

Farook Rahaman

# The Special Theory of Relativity

A Mathematical Approach

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*To my parents  
Majeda Rahaman and Late Obaidur  
Rahaman  
and my son Rahil Miraj*

# Preface

In 1905, Albert Einstein wrote three papers which started three new branches in physics; one of them was the special theory of relativity. The subject was soon included as a compulsory subject in graduate and postgraduate courses of physics and applied mathematics across the world. And then various eminent scientists wrote several books on the topic.

This book is an outcome of a series of lectures delivered by the author to graduate students in mathematics at Jadavpur University. During his lectures, many students asked several questions which helped him know the needs of students. It is, therefore, a well-planned textbook, whose contents are organized in a logical order where every topic has been dealt with in a simple and lucid manner. Suitable problems with hints are included in each chapter mostly taken from question papers of several universities.

In Chap. 1, the author has discussed the limitations of Newtonian mechanics and Galilean transformations. Some examples related to Galilean transformations have been provided. It is shown that laws of mechanics are invariant but laws of electrodynamics are not invariant under Galilean transformation. Now, assuming both the Galilean transformation and electrodynamics governed by Maxwell's equations are true, it is obvious to believe there exists a unique privileged frame of reference in which Maxwell's equations are valid and in which light propagates with constant velocity in all directions. Scientists tried to determine relative velocity of light with respect to earth. For this purpose, the velocity of earth relative to some privileged frame, known as ether frame, was essential. Therefore, scientists performed various experiments involving electromagnetic waves. In this regard, the best and most important experiment was performed by A.A. Michelson in 1881.

In Chap. 2, the author has described the Michelson–Morley experiment. In 1817, J.A. Fresnel proposed that light is partially dragged along by a moving medium. Using the ether hypothesis, Fresnel provided the formula for the effect of moving medium on the velocity of light. In 1951, Fizeau experimentally confirmed this effect. The author has discussed Fizeau's experiment. Later, the author provided the relativistic concept of space and time. To develop the special theory of relativity, Einstein used two fundamental postulates: the principle of relativity or equivalence

and the principle of constancy of the speed of light. The principle of relativity states that the laws of nature are invariant under a particular group of space-time coordinate transformations. Newton's laws of motion are invariant under Galilean transformation, but Maxwell's equations are not, and Einstein resolved this conflict by replacing Galilean transformation with Lorentz transformation.

In Chap. 3, the author has provided three different methods for constructing Lorentz transformation between two inertial frames of reference. The Lorentz transformations have some interesting mathematical properties, particularly in measurements of length and time. These new properties are not realized before. Mathematical properties of Lorentz transformations—length contraction, time dilation, relativity of simultaneity and their consequences—are discussed in Chaps. 4 and 5. Also, some paradoxes like twin paradox, car–garage paradox have been discussed in these chapters.

Various types of intervals as well as trajectories of particles in Minkowski's four-dimensional world can be described by diagrams known as space–time diagrams. To draw the diagram, one has to use a specific inertial frame in which one axis indicates the space coordinate and the other effectively the time axis. Chapter 6 includes the geometrical representation of space–time.

Chapter 7 discusses the relativistic velocity transformations, relativistic acceleration transformations and relativistic transformations of the direction cosines. The author has also discussed some applications of relativistic velocity and velocity addition law: (i) explanation of index of refraction of moving bodies (Fizeau effect), (ii) aberration of light and (iii) relativistic Doppler effect.

In Newtonian mechanics, physical parameters like position, velocity, momentum, acceleration and force have three-vector forms. It is natural to extend the three-vector form to a four-vector form in the relativistic mechanics in order to satisfy the principle of relativity. Therefore, a fourth component of all physical parameters should be introduced. In Euclidean geometry, a vector has three components. In relativity, a vector has four components.

The world velocity or four velocity, Minkowski force or four force, four momentum and relativistic kinetic energy have been discussed in Chaps. 8 and 10. It is known that the fundamental quantities, length and time are dependent on the observer. Therefore, it is expected that mass would be an observer-dependent quantity. The author provides three methods for defining relativistic mass formula in Chap. 9. The relativistic mass formula has been verified by many scientists; however, he has discussed the experiment done by Guye and Lavanchy.

A short discussion on photon and Compton effect is provided in Chap. 11. A force, in general, is not proportional to acceleration in case of the special theory of relativity. The author has discussed force in the special theory of relativity and covariant formula of the Newton's law. Relativistic Lagrangian and Hamiltonian are discussed in Chap. 12. He has also explained Lorentz transformation of force and relativistic harmonic oscillation. When charges are in motion, the electric and magnetic fields are associated with this motion, which will have space and time variation. This phenomenon is called *electromagnetism*. The study involves time-dependent electromagnetic fields and the behaviour of which is described by a set of

equations called *Maxwell's equations*. In Chap. 13, the author has discussed relativistic electrodynamics of continuous medium. In Chap. 14, the author has provided a short note on relativistic mechanics of continuous medium (continua). Finally, in appendixes, the author has provided some preliminary concepts of tensor algebra and action principle.

The author expresses his sincere gratitude to his wife Mrs. Pakizah Yasmin for her patience and support during the gestation period of the manuscript. It is a pleasure to thank Dr. M. Kalam, Indrani Karar, Iftikar Hossain Sardar and Mosiur Rahaman for their technical assistance in preparation of the book. Finally, he is thankful to painter Ibrahim Sardar for drawing the plots.

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## About the Author

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# Chapter 1

## Pre-relativity and Galilean Transformation

### 1.1 Failure of Newtonian Mechanics

Isaac Newton proposed three laws of motion in the year 1687. Newton's first law states: Every body continues in its state of rest or of uniform motion in a straight line unless it is compelled by external forces to change that state. The law requires no definition of the internal structure of the body and force, but simply refers to the behaviour of the body. I can remember, my teacher's example. Suppose, one puts a time bomb in a place. It is at rest and according to Newton's first law, it remains at rest. But after some time, it explodes and the splinters run here and there. Now, we cannot say the body (i.e. the bomb) is at rest as it was some time ago and no external force was applied. This is not a breakdown of Newton's first law, rather we should define the internal structure of the body. Also, we note that uniform motion is not an absolute property of the body but a joint property of the body and reference system used to observe it. Therefore, one should specify the frame of references relative to which the position of the body is determined. Moreover, Newton's first law is true only for some observers and their associated reference frame and not for others. Assume we have an observer  $A_1$  sitting in a reference frame  $S_1$  for which Newton's first law is true, i.e. he determines that the motion of a forceless body is uniform and rectilinear. Now, we consider an observer  $A_2$  sitting in another reference frame  $S_2$  moving with an acceleration with respect to the former. For him, the same body (not acted upon by external forces) will appear to have an accelerated motion. He concludes that Newton's first law is not true. Therefore, one should mention the frame of reference in developing Mechanics based on Newton's laws of motion. In the macroscopic world, the velocity of particles ( $v$ ) with respect to any observer is always less than the velocity of light  $c$  ( $=3 \times 10^8$  m/s). In fact,  $\frac{v}{c} \ll 1$  in our ordinary experiences. However, in the microscopic world, particles can be found frequently whose velocities are very close to the velocity of light. Experiments show that these high-speed particles do not follow Newtonian mechanics. In Newtonian mechanics, there is no upper bound of the particle's speed. As a result, speed of light plays no special role at all in Newtonian mechanics.

*For example:* An electron accelerated through a 10 million-volt (MeV) potential difference, then its speed 'v' equal to  $0.9988c$ . If this electron is accelerated through a 40 MeV potential difference, experiment shows that its speed equals  $.9999c$ . The speed is increased from  $0.9988c$  to  $0.9999c$ , a change of 11 % and remains below  $c$ . However, according to Newtonian mechanics ( $K = \frac{1}{2}mv^2$ ), we expect its velocity should be doubled to  $1.9976c$ . Hence, Newtonian mechanics may work at low speeds, it fails badly as  $\frac{v}{c} \rightarrow 1$ . Einstein extended and generalized Newtonian mechanics. He correctly predicted the results of mechanical experiments over the complete range of speeds from  $\frac{v}{c} \rightarrow 0$  to  $\frac{v}{c} \rightarrow 1$ .

*Inertial Frame:* The frames relative to which an unaccelerated body appears to be unaccelerated are called Inertial Frames. In other words, the frame in which the Newton's laws of motion are valid.

*Non-inertial Frame:* The frames relative to which an unaccelerated body appears to be accelerated are called Non-inertial Frames. In other words, the frame in which the Newton's laws of motion are not valid.

Loosely speaking, Newtonian frame of reference fixed in the stars is an inertial frame. The coordinate system fixed in earth is not an inertial frame since earth rotates about its axis and also about the sun. In fact, rotational frame is not an inertial frame.

The laws of Newtonian mechanics are applicable only for the inertial frames. Galileo pointed out that it is impossible to distinguish between rest and motion in the case of bodies moving with uniform velocities along any straight line relative to each other. In other words, one inertial frame is as good as any other.

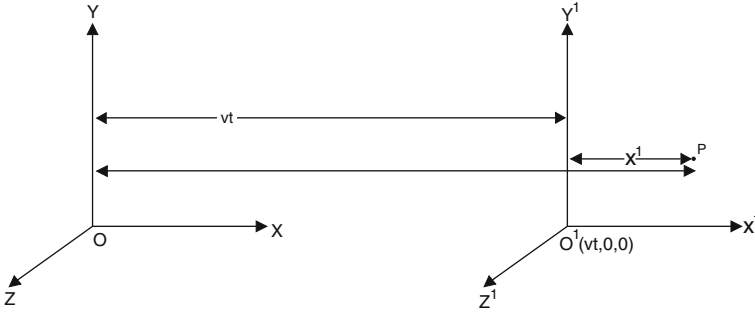
## 1.2 Galilean Transformations

We try to find the relationship of the motions as observed by two observers in two different inertial frames. Let us consider two different Cartesian frame of references  $S$  and  $S_1$  comprising the coordinate axes  $(x, y, z)$  and  $(x_1, y_1, z_1)$  which are inertial. Here, the axes  $(x_1, y_1, z_1)$  of  $S_1$  are parallel to the axes  $(x, y, z)$  of  $S$ . We further assume that at time  $t = 0$ , the origin  $O_1$  of  $S_1$  system coincides with the origin  $O$  of  $S$  system and all axes overlapped. The frame  $S_1$  is supposed to move at a uniform velocity  $v$  along the  $x$ -axis. Let an event occurs at a point  $P$  whose coordinates are  $(x, y, z)$  and  $(x_1, y_1, z_1)$  in  $S$  and  $S_1$  frames, respectively. Let  $t$  and  $t_1$  be the time of occurrence of  $P$  that observer  $S$  and  $S_1$  record their clocks. Then, the coordinates of two systems can be combined to each other as (from Fig. 1.1)

$$x_1 = x - vt \tag{1.1}$$

$$y_1 = y \tag{1.2}$$

$$z_1 = z \tag{1.3}$$



**Fig. 1.1**  $S_1$  frame is moving with velocity  $v$  along  $x$ -direction

It is assumed that ‘time’ can be defined independently of any particular frame of reference. As a result, time of the event recorded by the clocks of both systems will be the same. Hence, one can write the equation of transformation of time:

$$t_1 = t \tag{1.4}$$

Actually, our own common senses confirm this conclusion as we do not detect any influence of the motion on clock rate in daily life. The above equation indicates that the time interval between occurrence of two events, say A and B, remains the same for each observer, i.e.

$$t_1^A - t_1^B = t^A - t^B$$

The transformation Eqs. (1.1)–(1.4) are valid only for the relative position of the reference frame S and  $S_1$  and known as Galilean Transformations. Note that these are the transformations of the event coordinates from reference S to the reference  $S_1$ . To describe an event, one has to say where and when and this occurs. As a result, time turns out to be the fourth coordinate. Therefore, the coordinates of an event are defined by four numbers  $(x, y, z, t)$ .

