of electronic localization effects in macroscopic samples is almost always accompanied by inelastic scattering.

# 10.1.1 Inelastic Scattering in the Classical Wave Case

For classical waves, it is interesting to point out that inelastic scattering does not necessarily diminish the localization effects. In Chap. 6, it has been argued that the effect of inelastic scattering is analogous to that of finite sample size, i.e., cutting off the long scattering paths. However, an important distinction is that inelastic scattering is inherently a probabilistic process. Therefore, even in a semi-infinite dissipative sample there can still be very long backscattering paths. Since for classical waves inelastic scattering is just energy dissipation (frequency down-conversion with no attendant up-conversion), those

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long paths can be easily picked up by measurement at the incident wave frequency, and they are expected to display the full coherent backscattering effect because dissipation acts *equally* on both the scattering path and its time-reversed counterpart, as pointed out by Weaver (1993). However, the amplitudes of such long paths are necessarily diminished, and consequently when they are considered *together* with the shorter paths (which have higher amplitudes) as in a static backscattering experiment, where the consideration is for a single frequency wave (which may be regarded as a Fourier addition of paths of all different travel times) the result is similar to that of finite sample size, as shown in Chap. 6. But if these long scattering paths were separated out from the short scattering paths, e.g., by time-of-flight experiments instead of experiments in the frequency domain, then the localization effect should be observable even in a dissipative random medium.

#### 10.1.2 Inelastic Scattering in the Electronic Case

The same is not true for the electronic systems. Here the effect of inelastic scattering is to establish a *thermal equilibrium* in which energy down-conversion is balanced with energy up-conversion. Therefore, even if the long scattering paths could be separated out from the short scattering paths, these long paths would very likely contain inelastic scattering contributions from other energies and hence have random phase relations. For the electronic case the analogy between inelastic scattering and finite sample size is thus fairly accurate. The intention of this chapter is to describe a few electronic phenomena that result from the localization effect.

It should be noted that there can be two types of electronic inelastic scattering. Besides the electron-phonon scattering, which is caused by electrons losing (or gaining) energy to (from) phonons, there is also the electron-electron scattering. The electron-electron interaction is an extensive subject in itself and will not be treated here. However, the effect of the electron-electron interaction in 2D disordered films is noted to be very similar to that of coherent backscattering. Below we will note the effect of electron-electron interactions only when it is relevant.

# 10.2 Temperature Dependence of the Resistance in 2D Disordered Films

From (6.18) a quantitative correspondence between sample size L and  $l_{in}$  is given by

$$L_{\rm in} = \sqrt{\frac{l^* l_{\rm in}}{d}},\tag{10.1}$$

where the subscript "in" on L is intended to emphasize its inelastic scattering origin, and  $l^*$  is the elastic transport mean free path. From the temperature dependence of  $l_{in}$ , it follows that the effect of inelastic scattering may be 10.2 Temperature Dependence of the Resistance in 2D Disordered Films 283

translated into a temperature-dependent  $L_{in}$ :

$$L_{\rm in}(T) \propto T^{-p/2}.$$
 (10.2)

The value of p is ~2 if the inelastic scattering is caused by electron-phonon interaction. Below 5 K, however, there is evidence that the electron-electron interaction is the dominant inelastic scattering mechanism, for which p has a lower value. From the sample-size dependence of renormalized  $\sigma_{\rm E}$ , which has been addressed in both Chapters 7 and 9, (10.2) directly implies a temperature dependence for  $\sigma_{\rm E}$  which could be observed.

The rationale underlying the substitution of  $L_{in}(T)$  for the sample size L is that, in a macroscopic sample,  $L_{in}^d$  is the basic volume beyond which all phase coherence effects are lost. That means the macroscopically measured value of  $\sigma_E$  may be obtained from  $\{\sigma_E(L_{in})\}$ , defined as the values of  $\sigma_E$  measured on the scale of  $L_{in}$ , through a simple resistor-network calculation. If elastic scattering is weak so that the values of  $\sigma_E(L_{in})$  are narrowly distributed, then

 $\sigma_{\rm E} \sim \text{mean of } \{\sigma_{\rm E}(L_{\rm in})\}.$ 

Therefore, although the actual sample size L is macroscopic, nevertheless  $\sigma_{\rm E}$  corresponds to the mean of  $\{\sigma_{\rm E}(L_{\rm in})\}$ . In other words,  $L_{\rm in}$  defines the scale beyond which the intensive nature of conductivity is preserved. However, this picture is valid only when the sample is fairly conducting. When the states are strongly localized, an alternative picture emerges. This is discussed in Sec. 10.4.

# 10.2.1 Manifestation of Localization Correction in 2D Conductivity

Consider a film made by sputtering or evaporation of metal atoms onto an insulating substrate. The film is generally disordered due to the presence of either impurities or structural inhomogeneities. In regard to the localization effect, the film may be regarded as 2D if the thickness L is much less than  $L_{\rm in}$  because the coherent backscattering effect is cut off in the spatial dimension perpendicular to the film. For most metallic films, L < 100 Å may be regarded as a 2D film at T < 20 K. From (9.74) and the Einstein relation, one can write

$$\sigma_{\rm E} = \frac{2e^2}{\hbar} \frac{1}{4\pi^2} \ln\left[1 + \frac{1}{(\nu L_{\rm in})^2}\right]$$
$$\approx -\frac{e^2}{\hbar} \frac{1}{\pi^2} \ln(L_{\rm in}\nu), \tag{10.3}$$

where a factor of 2 has been included to account for the electronic spin degeneracy, and the factor 1 in the argument of  $\ln[1+(\nu L_{\rm in})^{-2}]$  is neglected because



Fig. 10.1. Measured resistance of a disordered Au–Pd film plotted as a function of  $\ln T$ . The data are from Dolan and Osheroff (1979). Here V/I is the measured resistance and  $R_0$  is the wire resistance in series with the sample and therefore has to be subtracted from the measurements.

it is assumed that the localization length is much larger than  $L_{in}$ ; i.e., the system is weakly localized at those temperatures considered. Combining (10.3) with (10.2) yields

$$\sigma_{\rm E} \cong \frac{e^2}{\hbar} \frac{p}{2\pi^2} \ln T + \text{constant.}$$
(10.4)

As  $T \to 0$ ,  $\sigma_{\rm E}$  decreases as  $\ln T$ . Since the constant term is usually larger than the  $\ln T$  term,  $\rho_{\rm E} = 1/\sigma_{\rm E} \propto -\ln T$ . Therefore, at T < 10 K, the resistance of a thin disordered film is expected to increase logarithmically. In fact, experimentally this is a fairly general phenomenon, as shown in Fig. 10.1. As Tincreases, however, a transition to (delocalized) 3D behavior occurs, and the role of the coherent backscattering becomes that of a correction to the classical conductivity expression. In that regime the temperature dependence of  $\sigma_{\rm E}$ is dominated by the temperature dependence of the usual metallic behavior, in which  $\sigma_{\rm E}$  decreases as T increases due to increased electron-phonon scattering. Thus, disordered metallic films generally display a shallow resistance minimum at  $T \approx 10$  K.

From (10.4), the slope of the  $\ln T$  behavior is seen to be a constant. However, in actual experiments the slope of the  $\ln T$  behavior can deviate from that predicted by (10.4). This could be an indication that the finite thickness of the film plays a role in determining the value of the coefficient, or that the electron-electron scattering mechanism is operative.

# 10.3 Magnetoresistance of Disordered Metallic Films

In a metallic film, the normal magnetoresistance is positive because when a magnetic field is applied perpendicular to the film (with an electric field applied in the plane of the film), the electrons tend to move normal to the electric

field direction (in the plane of the film). Since the current is deflected away from the electric field direction, the measured resistance increases. Magnetoresistance is generally expressed as  $\Delta \rho_{\rm E}(H)/\rho_{\rm E}(0)$ , where  $\Delta \rho_{\rm E}(H)$  is the change in resistance as a function of applied magnetic field H, and  $\rho_{\rm E}(0) = [\sigma_{\rm E}(0)]^{-1}$ is the resistivity at zero magnetic field. Magnetoresistance in metallic films is typically on the order of 1% or less for magnetic field up to 5–10 T.

# 10.3.1 Effect of a Magnetic Field

Just as inelastic scattering tends to weaken the localization effect, the presence of a magnetic field also tends to lessen the coherent backscattering effect by destroying the time-reversal symmetry. Since the localization effect increases the resistance, the renormalized resistance of a disordered film must be higher than that without the coherent backscattering correction. By using the magnetic field to quench the coherent backscattering correction, one should therefore see a decrease in the measured resistance, i.e., negative magnetoresistance (Kawabata 1980), just opposite to what is expected for normal metallic behavior.

To calculate quantitatively the effect of a magnetic field, it is necessary to observe that, in quantum mechanics, the wave vector (momentum) of an electron in a magnetic field is given by

$$\mathbf{k} - \frac{e\mathbf{A}}{\hbar c},$$

where c is the speed of light and **A** is the vector potential, related to **H** by

$$\mathbf{H} = \nabla \times \mathbf{A}.\tag{10.5}$$

If  $\mathbf{H} = H \, \hat{\mathbf{k}}$ , then one choice of  $\mathbf{A}$  would be

$$\mathbf{A} = (-Hy, 0, 0). \tag{10.6}$$

In the present case, where we are interested in monitoring the transport behavior of wave intensity, it is noted that the momenta of both the + and - channels are modified:

$$egin{aligned} \mathbf{k}_+ &
ightarrow \mathbf{k}_+ - rac{e \mathbf{A}}{\hbar c} \ \mathbf{k}_- &
ightarrow \mathbf{k}_- + rac{e \mathbf{A}}{\hbar c}, \end{aligned}$$

where the sign of the vector potential in the negative channel is reversed in accordance with the time arrow. The net result for  $\Delta \mathbf{k}$  is

$$\Delta \mathbf{k} = \mathbf{k}_{+} - \mathbf{k}_{-} \to \Delta \mathbf{k} - \frac{2e\mathbf{A}}{\hbar c}, \qquad (10.7)$$

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i.e., there is an additional factor of 2 associated with **A** for  $\Delta \mathbf{k}$ . In Chap. 5, it has been shown that  $\Delta \mathbf{k}$  is the conjugate variable to the propagation distance  $\mathbf{r} - \mathbf{r}'$ , so that in real space it corresponds to the operator  $-i\nabla$ . It follows that whereas  $|\Delta \mathbf{k}|^2$  may be regarded as the eigenvalue of the operator  $-\nabla^2$ , which takes on continuous values, now the corresponding operator is

$$-\nabla^2 \rightarrow \left(i\nabla + \frac{2e\mathbf{A}}{\hbar c}\right)^2,$$
 (10.8)

which has a discrete set of eigenvalues for the momenta perpendicular to the magnetic field as shown in the solution to Problem P10.1. These discrete levels correspond to the well-known Landau levels.

# 10.3.2 Magnetic Field Correction to the Coherent Backscattering Effect

In (7.17), the coherent backscattering correction to the diffusion constant is expressed as an integral over  $\mathbf{Q}$ , where  $\mathbf{Q}$  plays exactly the role of  $\Delta \mathbf{k}$  (see Chap. 7). In view of the quantized Landau levels, the integration over the  $Q^2$ values should be replaced by a sum over the eigenvalues of the new operator, given by

$$Q^2 \to \left(n + \frac{1}{2}\right) \frac{4}{l_H^2} + k_z^2, \qquad n = 0, 1, 2, \dots,$$
 (10.9)

where  $k_z$  is the wave vector in the direction normal to the film, and

$$l_H = \sqrt{\frac{\hbar c}{eH}}$$

is a magnetic length scale which measures the size of the quantized orbits in the magnetic field. It is on the order of 400 Å for H = 1 T. We will treat the magnetoresistance in both the thick- and the thin-film limits when a magnetic field is applied normal to the film.

# 10.3.3 Thick Film Case

From (7.17), it is easy to deduce that when the film is thick, the  $\delta \sigma_H$  due to the coherent backscattering is given by

$$\delta\sigma_H = -\frac{e^2/\hbar}{\pi} \frac{2}{(2\pi)^3} \sum_{n=0}^{N_0} \int_{-1/l}^{1/l} \mathrm{d}Q_z \frac{\pi(4/l_H^2)}{-\frac{i\Delta\omega}{D(\omega)} + \frac{4}{l_H^2} \left(n + \frac{1}{2}\right) + Q_z^2}.$$
 (10.10)

Here  $N_0 = (l_H/l)^2$ , l being the mean free path,  $dQ_x dQ_y$  is converted into  $\pi dQ_{\parallel}^2$ , where  $Q_{\parallel}^2$  means the component of  $Q^2$  perpendicular to the magnetic field (parallel to the film), and since the eigenvalues of  $Q_{\parallel}^2$  are quantized in units of  $4/l_H^2$ , the integral is replaced by summation. A factor of 2 has been included to account for the spin degeneracy, and the integral and summation are both cut off at  $|Q_z|, |Q_{\parallel}| = 1/l$ .

# 10.3.4 Thin Film Case

If the disordered material is in the form of a thin film with thickness L, then the  $dQ_z$  integral would have to be split into two parts: one from -1/l to -1/Land another one from 1/L to 1/l. As L approaches l, however, wave quantization in the z direction is expected where each quantized state corresponds to a standing wave. In the limit where the film is so thin that the z variation of the wave may be regarded as uniform, one can replace  $Q_z^2$  in the denominator of the integrand by zero and do away with the  $dQ_z$  integration. The resulting expression involves only a simple summation, as shown below by (10.11b).