Differentiation Eqs. (1.1)–(1.3) with respect to ‘time’, we get

$$u_x^1 = u_x - v \tag{1.5}$$

$$u_y^1 = u_y \tag{1.6}$$

$$u_z^1 = u_z \tag{1.7}$$

A second differentiation gives

$$a_x^1 = a_x \tag{1.8}$$

$$a_y^1 = a_y \tag{1.9}$$

$$a_z^1 = a_z \tag{1.10}$$

because  $v$  is a constant.



We find that components of acceleration are the same in the two frames. Hence, the Newton's law ( $P = mf$ ) is valid for both the frames S and  $S_1$ . Hence according to the definition, both S and  $S_1$  frames are inertial frames. Note that not only S and  $S_1$  frames but also all frames of reference having uniform motion relative to them are inertial. In other words, there exists infinite number of inertial frames.

Inverse of these transformations will be

$$x = x_1 + vt \quad (1.11)$$

$$y = y_1 \quad (1.12)$$

$$z = z_1 \quad (1.13)$$

$$t = t_1 \quad (1.14)$$

### 1.3 Galilean Transformations in Vector Form

Consider two systems S and  $S_1$  where  $S_1$  moving with velocity  $v$  relative to S. Initially, origins of two systems coincide. Let  $\vec{r}$  and  $\vec{r}_1$  be the position vectors of any particle P relative to origins O and  $O_1$  of systems S and  $S_1$ , respectively, after time  $t$ . Then,  $\vec{OO}_1 = \vec{v}t$ . So, from Fig. 1.2, using the law of triangle of vector addition

$$\vec{r}_1 = \vec{r} - \vec{v}t \quad (1.15)$$

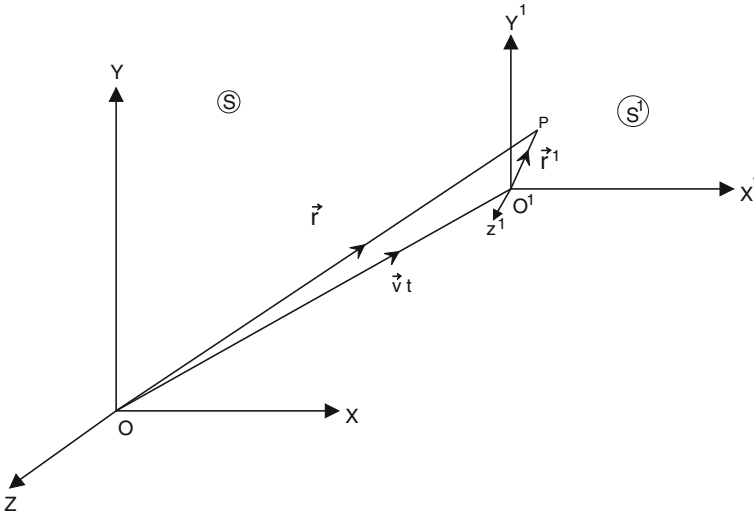


Fig. 1.2  $S_1$  frame is moving with velocity  $v$  in arbitrary direction

Also

$$t_1 = t \quad (1.16)$$

The inverse of Galilean Transformation in vectorial form is

$$\vec{\mathbf{r}} = \vec{\mathbf{r}}_1 + \vec{\mathbf{v}} t \quad (1.17)$$

Also

$$t = t_1 \quad (1.18)$$

Since Newtonian laws of motion remain the same in any two inertial frames  $S$  and  $S_1$  (where one frame is moving with uniform velocity with respect to other), therefore, one could not able to distinguish mechanical experiments following Newtonian laws of motion.

Since at the time of Galileo, the laws of mechanics were the only known physical laws, therefore, the following principle is known as *Principle of Galilean Relativity*:

*If an inertial system is given, then any other system that moves uniformly and rectilinear with respect to it is itself inertial.*

**Example 1.1** Using Galilean Transformation, show that the length of a rod remains the same as measured by the observers in frames  $S$  and  $S^1$ .

**Hint:** Let coordinates of the end points  $P$  and  $F$  of a rod in  $S$  frame are  $(x_1, y_1, z_1)$ ,  $(x_2, y_2, z_2)$  and coordinates of the end points  $P$  and  $F$  of the same rod in  $S^1$  frame are  $(x_1^1, y_1^1, z_1^1)$ ,  $(x_2^1, y_2^1, z_2^1)$ .

Then according to G.T.,

$$\begin{aligned} x_2^1 - x_1^1 &= (x_2 - vt) - (x_1 - vt) = x_2 - x_1; \\ y_2^1 - y_1^1 &= y_2 - y_1; \\ z_2^1 - z_1^1 &= z_2 - z_1; \end{aligned}$$

Therefore,

$$\begin{aligned} l^{1^2} &= [(x_2^1 - x_1^1)^2 + (y_2^1 - y_1^1)^2 + (z_2^1 - z_1^1)^2] \\ &= [(x_2 - x_1)^2 + (y_2 - y_1)^2 + (z_2 - z_1)^2] = l^2 \end{aligned}$$

In general,  $(\vec{\mathbf{r}}_2^1 - \vec{\mathbf{r}}_1^1)^2 = (\vec{\mathbf{r}}_2 - \vec{\mathbf{r}}_1)^2$ .

**Example 1.2** The distance between two points is invariant under Galilean Transformation.

*Example 1.3* Law of conservation of momentum is invariant under Galilean Transformation if the law of conservation of energy holds.

**Hint:** We prove it by means of collision of two bodies. Conservation of momentum states that in the absence of external forces, sum of momenta before and after collision remains the same. Let two bodies of masses  $m, M$  move with respective velocities  $u, U$  before impact and let  $u_1, U_1$  be their velocities after impact. According to the law of conservation of energy

$$\frac{1}{2}mu^2 + \frac{1}{2}MU^2 = \frac{1}{2}mu_1^2 + \frac{1}{2}MU_1^2 + \Delta W \quad (1)$$

where  $\Delta W$  is the change of internal excitation energy due to the collision. Now, one can observe the same collision from the  $S_1$  system which is moving with  $v$  with respect to  $S$  system. Therefore, according to the Galilean Transformation,

$$\begin{aligned} u^1 &= u - v \\ U^1 &= U - v \\ u_1^1 &= u_1 - v \\ U_1^1 &= U_1 - v \end{aligned} \quad (2)$$

Since law of conservation of energy holds in both system, then we have

$$\frac{1}{2}m(u^1)^2 + \frac{1}{2}M(U^1)^2 = \frac{1}{2}m(u_1^1)^2 + \frac{1}{2}M(U_1^1)^2 + \Delta W \quad (3)$$

where,  $u^1, U^1, u_1^1, U_1^1$  represent the corresponding velocities in  $S^1$  system. Here, internal excitation energy  $\Delta W$  remains the same in both systems. Using the values of the velocities in  $S_1$  system in (3), we get,

$$\frac{1}{2}m(u - v)^2 + \frac{1}{2}M(U - v)^2 = \frac{1}{2}m(u_1 - v)^2 + \frac{1}{2}M(U_1 - v)^2 + \Delta W \quad (4)$$

Using Eqs. (1), and (4) yields

$$mu + MU = mu_1 + MU_1 \quad (5)$$

Hence the proof.

It is important to note that in an elastic collision,  $\Delta W = 0$ . Therefore, if a collision is either elastic or inelastic, the conservation law of momentum holds good.

*Example 1.4* Derive Galilean Transformation for total momentum.

**Hint:** Galilean Transformation of momentum  $\vec{p}$  is

$$\vec{p}^1 = m \vec{u}^1 = m(\vec{u} - \vec{v}) = \vec{p} - m \vec{v}$$

Let us consider a system of particles. We define the total momentum and total mass as

$$P = \sum p_i, \quad M = \sum m_i$$

Galilean Transformation of total momentum  $\vec{P}$  is

$$\vec{P}^1 = \sum \vec{p}_i^1 = \sum \vec{p}_i - \vec{v} \sum m_i = \vec{P} - M \vec{v}$$

*Example 1.5* Let a reference frame  $S^1$  is moving with velocity  $v$  relative to  $S$  frame which is at rest. A rod is at rest relative to  $S$ , at an angle  $\theta$  to the  $x$ -axis. Find the angle  $\theta^1$  that makes with the  $x^1$  axis of  $S^1$ .

**Hint:** Let at any time the rod lies in  $xy$  plane with its ends  $(x_1, y_1), (x_2, y_2)$ . Then the angle  $\theta$  is given by

$$\theta = \tan^{-1} \left[ \frac{y_2 - y_1}{x_2 - x_1} \right]$$

Similarly, the angle  $\theta^1$  is defined by

$$\theta^1 = \tan^{-1} \left[ \frac{y_2^1 - y_1^1}{x_2^1 - x_1^1} \right]$$

From Galilean Transformation, we have

$$x_2^1 - x_1^1 = x_2 - x_1, \quad y_2^1 - y_1^1 = y_2 - y_1$$

Therefore, we get,

$$\theta^1 = \tan^{-1} \left[ \frac{y_2^1 - y_1^1}{x_2^1 - x_1^1} \right] = \tan^{-1} \left[ \frac{y_2 - y_1}{x_2 - x_1} \right] = \theta$$

*Example 1.6* Show that the electromagnetic wave equation

$$\frac{\partial^2 \phi}{\partial x^2} + \frac{\partial^2 \phi}{\partial y^2} + \frac{\partial^2 \phi}{\partial z^2} - \frac{1}{c^2} \frac{\partial^2 \phi}{\partial t^2} = 0$$

is not invariant under Galilean Transformation.

**Hint:** Applying chain rule, we get

$$\frac{\partial \phi}{\partial x} = \frac{\partial \phi}{\partial x^1} \frac{\partial x^1}{\partial x} + \frac{\partial \phi}{\partial y^1} \frac{\partial y^1}{\partial x} + \frac{\partial \phi}{\partial z^1} \frac{\partial z^1}{\partial x} + \frac{\partial \phi}{\partial t^1} \frac{\partial t^1}{\partial x}$$

Using G.T., one can get,

$$\frac{\partial \phi}{\partial x} = \frac{\partial \phi}{\partial x^1}$$

In a similar manner, one can find,

$$\frac{\partial \phi}{\partial y} = \frac{\partial \phi}{\partial y^1}, \quad \frac{\partial \phi}{\partial z} = \frac{\partial \phi}{\partial z^1}, \quad \frac{\partial \phi}{\partial t} = \frac{\partial \phi}{\partial t^1} - v \frac{\partial \phi}{\partial x^1}$$

The second derivative yields,

$$\begin{aligned} \frac{\partial^2 \phi}{\partial x^2} &= \frac{\partial^2 \phi}{\partial x^{12}}, & \frac{\partial^2 \phi}{\partial y^2} &= \frac{\partial^2 \phi}{\partial y^{12}}, & \frac{\partial^2 \phi}{\partial z^2} &= \frac{\partial^2 \phi}{\partial z^{12}}, \\ \frac{\partial^2 \phi}{\partial t^2} &= \frac{\partial^2 \phi}{\partial t^{12}} - 2v \frac{\partial^2 \phi}{\partial x^1 \partial t^1} + v^2 \frac{\partial^2 \phi}{\partial x^{12}} \end{aligned}$$

etc.

## 1.4 Non-inertial Frames

Consider two systems S and  $S^1$ , the latter moving with an acceleration  $f$  relative to the former. If the first frame to be inertial, then the second frame will be non-inertial and vice versa. Let us take a Cartesian frame of reference S, which is inertial. Let  $S^1$  be another Cartesian frame whose axes  $x^1, y^1, z^1$  are parallel to axes  $x, y, z$  of S and is moving with an acceleration  $f$  along the x-axis. We further assume that at the origin at time  $t = 0$ , the point O coincides with  $O^1$  and all axes overlapped. Then, the  $x^1$  coordinate of a particle P will be related to the x-coordinate by the relation:

$$x^1 = x - \frac{1}{2}ft^2$$

The  $y, y^1$  and  $z, z^1$  coordinates will be identical, i.e.

$$y^1 = y, \quad z^1 = z$$

It is assumed that ‘time’ can be defined independently of any particular frame of reference. That is, time of the event recorded by the clocks of both systems will be the same. Hence, one can write the equation of transformation of time:

$$t^1 = t$$

Differentiation twice of the above equations with respect to time, we get

$$a_x^1 = a_x - f \quad ; \quad a_y^1 = a_y \quad ; \quad a_z^1 = a_z$$

Thus total force acting on the particle as observed in  $S^1$  is given by

$$F^1 = m(a_x - f) = ma_x - mf$$

Clearly additional force on particle due to acceleration of frame  $S_1$  is given by

$$F_p = -mf = -\text{mass} \times \text{acceleration of non-inertial frame}$$

The force experienced by the particle due to acceleration of non-inertial frame is called *fictitious force*.