#### 10.3.5 Diffusion and Inelastic Relaxation

An important point to be considered at this point is the factor  $-i\Delta\omega/D(\omega)$ [following the self-consistent approximation,  $D^{(B)}(\omega)$  is replaced by  $D(\omega)$ ], which was argued to be proportional to  $1/\xi^2$  in the last chapter. In the presence of inelastic scattering, if  $L_{\rm in} \ll \xi$ , then  $L_{\rm in}$  plays a much more dominant role. That is, in addition to the  $-i\Delta\omega$  term, there is another term,  $1/\tau_{\rm in}$ , whose origin lies in the presence of the additional relaxation term in the continuity equation for S [see (5.44b)]:

$$\frac{1}{2} \frac{\mathrm{d}\kappa_0^2}{\mathrm{d}\omega} \left( \frac{\partial S}{\partial t} + \frac{S}{\tau_{\mathrm{in}}} \right) + \nabla \cdot \mathbf{J} = \text{source term.}$$

Here  $\tau_{in}$  denotes the inelastic scattering time. Therefore,

$$-\frac{\mathrm{i}\Delta\omega}{D(\omega)} \to \nu^2 + \frac{1}{\tau_{\mathrm{in}}D(\omega)}$$

By substituting  $lv_t/3$  for  $D(\omega)$  and defining  $l_{\rm in} = \tau_{\rm in} v_t$ , we get

$$\tau_{\rm in} D(\omega) = l_{\rm in} l/3 = L_{\rm in}^2$$

exactly. Under the assumption that  $L_{\rm in} \ll \xi$ , the  $\nu^2 \propto 1/\xi^2$  term can be neglected, and the thick-film expression for  $\delta\sigma_H$  is given by

$$\delta\sigma_{H} \simeq -\frac{e^{2}}{\hbar} \frac{1}{\pi^{3}} \frac{1}{l_{H}^{2}} \sum_{n=0}^{N_{0}} \int_{-1/l}^{1/l} \mathrm{d}Q_{z} \frac{1}{\frac{1}{L_{\mathrm{in}}^{2}} + \frac{4}{l_{H}^{2}} \left(n + \frac{1}{2}\right) + Q_{z}^{2}}$$

$$= -\frac{e^{2}}{\hbar} \frac{1}{l_{H}\pi^{3}} \sum_{n=0}^{N_{0}} \int_{-l_{H}/l}^{l_{H}/l} \mathrm{d}p \frac{1}{\left(l_{H}/L_{\mathrm{in}}\right)^{2} + 4\left(n + \frac{1}{2}\right) + p^{2}}$$

$$= -\frac{e^{2}}{\hbar} \frac{2}{l\pi^{3}} \sum_{n=0}^{N_{0}} \frac{l/l_{H}}{\sqrt{\left(l_{H}/L_{\mathrm{in}}\right)^{2} + 4\left(n + \frac{1}{2}\right)}}$$

$$\times \tan^{-1} \frac{l_{H}/l}{\sqrt{\left(l_{H}/L_{\mathrm{in}}\right)^{2} + 4\left(n + \frac{1}{2}\right)}}.$$
(10.11a)

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Equation (10.11a) gives the amount of conductivity correction arising from coherent backscattering, in the presence of a magnetic field. The thin-film, 2D version of  $\delta\sigma_H$  is given by

$$\delta\sigma_H = -\frac{2e^2}{\hbar} \frac{1}{\pi^3} \sum_{n=0}^{N_0} \frac{1}{(l_H/L_{\rm in})^2 + 4\left(n + \frac{1}{2}\right)}.$$
 (10.11b)

The magnetoresistance is defined relative to  $\delta \sigma_H$  at H = 0. Therefore, let

$$\Delta \sigma_{\rm E}(H) = \delta \sigma_H - \delta \sigma_0$$

where  $\delta \sigma_0$  may be obtained from (10.11a) by converting it into an integral in the limit of H = 0, as shown in the solution to Problem P10.2:

$$\delta\sigma_{0} = -\frac{e^{2}}{\hbar} \frac{1}{l\pi^{3}} \left( \sqrt{4 + \left(\frac{l}{L_{\rm in}}\right)^{2}} \tan^{-1} \left[4 + \left(\frac{l}{L_{\rm in}}\right)^{2}\right]^{-1/2} - \left(\frac{l}{L_{\rm in}}\right) \tan^{-1} \left(\frac{L_{\rm in}}{l}\right) + \frac{1}{2} \ln \frac{5L_{\rm in}^{2} + l^{2}}{L_{\rm in}^{2} + l^{2}} \right).$$
(10.12)

# 10.3.6 Magnetoresistance and the Distance to Localization

It follows that the magnetoresistance  $\Delta \rho_{\rm E}(H)/\rho_{\rm E} \simeq -\Delta \sigma_{\rm E}(H)/\sigma_{\rm E}$  is given by

$$\frac{\Delta\rho_{\rm E}(H)}{\rho_{\rm E}} = -\frac{\Delta\sigma_{\rm E}(H)}{\sigma_{\rm E}} 
= -\frac{3}{\pi(\kappa_{\rm e}l)^2} 
\times \left\{ \sqrt{4+a^{-2}} \tan^{-1}\frac{1}{\sqrt{4+a^{-2}}} + \frac{1}{2}\ln\frac{5a^2+1}{a^2+1} - \frac{1}{a}\tan^{-1}a 
-\sum_{n=0}^{1/\bar{H}}\frac{2\sqrt{\bar{H}}}{\sqrt{(\bar{H}a^2)^{-1}+4\left(n+\frac{1}{2}\right)}} \tan^{-1}\frac{1/\sqrt{\bar{H}}}{\sqrt{(\bar{H}a^2)^{-1}+4\left(n+\frac{1}{2}\right)}} \right\}.$$
(10.13a)

Here

$$a = \frac{L_{\rm in}}{l}, \qquad \bar{H} = \left(\frac{l}{l_H}\right)^2,$$

and we have used the relation

$$\sigma_{\rm E} = 2 \frac{e^2}{\hbar} \rho_{\rm e}(\omega) \frac{1}{3} v l = \frac{e^2}{\hbar} \frac{1}{3\pi^2} (\kappa_{\rm e} l)^2 \frac{1}{l},$$

where the spherical Brillouin zone approximation is used to get  $\rho_{\rm e}(\omega) = 4\pi\kappa_{\rm e}^2(\mathrm{d}\kappa_{\rm e}/\mathrm{d}\omega)/(2\pi)^3 = \kappa_{\rm e}^2/2\pi^2 v$ . When  $1/\bar{H}$  is not an integer, the last term



Fig. 10.2. Calculated magnetoresistance for  $\kappa_e l = 10$ , plotted as a function of the dimensionless magnetic field  $\overline{H}$ . Four cases of  $a = 2, 10, 50, \infty$  are considered, where  $a = \infty$  corresponds to the T = 0 case.

in the summation of (10.13a) is understood to be multiplied by the fraction of  $1/\bar{H}$  that is over the largest integer smaller than  $1/\bar{H}$ . What is clear from (10.13a) is that the maximum amount of magnetoresistance is  $\sim -(\kappa_e l)^{-2}$ . Since in 3D the value of  $\kappa_e l$  is a measure of the proximity to localization, for a disordered conducting film with  $\kappa_e l \ge 10$  the maximum magnetoresistance is on the order of -1%. A plot of  $\Delta \rho_{\rm E}(H)/\rho_{\rm E}$  is shown in Fig. 10.2 for  $\kappa_e l = 10$ and  $a = 2, 10, 50, \infty$ . The magnetoresistance is seen to vanish as the constant a becomes smaller, i.e., as T increases.

Similar considerations for the 2D film yield

$$\frac{\Delta\rho_{\rm E}(H)}{\rho_{\rm E}} = -\frac{1}{\pi^2(\kappa_{\rm e}l)} \left[ \ln(1+4a^2) - \sum_{n=0}^{1/\bar{H}} \frac{1}{(4a^2\bar{H})^{-1} + \left(n + \frac{1}{2}\right)} \right], \quad (10.13b)$$

i.e., the magnitude of the magnetoresistance is governed by  $(\kappa_{\rm e}l)^{-1}$ . Figure 10.3 shows the measured variation of resistance versus the magnetic field for a thin Mg film (top) and the resistance variation of the same film after a partial layer of Au is superposed (bottom) (Bergmann 1984). The top set of curves has been shown to be well fitted by (10.11b), whereas the bottom set of curves illustrates another effect, the spin–orbit coupling.

# 10.3.7 Spin-Orbit Coupling

The additional magnetoresistance effect is due to the coupling of the electronic spin to the wave vector  $\mathbf{k}$ , called the spin-orbit coupling. When the spin-orbit coupling is strong, the sign of the spin switches as  $\mathbf{k} \to -\mathbf{k}$  (i.e., spin follows the orbital magnetic moment) so that the overall sign of the coupling energy



Fig. 10.3. Measured magnetoresistance of an Mg film at different temperatures, from Bergmann (1984). Both the magnitude of the effect and its variations with the temperature and the magnetic field are similar to those shown in Fig. 10.2. The *top* part shows the magneto resistance to be negative, as predicted. The curves have been quantitatively fitted by (10.11b). The *bottom part* shows that by adding only a partial layer of Au on top of the Mg film, the magneto resistance changes sign at small magnetic field. Here the magnetic field is denoted by B, with T denoting tesla.

is not altered. That means for a scattering path and its time-reversed counterpart the original constructive interference in the backward direction now becomes destructive (due to the reversal of the sign of the time-reversed path), thus implying a positive magnetoresistance since the effect of a magnetic field becomes just the opposite - reducing the coherent backscattering means increasing the resistance. As the magnetic field increases in strength, however, the spins eventually have to be aligned with the magnetic field (the spin–orbit effect no longer dominates), and the magnetoresistance turns negative again. The magnitude of the spin–orbit coupling generally increases with the atomic charge. That is why the presence of a heavy metal, e.g., Au, alters the sign of the magnetoresistance at small magnetic field.

# 10.4 Transport of Localized States at Finite Temperatures: Hopping Conduction

Suppose for whatever reason an electron becomes strongly localized; then at T = 0 it is by definition nonmobile and therefore cannot contribute to current transport beyond its localization length. This situation is altered at finite temperatures, because although the localized states are separated in both their energies and spatial locations, an electron can nevertheless "hop" from one

localized state to another. The term "hopping" is coined to denote the combination of two mechanisms - thermal activation (up in energy) and tunneling (across in space) - which makes current transport possible for localized states.

Although thermal activation and tunneling are two nominally independent processes, Mott (1968) was the first to show that these two processes can actually couple to result in a nonactivated temperature dependence that is characteristic of hopping conduction. This coupling may be elucidated through the so-called critical-path method first proposed by Ambegaokar et al. (1971).

# 10.4.1 Constructing the Critical Percolating Path of Least Resistance

Consider the conductance  $\Gamma_{ij}$  between two localized sites *i* and *j*.  $\Gamma_{ij}$  may be expressed as the product of the probabilities for thermal activation and tunneling:

$$\Gamma_{ij} = \Gamma_0 \exp(-2\chi S_{ij} - E_{ij}/k_{\rm B}T),$$
 (10.14)

where  $\chi = \xi^{-1}$  is the tunneling constant,  $S_{ij}$  is the tunneling distance between the two sites,  $k_{\rm B}$  denotes Boltzmann constant, and

$$E_{ij} = \max[E_i, E_j] = \frac{1}{2}(|E_i - E_j| + |E_i| + |E_j|)$$
(10.15)

is the activation energy between the two sites. In (10.14), the indices i and j are not required to denote only the nearest-neighbor sites. Rather, they span all pairs of sites. The expression for  $E_{ij}$ , (10.15), requires some explanation. Suppose  $E_i > E_j$ , where  $E_{ij}$  is measured from the Fermi level; then for a charge to go from j to i it has to overcome an energy barrier  $E_i - E_j$  through thermal activation. However, for the charge to hop from i to j, there is no such barrier. In other words, the barrier is asymmetric. But this reasoning ignores the probability that a charge is on site i or j in the first place, before the hopping occurred. If that probability is taken into account, which in thermal equilibrium is proportional to  $\exp(-E_{i(j)}/k_{\rm B}T)$ , then the two conductances become symmetrical because going from j to i now means

$$\exp(-E_j/k_{\rm B}T)\exp[-(E_i - E_j)/k_{\rm B}T] = \exp(-E_i/k_{\rm B}T),$$

which is the same as going from i to j.

The conductances  $\Gamma_{ij}$  as expressed by (10.14) are noted to be correlated and not independent of each other. Consider  $\Gamma_{ij}$  and  $\Gamma_{ik}$  which share the same site *i*. If  $E_i$  is varied, then  $\Gamma_{ij}$  and  $\Gamma_{ik}$  would both vary in a correlated way. This correlation is an important character of the hopping conduction.

To apply the critical-path method, let us consider the following rule: Pick a  $\Gamma$  and consider the pair ij connected if  $\Gamma_{ij} \geq \Gamma$  and disconnected otherwise. If the value of  $\Gamma$  is sufficiently large, then only a few  $\Gamma_{ij}$ 's can satisfy the connection criterion, and the connected sites form only disjoint clusters. Now lower the value of  $\Gamma$  by a small amount and apply the rule again. More links

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will be added to the existing clusters. By lowering  $\Gamma$  continuously, the connected clusters are expected to increase in size until at  $\Gamma = \Gamma_c$ , the percolation conductance, an infinite network of connected sites is formed. The fact that there must be such a  $\Gamma_c$  is based on the phenomenon of percolation, in which an infinite conduction network must form at some critical concentration of bonds connecting the sites.