## 1.5 Galilean Transformation and Laws of Electrodynamics

The only physical laws known at the time of Galileo were the laws of mechanics. And the laws of physics (laws of mechanics) are invariant in Galilean Transformation (inertial frame). Lately, the laws of electrodynamics were discovered and are not invariant under Galilean Transformation. Electrodynamics is governed by Maxwell's equations. According to Maxwell's theory of light is identified as electromagnetic wave ( $\psi$ ) and electromagnetic waves propagate in empty space with constant velocity  $c = (3 \times 10^8 \text{ m/s})$ . In S frame,

$$\psi = f[\vec{\mathbf{k}} \cdot \vec{\mathbf{r}} - ct]$$

where  $\vec{\mathbf{k}} = l\hat{i} + m\hat{j} + n\hat{k}$  is unit vector along the wave normal. Now, we consider another inertial frame  $S^1$  which is moving with velocity  $\vec{v}$  relative to S. Therefore, in  $S^1$  frame,

$$\psi^1 = f(\vec{\mathbf{k}} \cdot (\vec{\mathbf{r}}^1 + \vec{v}t) - ct) = f[\vec{\mathbf{k}} \cdot \vec{\mathbf{r}}^1 - (c - \vec{\mathbf{k}} \cdot \vec{v})t]$$

Hence in  $S^1$  frame, the speed of the wave will be  $(c - \vec{\mathbf{k}} \cdot \vec{v})$ . Therefore, the velocity will be different in different directions. Therefore, Maxwell's equations are not preserved in form by the Galilean Transformation, i.e. Maxwell's equations are not invariant under Galilean Transformation.

## 1.6 Attempts to Locate the Absolute Frame

It is shown that Laws of Mechanics are invariant but Laws of Electrodynamics are not invariant under Galilean Transformation. Now, if we assume, both the Galilean Transformation and Electrodynamics governed by Maxwell's equations are true, then it is obvious to believe there should exist a unique privileged frame of reference in

which Maxwell's equations are valid and in which light propagates with constant velocity  $c$  in all directions.

*The theory of Ether:* After the advancement of modern techniques at the beginning of the twentieth century, scientists were trying to measure the speed of light in refined manner. Then, it was questioned what was the reference system or coordinate frame with respect to which the speed of light was measured? Also it was once believed that wave motion required a medium through which it propagates as sound in air. Since light evidently travels through a vacuum between the stars, it was believed that the vacuum must be filled by a medium for wave propagation. The imaginary medium filled the entire Universe is called ETHER (weight less). The obvious hypothesis that arises from this supposition is that the 'privileged' observers are those at rest relative to the ether, i.e. Scientists thought that a frame of reference fixed relative to this ether medium was the *special privileged frame of reference*. Thus in their opinion, ETHER was the frame absolutely at rest. The preferred medium for light was ether, which filled all space uniformly. It is assumed that the ether is perfectly transparent to light and it exerts no resistance on material bodies passing through it. The rotational velocity of earth is about  $3 \times 10^4$  m/s, therefore it was supposed that according to Galilean Transformation, the speed of light should depend upon the direction of motion of light with respect to earth's motion through the ether. Since ether remains fixed in space, therefore it was considered to be possible to find the absolute velocity of any body moving through it. Several Physicists such as Fizeau, Michelson and Morley, Trouton and Noble, and many others conducted a number of experiments to find the absolute velocity of earth through ether. However, their attempts were failed. Later Hertz proposed that ether was not at rest rather it is moving with the same velocity of body moving through it. This assumption was inconsistent with the phenomenon of aberration and hence the assumption was rejected.

*Aberration:* The phenomenon of apparent displacement in the sky is due to finite speed of light and the motion of earth in its orbit about the sun is called Aberration.

After this, Fizeau and Fresnel predicted that light would be partially dragged along a moving medium, i.e. ether was partially dragged by the body moving through it.

## Chapter 2

# Michelson–Morley Experiment and Velocity of Light

### 2.1 Attempts to Locate Special Privileged Frame

Scientists tried to determine relative velocity of light with respect to earth. For this purpose, the velocity of earth relative to ether frame of special privileged frame was essential. Therefore, scientists performed various experiments involving electromagnetic waves. In this regard, the best and most important experiment was first performed by Michelson in 1881 and repeated the experiment more carefully in collaboration with Morley in 1887 and also reiterated many times thereafter. Actually, Michelson and Morley experiments laid the experimental foundations of special relativity. Michelson was awarded the Nobel prize for his experiment in 1907.

### 2.2 The Michelson–Morley Experiment (M–M)

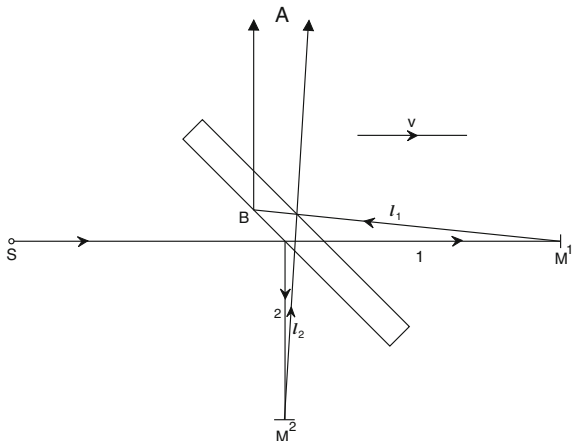
The experiment of M–M was an attempt to measure the velocity of earth through the ether. The experiment failed in the sense that it showed conclusively that ether does not exist. More generally, it demonstrated that there are no privileged observer that the velocity of light is  $c$  irrespectively of the state of motion of the observer who measures it.

*The apparatus of the experiment are:* Monochromatic source S of light, Interferometer A, one semi-silvered plate B, two mirrors,  $M_1$ ,  $M_2$  (see Fig. 2.1).

The light from a source at S falls on a semi-silvered plate B inclined at  $45^\circ$  to the direction of propagation. The plate B, due to semi-silvered, divides the beam of light into two parts namely reflected and transmitted beams. The two beams, after reflecting at mirrors  $M_1$  and  $M_2$  are brought together at A, where interference fringes are formed due to the small difference between the path lengths of two beams. Let us assume that earth (carrying the apparatus) is moving with velocity  $v$  along  $BM_1 = l_1$  relative to the privileged frame. The time ( $t_1$ ) required for the transmitted light ray



**Fig. 2.1** Experimental set-up of Michelson–Morley Experiment



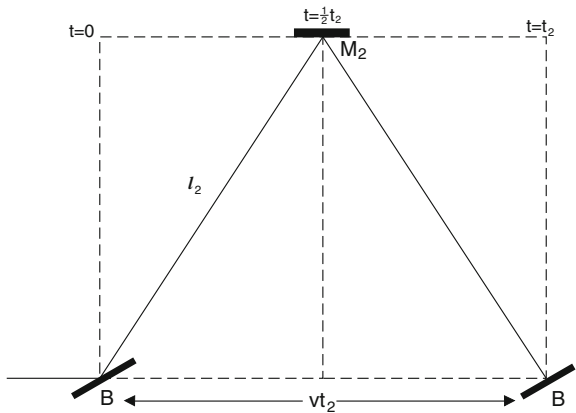
to traverse the length  $l_1$  both ways, i.e. along  $BM_1B$  is

$$t_1 = \frac{l_1}{(c - v)} + \frac{l_1}{(c + v)} = \frac{(\frac{2l_1}{c})}{(1 - \frac{v^2}{c^2})} \tag{2.1}$$

[ $c$  is the velocity of light in the ether. So, downstream speed is  $(c + v)$  and upstream speed is  $(c - v)$ ]

Now, consider the path of the reflected light ray, i.e. the path  $BM_2B$ : As the light travels from  $B$  to  $M_2$  ( $BM_2 = l_2$ ), the whole apparatus has moved a distance  $x$  along  $BM_1$ . The light has therefore travelled a distance  $\sqrt{(l_2^2 + x^2)}$ . We have seen that earth moves some distance during the same time. According to Fig. 2.2, the mirror moves the distance  $x = vt$  in time  $t$  and light moves in time  $t$  is  $ct = \sqrt{(l_2^2 + x^2)}$ .

**Fig. 2.2** Mirror moves some distance in time  $t$



Hence,

$$t = \frac{x}{v} = \frac{\sqrt{(l_2^2 + x^2)}}{c}$$

Or,

$$\sqrt{(l_2^2 + x^2)} = \frac{l_2}{\sqrt{(1 - \frac{v^2}{c^2})}} \quad (2.2)$$

On the back, it travels an equal distance, so the time for to and fro travel of the light beam ( i.e. the path  $BM_2B$  ) is

$$t_2 = \frac{(\frac{2l_2}{c})}{\sqrt{(1 - \frac{v^2}{c^2})}} \quad (2.3)$$

Interference fringes are created by the difference between  $t_1$  and  $t_2$ . The whole apparatus is now swung through  $90^\circ$ , so the roles of the two arms are interchanged. Therefore,  $l_1$  takes the place of  $l_2$  and vice versa. We have,

$$\bar{t}_1 = \frac{(\frac{2l_1}{c})}{\sqrt{(1 - \frac{v^2}{c^2})}} \quad (2.4)$$

$$\bar{t}_2 = \frac{(\frac{2l_2}{c})}{(1 - \frac{v^2}{c^2})} \quad (2.5)$$

The time differences between the two paths, before and after the apparatus is swung round, are

$$\Delta t = t_1 - t_2 = \frac{(\frac{2}{c})}{\sqrt{(1 - \frac{v^2}{c^2})}} \left( \frac{l_1}{\sqrt{(1 - \frac{v^2}{c^2})}} - l_2 \right) \quad (2.6)$$

$$\underline{\Delta t} = \underline{t}_1 - \underline{t}_2 = \frac{(\frac{2}{c})}{\sqrt{(1 - \frac{v^2}{c^2})}} \left( l_1 - \frac{l_2}{\sqrt{(1 - \frac{v^2}{c^2})}} \right) \quad (2.7)$$

The change in  $t$  brought about by rotating the apparatus is

$$\underline{\Delta t} - \Delta t = -\frac{(\frac{2}{c})(l_1 + l_2)}{\sqrt{(1 - \frac{v^2}{c^2})}} \left( \frac{1}{\sqrt{(1 - \frac{v^2}{c^2})}} - 1 \right) \quad (2.8)$$

We expand the right-hand side in a Binomial series, retaining terms up to the second power of  $(v/c)$  to yield

$$\underline{\Delta t} - \Delta t \approx -\frac{1}{c}(l_1 + l_2)\frac{v^2}{c^2} \quad (2.9)$$

Hence, the shift in the number of fringes is

$$\Delta N = \frac{\underline{\Delta t} - \Delta t}{T} = -\frac{\gamma}{c}(l_1 + l_2)\frac{v^2}{c^2} \quad (2.10)$$

[Period of one vibration is  $T = (\frac{1}{\gamma}) = (\frac{\lambda}{c})$ , where  $\gamma$  and  $\lambda$  are frequency and wave length of the light]

Earth's velocity is presumed to be  $30 \text{ km/s} = 3 \times 10^6 \text{ cm/s}$ ,  $(\frac{v}{c})^2 = 10^{-8}$  and  $(\frac{\gamma}{c}) \approx 10^4 \text{ cm}^{-1}$  for visible light, so

$$\Delta N \approx -2(l_1 + l_2) \times 10^{-4} \quad (2.11)$$

Since  $l_1$  and  $l_2$  were several metres, such a shift would be easily detectable and measurable. However, no fringe shift was observed. This experiment was repeated with multiple mirrors to increase the lengths  $l_1$  and  $l_2$ , however, the experimental conclusion was that there was no fringe shift at all. In the year 1904, Trouton and Noble done again the experiment using electromagnetic waves instead of visible light, however, no fringe shift was observed. Since the velocity of earth relative to ether cannot always be zero, therefore experiment does not depend solely on an absolute velocity of earth through an ether, but also depends on changing velocity of earth with respect to 'ether'. Such a changing motion through the ether would be easily detected and measured by the precision of experiments, if there was an ether frame.