The value of  $\Gamma_{\rm c}$  may be regarded as the hopping conductance of the macroscopic sample for the following reasons. First, since at  $\Gamma > \Gamma_{\rm c}$  the network is not percolating, the last added resistance with the value of  $(\Gamma_c)^{-1}$  must be in series with the resistance of the rest of the network. Since (10.14) gives an exponential form for  $\Gamma_{ij}$ , even a relatively narrow distribution in the values of  $S_{ij}$  and  $E_{ij}$  can translate into an exponentially broad distribution for  $\Gamma_{ij}$ . As a result, most of the resistances in the percolating network may be regarded as having zero resistance compared with  $(\Gamma_c)^{-1}$ , and only an exponentially small number of resistances are comparable with  $(\Gamma_c)^{-1}$ . The overall network resistance should, therefore, be on the order of  $(\Gamma_c)^{-1}$ . Second, those conductances  $\Gamma_{ij} < \Gamma_{\rm c}$  can add additional parallel conduction channels; however, here again the exponentially broad distribution of  $\Gamma_{ij}$  means the additional parallel channels will mostly have conductances exponentially smaller than  $\Gamma_{\rm c}$ , and only very few are comparable to  $\Gamma_{\rm c}$ . The net result of these two considerations is that  $\Gamma_{\rm c}$  characterizes the overall hopping conductance, if not accurately in absolute value, certainly in qualitative features such as the temperature dependence.

#### 10.4.2 Temperature Dependence of the Critical Path Conductance

To examine the temperature dependence of  $\Gamma_c$ , let us start with the inequality  $\Gamma_{ij} \geq \Gamma_c$ , or

$$\Gamma_0 \exp(-2\chi S_{ij} - E_{ij}/k_{\rm B}T) \ge \Gamma_{\rm c}.$$
(10.16)

By taking the logarithms of both sides of the inequality and substituting (10.15) for  $E_{ij}$ , one gets

$$\frac{S_{ij}}{S_m} + \frac{|E_i| + |E_j| + |E_i - E_j|}{E_m} \le 1,$$
(10.17)

where

$$S_m = \frac{1}{2\chi} \ln \frac{\Gamma_0}{\Gamma_c},\tag{10.18a}$$

$$E_m = 2k_{\rm B}T\ln\frac{\Gamma_0}{\Gamma_{\rm c}}.$$
(10.18b)

Equation (10.17) makes clear that there is indeed a correlation between thermal activation and tunneling. This correlation arises from the global optimization of the hopping conduction paths based on the concept of percolation. Besides relating  $S_{ij}$  to  $E_{ij}$ , (10.17) also tells us that in order for the inequality to hold,  $S_{ij}$  must be less than  $S_m$ , and  $E_i$ ,  $E_j$  must be less than  $E_m$ . Therefore,  $S_m$  and  $E_m$  are the respective upper bounds to  $S_{ij}$  and  $E_i$  (or  $E_j$ ).

The requirement for the formation of a critical percolation network may be cast in the form that the average number of bonds emanating from a given site must exceed a critical number (Shante 1973). Since the actual value of this number is not important, let us denote it as  $b_c$ . This condition must be an equality at the percolation threshold, and may be written in the following form for the site *i*:

$$b_{\rm c} = \rho_0 \int_{-E_m}^{E_m} \mathrm{d}E_j \int_0^{S_m} \mathrm{d}S_{ij} \, S_d \, S_{ij}^{d-1}.$$
 (10.19)

Here the right-hand side expresses the average number of bonds emanating from site *i* that satisfy the condition of (10.17), under the assumption of homogeneous distribution of the sites and a constant density of states  $\rho_0$ .  $S_d = 4\pi$ ,  $2\pi$ , 2 denotes the area of a *d*-dimensional sphere of unit radius. By performing the integral, it is seen that

$$b_{\rm c} = \left(\frac{2\rho_0}{d}S_d\right)S_m^d E_m. \tag{10.20}$$

By substituting (10.18a) and (10.18b) into (10.20), it is seen that

$$\left[\ln\frac{\Gamma_0}{\Gamma_c}\right]^{d+1} \frac{k_{\rm B}T}{(2\chi)^d} \frac{4S_d}{d} \rho_0 = b_{\rm c},\tag{10.21}$$

so that

$$\Gamma_{\rm c} = \Gamma_0 \exp[-(T_0/T)^{1/(d+1)}],$$
(10.22)

where

$$T_0 = \frac{(2\chi)^d b_{\rm c} d}{4S_d \rho_0 k_{\rm B}}.$$
 (10.23)

Equation (10.22) is called Mott's law, which has been widely observed in amorphous materials (Mott 1968).

## 10.4.3 Variable Range Hopping

Physically, the fractional temperature dependence comes from the continuous shifting of the globally optimal conduction paths as the temperature is varied. This can best be described in terms of Mott's argument. At high temperatures, activation is easy, so the limiting factor is the tunneling distance. The optimal path is thus the one with the smallest tunneling distances between the successive sites, and the temperature dependence is activated, with the activation energy given by the largest  $E_i$  of the conduction path sites. As temperature is lowered, activation becomes more of a limiting factor, and the

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trade-off between a lower activation energy and a larger tunneling distance becomes favorable to the latter. Another way of looking at the trade-off is that as T is lowered, the number of thermally accessible states becomes smaller, and it forces the tunneling distances to be larger so that there can still be a percolating network for conduction. Since the tunneling distance can be continuously increased as T decreases, this type of conduction is referred to as "variable-range hopping." A heuristic way of estimating the optimal hopping distance is to minimize the exponent,

$$\frac{E_{ij}}{k_{\rm B}T} + 2\chi S_{ij},$$

by letting

$$E_{ij} \approx (C_d S^d_{ij} \rho_0)^{-1}$$

where  $C_d = 4\pi/3$ ,  $\pi$ , 2 for d = 3, 2, 1, respectively. That is,  $E_{ij}$  is approximated by the inverse of the average level separation within a sphere of radius  $S_{ij}$ . By differentiating the exponent with respect to  $S_{ij}$  and setting the result to zero, one gets the optimal tunneling distance  $(S_{ij})_o$  as

$$(S_{ij})_{\rm o} = \left(\frac{d}{2\chi C_d \rho_0 k_{\rm B} T}\right)^{1/(d+1)}.$$
 (10.24)

Substitution of (10.24) back into the exponent yields the same fractional temperature dependence of the hopping conduction as that given by (10.22). From  $(S_{ij})_{\rm o}$  it is seen that the tunneling distance varies as  $T^{-1/4}$  in 3D.

The spatial dimensionality dependence of the temperature variation is an inherent characteristic of the hopping conduction. However, in 1D this behavior has to be qualified. This is so because if there is a continuous distribution of  $S_{ij}$ , then as the charge is transported along the chain it will inevitably encounter a particularly large value of  $S_{ij}$ . In 1D, there is no chance that a barrier can be circumvented as in 2D or 3D. Therefore, if the distribution of  $S_{ij}$  is continuous, the tunneling part of the hopping mechanism will always be stopped by a barrier with  $S_{ij} \to \infty$ . In that case, simple activation is the only allowed mechanism possible, and the temperature dependence must therefore be activated. However, if the distribution of  $S_{ij}$  has an upper bound then the hopping behavior would be recovered when the temperature is low enough so that  $(S_{ij})_0$  exceeds the upper bound.

# **Problems and Solutions**

**Problem 10.1** Calculate the eigenvalue spectrum of the operator  $(i\nabla + 2e\mathbf{A}/\hbar c)^2$ , where  $\mathbf{A} = (-Hy, 0, 0)$ .

**Solution** By writing the operator in its component form and denoting the eigenvalue by  $K^2$ , we have

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$$\left[\left(\mathrm{i}\frac{\partial}{\partial x} - \frac{2eHy}{\hbar c}\right)^2 - \frac{\partial^2}{\partial y^2} - \frac{\partial^2}{\partial z^2}\right]\phi(x, y, z) = K^2\phi(x, y, z).$$
(P10.1)

Since the operator  $2e\mathbf{A}/\hbar c$  does not have a x or z coordinate, one can use the following form for  $\phi$ :

$$\phi(x, y, z) = \exp(\mathrm{i}k_x x + \mathrm{i}k_z z)\psi(y).$$
(P10.2)

Substitution of (P10.2) into (P10.1) yields

$$-\frac{\mathrm{d}^2\psi(y)}{\mathrm{d}y^2} + \left[k_z^2 + \left(k_x + \frac{2eHy}{\hbar c}\right)^2\right]\psi(y) = K^2\psi(y).$$
(P10.3)

By defining

$$l_H = \sqrt{\frac{\hbar c}{eH}},$$

one gets

$$-\frac{\mathrm{d}^2\psi(y)}{\mathrm{d}y^2} + \frac{4}{l_H^4} \left(y + \frac{1}{2}k_x l_H^2\right)^2 \psi(y) + k_z^2 \psi(y) = K^2 \psi(y).$$
(P10.4)

The first two terms correspond exactly to the linear oscillator, so the eigenvalue  $K^2$  is given by

$$K^{2} = \left(n + \frac{1}{2}\right)\frac{4eH}{\hbar c} + k_{z}^{2}, \qquad n = 0, 1, 2, \dots.$$
(P10.5)

**Problem 10.2** Derive an analytical expression for  $\delta \sigma_H$  in the limit of  $H \to 0$ .

**Solution** From (10.11),  $\delta \sigma_H$  may be written as

$$\delta\sigma_{H} = -\frac{e^{2}}{\hbar} \frac{1}{l\pi^{3}} \sum_{n=0}^{1/\bar{H}} \frac{2\sqrt{\bar{H}}}{\sqrt{(\bar{H}a^{2})^{-1} + 4\left(n + \frac{1}{2}\right)}} \tan^{-1} \frac{1}{\sqrt{a^{-2} + 4\bar{H}\left(n + \frac{1}{2}\right)}}$$
$$= -\frac{e^{2}}{\hbar} \frac{1}{l\pi^{3}} \sum_{n=0}^{1/\bar{H}} \frac{2\bar{H}}{\sqrt{a^{-2} + 4\bar{H}\left(n + \frac{1}{2}\right)}} \tan^{-1} \frac{1}{\sqrt{a^{-2} + 4\bar{H}\left(n + \frac{1}{2}\right)}}.$$
 (P10.6)

Let  $x = \bar{H}n$ . As  $\bar{H} \to 0$  and  $n \to \infty$ , x becomes a continuous variable. We thus obtain

$$\delta\sigma_0 = \lim_{H \to 0} \delta\sigma_H = -\frac{e^2}{\hbar} \frac{2}{l\pi^3} \int_0^1 \frac{\mathrm{d}x}{\sqrt{a^{-2} + 4x}} \tan^{-1} \frac{1}{\sqrt{a^{-2} + 4x}}$$
$$= -\frac{e^2}{\hbar} \frac{1}{l\pi^3} \frac{1}{2} \int_{a^{-2}}^{4+a^{-2}} \frac{\mathrm{d}y}{\sqrt{y}} \tan^{-1} \frac{1}{\sqrt{y}}.$$
(P10.7)

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By using integration by parts, this integral may be explicitly evaluated, giving

$$\delta\sigma_0 = -\frac{e^2}{\hbar} \frac{1}{l\pi^3} \left[ \sqrt{4+a^{-2}} \tan^{-1} \frac{1}{\sqrt{4+a^{-2}}} - \frac{1}{a} \tan^{-1} a + \frac{1}{2} \ln \frac{5a^2 + 1}{a^2 + 1} \right].$$
(P10.8)

When  $T \to 0$  so that  $a \to \infty$ , we have

$$\delta\sigma_0 = -\frac{e^2}{\hbar} \frac{1}{l\pi^3} \left[ 2\tan^{-1}\frac{1}{2} + \frac{1}{2}\ln 5 \right].$$
 (P10.9)

# 11.1 What is "Mesoscopic"?

In the literature, the term "mesoscopic" generally denotes a sample size regime intermediate between the molecular and the bulk. However, in this chapter we wish to define the term more precisely as the regime where  $L < L_{\rm in}$ . For electronic systems, this more precise definition coincides with the usual definition at low temperatures ( $T \leq 10 \,\mathrm{K}$ ). But for classical waves  $L_{\rm in}$  is generally macroscopic in size, and a mesoscopic sample for classical waves is thus equivalent to a bulk sample.

What is special about a mesoscopic sample is that the wave transport behavior may be observed without being filtered by inelastic scattering. Localization is therefore a mesoscopic phenomenon observable when  $\xi < L < L_{in}$ . However, even when waves are delocalized, or localized but with  $L < L_{in} < \xi$ , the behavior of mesoscopic samples can still differ significantly from that of bulk samples, where inelastic scattering plays a role. In particular, mesoscopic phenomena are related in one way or the other to the preservation of wave characteristics, such as interference and phase memory, even after strong multiple (elastic) scatterings. A striking example in this respect is the speckle pattern formed by scattered light in transmission through a random medium, alluded to in Chap.6 (speckle pattern also exists in mesoscopic electronic systems, but it is much more difficult to observe). The spatial intensity fluctuation in this case is a direct result of wave interference. These fluctuations have a component which remains constant over a range of wave propagation distances, a characteristic which is called "universal conductance fluctuations" (UCF) (Lee 1987).

# 11.1.1 Dissipation and Mesoscopic Conductance

An intriguing question for mesoscopic electronic samples is the definition of electrical conductivity. The traditional definition of  $\sigma_{\rm E}$  is associated with dissipation:  $\sigma_{\rm E} E^2$  is the power dissipated per unit volume. Since in a mesoscopic



Fig. 11.1. A schematic picture of the two-step conduction process across a tunnel junction. In the first step, an electron tunnels elastically through the barrier. In the second step, the electron gives up its excess energy. Dissipation occurs in the second step.

sample the absence of inelastic scattering implies the absence of dissipation, a new definition of electrical conductivity must be devised. Based on the analogy with a tunnel junction, Landauer (1957) proposed that the electrical conduction in mesoscopic samples be viewed as a two-step process. Consider a voltage difference V applied across a tunnel junction shown schematically in Fig. 11.1. An electron tunneling elastically through the barrier will end up on the other side with energy eV above the Fermi level. The subsequent decay to the Fermi level completes the conduction process across the junction. In this picture the electrical conduction consists of two separate steps: tunneling and energy dissipation. If the energy barrier is replaced by a mesoscopic sample, it is clear that one can identify the wave transmission probability across the sample as something that is *proportional* to the electrical conductance. In this definition of  $\sigma_{\rm E}$  the element of dissipation still exists. However, instead of occurring inside the sample, dissipation now occurs in the leads attached to the sample; i.e., it is nonlocal. Landauer's formula of electrical conductance for mesoscopic samples is the result of these considerations.

Detailed mathematical descriptions of a number of mesoscopic phenomena are given below.

# 11.2 Intensity Distribution of the Speckle Pattern

In contrast to classical diffusion, it is noted at the beginning of Chap. 6 that wave diffusion can lead to inhomogeneous intensity distributions, called speckle pattern. A speckle pattern results from the interference of randomly scattered waves in the absence of inelastic scattering. The existence of intensity fluctuations is thus an inherently mesoscopic phenomenon which is easily observable with classical waves. For a given configuration of random scatterers, the intensity at a given observation point is the absolute square of the sum of amplitudes for all the different wave paths emanating from the source:

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$$I(\mathbf{r}) = |A(\mathbf{r})|^2 = \frac{1}{M} \left| \sum_{k=1}^M a_k \exp(i\phi_k) \right|^2,$$
(11.1)

where for definitiveness the total number of path is given by the number M. The factor 1/M is present so that the defined amplitude  $a_k$  remains finite as  $M \to \infty$ . Let us make the plausible assumption that  $a_k$  and  $\phi_k$  are statistically independent, with the mean and second moment of  $a_k$  given by a and  $a^2$ , respectively, and that  $\phi_k$  is distributed uniformly between 0 and  $2\pi$ . Equation (11.1) can be rewritten as

$$I(\mathbf{r}) = r^2 + i^2, \tag{11.2}$$

with

$$r = \frac{1}{\sqrt{M}} \sum_{k=1}^{M} a_k \cos \phi_k, \qquad (11.3a)$$

$$i = \frac{1}{\sqrt{M}} \sum_{k=1}^{M} a_k \sin \phi_k, \qquad (11.3b)$$

It is clear that  $\langle r \rangle_{\rm c} = \langle i \rangle_{\rm c} = 0$ . Moreover,

$$\langle r^2 \rangle_{\rm c} = \frac{1}{M} \sum_{k=1}^{M} \sum_{n=1}^{M} \langle a_k a_n \rangle_{\rm c} \langle \cos \phi_k \cos \phi_n \rangle_{\rm c}$$

$$= \frac{1}{M} \sum_{k=1}^{M} \sum_{n=1}^{M} \langle a_k a_n \rangle_{\rm c} \frac{1}{2} \delta_{k,n}$$

$$= \frac{a^2}{2},$$
(11.4)

and  $\left\langle i^2 \right\rangle_{\rm c} = \left\langle r^2 \right\rangle_{\rm c}.$ 

# 11.2.1 Statistical Independence and Gaussian Distribution

It is also easy to see that  $\langle ri \rangle_c = 0$ , so the variables r and i are uncorrelated. From (11.3a) and (11.3b), it is seen that r and i are sums of independent random variables. The central limit theorem states that the distributions of r and i must each be Gaussian (Feller 1957), so that the joint distribution is given by

$$D(r,i) = \frac{1}{2\pi(a^2/2)} \exp\left(-\frac{r^2 + i^2}{2(a^2/2)}\right)$$
$$= \frac{1}{\pi a^2} \exp\left(-\frac{r^2 + i^2}{a^2}\right).$$
(11.5)

By changing the variable from r, i to I, it is immediately clear that

$$D(I) = \frac{1}{\langle I \rangle_c} \exp\left(-\frac{I}{\langle I \rangle_c}\right), \qquad (11.6)$$

where  $I = r^2 + i^2$ ,  $\langle I \rangle_c = a^2$ , and  $drdi = 2\pi\sqrt{I}d\sqrt{I} = \pi dI$ . Equation (11.6) is the main result of this section. By assuming that the intensity distribution of the speckle pattern for a given configuration is the same as that for a given observation point varied over different configurations, (11.6) states that the intensity has an *exponential distribution*. A special feature of the exponential distribution is that its mean,  $\langle I \rangle_c$ , is equal to its standard deviation  $[\langle I^2 \rangle_c - \langle I \rangle_c^2]^{1/2}$ , a fact which will be used later. The intensity fluctuations are therefore not small. Below we calculate the frequency and spatial correlations in a speckle pattern.

# 11.3 Correlations in the Diffusive Intensity

Let us first calculate the average intensity at **r** resulting from a point source at **r'**. This is given by  $\langle |G^+(\omega, \mathbf{r}, \mathbf{r'})|^2 \rangle_c$ . From Chap. 5, Equation (5.8), this is expressible as

$$\langle I(\omega, |\mathbf{r} - \mathbf{r}'|) \rangle_{c} = S(\Delta \omega = 0, \mathbf{r} - \mathbf{r}'|\omega)$$
  
=  $\frac{L^{d}}{N} \int \frac{d\Delta \mathbf{k}}{(2\pi)^{d}} S(\Delta \omega = 0, \Delta \mathbf{k}|\omega) \exp[i\Delta \mathbf{k} \cdot (\mathbf{r} - \mathbf{r}')]^{(11.7)}$ 

Since from (5.54) it is known that

$$S(\Delta\omega = 0, \Delta \mathbf{k}|\omega) = \frac{2(\mathrm{d}\omega/\mathrm{d}\kappa_0^2)(\mathrm{d}\omega/\mathrm{d}\kappa_\mathrm{e}^2)\rho_\mathrm{e}(\omega)\pi}{D(\omega)|\Delta \mathbf{k}|^2}\frac{N}{L^d},$$
(11.8)

it follows that

$$\langle I(\omega, |\mathbf{r} - \mathbf{r}'|) \rangle_{c} = \frac{2\pi\rho_{e}(\omega)(d\omega/d\kappa_{0}^{2})(d\omega/d\kappa_{e}^{2})}{(v_{t}l/3)(2\pi)^{d}} \\ \times \int \frac{d\Delta \mathbf{k}}{|\Delta \mathbf{k}|^{2}} \exp[i\Delta \mathbf{k} \cdot (\mathbf{r} - \mathbf{r}')] \\ = \frac{3(d\omega/d\kappa_{0}^{2})(d\omega/d\kappa_{e}^{2})\rho_{e}(\omega)}{2v_{t}l|\mathbf{r} - \mathbf{r}'|}$$
(11.9)

for d = 3. The integral is noted to diverge for  $d \leq 2$ , which is an indication that the diffusive wave intensity in 1D or 2D must depend on the boundary conditions, however large the sample. The dependence of the intensity on  $|\mathbf{r} - \mathbf{r}'|^{-1}$  is also noted to differ from the wave behavior (which should give a  $|\mathbf{r} - \mathbf{r}'|^{-2}$  dependence) but to be characteristic of diffusion, as expected. For classical scalar waves with  $d\omega/d\kappa_0^2 = v_0^2/2\omega$ ,  $d\omega/d\kappa_e^2 = v_e^2/2\omega$ , and  $\rho_e(\omega) = 4\pi\omega^2/v_e^3$ , we get

$$\langle I(\omega, |\mathbf{r} - \mathbf{r}'|) \rangle_{c} = \frac{3\pi v_{0}^{2}}{2v_{t}v_{e}l|\mathbf{r} - \mathbf{r}'|}$$
$$\cong \frac{3\pi \bar{\varepsilon}}{2l|\mathbf{r} - \mathbf{r}'|}$$
(11.10)

in the low-frequency limit, where  $v_{\rm e} \cong v_t$ , and  $\bar{\varepsilon} = v_0^2/v_{\rm e}^2$ .

# 11.3.1 Frequency Correlation of the Diffusive Intensity

From the calculation of the average diffusive intensity at a point, it is only a small step to calculate the frequency correlation of the intensity at a given observation point. Let us define the frequency correlation function as

$$C(\Delta\omega, |\mathbf{r} - \mathbf{r}'|) = \langle I(\omega, |\mathbf{r} - \mathbf{r}'|)I(\omega + \Delta\omega, |\mathbf{r} - \mathbf{r}'|)\rangle_{c} - \langle I(\omega, |\mathbf{r} - \mathbf{r}'|)\rangle_{c} \langle I(\omega + \Delta\omega, |\mathbf{r} - \mathbf{r}'|)\rangle_{c}. \quad (11.11a)$$

The first term on the right-hand side can be written alternatively as

$$\langle G^+(\omega,\mathbf{r},\mathbf{r}')G^-(\omega,\mathbf{r}',\mathbf{r})G^+(\omega+\Delta\omega,\mathbf{r},\mathbf{r}')G^-(\omega+\Delta\omega,\mathbf{r}',\mathbf{r})\rangle_{\alpha}$$

We look for pairings involving  $G^+$  and  $G^-$ , which would give the largest contribution. The configurational averaging can be done in subsets of  $\langle G^+G^-\rangle_c$ in two ways. One is to pair the first two and the last two. This pairing leads to a term which is exactly canceled by the second term on the right-hand side of (11.11a). The other remaining pairing choice is given by

$$\left\langle G^{+}(\omega, \mathbf{r}, \mathbf{r}')G^{-}(\omega + \Delta\omega, \mathbf{r}', \mathbf{r}) \right\rangle_{c} \times \left\langle G^{+}(\omega + \Delta\omega, \mathbf{r}, \mathbf{r}')G^{-}(\omega, \mathbf{r}', \mathbf{r}) \right\rangle_{c} = |S(\Delta\omega, \mathbf{r} - \mathbf{r}'|\omega)|^{2}.$$
(11.11b)

It follows that

$$C(\Delta\omega, |\mathbf{r} - \mathbf{r}'|) \cong |S(\Delta\omega, \mathbf{r} - \mathbf{r}'|\omega)|^2.$$
(11.12)

Since in the  $\Delta \omega, \Delta \mathbf{k}$  representation S has a diffusive pole, it is seen that

$$S(\Delta\omega, \mathbf{r} - \mathbf{r}'|\omega) = \frac{2\pi\rho_{\rm e}(\omega)(\mathrm{d}\omega/\mathrm{d}\kappa_0^2)(\mathrm{d}\omega/\mathrm{d}\kappa_{\rm e}^2)}{(2\pi^d)} \int \frac{\mathrm{d}(\Delta\mathbf{k})\exp[\mathrm{i}\Delta\mathbf{k}\cdot(\mathbf{r} - \mathbf{r}')]}{-\mathrm{i}\Delta\omega + v_t l|\Delta\mathbf{k}|^2/3}$$

We will consider only the case of d = 3, where there is a finite limit at  $\Delta \omega \to 0$ . The integral is done in the solution to Problem 11.1. The result is

$$S(\Delta\omega, \mathbf{r} - \mathbf{r}'|\omega) = \frac{3\rho_{\rm e}(\omega)(\mathrm{d}\omega/\mathrm{d}\kappa_0^2)(\mathrm{d}\omega/\mathrm{d}\kappa_{\rm e}^2)}{2v_t l|\mathbf{r} - \mathbf{r}'|} \exp\left[(\mathrm{i} - 1)\sqrt{\frac{3\Delta\omega}{2v_t l}}|\mathbf{r} - \mathbf{r}'|\right].$$
(11.13)

From (11.9) and (11.12), it follows that

$$C(\Delta\omega, |\mathbf{r} - \mathbf{r}'|) \cong \langle I(\omega, |\mathbf{r} - \mathbf{r}'|) \rangle_{c}^{2} \exp\left[-\sqrt{\frac{6\Delta\omega}{v_{t}l}}|\mathbf{r} - \mathbf{r}'|\right].$$
(11.14)

At  $\Delta \omega = 0$ ,  $C(\Delta \omega = 0, |\mathbf{r} - \mathbf{r}'|)$  is exactly the variance of the diffusive intensity fluctuations in the speckle pattern. This is consistent with the result of our earlier calculation, which predicts the standard deviation to be the same as the mean. Equation (11.14) states that at a given observation point, a variation in the frequency of the wave will cause a corresponding variation in the intensity, with a correlation which decays exponentially as the square root of the frequency variation (Shapiro 1986).