The null result seems to rule out an ether (absolute) frame. Thus, no preferred inertial frame exists. Also, one way to interpret the null result of M–M experiment is to conclude simply that the measured speed of light is the same, i.e.  $c$  for all directions in every inertial system. In the experiment, the downstream speed and upstream speed being  $c$  rather than  $c + v$  or  $c - v$  in any frame.

*Note 1:* In 1882, Fitzgerald and Lorentz independently proposed a hypothesis to explain null results of Michelson–Morley experiment. Their hypothesis is "Every body moving with a velocity  $v$  through the ether has its length in the direction of motion contracted by a factor  $\frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$ , while the dimensions in directions perpendicular to motion remain unchanged".

Let us consider in Michelson–Morley experiment, the arms are equal, i.e.  $l_1 = l_2 = l$  at rest relative to ether and according to Fitzgerald and Lorentz hypothesis its length  $l$  when it is in motion with velocity  $v$  relative to ether will be  $l\sqrt{(1 - \frac{v^2}{c^2})}$ .

By this, they interpreted the null result of Michelson–Morley experiment as follows: From, Eqs. (2.1) and (2.3), we have

$$t_1 = \frac{\left(\frac{2l}{c}\right)}{\left(1 - \frac{v^2}{c^2}\right)} \approx \left(\frac{2l}{c}\right) \left(1 + \frac{v^2}{c^2}\right), \quad t_2 = \frac{\left(\frac{2l}{c}\right)}{\sqrt{\left(1 - \frac{v^2}{c^2}\right)}} \approx \left(\frac{2l}{c}\right) \left(1 + \frac{v^2}{2c^2}\right)$$

Using Fitzgerald and Lorentz hypothesis, we get by replacing  $l$  by  $l\sqrt{\left(1 - \frac{v^2}{c^2}\right)}$ ,

$$t_1 = \left(\frac{2l}{c}\right) \left(1 + \frac{v^2}{c^2}\right) = \left(\frac{2l}{c}\right) \sqrt{1 - \frac{v^2}{c^2}} \left(1 + \frac{v^2}{c^2}\right)$$

Expanding binomially and neglecting  $\frac{v^4}{c^4}$ , we get,

$$t_1 = \left(\frac{2l}{c}\right) \left(1 + \frac{v^2}{2c^2}\right) = t_2$$

Therefore, according to Fitzgerald and Lorentz hypothesis, the times taken by reflected and transmitted beams are the same and hence no shift of fringe is observed. However, Fitzgerald and Lorentz hypothesis fails to give correct explanation of null results of Michelson–Morley experiment when the two ears of the interferometer are not equal.

*Note 2:* If some instant, the velocity of earth were zero with respect to ether, no fringe shift would be expected. From, Eqs. (2.1) and (2.3), we have, by putting  $v = 0$ ,

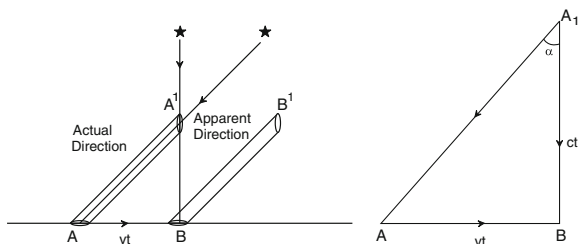
$$t_1 = \left(\frac{2l}{c}\right) = t_2$$

Then, there is no relative velocity between earth and the ether, i.e. earth drags the ether with the same velocity as earth during its motion. However, this explanation is in contradiction with various experiments like Bradley's result of aberration, etc.

### 2.3 Phenomena of Aberration: Bradley's Observation

In the year 1727, Bradley observed that the star,  $\gamma$  Draconis revolves in nearly circular orbit. He found that angular diameter of this circular orbit is about 41 s of arc and other stars appeared to move in elliptical orbit have almost same period. This phenomena of apparent displacement is known as aberration. This can be explained as follows. Let a light ray comes from a star at zenith. An observer together with earth has a velocity  $v$  will see the direction of light ray will not be vertical. This is

**Fig. 2.3** Telescope will have to tilt to see the star



due to relative motion of the observer and light ray. To see the star, the telescope will have to tilt, i.e. direction slightly away from the vertical as shown in the Fig. 2.3.

The angle  $\alpha$  between the actual and apparent directions of the star is called aberration.

Let  $v$  be the velocity of earth in the direction AB. The light ray enters the telescope tube at  $A^1$  and reaches the observer's eye at B in time  $\Delta t$ . As the light reaches from  $A^1$  to B in the mean time eyepiece moves from A to B. Hence,  $A^1B = c\Delta t$ ,  $AB = v\Delta t$ . The tilt of the telescope  $\alpha$  is given by (see Fig. 2.3)

$$\tan \alpha = \frac{v}{c}$$

or

$$\alpha = \tan^{-1} \frac{v}{c}$$

Using the velocities of earth and light as 30 km/s and  $3 \times 10^5$  km/s, respectively, one gets,

$$\alpha = \tan^{-1} \frac{v}{c} = \tan^{-1} \frac{3 \times 10^4}{3 \times 10^8} = \tan^{-1}(10^{-4}) = 20.5 \text{ arc s}$$

Since earth's motion is nearly circular, therefore every six months, the direction of aberration will reverse and telescope axis is tracing a cone of aberration during the year. Hence the angular diameter of the cone would be  $2\alpha = 41$  arc s. In 1727, Bradley observed this motion and found the angular diameter was 41 arc s for the star  $\gamma$  Draconis. This experiment apparently showed the absolute velocity of earth is 30 km/s.

The outcome of this experiment is that ether is not dragged around with earth. If it were, then the phenomena of aberration would not be happened.

## 2.4 Fizeau's Experiment

In 1817, Fresnel proposed that light would be partially dragged along by a moving medium. Using ether hypothesis, he provided the formula for the effect of moving medium on the velocity of light. In 1951, Fizeau confirmed this effect experimentally.

Figure 2.4 shows the Fizeau's experimental arrangement. Fizeau used water as the medium inside the tubes. Water flows through the tubes with velocity  $u$ .

The S is the source of light. The light ray from S falls on a semi-silver plate inclined  $45^\circ$  from the horizontal is divided into two parts. One is reflected in the direction  $M_1M_2$  and other part is transmitted in the direction  $M_1D$ . Here, one beam of light is entering towards the velocity of water and other opposite to the motion of water. One can note that reflected beam travels the path  $M_1M_2CDM_1$  in the direction of motion of water and the transmitted beam travels the path  $M_1EDCM_2M_1$  opposite to the motion of water. Both beams finally enter the telescope.

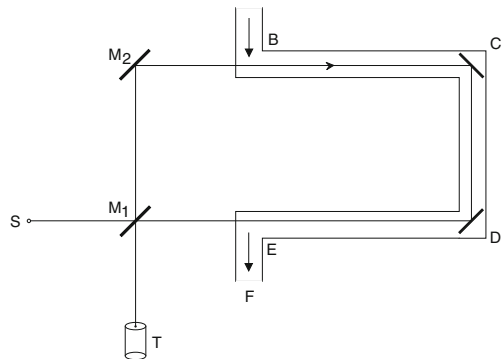
Difference in time taken by both the beams travel equal distance  $d$  due to the motion of the water and velocity  $v$  of earth. It was assumed that space is filled uniformly with the ether and ether to be partially dragged. Now, the time difference is

$$\Delta t = \frac{d}{\frac{c}{\mu} + \left(\frac{1}{\mu^2}\right)v + u \left(1 - \frac{1}{\mu^2}\right)} - \frac{d}{\frac{c}{\mu} + \left(\frac{1}{\mu^2}\right)v - u \left(1 - \frac{1}{\mu^2}\right)}$$

It was assumed that ether to be partially dragged and as a result component of the velocity in the tube  $\left(\frac{1}{\mu^2}\right)v$ . It follows also that if a material body is moving with velocity  $u$  through ether, the velocity of ether will be  $u \left(1 - \frac{1}{\mu^2}\right)$ . Here,  $\left(\frac{1}{\mu^2}\right) =$  constant. Neglecting higher order terms, we get the time difference as

$$\Delta t = \frac{2ud\mu^2}{c^2} \left(1 - \frac{1}{\mu^2}\right)$$

**Fig. 2.4** Experimental set-up of Fizeau's experiment



The phase difference is given by

$$\Delta f = \frac{\frac{2ud\mu^2}{c^2} \left(1 - \frac{1}{\mu^2}\right)}{T} = \frac{2udn\mu^2}{c^2} \left(1 - \frac{1}{\mu^2}\right)$$

where,  $n$  be the frequency of the oscillation and  $T$  is time period of one oscillation ( $nT = 1$ ). Fizeau observed the shift in the fringes and found exactly the same as given by the above equation. This result also confirms the result that the velocity of the light in any medium of refractive index  $\mu$  is not simply  $\frac{c}{\mu}$  but  $\frac{c}{\mu} \pm \left(1 - \frac{1}{\mu^2}\right)u$ ,  $u$  is the velocity of the medium. The factor  $\left(1 - \frac{1}{\mu^2}\right)$  is called *Fresnel’s ether dragging coefficient*. Thus, Fresnel’s formula for ether drag was verified by Fizeau.

### 2.5 The Relativistic Concept of Space and Time

The null results of Michelson–Morley Experiment ruled out the absolute concept of space and laid Einstein to formulate the new concept of space and time.

Consider the following two contradictory statements: [1] In classical mechanics, the velocity of any motion has different values for observers moving relative to each other. [2] The null results of Michelson–Morley experiment concludes that the velocity of light is not affected by the motion of the frame of reference. The genius Einstein applied his intuition and realized that this contradiction comes from the imperfection of the classical ideas about measuring space and time. He ruled out the Newton’s concept of absolute time by criticizing our preconceived ideas about simultaneity. He commented ‘time is indeed doubtful’. His intuition may be supported as follows:



Suppose an observer sitting at O where O is the middle point of the straight line XY. Suppose that two events occurring at X and Y which are fixed in an inertial frame S and the two clocks giving absolute time are kept at X and Y. In the inertial frame S, the events occurring at X and Y are said to be simultaneous, if the clocks placed at X and Y indicate the same time when the events occur. Let a light signal is giving at X and Y simultaneously. Suppose XOY is moving with  $v$  velocity in the direction XY. Then, one can show that these light signals will not reach simultaneously the observer at O which is moving with velocity  $v$ . Here, the moving O seems to lie in another inertial frame. The light signal takes  $t_1$  time to reach from Y to O as  $\frac{OY}{(c-v)}$  and the time  $t_2$  taken by the light signal from X to O is  $\frac{OX}{(c+v)}$ .

[ $c$  is the velocity of light in all direction. So, according to Newtonian mechanics, downstream speed is  $(c + v)$  and upstream speed is  $(c - v)$ ]

One can note that

$$t_1 = \frac{OY}{(c - v)} \neq \frac{OX}{(c + v)} = t_2, \text{ as } c + v \neq c - v, \quad OX = OY$$

Therefore, time is not absolute. It varies from one inertial frame to another inertial frame, i.e.  $t_1 \neq t_2$ . In other words, every reference frame has its own particular time. Therefore, concept of absolute time was ruled out by Einstein.

The space coordinate is transformed due to Galilean Transformation as  $x^1 = x - vt$ . Since, time depends on frame of reference, therefore, distance between two points will also depend on the frame of reference. In other words, distance between two points is not absolute. Hence, Einstein also ruled out the concept of absolute space.