#### 11.3.2 Spatial Correlation in the Diffusive Intensity

Suppose the frequency is held fixed but the observation point is varied; what would be the *spatial* correlation? That is, what is the spatial correlation in the speckle pattern? To calculate that quantity, let us define

 $\mathbf{r}_1 - \mathbf{r}' = \mathbf{R}$ 

to be the separation vector between the source and the observation point  $\mathbf{r}_1$ , and  $\Delta \mathbf{R} = \mathbf{r}_2 - \mathbf{r}_1$  to be the vector defining the separations between the two observation points  $\mathbf{r}_2$  and  $\mathbf{r}_1$ . It is assumed that  $R \gg \Delta R$ . The spatial correlation  $C^{(1)}(\omega, \mathbf{R}, \Delta \mathbf{R})$  is defined by

$$\langle I(\omega, \mathbf{R}) I(\omega, \mathbf{R} + \Delta \mathbf{R}) \rangle_{c}$$

$$= \langle G^{+}(\omega, \mathbf{R}) G^{-}(\omega, \mathbf{R}) G^{+}(\omega, \mathbf{R} + \Delta \mathbf{R}) G^{-}(\omega, \mathbf{R} + \Delta \mathbf{R}) \rangle_{c}$$

$$= \langle I(\omega, \mathbf{R}) \rangle_{c} \langle I(\omega, \mathbf{R} + \Delta \mathbf{R}) \rangle_{c} + C^{(1)}(\omega, \mathbf{R}, \Delta \mathbf{R}).$$
(11.15)

Here the superscript (1) is to distinguish this *short-range* spatial correlation from the anticipated long-range spatial correlation to be introduced later, denoted  $C^{(2)}$ . To calculate  $C^{(1)}(\omega, \mathbf{R}, \Delta \mathbf{R})$ , defined as the difference between the left hand side and the first term on the right hand side of (11.15), it is noted that in the second line of (11.15), besides the  $\langle G^+G^-\rangle_c$  pairing that gives the  $\langle I(\omega, \mathbf{R})\rangle_c \langle I(\omega, \mathbf{R} + \Delta \mathbf{R})\rangle_c$  term, there can also be the pairing choice

$$\langle G^+(\omega, \mathbf{R})G^-(\omega, \mathbf{R} + \Delta \mathbf{R}) \rangle_{\rm c} \langle G^+(\omega, \mathbf{R} + \Delta \mathbf{R})G^-(\omega, \mathbf{R}) \rangle_{\rm c}$$
.

That means to a good approximation,

$$C^{(1)}(\omega, \mathbf{R}, \Delta \mathbf{R}) \cong \left\langle G^{+}(\omega, \mathbf{R}) G^{-}(\omega, \mathbf{R} + \Delta \mathbf{R}) \right\rangle_{c} \times \left\langle G^{+}(\omega, \mathbf{R} + \Delta \mathbf{R}) G^{-}(\omega, \mathbf{R}) \right\rangle_{c}.$$
 (11.16)

Let us recall from Chap. 5 that [see (5.16)]

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$$\langle \mathbf{G}^+ \otimes \mathbf{G}^- \rangle_{\mathrm{c}} = \mathbf{G}_{\mathrm{e}}^+ \otimes \mathbf{G}_{\mathrm{e}}^- + (\mathbf{G}_{\mathrm{e}}^+ \otimes \mathbf{G}_{\mathrm{e}}^-) : \mathbf{\Gamma} : (\mathbf{G}_{\mathrm{e}}^+ \otimes \mathbf{G}_{\mathrm{e}}^-)$$
(11.17)

In the present case, the output channels in real space are noted to be different from before. That is, there are now two different observation points. In (11.17), the first term on the right-hand side is negligible, because both  $\mathbf{G}_{e}^{+}$  and  $\mathbf{G}_{e}^{-}$ represent the coherent component of the wave, which decays exponentially with a short decay length l. The second term, however, can be considerably larger in magnitude. To see how this can come about, we recall that there are two channels (+ and – channels) of propagation. The presence of the  $\Gamma$ operator means that the two channels are coupled through the scatterers. It was shown in Chap. 5 that the ladder-diagram couplings result in diffusive transport, since in this case we have *intensity* transport which is insensitive to phase coherence. Otherwise, if the two channels propagate independently then both would decay exponentially just as the first term of (11.17), owing to the sensitivity to phase coherence. Thus for the second term of (11.17) we would like to have diffusive transport from the source point  $\mathbf{r'}$  to one of the two measurement points  $\mathbf{r}_1$ ,  $\mathbf{r}_2$  (with only algebraic attenuation), and then coherent transport in the last step (between the two observation points) to account for phase correlation. In mathematical terms, that means the second term of (11.17) can be separated into two parts. The first part,  $(\mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-})$ :  $\Gamma$ , is diffusive in character. The second part is just the coherent transport represented by  $\mathbf{G}_e^+ \otimes \mathbf{G}_e^-$ . From (5.18),

$$\Gamma = \mathbf{U} + \mathbf{U} : (\mathbf{G}_{\mathbf{e}}^+ \otimes \mathbf{G}_{\mathbf{e}}^-) : \Gamma,$$

and in the Boltzmann limit (ladder diagrams with isotropic scattering approximation) the inner product of  $(\mathbf{G}_{\rm e}^+ \otimes \mathbf{G}_{\rm e}^-)$  with  $\mathbf{\Gamma}^{({\rm B})}$  is represented by the operations of (1) setting the observation points of  $\mathbf{G}_{\rm e}^+$  and  $\mathbf{G}_{\rm e}^-$  the same (same scatterer) and (2) summing over all scatterer positions [see (5.69)], it follows that in the k-representation ( $U^{({\rm B})}$  is treated to be k independent), we have

$$\Gamma^{(B)} = \frac{U^{(B)}}{1 - (U^{(B)}/N)\Sigma_{\mathbf{k}}(G_{e}^{+}G_{e}^{-})_{\mathbf{k}}},$$
(11.18)

where the summation over  $\mathbf{k}$  follows from the summation over all scatterers (since in the effective medium picture scatterers fill all space, this summation is equivalent to summation over all space). Comparison with (5.30) shows that

$$(\mathbf{G}_{\mathrm{e}}^{+} \otimes \mathbf{G}_{\mathrm{e}}^{-}) : \mathbf{\Gamma} \to \frac{(1/N)\Sigma_{\mathbf{k}}(G_{\mathrm{e}}^{+}G_{\mathrm{e}}^{-})_{\mathbf{k}}}{1 - (U^{(\mathrm{B})}/N)\Sigma_{\mathbf{k}}(G_{\mathrm{e}}^{+}G_{\mathrm{e}}^{-})_{\mathbf{k}}} U^{(B)},$$
$$= S^{(\mathrm{B})}U^{(\mathrm{B})}.$$
(11.19)

Since  $S^{(B)}$  is diffusive in character, the incoherent component of the wave thus decays algebraically from the source point  $\mathbf{r}'$  to the observation point  $\mathbf{r}_1 = \mathbf{r}' + \mathbf{R}$ .

The second part of the term,  $(\mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-})$ , has the input point at  $\mathbf{r}_{1}$  instead of  $\mathbf{r}'$ , and the output points at both  $\mathbf{r}_{1}$  and  $\mathbf{r}_{2}$ , separated by  $\Delta \mathbf{R}$ . Since the inner product with  $\Gamma$  means summation over all input channels  $\mathbf{r}_{1}$ ,  $(\mathbf{G}_{e}^{+} \otimes \mathbf{G}_{e}^{-})$  can be written approximately as

$$\sum_{\mathbf{r}_1} G_{\mathbf{e}}^+(\omega, \ \mathbf{r}_1 - \mathbf{r}_1) G_{\mathbf{e}}^-(\omega, \ \mathbf{r}_2 - \mathbf{r}_1)$$
$$= \frac{L^d}{N(2\pi)^d} \int d\mathbf{k} \ (G_{\mathbf{e}}^+ G_{\mathbf{e}}^-)_{\mathbf{k}} \exp(i\mathbf{k} \cdot \Delta \mathbf{R}).$$

By noting that

$$(G_{\mathrm{e}}^{+}G_{\mathrm{e}}^{-})_{\mathbf{k}} \cong (G_{\mathrm{e}}^{+})_{\mathbf{k}}(G_{\mathrm{e}}^{-})_{\mathbf{k}} \cong \frac{(\Delta G_{e})_{\mathbf{k}}}{2\mathrm{i}\mathrm{Im}\Sigma^{+}}\frac{N}{L^{d}}$$

it follows from the approximation that  $\Sigma^+$  being **k**-independent, that

$$\frac{L^{d}}{N(2\pi)^{d}} \int d\mathbf{k} (G_{e}^{+}G_{e}^{-})_{\mathbf{k}} \exp(i\mathbf{k}\cdot\Delta\mathbf{R})$$

$$\approx \frac{1}{2i\mathrm{Im}\Sigma^{+}} \int \frac{d\mathbf{k}}{(2\pi)^{d}} (\Delta G_{e})_{\mathbf{k}} \exp(i\mathbf{k}\cdot\Delta\mathbf{R})$$

$$= \frac{\Delta G_{e}(\Delta\mathbf{R})}{2i\mathrm{Im}\Sigma^{+}} \frac{N}{L^{d}}.$$
(11.20)

Combining (11.20) with (11.19) results in

$$\langle G^+(\omega, \mathbf{R})G^-(\omega, \mathbf{R} + \Delta \mathbf{R}) \rangle_{c} \cong S^{(B)}(\Delta \omega = 0, \mathbf{R}|\omega) \frac{\Delta G_{e}(\Delta \mathbf{R})}{\Delta G_{e}(\Delta \mathbf{R} = 0)},$$
(11.21)

where we have used the identity  $U^{(B)} = [2iIm\Sigma^+/\Delta G_e(\Delta R = 0)]L^d/N$ , and the coherent component [first term on the right hand side of (11.17)] has been neglected.

# 11.3.3 Intensity Fluctuations for Classical Waves in 3D

For the case of classical waves in 3D, one can go a step further and write

$$G_{\rm e}^{\pm}(\Delta \mathbf{R}) = -\frac{\exp(\pm \mathrm{i}\kappa\,\Delta R)\exp(-\Delta R/2l)}{4\pi\,\Delta R}$$

from which it follows that

$$\Delta G_{\rm e}(\Delta \mathbf{R}) = -\frac{\mathrm{i}}{2\pi \,\Delta R} \sin(\kappa \,\Delta R) \exp(-\Delta R/2l). \tag{11.22}$$

Since  $\text{Im}\Sigma^+ = -\kappa/l$  [(9.41b)] for classical waves, one gets

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$$\sum_{\mathbf{r}_{1}} \left\langle G^{+}(\omega, \mathbf{r}_{1}, \mathbf{r}_{1}) G^{-}(\omega, \mathbf{r}_{1}, \mathbf{r}_{2}) \right\rangle_{c}$$
$$\cong \frac{l}{4\pi\kappa \,\Delta R} \sin(\kappa \,\Delta R) \exp(-\Delta R/2l) \frac{N}{L^{d}}.$$
(11.23)

Therefore,

$$\left\langle G^{+}(\omega, \mathbf{R})G^{-}(\omega, \mathbf{R} + \Delta \mathbf{R})\right\rangle_{c} \cong S^{(B)}(\Delta \omega = 0, \mathbf{R}|\omega) \ U^{(B)} \ \frac{l\sin(\kappa \,\Delta R)}{4\pi\kappa \,\Delta R} \exp(-\Delta R/2l)\frac{N}{L^{d}}.$$
 (11.24)

By using the Ward identity,

$$U^{(B)}\frac{N}{L^{d}} = \frac{2iIm\Sigma^{+}}{\Delta G_{e}(\Delta R = 0)} = \frac{2iIm\Sigma^{+}}{-i\kappa/2\pi} = \frac{4\pi}{l},$$
 (11.25)

(11.24) becomes, for classical scalar waves,

$$\left\langle G^{+}(\omega, \mathbf{R})G^{-}(\omega, \mathbf{R} + \Delta \mathbf{R})\right\rangle_{c} \\ \cong \left\langle I(\omega, \mathbf{R})\right\rangle_{c} \frac{\sin(\kappa \,\Delta R)}{\kappa \,\Delta R} \exp(-\Delta R/2l).$$
(11.26)

Here we have identified  $S^{(B)}(\Delta \omega = 0, \mathbf{R}|\omega)$  as  $\langle I(\omega, \mathbf{R}) \rangle_c$ . Since the other term,  $\langle G^+(\omega, \mathbf{R} + \Delta \mathbf{R}) G^-(\omega, \mathbf{R}) \rangle_c$ , can be obtained from (11.26) by exchanging  $\mathbf{r}_1$  and  $\mathbf{r}_2$ , it follows from (11.16) and (11.21) that

$$C^{(1)}(\omega, \mathbf{R}, \Delta \mathbf{R}) \cong \langle I(\omega, \mathbf{R}) \rangle_{c} \langle I(\omega, \mathbf{R} + \Delta \mathbf{R}) \rangle_{c} \left( \frac{\Delta G_{e}(\Delta \mathbf{R})}{\Delta G_{e}(\Delta R = 0)} \right)^{2}$$
(11.27a)

in the general case and

$$C^{(1)}(\omega, \mathbf{R}, \Delta \mathbf{R}) \cong \langle I(\omega, |\mathbf{R}|) \rangle_{c} \langle I(\omega, |\mathbf{R} + \Delta \mathbf{R}|) \rangle_{c} \\ \times \frac{\sin^{2}(\kappa \Delta R)}{(\kappa \Delta R)^{2}} \exp\left(-\frac{\Delta R}{l}\right)$$
(11.27b)

in the particular case of 3D classical waves. Again, when  $\Delta R = 0$ ,  $C^{(1)}(\omega, \mathbf{R}, \Delta R = 0)$  is noted to give the variance of the intensity fluctuations as expected.

Equation (11.27b) tells us that the spots in the speckle pattern are correlated over a distance l (Shapiro 1986), and such correlation exists no matter how far the wave travels. The result can be viewed physically as the consequence of diffusion plus interference. Since the effect of interference is to produce rapid spatial fluctuations in intensity, its effect on diffusion may be approximated by a random spatial source term in the diffusive process. Such a viewpoint was first put forth by Spivak and Zyuzin (1988). Let us follow this approach and examine the *long-range* implications of the short-range interference effect.

# 11.4 Long-Range Correlation in Intensity Fluctuations

Diffusion is inherently a smoothing process, driven by gradient in the spatial distribution. Due to the wave interference effect in the absence of inelastic scattering, wave diffusion is accompanied by intensity fluctuations  $\Delta I(\mathbf{r})$ , as seen above. Such intensity fluctuations naturally serve as secondary diffusive sources, due to the intensity gradient that accompanies such inhomogeneities. We show that such consideration can give rise to the prediction of long-range intensity fluctuations.