Einstein proposed his new relativistic concepts of space and time based on the fact that observers who are moving relative to each other measure the speed of light to be the same. This new concept of space and time is not only valid for mechanical phenomenon but also for all optical and electromagnetic phenomenon and known as theory of relativity. To develop these new concepts of space and time, Einstein replaced Galilean Transformation equations by a new type of transformation equations called *Lorentz transformation*, which was based on the invariant character of the speed of light.

The theory of relativity having new concept of space and time is divided into two parts. [1] *Special theory of relativity (STR)* [2] *General theory of relativity (GTR)*.

The STR deals with inertial systems, i.e. the systems which move in uniform rectilinear motion relative to one another. GTR deals with non-inertial system, i.e. it is an extension of inertial system to accelerated systems, i.e. the systems moving with accelerated velocity relative to one another. The GTR is applied to the area of gravitational field and based on which laws of gravitation can be explained in a more perfect manner than given by Newton.



# Chapter 3

## Lorentz Transformations

### 3.1 Postulates of Special Theory of Relativity

To develop the special theory of relativity, Einstein used two fundamental postulates.

- [1] *The principle of relativity or equivalence:* The laws of physics are the same in all inertial systems. No preferred inertial system exists.
- [2] *The principle of constancy of the speed of light:* The velocity of light in an empty space is a constant, independent not only of the direction of propagation but also of the relative velocity between the source of the light and observer.

Note that the first principle has not changed its Newtonian form except the Galilean Transformation equations is replaced by a new type of transformation equations known as Lorentz transformation. The second principle is not consistent with Galilean Transformation and it indicates a clear division between Newtonian theory and Einstein theory. Therefore, we have to replace Galilean Transformation equations by Lorentz transformation equations which fulfil the above principles.

### 3.2 Lorentz Transformations

The principle of relativity states that the laws of nature are invariant under a particular group of space-time coordinate transformations. Newton's laws of motion are invariant under G.T. but Maxwell's equations are not, and that Einstein resolved this conflict by replacing Galilean transformation with Lorentz transformation.

#### 3.2.1 Lorentz Transformation Between Two Inertial Frames of Reference (Non-axiomatic Approach)

Let us suppose that an event occurred at a point P in an inertial frame S specifying the coordinates  $x, y, z$  and  $t$ . In another inertial frame  $S_1$ , which is moving with a

constant velocity  $v$  along  $x$ -axis, the same event will be specified by  $x_1, y_1, z_1$  and  $t_1$ . Let the observers in the two systems be situated at the origin  $O$  and  $O_1$  of  $S$  and  $S_1$  frames. The rectangular axes  $X_1, Y_1, Z_1$  are parallel to  $X, Y, Z$ , respectively. Therefore, the coordinates  $x_1, y_1, z_1$  and  $t_1$  are related to  $x, y, z$  and  $t$  by a functional relationship. Hence, we must have

$$x_1 = x_1(x, y, z, t); \quad y_1 = y_1(x, y, z, t); \quad z_1 = z_1(x, y, z, t); \quad t_1 = t_1(x, y, z, t) \quad (3.1)$$

Since  $S$  and  $S_1$  are both inertial frames, therefore laws of mechanics remain the same in both the frames. As a result, a particle is moving with a constant velocity along a straight line relative to an observer in  $S$ , then an observer in  $S_1$  will see also that that particle is moving with uniform rectilinear motion relative to him. Under these conditions, the transformation Eq. (3.1) must be linear and assume the most general form as

$$\begin{aligned} x_1 &= a_{11}x + a_{12}y + a_{13}z + a_{14}t + a_{15} \\ y_1 &= a_{21}x + a_{22}y + a_{23}z + a_{24}t + a_{25} \\ z_1 &= a_{31}x + a_{32}y + a_{33}z + a_{34}t + a_{35} \\ t_1 &= a_{41}x + a_{42}y + a_{43}z + a_{44}t + a_{45} \end{aligned} \quad (3.2)$$

[where  $a_{ij}$  are constants]

Note that we have chosen the relative velocity  $v$  of the  $S$  and  $S_1$  frames to be along a common  $x - x_1$  axis to keep corresponding planes parallel. This assumption simplifies the algebra without any loss of generality. Also at the instant, when the origins  $O$  ( $x = y = z = 0$ ) and  $O_1$  ( $x_1 = y_1 = z_1 = 0$ ) coincide, the clocks read  $t = 0$  and  $t_1 = 0$ , respectively. These imply,  $a_{15} = a_{25} = a_{35} = a_{45} = 0$ . Now, regarding the remaining 16 coefficients, it is expected that their values will depend on the relative velocity  $v$  of the two inertial frames. Now, we use the postulates of the STR (1. The principle of relativity 2. The principle of constancy of the speed of light) to determine the values of the 16 coefficients. Here, we note that  $x$ -axis coincides continuously with  $x_1$ -axis. Therefore, points on  $x$ -axis and  $x_1$ -axis are characterized by  $y = z = 0$  and  $y_1 = z_1 = 0$ . As a result, the transformation formulas for  $y$  and  $z$  should be of the form

$$y_1 = a_{22}y + a_{23}z \quad \text{and} \quad z_1 = a_{32}y + a_{33}z \quad (3.3)$$

That is, coefficients  $a_{21}, a_{24}, a_{31}$  and  $a_{34}$  must be zero. In a similar manner, the  $x - y$  and  $x - z$  planes (which are characterized by  $z = 0$  and  $y = 0$  respectively) should transform over to the  $x_1 - y_1$  and  $x_1 - z_1$  planes (which are characterized by  $z_1 = 0$  and  $y_1 = 0$ , respectively). Therefore, in Eq. (3.3), the coefficients  $a_{23}$  and  $a_{32}$  are zero to yield

$$y_1 = a_{22}y \quad \text{and} \quad z_1 = a_{33}z \quad (3.4)$$

Now, we try to find the coefficients  $a_{22}$  and  $a_{33}$ . Suppose we put a rod of unit length (measured by observer in  $S$  frame) along  $y$ -axis. According to Eq. (3.4), the observer in  $S_1$  frame measures the rod's length as  $a_{22}$ . Now, we put a rod of unit length (measured by observer in  $S_1$  frame) along  $y_1$  axis. The observer in  $S$  frame measures this rod as  $\frac{1}{a_{22}}$ . The first postulate of STR requires that these measurements must be same. Hence, we must have  $a_{22} = \frac{1}{a_{22}}$ , i.e.  $a_{22} = 1$ . Using similar argument, one can determine  $a_{33} = 1$ . Therefore, we get

$$y_1 = y \quad \text{and} \quad z_1 = z \quad (3.5)$$

For the reason of symmetry, we assume that  $t_1$  does not depend on  $y$  and  $z$ . Hence,  $a_{42} = 0$  and  $a_{43} = 0$ . Thus,  $t_1$  takes the form as

$$t_1 = a_{41}x + a_{44}t \quad (3.6)$$

We know points which are at rest relative to  $S_1$  frame will move with velocity  $v$  relative to  $S$  frame in  $x$ -direction. Therefore, the statement  $x_1 = 0$  must be identical to the statement  $x = vt$ . As a result, the correct transformation equation for  $x_1$  will be

$$x_1 = a_{11}(x - vt) \quad (3.7)$$

Still we will have to determine the three coefficients, namely,  $a_{11}$ ,  $a_{41}$ ,  $a_{44}$ . These can be determined by using the principle of the constancy of the velocity of light. Let us assume that a light pulse produced at the time  $t = 0$  and will propagate in all direction with same velocity  $c$ . Initially, it coincides with the origin of  $S_1$ . Due to the second postulate, the wave propagates with a speed  $c$  in all directions in each inertial frame. Hence, equations of motion of this wave with respect to both frames  $S$  and  $S_1$  are given as

$$x^2 + y^2 + z^2 = c^2t^2 \quad (3.8)$$

$$x_1^2 + y_1^2 + z_1^2 = c^2t_1^2 \quad (3.9)$$

Now, we substitute the transformation equations for  $x_1$ ,  $y_1$ ,  $z_1$  and  $t_1$  into Eq. (3.9) and get

$$[a_{11}^2 - c^2a_{41}^2]x^2 + y^2 + z^2 - 2[va_{11}^2 + c^2a_{41}a_{44}]xt = [c^2a_{44}^2 - v^2a_{11}^2]t^2 \quad (3.10)$$

But this equation and Eq. (3.8) are identical, so we must have

$$[c^2a_{44}^2 - v^2a_{11}^2] = c^2; \quad [a_{11}^2 - c^2a_{41}^2] = 1; \quad [va_{11}^2 + c^2a_{41}a_{44}] = 0$$

Solving these equations, we get

$$a_{44} = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}; \quad a_{11} = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}; \quad a_{41} = -\frac{v}{c^2 \sqrt{1 - \frac{v^2}{c^2}}}.$$

So finally we get, the Lorentz Transformation as

$$x_1 = a(x - vt); \quad y_1 = y; \quad z_1 = z; \quad t_1 = a \left[ t - \left( \frac{vx}{c^2} \right) \right]. \quad (3.11)$$

where

$$a = \frac{1}{\sqrt{1 - \beta^2}} \quad \text{and} \quad \beta = \frac{v}{c}$$

The inverse relations can be obtained by solving  $x, y, z, t$  in terms of  $x_1, y_1, z_1, t_1$  as

$$x = a(x_1 + vt_1); \quad y = y_1; \quad z = z_1; \quad t = a \left[ t_1 + \left( \frac{vx_1}{c^2} \right) \right] \quad (3.12)$$

Actually, Eq.(3.12) are obtained from Eq.(3.11) by replacing  $v$  by  $-v$  and subscripted quantities by unsubscripted ones and vice versa. In particular, for speed small compared to  $c$ , i.e.  $\frac{v}{c} \ll 1$ , the Lorentz Transformations reduce to Galilean Transformations. Thus,

$$x_1 = x - vt; \quad y_1 = y; \quad z_1 = z; \quad t_1 = t$$

**Note 1:** If the velocity of light would be infinity, then Lorentz Transformation will never exist and we have only Galilean Transformation.

*Example 3.1* Show that the expressions

$$[i] x^2 + y^2 + z^2 - c^2 t^2$$

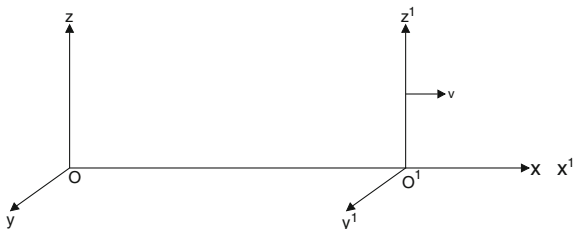
$$[ii] dx^2 + dy^2 + dz^2 - c^2 dt^2 \equiv dx_1^2 + dx_2^2 + dx_3^2 + dx_4^2$$

where,  $x_1 = x, \quad x_2 = y, \quad x_3 = z, \quad x_4 = ict$ , are invariant under Lorentz Transformations.

### 3.2.2 Axiomatic Derivation of Lorentz Transformation

Let there be two reference system  $S, S_1$ .  $S_1$  is moving with velocity  $v$  with respect to  $S$  along  $x$ -axis (see Fig. 3.1). At time,  $t = 0$ , two origins  $O$  and  $O_1$  coincide. If  $(x, y, z, t)$  be any event in  $S$  frame, then this event is represented by  $(x_1, y_1, z_1, t_1)$  in  $S_1$  frame.

**Fig. 3.1** The frame  $S_1$  is moving with velocity  $v$  along  $x$ -direction



To find the relation between the coordinates of  $S$  and  $S_1$  frames, we use the following **axioms**:

- [i] The direct transformation (i.e.  $S$  to  $S_1$ ) and reverse transformation (i.e.  $S_1$  to  $S$ ) should be symmetrical with respect to both systems, i.e. one should be derivable from the other by replacing  $v$  by  $-v$  and subscripted quantities by unsubscripted ones and vice versa.
- [ii] If  $(x, y, z, t)$  are finite, then  $(x_1, y_1, z_1, t_1)$  are also finite.
- [iii] In the limit  $v \rightarrow 0$ , the transformations should be reduced to identity transformations, i.e.  $x_1 = x, y_1 = y, z_1 = z, t_1 = t$ .
- [iv] In the limit  $c \rightarrow \infty$ , the Lorentz Transformations reduce to Galilean Transformations. Thus,

$$x_1 = x - vt; \quad y_1 = y; \quad z_1 = z; \quad t_1 = t$$

- [v] The velocity of light is same in two reference frames, i.e.  $c_1=c$ .

Now, we search the possible transformations between the coordinates of two frames. If the transformation functions are quadratic or higher, then the inverse transformations should be irrational functions. Thus, transformations cannot be quadratic or higher. Another choice is linear fractional transformation function,  $x_1 = \frac{ax+b}{cx+d}$ . Then, the inverse transformation gives,  $x = \frac{b-dx_1}{cx_1-a}$ . But this function for  $x_1 = \frac{a}{c}$  becomes infinite and this violates the condition (ii). Hence, linear transformation is the only acceptable one. Since the relative motion between the reference frames is along the  $x$ -axis, we get,

$$y_1 = y; \quad z_1 = z \tag{3.13}$$

Now, we choose the linear transformations for  $x$  and  $t$  as

$$x_1 = \lambda x + \mu t + A \tag{3.14}$$

$$t_1 = \psi x + \phi t + B \tag{3.15}$$

Since origins coincide initially, i.e.  $x = 0, t = 0, x_1 = 0, t_1 = 0$ , so the constants present in Eqs. (3.14) and (3.15) namely,  $A = B = 0$  should be zero. Since the point  $O_1$  (its coordinate  $x_1 = 0$ ) moves with a velocity  $v$  relative to  $O$ , we get

$$x = vt.$$

Now, we put together  $x_1 = 0$  and  $x = vt$  in (3.14), we get,

$$\lambda v + \mu = 0 \quad (3.16)$$

Solving (3.14) and (3.15), we get,

$$x = \frac{\phi x_1 - \mu t_1}{\lambda \phi - \mu \psi} \quad (3.17)$$

$$t = \frac{\lambda t_1 - \psi x_1}{\lambda \phi - \mu \psi} \quad (3.18)$$

From condition (i), we can write the inverse transformation by replacing  $v$  by  $-v$  and subscripted quantities by unsubscripted ones as,

$$x = \lambda x_1 - \mu t_1 \quad (3.19)$$

$$t = -\psi x_1 + \phi t_1 \quad (3.20)$$

Since the coefficients  $\mu$  and  $\psi$  are related with  $x$  and  $t$  coordinates and therefore changes its sign when  $v \rightarrow -v$  (one can also take  $\lambda$  and  $\phi$ ).

Since Eqs. (3.17) and (3.19) are identical, therefore, comparing the coefficients, we have,

$$\lambda = \frac{\phi}{\lambda \phi - \mu \psi}; \quad -\mu = \frac{-\mu}{\lambda \phi - \mu \psi}$$

These imply,

$$\lambda = \phi \quad (3.21)$$

$$\lambda \phi - \mu \psi = 1 \quad (3.22)$$

From Eqs. (3.14) and (3.15), we have,

$$\frac{x_1}{t_1} = \frac{\lambda(\frac{x}{t}) + \mu}{\psi(\frac{x}{t}) + \phi} \quad (3.23)$$

Let a light signal emitted from the origin in S frame, then  $\frac{x}{t} = c$ . But condition (v) implies  $\frac{x_1}{t_1} = c$  also. Hence, from Eq. (3.23), we have,

$$c = \frac{\lambda c + \mu}{\psi c + \phi} \quad (3.24)$$

From Eqs. (3.16), (3.21), (3.22) and (3.24), we get the following solutions as

$$\psi = -\frac{\lambda v}{c^2} \quad (3.25)$$

$$\lambda = \pm \frac{1}{\sqrt{1 - \left(\frac{v^2}{c^2}\right)}} \quad (3.26)$$

Condition (ii) indicates that we must get back  $t_1 = t$  when  $c \rightarrow 0$ . Therefore, we should select + sign.

Thus, finally we the Lorentz Transformation as,

$$x_1 = a(x - vt); \quad y_1 = y; \quad z_1 = z; \quad t_1 = a \left[ t - \left( \frac{vx}{c^2} \right) \right] \quad (3.27)$$

where

$$a = \frac{1}{\sqrt{1 - \beta^2}} \quad \text{and} \quad \beta = \frac{v}{c}$$

### 3.2.3 Lorentz Transformation Based on the Postulates of Special Theory of Relativity

Let us suppose that an event occurred at a point P in an inertial frame S specifying the coordinates  $x, y, z$  and  $t$ . In another inertial frame  $S_1$ , which is moving with a constant velocity  $v$  along  $x$  axis, the same event will be specified by  $x_1, y_1, z_1$  and  $t_1$ . Let the observers in the two systems be situated at the origin O and  $O_1$  of S and  $S_1$  frames. The Galilean transformation between  $x$  and  $x^1$  is  $x^1 = x - vt$ .

We know Galilean transformation fails, so a reasonable guess about the nature of the correct relationship is

$$x^1 = k(x - vt) \quad (3.28)$$

Here  $k$  is the factor that does not depend upon either  $x$  or  $t$  but may be a function of  $v$ .

The first postulate of Special Theory of Relativity: All laws of physics are the same on all inertial frames of reference. Therefore, Eq. (3.28) must have the same form on both S and  $S^1$ . We need only change the sign of  $v$  (in order to take into account the difference of the direction of relative motion) to write the corresponding equation for  $x$  in terms of  $x^1$  and  $t^1$ . Therefore,

$$x = k(x^1 + vt^1) \quad (3.29)$$

It is obtained that  $y, y^1$  and  $z, z^1$  are equal to each other as they are perpendicular to velocity  $v$ . Therefore,

$$y^1 = y \quad (3.30)$$

and

$$z^1 = z. \quad (3.31)$$

Remember  $t^1$  may not be equal to  $t$ . Putting the value of  $x^1$  in (3.29) we get,

$$x = k[k(x - vt) + vt^1] = k^2(x - vt) + kv t^1$$

or,

$$t^1 = kt + \left( \frac{1 - k^2}{kv} \right) x \quad (3.32)$$

The second postulates of relativity gives us a way to evaluate  $k$ . At the instant  $t = 0$ , the origins of the two frames of reference S and  $S^1$  are on the same place. Therefore, according to our initial condition  $t^1 = 0$  when  $t = 0$ . Suppose that a light signal from a light source emits a pulse at the common origin of S and  $S^1$  at time  $t = t^1 = 0$  and the observers in each system measures the speed of the propagate of light wave. Both observes must find the same speed  $c$ , which means that in the S frame,

$$x = ct \quad (3.33)$$

and in  $S^1$  frame,

$$x^1 = ct^1 \quad (3.34)$$

Now from (3.33), we get

$$k(x - vt) = c \left[ kt + \left( \frac{1 - k^2}{kv} \right) x \right]$$

Therefore,

$$x = \frac{ct \left( 1 + \frac{v}{c} \right)}{\left[ 1 - \left( \frac{1 - k^2}{k^2} \right) \frac{c}{v} \right]}$$

i.e.

$$ct = \frac{ct \left( 1 + \frac{v}{c} \right)}{1 - \left( \frac{1}{k^2} - 1 \right) \frac{c}{v}} \quad [\text{since from (3.33), } x = ct]$$



Hence,

$$k = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (3.35)$$

Using the value of  $k$  from (3.35) in Eqs. (3.28) and (3.32), we get the values of  $x^1$  and  $t^1$  as

$$x^1 = \frac{(x - vt)}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (3.36)$$

$$t^1 = \frac{t - \frac{vx}{c^2}}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (3.37)$$

Thus Eqs. (3.30), (3.31), (3.36) and (3.37) give the required Lorentz Transformations.

### 3.3 The General Lorentz Transformations

Now, we search the Lorentz Transformations for an arbitrary direction of the relative velocity  $\vec{v}$  between  $S$  and  $S_1$  frames. Let us consider the position vector  $\vec{r}$  of a particle separate into two components, one parallel and another perpendicular to  $\vec{v}$ . That is,

$$\vec{r} = \vec{r}_\perp + \vec{r}_\parallel \quad (3.38)$$

$$[\text{here, } \vec{r}_\perp \perp \vec{v} \text{ and } \vec{r}_\parallel \parallel \vec{v} ]$$

The component  $\vec{r}_\parallel$  assumes the following value

$$\vec{r}_\parallel = \frac{(\vec{r} \cdot \vec{v}) \vec{v}}{v^2} \quad (3.39)$$

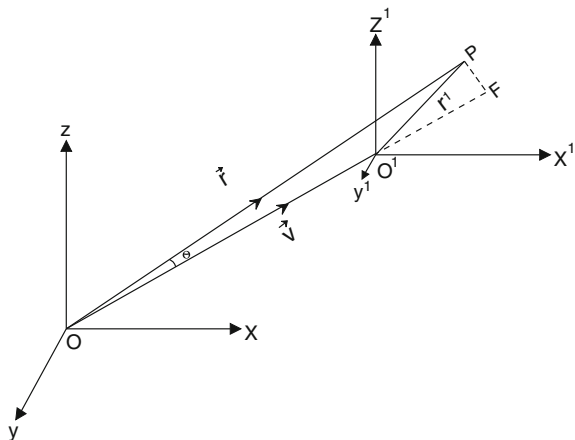
[from Fig. 3.2, we have  $\vec{r} \cdot \vec{v} = r v \cos \theta$   $|\vec{OF}| = |\vec{OP}| \cos \theta$  ;  $|\vec{r}| = r$  ,  $|\vec{v}| = v$  ;  $\vec{OF} = \vec{r}_\parallel = |\vec{OF}| \hat{n} = (\frac{\vec{r} \cdot \vec{v}}{v})(\frac{\vec{v}}{v}) = (\frac{\vec{r} \cdot \vec{v}}{v^2}) \vec{v}$  ]

The other component  $\vec{r}_\perp$  takes the value

$$\vec{r}_\perp = \vec{r} - \vec{r}_\parallel = \vec{r} - \frac{(\vec{r} \cdot \vec{v}) \vec{v}}{v^2} \quad (3.40)$$

Since coordinates change in the direction of motion, therefore, the component  $\vec{r}_\parallel$  will transform like the x-coordinate. Hence

**Fig. 3.2** The frame  $S_1$  is moving with velocity  $v$  in arbitrary direction



$$\vec{\mathbf{r}}'_{\parallel} = a(\vec{\mathbf{r}}_{\parallel} - \vec{\mathbf{v}}t) \quad (3.41)$$

where

$$a = \frac{1}{\sqrt{1 - \beta^2}} \quad \text{and} \quad \beta = \frac{v}{c}$$

The coordinates remain invariable in a direction perpendicular to motion as a result component  $\vec{\mathbf{r}}_{\perp}$  will transform like  $y$  or  $z$  coordinates. Hence

$$\vec{\mathbf{r}}'_{\perp} = \vec{\mathbf{r}}_{\perp} \quad (3.42)$$

$$\begin{aligned} \text{Now, } \vec{\mathbf{r}}' &= \vec{\mathbf{r}}'_{\parallel} + \vec{\mathbf{r}}'_{\perp} = a(\vec{\mathbf{r}}_{\parallel} - vt) + \vec{\mathbf{r}}_{\perp} \\ &= a \left( \frac{(\vec{\mathbf{r}} \cdot \vec{\mathbf{v}}) \vec{\mathbf{v}}}{v^2} - \vec{\mathbf{v}}t \right) + \left( \vec{\mathbf{r}} - \frac{(\vec{\mathbf{r}} \cdot \vec{\mathbf{v}}) \vec{\mathbf{v}}}{v^2} \right) \end{aligned}$$

Rearranging the above expression, we get,

$$\vec{\mathbf{r}}' = \vec{\mathbf{r}} - \vec{\mathbf{v}}t + (a - 1) \left( \frac{\vec{\mathbf{v}}}{v^2} \right) [(\vec{\mathbf{r}} \cdot \vec{\mathbf{v}}) - v^2t] \quad (3.43)$$

Similarly,

$$t' = a \left[ t - \frac{(\vec{\mathbf{r}} \cdot \vec{\mathbf{v}})}{c^2} \right] \quad (3.44)$$

The inverse transformations are

$$\vec{\mathbf{r}} = \vec{\mathbf{r}}' + \vec{\mathbf{v}}t' + (a-1)\left(\frac{\vec{\mathbf{v}}}{v^2}\right)[(\vec{\mathbf{r}}' \cdot \vec{\mathbf{v}}) + v^2t']$$

$$t = a \left[ t' + \frac{(\vec{\mathbf{r}}' \cdot \vec{\mathbf{v}})}{c^2} \right]$$

When  $v \ll c$ ,  $(a-1) \rightarrow 0$ , then Eqs.(3.43) and(3.44) reduce to

$$\vec{\mathbf{r}}' = \vec{\mathbf{r}} - \vec{\mathbf{v}}t$$

$$t' = t$$

This is the generalized form of the Galilean Transformation.