Consider the division of  $I(\mathbf{r})$  into

$$I(\mathbf{r}) = \langle I(\mathbf{r}) \rangle_{c} + \delta I(\mathbf{r}) + \Delta I(\mathbf{r}), \qquad (11.28)$$

where the frequency dependence is suppressed. The separation of the intensity fluctuations into  $\delta I(\mathbf{r})$  and  $\Delta I(\mathbf{r})$  is based on the difference in spatial scales: Whereas  $\delta I(\mathbf{r})$  denotes the diffusive relaxation of the intensity over distances  $\gg l$ ,  $\Delta I(\mathbf{r})$  represents the short-range fluctuations manifest as the speckle pattern. One can also view  $\delta I(\mathbf{r})$  to be that part of the diffusive flux with  $\Delta I(\mathbf{r})$  as the secondary source. The approach of Spivak and Zyuzin is to treat  $\Delta I(\mathbf{r})$  as random, with  $\langle \Delta I(\mathbf{r}_1) \Delta I(\mathbf{r}_2) \rangle$  given by the correlation function (11.27). Below we treat the long-range correlation only in the classical scalar wave case.

# 11.4.1 Reduction to a Diffusion Problem with Random Sources

In the calculation of long-range correlation, a reasonable approximation is to treat the short-range correlation as deita-function-like, i.e., with  $|\mathbf{r}_2 - \mathbf{r}_1| = \Delta R \rightarrow 0$ ,

$$\begin{split} \left\langle \Delta I(\mathbf{r}_1) \Delta I(\mathbf{r}_2) \right\rangle_{\rm c} &\cong \delta(\mathbf{r}_2 - \mathbf{r}_1) \left\langle I(\mathbf{r}_1) \right\rangle_{\rm c}^2 4\pi \\ &\times \int_0^\infty (\Delta R)^2 {\rm d}(\Delta R) \frac{\sin^2(\kappa \,\Delta R)}{(\kappa \,\Delta R)^2} \exp\left(-\frac{\Delta R}{l}\right) \\ &\cong \delta(\mathbf{r}_2 - \mathbf{r}_1) \left\langle I(\mathbf{r}_1) \right\rangle_{\rm c}^2 \frac{2\pi l}{\kappa^2}. \end{split}$$
(11.29)

In terms of wave flux, we have

$$\mathbf{J}(\mathbf{r}) = -\frac{v_t l}{3} \nabla \left[ \langle I(\mathbf{r}) \rangle_{\mathbf{c}} + \delta I(\mathbf{r}) \right] + v_t \Delta I(\mathbf{r}), \qquad (11.30)$$

where the first term on the right-hand side is the diffusive flux, whereas the second term is the random flux source

$$\mathbf{j}_s(\mathbf{r}) = v_t \Delta I(\mathbf{r}),\tag{11.31}$$

with

$$\left\langle j_s^{(i)}(\mathbf{r}_1) j_s^{(k)}(\mathbf{r}_2) \right\rangle_{\rm c} = \frac{1}{3} \delta_{ik} \frac{2\pi v_t^2 l}{\kappa^2} \left\langle I(\mathbf{r}_1) \right\rangle_{\rm c}^2 \delta(\mathbf{r}_2 - \mathbf{r}_1).$$
(11.32)

Here i, k denote the vectorial components of  $\mathbf{j}_s$ . In steady state, the continuity condition requires

$$\nabla \cdot \mathbf{J}(\mathbf{r}) = 0. \tag{11.33}$$

Since  $\nabla^2 \left< I({\bf r}) \right>_{\rm c} = 0$  in the steady state, it follows from (11.30) that

$$\frac{v_t l}{3} \nabla^2 \delta I(\mathbf{r}) = \nabla \cdot \mathbf{j}_s(\mathbf{r}). \tag{11.34}$$

#### 11.4.2 Solution in the Slab Geometry

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To calculate the correlation of  $\delta I(\mathbf{r})$ , let us specialize to the slab geometry where z is the sample thickness direction, L is the sample thickness, and the desired quantity is the correlation

$$C^{(2)}(\boldsymbol{\rho}_1, \boldsymbol{\rho}_2) = \left\langle \delta I(\boldsymbol{\rho}_1, L-l) \delta I(\boldsymbol{\rho}_2, L-l) \right\rangle_{\rm c}, \qquad (11.35)$$

where  $\rho_1$ ,  $\rho_2$  are the coordinates on the (x, y) plane. The z position is set at L - l because we want to impose the boundary condition at z = L. Here we want  $\Delta \rho = |\rho_1 - \rho_2| \gg l$  so as to distinguish  $C^{(2)}$  from  $C^{(1)}$ . Through Fourier transform in the (x, y) plane, we obtain the equation for  $\delta I(\mathbf{k}, z)$ , where  $\mathbf{k}$  is the 2D wave vector, as

$$\frac{\mathrm{d}}{\mathrm{d}z^2} \delta I(\mathbf{k}, z) - k^2 \delta I(\mathbf{k}, z) 
= \frac{3}{v_t l} \left[ \mathrm{i} \left( k_x j_s^{(x)}(\mathbf{k}, z) + k_y j_s^{(y)}(\mathbf{k}, z) \right) + \frac{\mathrm{d}}{\mathrm{d}z} j_s^{(z)}(\mathbf{k}, z) \right].$$
(11.36)

The boundary conditions are  $\delta I(\mathbf{k}, z) = 0$  at z = 0, L. That is,  $I(z = 0) = \langle I \rangle_c$ , and at z = L the condition I = 0 is imposed. One can be more general by specifying an extrapolation length (see Sect. 5.8); however, this would not alter the qualitative conclusions to be reached. In the present case we just have to examine the long-range correlation at z = L - l [see (11.35)]. The Green function for (11.36) is given by

$$G(\mathbf{k}, z, z') = -\frac{1}{k} \frac{\sinh(kz_{<})\sinh[k(L-z_{>})]}{\sinh(kL)},$$
(11.37)

where  $z_{<} = \min[z, z']$ , and  $z_{>} = \max[z, z']$ . The fact that  $G(\mathbf{k}, z, z')$  indeed satisfies

$$\frac{\mathrm{d}}{\mathrm{d}z^2}G(\mathbf{k}, z, z') - k^2 G(\mathbf{k}, z, z') = \delta(z, z')$$

is shown in the solution to Problem 11.2. With the knowledge of the Green function,  $\delta I(\mathbf{k}, z)$  is given by

$$\delta I(\mathbf{k}, z) = \frac{3}{v_t l} \left\{ \int_0^L dz' G(\mathbf{k}, z, z') \, \mathbf{i} \left[ k_x j_s^{(x)}(\mathbf{k}, z') + k_y j_s^{(y)}(\mathbf{k}, z') \right] - \int_0^L dz' \left[ \frac{d}{dz'} G(\mathbf{k}, z, z') \right] j_s^{(z)}(\mathbf{k}, z') \right\},$$
(11.38)

where the second term is obtained through integration by parts and the fact that  $G(\mathbf{k}, z, z')$  vanishes at  $z_{\leq} = 0$  and  $z_{\geq} = L$ . Now the quantity  $\langle \delta I(\mathbf{k}_1, z_1) \delta I^*(\mathbf{k}_2, z_2) \rangle_c$  may be written as

$$\left\langle \delta I(\mathbf{k}_{1}, z_{1}) \delta I^{*}(\mathbf{k}_{2}, z_{2}) \right\rangle_{c} = \left( \frac{3}{v_{t}l} \right)^{2} \int_{0}^{L} \int_{0}^{L} dz'_{1} dz'_{2} \left\langle j_{s}^{(i)}(\mathbf{k}_{1}, z'_{1}) j_{s}^{(i)*}(\mathbf{k}_{2}, z'_{2}) \right\rangle_{c} \\ \times \left[ (\mathbf{k}_{1} \cdot \mathbf{k}_{2}) G(\mathbf{k}_{1}, z_{1}, z'_{1}) G(\mathbf{k}_{2}, z_{2}, z'_{2}) + \frac{\mathrm{d}G(\mathbf{k}_{1}, z_{1}, z'_{1})}{\mathrm{d}z'_{1}} \frac{\mathrm{d}G(\mathbf{k}_{2}, z_{2}, z'_{2})}{\mathrm{d}z'_{2}} \right],$$

$$(11.39)$$

where the averages of the cross terms  $\left\langle j_s^{(i)} j_s^{(j)*} \right\rangle_c$ ,  $i \neq j$ , are noted to vanish. From (11.32) one can easily deduce

$$\left\langle j_{s}^{(i)}(\mathbf{k}_{1}, z_{1}') j_{s}^{(i)*}(\mathbf{k}_{2}, z_{2}') \right\rangle_{c} = \frac{N^{2}}{W^{4}} \frac{(2\pi)^{3} v_{t}^{2} l}{3\kappa^{2}} \delta(z_{1}' - z_{2}') \delta(\mathbf{k}_{1} - \mathbf{k}_{2}) \left\langle I(z_{1}') \right\rangle_{c}^{2}, \qquad (11.40)$$

where W, the lateral dimension of the sample, arises from the 2D Fourier transform of the correlation relation, (11.29). Here it is also assumed that the incident wave is a plane wave, so that  $\langle I(\mathbf{r}_1) \rangle_c$  depends only on the z coordinate. In that case the solution to the 1D diffusion equation,

$$\nabla^2 \left\langle I(z) \right\rangle_{\rm c} = 0,$$

with the boundary conditions that  $\langle I(z) \rangle_c = 1$  at z = 0 and 0 at z = L, gives the solution

$$\langle I(z)\rangle_{\rm c} = 1 - \frac{z}{L}.\tag{11.41}$$

By substituting (11.40) and (11.41) into (11.39), one gets zero for  $z_1 \neq z_2$ , and for  $z_1 = z_2$  the inplane correlation in the **k** representation is given by (with  $z' = z_{\leq}$ )

$$\langle \delta I(\mathbf{k}_{1}, z_{1}) \delta I^{*}(\mathbf{k}_{2}, z_{1}) \rangle_{c} = \frac{N^{2}}{W^{4}} \frac{3(2\pi)^{3}}{\kappa^{2}l} \delta(\mathbf{k}_{1} - \mathbf{k}_{2}) \times \frac{\sinh^{2} k_{1}(L - z_{1})}{\sinh^{2} k_{1}L} \int_{0}^{L} \cosh 2k_{1}z' \left(1 - \frac{z'}{L}\right)^{2} dz' = \frac{N^{2}}{W^{4}} \frac{3(2\pi)^{3}}{\kappa^{2}l} \delta(\mathbf{k}_{1} - \mathbf{k}_{2}) \frac{\sinh^{2} k_{1}(L - z_{1})}{\sinh^{2} k_{1}L} \times \frac{\sinh 2k_{1}L - 2k_{1}L}{4k_{1}^{3}L^{2}}.$$
(11.42)

By choosing  $z_1 = L - l$  instead of  $z_1 = L$  (at which I = 0 by boundary condition), one gets the long-range correlation in real space as

$$C^{(2)}(\rho_1, \rho_2) = C^{(2)}(\Delta \rho)$$

$$= \frac{3}{2\pi\kappa^2 l} \int_0^\infty \mathrm{d}k \, k \int_0^{2\pi} \mathrm{d}\theta \exp(\mathrm{i}k\Delta\rho\cos\theta) \frac{\sinh^2 kl}{\sinh^2 kL}$$

$$\times \frac{\sinh 2kL - 2kL}{4k^3L^2}$$

$$\cong \frac{3l^3}{4(\kappa^2 l^2)L^3} \int_0^\infty J_0\left(x\frac{\Delta\rho}{L}\right) \frac{\sinh 2x - 2x}{\sinh^2 x} \mathrm{d}x, \qquad (11.43)$$

where we have made the approximation  $\sinh(kl) \cong kl$ , which is consistent with the assumption expressed by (11.29). It is seen that the correlation  $C^{(2)}$ is measured in terms of sample thickness L, which shows its long-range nature (Stephen and Cwilich 1987)). The function

$$F(\Delta \rho/L) = \int_0^\infty J_0\left(x\frac{\Delta \rho}{L}\right) \frac{\sinh 2x - 2x}{\sinh^2 x} \mathrm{d}x$$

is plotted in Fig. 11.2.

# 11.4.3 Slow Decay of Transmitted Intensity Fluctuations

There can also be fluctuations in the total transmitted intensity. This may be obtained from (11.42) by integrating over  $\mathbf{k}_2$  and then setting  $k_1 \to 0$ . Since **k** is a 2D wavevector, setting  $k_1 = 0$  means taking the cross sectional average. The result is  $\left\langle \delta I^2(L, \ k_1=0) \right\rangle_{\rm c} = \frac{1}{W^2} \frac{2\pi}{\kappa^2} \frac{l}{L}.$ 

(11.44)



Fig. 11.2. Variation of the long-range correlation in the intensity fluctuations as a function of  $\Delta \rho / L$ .

Here a factor of N has been divided out by normalizing the result by  $\langle \mathbf{k}_1 | \mathbf{k}_1 \rangle$ . Equation (11.44) is interesting because it tells us that the decay of the longrange fluctuations is much slower as a function of L than that predicted by the simple-minded picture of  $\langle \delta I^2 \rangle_c = \langle I \rangle_c^2$ , which follows a much faster  $1/L^2$ decay. The reason for the larger-than-expected fluctuations is easy to understand: The effect of interference causes random intensity variation everywhere in space, which counters the smoothing effect of the diffusion process. Therefore, intensity fluctuations contain both a fast varying component, given by  $C^{(1)}$ , and a slow varying component, given by  $C^{(2)}$ , arising from the intensity inhomogeneities acting as the secondary diffusive sources.