One can write the generalized Lorentz Transformation equations in component form:

$$x^1 = x\left\{1 + (a-1)\frac{v_x^2}{v^2}\right\} + y(a-1)\frac{v_x v_y}{v^2} + z(a-1)\frac{v_x v_z}{v^2} - av_x t \quad (3.45)$$

$$y^1 = x(a-1)\frac{v_x v_y}{v^2} + y\left\{1 + (a-1)\frac{v_y^2}{v^2}\right\} + z(a-1)\frac{v_y v_z}{v^2} - av_y t \quad (3.46)$$

$$z^1 = x(a-1)\frac{v_x v_z}{v^2} + y(a-1)\frac{v_y v_z}{v^2} + z\left\{1 + (a-1)\frac{v_z^2}{v^2}\right\} - av_z t \quad (3.47)$$

$$t^1 = a \left( t - \frac{xv_x + yv_y + zv_z}{c^2} \right) \quad (3.48)$$

[here,  $v^2 = v_x^2 + v_y^2 + v_z^2$ ]

If the relative velocity  $v$  between two frames to be parallel to  $x$ -axis, then,  $v_x = v$  and  $v_y = v_z = 0$ . Therefore the above transformation equations reduce to

$$x_1 = a(x - v_x t); \quad y_1 = y; \quad z_1 = z; \quad t_1 = a \left[ t - \left( \frac{xv_x}{c^2} \right) \right]$$

If the relative velocity  $v$  between two frames to be parallel to  $y$  axis, then,  $v_y = v$  and  $v_x = v_z = 0$ . Therefore the above transformation equations reduce to

$$y_1 = a(y - v_y t); \quad x_1 = x; \quad z_1 = z; \quad t_1 = a \left[ t - \left( \frac{yv_y}{c^2} \right) \right]$$

### 3.4 Thomas Precession

Consider three inertial frames  $S$ ,  $S^1$  and  $S^2$ . Let  $S^1$  is moving with a constant velocity  $v_1$  along  $x$ -axis of  $S$  and  $S^2$  is moving with a constant velocity  $v_2$  along  $y$ -axis of  $S^1$ . Then, the relative velocity between  $S$  and  $S^2$  will not be parallel to any of the axes.

Thus, there exists some rotation of axes of  $S^2$  relative to those of S. The amount of rotation of axes is known as **Thomas Precession**.

The Transformation equations between S and  $S^1$  are given by

$$x^1 = a_1(x - v_1 t), \quad y^1 = y, \quad z^1 = z, \quad t^1 = a_1 \left( t - \frac{v_1 x}{c^2} \right) \quad (3.49)$$

where,  $a_1 = \frac{1}{\sqrt{1 - \frac{v_1^2}{c^2}}}$ .

Similarly, the transformation between  $S^1$  and  $S^2$  are given by

$$x^2 = x^1, \quad y^2 = a_2(y^1 - v_2 t^1), \quad z^2 = z^1, \quad t^2 = a_2 \left( t^1 - \frac{v_2 y^1}{c^2} \right) \quad (3.50)$$

where,  $a_2 = \frac{1}{\sqrt{1 - \frac{v_2^2}{c^2}}}$ .

Using (3.49) in to (3.50), we get the required transformation between S and  $S^2$  as

$$\begin{aligned} x^2 &= a_1(x - v_1 t), \quad y^2 = a_2 \left( y + \frac{v_1 v_2 a_1}{c^2} - v_2 a_1 t \right), \\ z^2 &= z^1, \quad t^2 = a_2 \left( a_1 t - \frac{v_1 a_1 x}{c^2} - \frac{v_2 y}{c^2} \right) \end{aligned} \quad (3.51)$$

Let us assume that the relative velocity vector  $\vec{u}$  between S and  $S^2$  frames makes angles  $90^\circ - \theta_1$  and  $90^\circ - \theta_2$  with x and  $x^2$  axes. This means normal to relative velocity vector  $\vec{u}$  makes  $\theta_1$  and  $\theta_2$  angles with x and  $x^2$  axes. Therefore, Thomas Precession is the angle  $\theta_2 - \theta_1$  (see Fig. 3.3). Due to Lorentz Transformation, the components of the position vectors  $\vec{r}$  and  $\vec{r}^2$  remain unchanged in a direction perpendicular to the relative velocity, therefore,

$$\begin{aligned} x \cos \theta_1 - y \sin \theta_1 &= x^2 \cos \theta_2 - y^2 \sin \theta_2 \\ &= a_1 x \left( \cos \theta_2 - \frac{v_1 v_2 a_2 \sin \theta_2}{c^2} \right) - a_2 y \sin \theta_2 \\ &\quad - a_1 t (v_1 \cos \theta_2 - v_2 a_2 \sin \theta_2) \end{aligned} \quad (3.52)$$

[using the values of  $x^2$  and  $y^2$  from (3.52)]

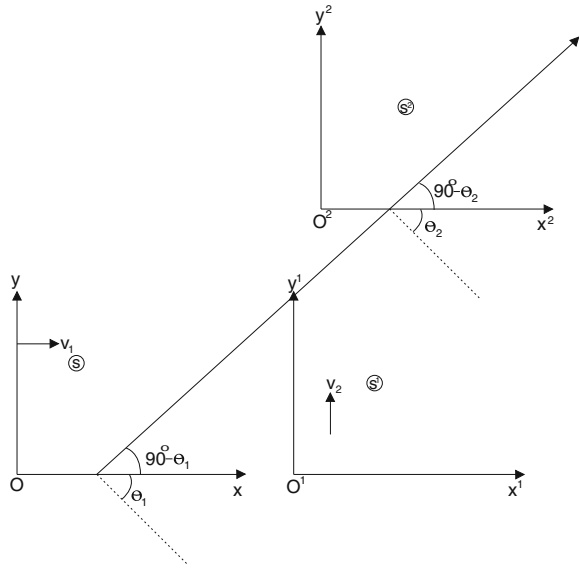
Definitely, coefficient of t should vanish. Hence we get,

$$\tan \theta_2 = \frac{v_1}{v_2 a_2} \quad (3.53)$$

This gives

$$\sin \theta_2 = \frac{v_1}{D a_2}, \quad \cos \theta_2 = \frac{v_2}{D} \quad (3.54)$$

Fig. 3.3 Thomas precession



where,  $D = v_1^2 + v_2^2 - \frac{v_1^2 v_2^2}{c^2}$ .  
 From (3.54), we have,

$$x \cos \theta_1 - y \sin \theta_1 = \frac{v_2 x}{D a_1} - \frac{v_1 y}{D}$$

This gives,

$$\tan \theta_1 = \frac{v_1 a_1}{v_2} \tag{3.55}$$

Now finally, we get Thomas Precession as

$$\tan(\theta_2 - \theta_1) = \frac{\tan \theta_2 - \tan \theta_1}{1 + \tan \theta_2 \tan \theta_1} = -\frac{v_1 v_2 (a_1 a_2 - 1)}{a_1 v_1^2 + v_2^2 a_2}$$

If  $v_1$  and  $v_2$  much less than  $c$ , then,  $\theta_2 - \theta_1 = -\frac{v_1 v_2}{2c^2}$ .

# Chapter 4

## Mathematical Properties of Lorentz Transformations

The Lorentz Transformations have some interesting mathematical properties, particularly in the measurements of length and time. These new properties are not realized before.

### 4.1 Length Contraction (Lorentz-Fitzgerald Contraction)

Suppose there is a rod rest on the  $x$ -axis of  $S$  frame (see Fig. 4.1). Therefore, the coordinates of its end points are  $A(x_1, 0, 0)$  and  $B(x_2, 0, 0)$ . Its length in this frame is  $l = x_2 - x_1$ . Let an observer fixed in the  $S^1$  system which is moving uniformly along  $x$ -axis w.r.t.  $S$  system wants to measure the length of the given rod. He must find the coordinates of the two end points  $x_1^1$  and  $x_2^1$  in his system at the same time  $t^1$ . Thus from the inverse Lorentz transformations equations, we have

$$x_1 = a(x_1^1 + vt^1) \quad ; \quad x_2 = a(x_2^1 + vt^1)$$

so that

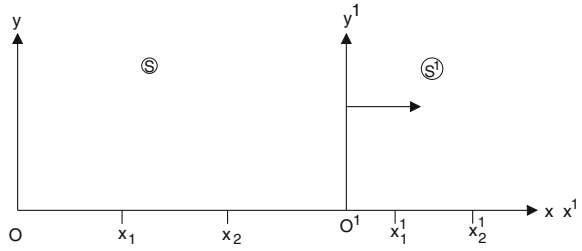
$$x_2 - x_1 = a(x_2^1 - x_1^1) \tag{4.1}$$

Therefore, the length of the rod in  $S^1$  system is

$$l^1 = (x_2^1 - x_1^1) = l\sqrt{(1 - \beta^2)} \quad \left[ \text{here, } a = \frac{1}{\sqrt{(1 - \beta^2)}} \quad , \quad \beta = \frac{v}{c} \right] \tag{4.2}$$

Thus to the moving observer, rod appears to be contracted. In other words, the length of the rod is greatest in the reference  $S$ , i.e. in the rest frame.

**Fig. 4.1** Rod's length in S frame is  $x_2 - x_1$  and in  $S^1$  frame is  $x_2^1 - x_1^1$



### Alternative

Now, we place the rod at rest in  $S^1$  frame with the space coordinates  $(x_1^1, 0, 0)$  and  $(x_2^1, 0, 0)$ . If the observer on S frame measures the length of this rod, then he finds the coordinates of the two end points  $x_1$  and  $x_2$  in his system at the same time  $t$ . Now, using direct Lorentz transformations equations, one gets

$$x_1^1 = a(x_1 - vt) \quad ; \quad x_2^1 = a(x_2 - vt)$$

so that

$$l^1 = x_2^1 - x_1^1 = a(x_2 - x_1) = al \quad (4.3)$$

Therefore, the length of the rod in S system is

$$l = l^1 \sqrt{1 - \beta^2} \quad (4.4)$$

This proves that the length of a rod, i.e. rigid body appears to be contracted in the ratio,  $\sqrt{1 - \beta^2} : 1$  in the direction of the relative motion in all other coordinate system,  $v$  being the relative velocity. This result is known as *Lorentz-Fitzgerald contraction* in special theory of relativity.

**Proper length:** The proper length of the rod is its length in a reference system in which it is at rest.

**Note 1:** The reader should note that we have used here inverse Lorentz transformations equations. It would not be convenient to use direct equation since events (location of end points) that are simultaneous in primed system (ends are measured at the same time  $t^1$ ) will not be simultaneous in unprimed system as they are at different points  $x_1$  and  $x_2$ . As we are measuring from  $S^1$  system, we do not know the value of  $x$  and  $t$  in S system. So, we have to use inverse Lorentz transformations equations.