The calculations so far follow the pattern of diffusion-interference giving rise to the speckle pattern, and diffusion-interference-diffusion giving rise to the long-range fluctuation correlation. This process can be continued by calculating the diffusion on a background of long-range correlated fluctuations. Due to the interaction that interference introduces into the propagation mechanism, one can expect the existence of even slower-decay fluctuations that dominate at longer distances. In fact, as we shall show later through numerical simulation results, there can be a range of propagation distances in which the intensity fluctuation is roughly a constant, i.e., independent of L. This phenomenon is called the UCF. A heuristic explanation of this unusual mesoscopic behavior is given at the end of this chapter. However, the universal conductance fluctuation phenomenon is not expected to extend to infinite propagation distance because the effect of coherent backscattering, i.e., the maximally crossed diagrams, has been neglected so far. In order to take all these effects into consideration, it is necessary to have a formula for the accurate evaluation of the wave diffusion constant, or the conductance, of a mesoscopic sample. This is provided by Landauer's formula, derived next in the context of electronic systems at zero temperature.

# 11.5 Landauer Formula and Quantized Conductance

To derive the Landauer formula for the calculation of mesoscopic conductance, let us refer to Fig. 11.1, where a tunneling barrier (the mesoscopic sample) is shown together with the two "perfect" leads connected to its right and left. The leads are in turn connected to the electron reservoirs with chemical potentials  $\mu_1$  and  $\mu_2$ . A voltage is applied across the sample, so  $\mu_1 > \mu_2$ . The electron current in this case is given by

$$j = \frac{ev_{\rm e}}{2\hbar} \rho_{\rm e}(\omega)(\mu_1 - \mu_2) |\tau|^2, \qquad (11.45)$$

where  $|\tau|^2$  is the intensity transmission coefficient, and the factor 1/2 arises because we are considering only the right-traveling electrons. If one writes for the quantum wave case  $v_{\rm e} = \sqrt{2\hbar\omega/m}$ , then  $\rho_{\rm e}(\omega) = (\pi v_{\rm e})^{-1}$  for both quantum and classical waves [see (2.26), (2.27)]. Therefore,

$$j = \frac{e^2}{h} \left(\frac{\mu_1 - \mu}{e}\right) |\tau|^2.$$

Since  $(\mu_1 - \mu_2)/e = V$ , we have  $j = \Gamma V$ , where

$$\Gamma = \frac{e^2}{h} |\tau|^2 \tag{11.46}$$

is the one-way conductance. If the 1D sample is broadened into a finite-width strip, the only change in the earlier formula is in the calculation of  $|\tau|^2$ .

# 11.5.1 Quantized Conductance

As seen in Problem 8.1, for the case of a strip with width M there can be a number  $M_0 < M$  of (nonevanescent) incident channels and the same number of exit channels. By defining

$$T_i = \sum_j |\tau_{ij}|^2$$

as the transmission coefficient for the ith output channel, we get

$$\Gamma = \frac{e^2}{h} \sum_{i=1}^{M_0} T_i.$$
(11.47)

If the mesoscopic sample is sufficiently small and ordered, then it is possible that  $T_i = 1$  for all  $i = 1, \ldots, M_0$  (total transmission, no reflection). In that case the conductance is given by

$$\Gamma = \frac{e^2}{h} M_0, \tag{11.48}$$

i.e.,  $\Gamma$  is quantized in units of  $e^2/h$ . Such quantization is seen to arise directly from the finite width of the conduction channel, which acts as an electronic waveguide. Quantized conductances have indeed been observed experimentally in mesoscopic structures (van Wees et al. 1988). Due to spin degeneracy, the observed unit of quantization is  $2e^2/h$  since each channel can accommodate two electrons.

#### 11.5.2 Two-Terminal versus Four-Terminal

The earlier situation applies in the case of two-terminal measurement, i.e., the current leads coincide with the voltage leads. However, this is not necessarily always the case experimentally. One can equally well have four-terminal

measurement where the voltage measurement is conducted via two separate leads, apart from the two current leads. For the same sample, the four-terminal measurement of mesoscopic conductance can differ from the two-terminal measurement; although their difference can be small when the sample is very conducting. To describe the four-terminal measurement, let us refer to the configuration depicted in Fig. 11.3. Besides the two current reservoirs at  $\mu_1$ and  $\mu_2$ , just as in the two-terminal case, we wish to measure the voltage difference across the sample by attaching two additional leads just outside the sample, to its left and right. The chemical potentials of the two voltage leads are denoted by  $\mu_A$  and  $\mu_B$ . Due to the current flow, the chemical potentials  $\mu_A$  and  $\mu_2$ .  $\mu_A$  and  $\mu_B$  are determined by the condition that the number of occupied states above them is the same as the number of unoccupied states below them.

## 11.5.3 The Single Channel (1D) Case

By applying the earlier principle to the right lead in the single channel case, we get the number of occupied states above  $\mu_{\rm B}$  to be

$$\frac{1}{2\hbar}\rho_{\rm e}(\omega)|\tau|^2(\mu_1-\mu_{\rm B}).$$

Since the total number of states between  $\mu_{\rm B}$  and  $\mu_2$  is



Fig. 11.3. A schematic depiction of a mesoscopic sample with four-leads connected to its left and right. The leads connected to electron reservoirs with chemical potentials  $\mu_1$  and  $\mu_2$  supply the current. Two additional leads are for the purpose of measuring the voltage across the sample. Due to the reflected and transmitted currents, the chemical potentials of the voltage leads are at  $\mu_A \neq \mu_1$  and  $\mu_B \neq \mu_2$ . The voltage difference across the sample is given by  $(\mu_A - \mu_B)/e$ .

it follows that the number of unoccupied states below  $\mu_{\rm B}$  is

$$\frac{1}{\hbar} \left( 1 - \frac{1}{2} |\tau|^2 \right) \rho_{\rm e}(\omega) (\mu_{\rm B} - \mu_2).$$

Thus  $\mu_{\rm B}$  is determined by the condition

$$\frac{1}{\hbar}\rho_{\rm e}(\omega)\frac{1}{2}|\tau|^2(\mu_1-\mu_{\rm B}) = \frac{1}{\hbar}\left(1-\frac{1}{2}|\tau|^2\right)\rho_{\rm e}(\omega)(\mu_{\rm B}-\mu_2),\tag{11.49}$$

or

$$\mu_{\rm B} = \frac{1}{2} |\tau|^2 (\mu_1 - \mu_2) + \mu_2.$$
(11.50)

To the left, the number of occupied states above  $\mu_A$  is given by the addition of incident and reflected electrons, i.e.,

$$\frac{1}{2\hbar}(1+|\rho|^2)\rho_{\rm e}(\omega)(\mu_1-\mu_{\rm A}),$$

where  $|\rho|^2$  denotes the intensity reflection coefficient. Similar consideration leads to the number of unoccupied states as

$$\frac{1}{\hbar} \left[ 1 - \frac{1}{2} (1 + |\rho|^2) \right] \rho_{\rm e}(\omega) (\mu_{\rm A} - \mu_2)$$

Equating the two gives

$$\mu_{\rm A} = \mu_1 - \frac{1}{2} (1 - |\rho|^2) (\mu_1 - \mu_2). \tag{11.51}$$

By writing  $|\tau|^2 = 1 - |\rho|^2$  in (11.50), it follows that

$$\mu_{\rm A} - \mu_{\rm B} = |\rho|^2 (\mu_1 - \mu_2). \tag{11.52}$$

Since the voltage  $V = (\mu_{\rm A} - \mu_{\rm B})/e$ , it follows from (11.45) that

$$\Gamma = \frac{e^2 v_{\rm e}}{2\hbar} \rho_{\rm e}(\omega) \frac{|\tau|^2}{|\rho|^2} = \frac{e^2}{h} \frac{|\tau|^2}{|\rho|^2},$$
(11.53)

where we have again used the expression  $\rho_{\rm e}(\omega) = 2(\mathrm{d}\kappa_{\rm e}/\mathrm{d}\omega)/2\pi = 1/\pi v_{\rm e}$ . The dimensionless conductance of a mesoscopic sample is thus given by

$$\frac{\Gamma}{(e^2/h)} = \gamma = \frac{|\tau|^2}{|\rho|^2}.$$
(11.54)

Equation (11.54) should be contrasted with (11.46) for the two terminal case. It is noted that (11.54) is what has been used in Chap. 8 (8.49) to obtain the 1D scaling function.

## 11.5.4 Landauer–Buttiker Formula for the Multichannel Case

To generalize this (four-terminal) formula to 2D and 3D samples, one simply notes that there are many more incident and transmission "channels," as defined in the solution to Problem 8.1. Here we will define the channels in the momentum representation. Let k be the channel index for the left lead and ibe the index for the right lead. If  $|\tau_{ik}|^2$  is the intensity transmission coefficient between the kth input channel and the ith output channel, the current in the ith channel is now given by

$$j_i = \frac{e^2}{h}(\mu_1 - \mu_2) \sum_{k=1}^{M_0} |\tau_{ik}|^2 = \frac{e^2}{h}(\mu_1 - \mu_2)T_i, \qquad (11.55)$$

where  $M_0$  is the number of propagating channels (see the solution to Problem 8.1), and  $T_i$  is the sum of  $|\tau_{ik}|^2$  over all k. The total transmitted current is thus

$$j = \sum_{i} j_{i} = \frac{e^{2}}{h} (\mu_{1} - \mu_{2}) \sum_{i=1}^{M_{0}} T_{i}.$$
 (11.56)

Another way of calculating the current is as follows. The current in channel k of left lead is given by the flux of right-going electrons (from the lead) in channel k, minus the reflected flux:

$$j_k = \frac{e^2}{h}(\mu_1 - \mu_2) \left[ 1 - \sum_{i=1}^{M_0} |\rho_{ik}|^2 \right] = \frac{e^2}{h}(\mu_1 - \mu_2)(1 - R_k),$$

where  $R_k$  is defined as the sum of  $|\rho_{ik}|^2$  over all *i*. The total current is therefore

$$j = \frac{e^2}{h}(\mu_1 - \mu_2) \sum_{k=1}^{M_0} (1 - R_k).$$
(11.57)

Since the total current must be conserved, we have the identity (expressing the unitarity condition)

$$\sum_{i=1}^{M_0} T_i = \sum_{k=1}^{M_0} (1 - R_k).$$
(11.58)

By following the same argument as the 1D case,  $\mu_{\rm B}$  is determined by the condition [see (11.49)]

$$\sum_{i=1}^{M_0} v_i^{-1} T_i(\mu_1 - \mu_{\rm B}) = \sum_{i=1}^{M_0} v_i^{-1} (2 - T_i)(\mu_{\rm B} - \mu_2), \qquad (11.59)$$

where the density of states for the *i*th channel is given by  $(\pi v_i)^{-1}$ . The only complication here, as compared with the 1D case, is that the density of states can differ from channel to channel. From (11.59) it follows that

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$$\mu_{\rm B} = \frac{\sum_{i=1}^{M_0} v_i^{-1} T_i}{2\sum_{i=1}^{M_0} v_i^{-1}} (\mu_1 - \mu_2) + \mu_2.$$
(11.60a)

Similar reasoning leads to the formula

$$\mu_{\rm A} = \mu_1 - \frac{\sum_{i=1}^{M_0} v_i^{-1} (1 - R_i)}{2\sum_{i=1}^{M_0} v_i^{-1}} (\mu_1 - \mu_2).$$
(11.60b)

From the fact  $V = (\mu_{\rm A} - \mu_{\rm B})/e$  and  $\Gamma = j/V$ , we have

$$\Gamma = \frac{e^2}{h} \frac{(\sum_{i=1}^{M_0} v_i^{-1}) \sum_{i=1}^{M_0} T_i}{\sum_{i=1}^{M_0} (1 + R_i - T_i) v_i^{-1}}.$$
(11.61)

This is the general form of the Landauer, or the Landauer–Buttiker, formula (Buttiker et al. 1985) for four-terminal measurement of mesoscopic conductances. It should be noted that spin degeneracy of the electrons is not taken into account in (11.61).

With Landauer's formula and the recursive Green function method as presented in the solution to Problem 8.1, the conductance of mesoscopic samples can be evaluated numerically. Some of their special characteristics are shown below.

# 11.6 Characteristics of Mesoscopic Conductance

Small electronic samples at low temperatures can realize the condition of  $L_{\rm in} > L$ . When this happens a number of interesting, nonclassical behaviors occur. First, the dimensionless  $\gamma$  can vary significantly from one sample to the next, even if the samples are made in the same batch. Second,  $\gamma$  can be sample size dependent. This is due to the coherent backscattering effect as discussed in earlier chapters. Moreover,  $\gamma \to 0$  in 1D and 2D as  $L \to \infty$ , since all states are localized in 1D and 2D.

# 11.6.1 Disordered 2D Metallic Film

To be more specific, it is instructive to look at a particular system, such as a thin disordered metallic film. In practice, a thin film may be regarded as 2D if its thickness is less than a few hundred Angstroms. We will thus model this system by a 2D square lattice which has a certain fraction p of the sites randomly occupied by metal particles and the rest by insulating particles. Just as in the Anderson model [see (2.53)], the wave equation may be written as

$$\mathbf{M} \left| \phi \right\rangle = \frac{2m}{\hbar^2} (\hbar \omega - \varepsilon_0) \left| \phi \right\rangle, \qquad (11.62)$$
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where

$$\mathbf{M} = \sum_{\mathbf{l}} \sigma(\mathbf{l}) |\mathbf{l}\rangle \langle \mathbf{l}| + t \sum_{\mathbf{l},\mathbf{n}} |\mathbf{l}\rangle \langle \mathbf{l} + \mathbf{n}|, \qquad (11.63)$$

t is a constant characterizing nearest-neighbor coupling, and **n** is the vector that points from **l** to one of its nearest neighbors. However, unlike the Anderson model, here  $\sigma(\mathbf{l})$  will take only two values:

$$\sigma(\mathbf{l}) = \begin{cases} 0 & \text{if site } \mathbf{l} \text{ is occupied by a metal particle} \\ \infty & \text{if site } \mathbf{l} \text{ is occupied by insulator.} \end{cases}$$
(11.64)

Since by setting  $\sigma(\mathbf{l}) = \infty$  means the electron will be prevented from entering the site, the randomness in this model is noted to be in the position of the metal particles, in contrast to the Anderson model, where the randomness is in the value of  $\sigma(\mathbf{l})$ . In the literature the present model is usually called the *quantum percolation model*. A finite-sized sample in this model may be depicted schematically as shown in Fig. 11.4.

#### 11.6.2 Distribution of the Dimensionless Conductances

By using the method of recursive Green function and (11.61), the conductance of the model may be evaluated numerically. For a fixed p, there can be many configurations of the sample as the arrangement of metal particles can vary randomly. As s result, the conductance will also vary randomly. Nevertheless, the distribution of the conductances is well defined. Figures 11.5a, b show the forms of the distribution for two samples with L = 60 (in units of the lattice constant) but different values of p. In Fig. 11.5a the value of p = 0.95, the sample's total transmission coefficient is close to 1, and the distribution is Gaussian in character. In Fig. 11.5b the value of p = 0.75, the sample's total transmission is rather low, and the distribution is well described by a lognormal distribution, i.e., Gaussian in terms of  $\ln \gamma$ . Between p = 0.95 and 0.75 the distribution can be difficult to describe analytically.



Fig. 11.4. A schematic picture of a 2D, L = 4, p = 0.5 quantum percolation model configuration. *Filled circles* denote metallic particles.



Fig. 11.5. Histogram of the conductance distribution. In (a), 500 random configurations are calculated with the parameters L = 60 (in units of lattice constant), p = 0.95, and  $2m(\hbar\omega - \varepsilon_0)/\hbar^2 t = 0.01$  (near the bottom of the band). In (b), 900 configurations are calculated with the parameters L = 60, p = 0.75, and  $2m(\hbar\omega - \varepsilon_0)/\hbar^2 t = 0.01$ . The *dashed lines* denote the Gaussian lineshape to guide the eye. It is seen that in (a), the distribution is Gaussian in terms of  $\gamma$ ; whereas in (b), the distribution is Gaussian in terms of ln  $\gamma$ . The figures are from Sheng and Zhang (1991).

#### 11.6.3 Localization Length, Conductance Fluctuations and Transport Regimes

The distribution of  $\gamma$  and its variation with p and L contain all the essential information about mesoscopic conductances. For example, by looking at the logarithmic mean of  $\gamma$ , i.e.,  $\langle \ln \gamma \rangle_c$ , as a function of L it is easy to deduce the localization length. This is shown in Fig. 11.6. Also, the standard deviation of the distribution, defined as  $\Delta \gamma = \sqrt{\langle \gamma^2 \rangle - \langle \gamma \rangle^2}$ , displays interesting variation with L as shown in Fig. 11.7. At small L,  $\Delta \gamma$  displays fast decay. But at a certain point  $\Delta \gamma$  becomes roughly a constant  $\tilde{=}1$ , independent of L. This is



Fig. 11.6. The average value of  $\ln \gamma$  plotted as a function of L. The data points show a linear behavior. The inverse of the slope gives the localization length. (a) is for p = 0.75, (b) is for p = 0.8, and (c) is for p = 0.85. The value of  $2m(\hbar\omega - \varepsilon_0)/\hbar^2 t = 0.01$ . The figure is from Sheng and Zhang (1991).



Fig. 11.7. The conductance fluctuations, expressed in terms of the standard deviation  $\Delta \gamma$  of the distribution, plotted as a function of L for various values of  $p : \Delta$ is for p = 0.9,  $\bigcirc$  is for p = 0.85,  $\blacktriangle$  is for p = 0.8,  $\bigstar$  is for p = 0.75,  $\bullet$  is for p = 0.7, and  $\Box$  is for p = 0.65. The value of  $2m(\hbar\omega - \varepsilon_0)/\hbar^2 t$  is 0.5. The figure is from Sheng and Zhang (1991).

precisely the phenomenon of UCF. However, at larger distances  $\Delta \gamma$  is expected to decay further. The region of UCF extends roughly over the regime  $10l < L < \xi$ .

#### 11.6.4 Heuristic Explanation of Universal Conductance Fluctuations

In terms of Landauer's formula and the prior discussion on spatial correlations, there is a heuristic explanation for the existence of the UCF regime. Take the two-terminal version of Landauer's formula,

$$\gamma = \sum_{i,k}^{M_0} |\tau_{ik}|^2,$$

and consider the evaluation of  $\Delta \gamma$ . From the discussion in Sect. 11.5 it is clear that due to interference effect, there can be strong correlation in the transmission coefficients  $\{\tau_{ik}\}$ , so they cannot be independent. However, Lee proposed (Lee 1986) that the reflection coefficients,  $\{\rho_{ik}\}$ , should be roughly independent because they are dominated by the short paths near the input lead-sample interface. Since the reflection coefficients are related to the transmission coefficients due to the unitary requirement, we have from (11.58)

$$\sum_{i,k}^{M_0} |\tau_{ik}|^2 = M_0 - \sum_{i,k}^{M_0} |\rho_{ik}|^2.$$
(11.65)

It follows that

$$\Delta \gamma = \left[ \left\langle \gamma^2 \right\rangle_c - \left\langle \gamma \right\rangle_c^2 \right]^{1/2} \cong M_0 \left[ \left\langle |\rho_{ik}|^4 \right\rangle_c - \left\langle |\rho_{ik}|^2 \right\rangle_c^2 \right]^{1/2}, \qquad (11.66)$$

where the reflection coefficients are assumed to be *independent* so the  $M_0^2 \cdot (M_0^2 - 1)$  terms like  $\langle |\rho_{ik}|^2 |\rho_{lm}|^2 \rangle_c - \langle |\rho_{ik}|^2 \rangle_c^2 = 0$  if  $i \neq l$  and/or  $k \neq m$ . Also, it should be noted that configuration-averaged quantities are independent of the indices. Therefore, e.g.,  $\langle \gamma \rangle_c \simeq M_0^2 \langle |\tau_{ik}|^2 \rangle_c$  due to the fact that there are roughly  $M_0^2$  terms in the *i*, *j* double summation. If  $|\rho_{ik}|^2$  is considered to be the intensity of many interfering paths, then the argument of Sect. 11.2 applies, and the standard deviation of the distribution is equal to its mean. That implies

$$\Delta \gamma \cong M_0 \left\langle \left| \rho_{ik} \right|^2 \right\rangle_{\rm c} = 1 - \frac{1}{M_0} \left\langle \sum_{i,k}^{M_0} \left| \tau_{ik} \right|^2 \right\rangle_{\rm c}, \tag{11.67}$$





Fig. 11.8. Transport regimes of the 2D quantum percolation model as a function of L and p. In both cases the *upper solid line* denotes the localization length ( $\xi = L$ ), the *lower solid line* denotes the mean free path (l = L). Dashed line denotes ten times the mean free path (10l = L), and  $2m(\hbar\omega - \varepsilon_0)/\hbar^2 t = 0.01$  in (**a**), 0.5 in (**b**). The figures are from Sheng and Zhang (1991).

where the last equality follows from (11.65) by configurational averaging both sides and dividing by  $M_0$ , and noting the last term of (11.65) to have  $\sim M_0^2$ terms in the double summation. Since the averaged total transmission varies as 1/L, it follows from (11.67) that for a *sufficiently thick sample* the leading order term in  $\Delta \gamma$  is independent of L, with the magnitude on the order of 1. This is precisely the universal conductance fluctuation.

The behavior of mesoscopic conductances can be summarized by the "phase" diagrams of Figs. 11.8a, b. The two external control variables are L and p. It is seen that mesoscopic systems have an entirely different set of characteristics from the bulk systems. However, it is interesting to observe that it is precisely at the mesoscopic level that quantum and classical waves display very similar behavior due to the common underlying physics.

### **Problems and Solutions**

**Problem 11.1** Evaluate  $S(\Delta \omega, \mathbf{r} - \mathbf{r}' | \omega)$  in 3D.

**Solution** From the diffusive form of  $S(\Delta \omega, \Delta \mathbf{k} | \omega)$ , we get

$$S(\Delta\omega, \mathbf{r} - \mathbf{r}'|\omega) = \frac{\rho_{\rm e}(\omega)}{2\pi} \frac{\mathrm{d}\omega}{\mathrm{d}\kappa_0^2} \frac{\mathrm{d}\omega}{\mathrm{d}\kappa_{\rm e}^2} \int_0^\infty (\Delta k)^2 \mathrm{d}(\Delta k) \times \int_{-1}^1 \mathrm{d}\cos\theta \frac{\exp\left[\mathrm{i}\Delta k|\mathbf{r} - \mathbf{r}'|\cos\theta\right]}{-\mathrm{i}\Delta\omega + v_t l(\Delta k)^2/3} = \frac{\rho_{\rm e}(\omega)}{2\pi} \frac{\mathrm{d}\omega}{\mathrm{d}\kappa_0^2} \frac{\mathrm{d}\omega}{\mathrm{d}\kappa_{\rm e}^2} \int_0^\infty \mathrm{d}(\Delta k) \frac{(\Delta k)2\sin\left[\Delta k|\mathbf{r} - \mathbf{r}'|\right]}{|\mathbf{r} - \mathbf{r}'|(v_t l/3)\left[(\Delta k)^2 - \mathrm{i}\frac{3\Delta\omega}{v_t l}\right]} = \frac{3\rho_{\rm e}(\omega)(\mathrm{d}\omega/\mathrm{d}\kappa_0^2)(\mathrm{d}\omega/\mathrm{d}\kappa_{\rm e}^2)}{2\pi v_t l|\mathbf{r} - \mathbf{r}'|} \int_{-\infty}^\infty \mathrm{d}x \frac{x\sin x}{x^2 - \mathrm{i}\left(\frac{3\Delta\omega|\mathbf{r} - \mathbf{r}'|^2}{v_t l}\right)}.$$
 (P11.1)

The poles of the integrand are at

$$x = x_{\pm} = \pm (1+\mathrm{i}) \sqrt{\frac{3\Delta\omega}{2v_t l}} |\mathbf{r} - \mathbf{r}'|.$$
(P11.2)

By carrying out the integration in the complex plane, we get

$$S(\Delta\omega, \mathbf{r} - \mathbf{r}'|\omega) = \frac{3\rho_{\rm e}(\omega)(\mathrm{d}\omega/\mathrm{d}\kappa_0^2)(\mathrm{d}\omega/\mathrm{d}\kappa_{\rm e}^2)}{4v_t l|\mathbf{r} - \mathbf{r}'|} \left[\exp(\mathrm{i}x_+) + \exp(-\mathrm{i}x_-)\right]$$
$$= \frac{3\rho_{\rm e}(\omega)(\mathrm{d}\omega/\mathrm{d}\kappa_0^2)(\mathrm{d}\omega/\mathrm{d}\kappa_{\rm e}^2)}{2v_t l|\mathbf{r} - \mathbf{r}'|}$$
$$\times \exp\left[(1+\mathrm{i})\sqrt{\frac{3\Delta\omega}{2v_t l}}|\mathbf{r} - \mathbf{r}'|\right].$$
(P11.3)

**Problem 11.2** Show that the  $G(\mathbf{k}, z, z')$  given by (11.37) solves (11.36).

Solution The Green function to (11.36) must satisfy

$$\frac{\mathrm{d}^2}{\mathrm{d}z^2}G(\mathbf{k}, z_<, z_>) - k^2 G(\mathbf{k}, z_<, z_>) = \delta(z_< - z_>).$$
(P11.4)

Let us first take the  $z_{<}, z_{>}$  derivatives of G:

$$\frac{\mathrm{d}G(\mathbf{k}, z_{<}, z_{>})}{\mathrm{d}z_{<}} = -\frac{\cosh(kz_{<})\sinh[k(L-z_{>})]}{\sinh(kL)}.$$
 (P11.5)

$$\frac{\mathrm{d}G(\mathbf{k}, z_{<}, z_{>})}{\mathrm{d}z_{>}} = +\frac{\sinh(kz_{<})\cosh[k(L-z_{>})]}{\sinh(kL)}.$$
 (P11.6)

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By taking the second derivatives with respect to  $z_{<}$  and  $z_{>}$ , one gets back exactly  $k^2G(\mathbf{k}, z_{<}, z_{>})$  everywhere except at  $z_{<} = z_{>}$ . Therefore, (P11.4) is clearly satisfied everywhere  $z_{<} \neq z_{>}$ . At  $z_{<} = z_{>}$ , we have

$$\frac{\mathrm{d}G(\mathbf{k}, z_{<}, z_{>})}{\mathrm{d}z_{>}} - \frac{\mathrm{d}G(\mathbf{k}, z_{<}, z_{>})}{\mathrm{d}z_{<}} = 1.$$
(P11.7)

Therefore, there is a unit jump at  $z_{<} = z_{>}$ , which means the second derivative must give a delta function at  $z_{<} = z_{>}$ . Moreover,  $G(\mathbf{k}, z_{<}, z_{>})$  satisfies the boundary conditions that  $G(\mathbf{k}, z_{<}, z_{>}) = 0$  at  $z_{<} = 0$ ,  $z_{>} = L$ .

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