**Note 2:** We note that the contraction depends on  $v^2$ . Therefore, if one replaces  $v$  by  $-v$ , the contraction result remains the same. Therefore, the Lorentz-Fitzgerald contraction law is independent of the direction of the relative velocity between the systems.

**Note 3:** The proper length is the maximum length, i.e. every rigid body appears to have its maximum length in the coordinate system in which it is at rest.

**Note 4:** When the rod moves with a velocity relative to the observer, then rod's length remains unaffected in the perpendicular direction of motion.

*Proof* Let,  $S^1$  is moving with velocity  $v$  along  $x$ -axis relative to a system  $S$ . Let a rod of length  $l$  is placed on  $y$ -axis. The coordinates of its ends are  $(0, y_1, 0)$  and  $(0, y_2, 0)$ . Suppose, in  $S^1$  system, the coordinates of its ends are  $(0, y_1^1, 0)$  and  $(0, y_2^1, 0)$ . Now, using the Lorentz transformations equations, one gets,

$$l^1 = y_2^1 - y_1^1 = y_2 - y_1 = l$$

This means length of the rod remains unaltered when an observer in  $S^1$  system measures it. In other words, the lengths perpendicular to the direction of motion remains unaffected.

## 4.2 Time Dilation

Suppose there is a clock A rest at point  $(x, 0, 0)$  in  $S$  frame. Let this clock gives a signal at time  $t_1$  in  $S$  frame. An observer fixed in the  $S^1$  system which is moving uniformly along  $x$ -axis w.r.t.  $S$  system measures this time as  $t_1^1$ . According to the Lorentz transformations equations,

$$t_1^1 = a \left[ t_1 - \frac{vx}{c^2} \right]$$

Let the clock A fixed at the same point  $(x, 0, 0)$  in  $S$  system gives another signal at time  $t_2$ . The observer in  $S^1$  system notes this time as

$$t_2^1 = a \left[ t_2 - \frac{vx}{c^2} \right]$$

Therefore, the apparent time interval is

$$t_2^1 - t_1^1 = a [t_2 - t_1] \tag{4.5}$$

This shows that

$$(t_2^1 - t_1^1) > (t_2 - t_1) \tag{4.6}$$

Hence, it appears to the moving observer that the stationary clock is moving at a slow rate. This effect is called *time dilation*.

### Alternative



Suppose, there is a clock B that is rigidly attached at point  $(x^1, 0, 0)$  in  $S^1$  frame which is moving uniformly along  $x$ -axis w.r.t. S system. Let this clock gives a signal at time  $t_1^1$  in  $S^1$  frame. A clock in S system measures this time as  $t_1$ . Thus from the inverse Lorentz transformations equations, we have

$$t_1 = a \left[ t_1^1 + \frac{vx^1}{c^2} \right]$$

Therefore, time interval  $(t_2 - t_1)$  recorded by the clock at S frame is related to the interval  $(t_2^1 - t_1^1)$ , recorded by the clock in  $S^1$  frame as follows:

$$(t_2 - t_1) = a(t_2^1 - t_1^1) \quad (4.7)$$

Obviously,

$$(t_2 - t_1) > (t_2^1 - t_1^1) \quad (4.8)$$

If  $(t_2^1 - t_1^1)$  is 1 s, then the observer in S judges that more than 1 s have elapsed between two events. Thus, the moving clock appears to be slowed down or retarded.

**Proper time:** The time recorded by a clock moving with a given system is called *proper time*.

**Note 5:** Note that if we fix the clock in the moving (primed) system or rest (unprimed) system even then the conclusion is true, i.e. observer reports that the clock (in unprimed or primed system) is running slow as compared to his own. Therefore by recording the time on a clock fixed in a certain system by an observer fixed another system, it is difficult to distinguish which one of the systems is stationary or in motion; only a relative motion can be guessed.

**Note 6:** A clock will be found to run more and more slowly with the increase of the relative motion between clock and observer.

**Note 7:** (i) If moving observer measures the time of a stationary clock, then use

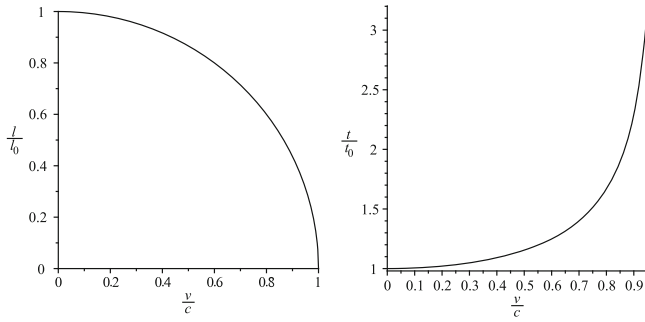
$$\Delta t^1 = a \Delta t$$

(ii) If stationary observer measures the time of a moving clock, then use

$$\Delta t = a \Delta t^1$$

In general, time dilation

$$t = \frac{t_0}{\sqrt{1 - \frac{v^2}{c^2}}}$$



**Fig. 4.2** Variations of length contraction (*left*) and time dilation (*right*) with respect to  $\frac{v}{c}$  are shown in the figures

$t_0$  = time interval on clock at rest relative to an observer = proper time.  
 $t$  = time interval on clock in motion relative to an observer.  
 $v$  = relative speed.

**Note 8:** If the moving observer observes the rod, then

$$l = l_0 \sqrt{1 - \frac{v^2}{c^2}}.$$

If the moving observer measures the time, then

$$t = \frac{t_0}{\sqrt{1 - \frac{v^2}{c^2}}}.$$

**Note 9:** Figure 4.2 indicates that for moving observer, the length is gradually decreasing with the increase of its velocity and when velocity is approaching to the velocity of light, the measurement of length is tending to zero. However, for time dilation the nature is just opposite to the length contraction. That is, time interval is gradually increasing with the increase of observer’s velocity and when velocity is approaching to the velocity of light, the time interval will be infinitely large.

### 4.3 Relativity of Simultaneity

Two events are said to be simultaneous if they occur at the same time. It can be shown that two events occur at a same time in one frame of reference  $S$  are not in general simultaneous w.r.t. another frame of reference  $S^1$  which is in relative motion w.r.t. the previous frame. Let us consider two events happening at times  $t_1$  and  $t_2$  in  $S$  frame at rest. Suppose an observer in  $S^1$  is observing these two events at two different points  $x_1$  and  $x_2$ , respectively. The observer in  $S^1$  records the corresponding times with

respect to his frame of reference are  $t_1^1$  and  $t_2^1$ . Then using Lorentz Transformation equations, one gets

$$t_1^1 = a \left[ t_1 - \left( \frac{vx_1}{c^2} \right) \right] \quad ; \quad t_2^1 = a \left[ t_2 - \left( \frac{vx_2}{c^2} \right) \right] \quad (4.9)$$

Thus,

$$t_2^1 - t_1^1 = a(t_2 - t_1) + a\left(\frac{v}{c^2}\right)(x_1 - x_2) \quad (4.10)$$

If the two events are simultaneous in the frame of reference S, then  $t_2 = t_1$ . Therefore the above expression gives,

$$t_2^1 - t_1^1 = a\left(\frac{v}{c^2}\right)(x_1 - x_2) \neq 0, \quad \text{since } x_1 \neq x_2 \quad (4.11)$$

Thus, if the events are simultaneous in S frame, then, in general, the same events are not simultaneous in  $S^1$  frame which is moving with velocity  $v$  relative to S. This indicates that simultaneity is not absolute, but relative.

#### 4.4 Twin Paradox in Special Theory of Relativity

Let Rina and Mina, each 20 years of age, are two twins. Suppose Mina travels within a spaceship with velocity,  $v = \frac{3}{5}c$  for four years along positive direction of  $x$ -axis, whereas Rina stays in Kolkata. Therefore, one can imagine, Rina is staying in a rest frame, say S and the spaceship containing Mina is moving a moving frame  $S^1$ . Note that Mina counts four year in her reference frame, i.e. in  $S^1$ . Then, Mina returns with same speed and reaches the starting point after the end of another four years of her time. Then according to Mina, the age of Mina is  $20 + 8 = 28$  years, but according to Rina (i.e. ith respect to rest frame S), the age of Mina is

$$20 + \frac{8}{\sqrt{1 - \left(\frac{3c}{5c}\right)^2}} = 30 \text{ years}$$

On the other hand, the age of Rina according to her (i.e. S frame) would be 30 years, whereas the age of Mina according to her (i.e. in  $S^1$  frame) would be 28 years. These two statements are contradict to each other. That is, one of the two twins appear younger than the other. This is known as *twin paradox* in special theory of relativity. According to special theory of relativity, Mina will be 28 years old, while Rina 30 years. That is Mina is younger than Rina.

**Note 10:** The two situations are not equivalent. Mina changed from one inertial frame S to another inertial frame  $S^1$  of her journey, however, Rina remained in the same

inertial frame during Mina's journey. Therefore, the time dilation formula applies to Mina's movement from Rina.

**Note 11:** After travelling four years, the velocity of Mina changes from  $v$  to  $-v$  instantly to reach the Kolkata. However, the calculation of time remains the same.

**Note 12:** The distance Mina covers should be shorten as compare to Rina, i.e. compare to the distance measured by S frame. Let us consider Mina reaches to a star 4 light years away from earth (at the beginning of her journey origins of both frames coincided). Mina covers the distance which is shorten to

$$L = L_0 \sqrt{\left(1 - \frac{v^2}{c^2}\right)} = (4 \text{ light years}) \sqrt{\left(1 - \frac{9}{25}\right)} = 3.2 \text{ light years}$$

**Note 13:** The clock could be substituted for any periodic natural phenomena, i.e. process of life such as heartbeat, respiration, etc. For, Rina who stays at rest, Mina's life (who is moving in a high speed) is slower than her by a factor

$$\sqrt{\left(1 - \frac{v^2}{c^2}\right)} = \sqrt{\left(1 - \frac{9}{25}\right)} = 80 \%$$

It follows that in a certain period, Mina's heart beats only 4 times, whereas Rina's heart beats 5 times within that period. Mina takes 4 breaths for every 5 of Rina. Mina thinks 4 thoughts for every 5 of Rina. Thus, during the travel, Mina's biological functions proceeding at the same slower rate as her physical clock.

## 4.5 Car-Garage Paradox in Special Theory of Relativity

Consider a car and a garage with same proper length,  $l_0$ . When at rest, car be parked exactly inside the garage. Now, a person driving the car towards the garage with speed  $v$ . For the gate man, the car appears to be of length

$$l = l_0 \left(1 - \frac{v^2}{c^2}\right) < l_0$$

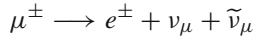
He realizes that the car smoothly enters the garage and the driver does not need to stop the car before the garage. Now, we see what will happen for the driver. For him, the car length is  $l_0$ , while the garage is of length

$$l = l_0 \left(1 - \frac{v^2}{c^2}\right) < l_0$$

Therefore, he realizes that the garage length is smaller than the car and he stops the car before the garage. This is known as car-garage paradox.

## 4.6 Real Example of Time Dilation

The elementary particle Muons are found in cosmic rays. It has a mass 207 times that of the electron ( $e$ ) and has a charge of either  $+e$  or  $-e$ . These Muons collide with atomic nuclei in earth's atmosphere and decay into an electron, a muon neutrino and an anti-neutrino:



The Muon has speed of about  $0.998c$  and mean lifetime in the rest frame is

$$t_0 = 2.2 \times 10^{-6} \text{ s} = 2.2 \mu\text{s}$$

These unstable elementary particles muons collide with atomic nuclei in earth's atmosphere at high altitudes and destroy within  $2.2 \mu\text{s}$ , however, they can be seen in the sea level in profusion. But in  $t_0 = 2.2 \mu\text{s}$ , muons can travel a distance of only

$$vt_0 = 0.998 \times 3 \times 10^8 \times 2.2 \times 10^{-6} = 0.66 \text{ km}$$

before decaying, whereas the collisions occurred at altitudes of 6 km or more.

Now, we have the following problems: (i) either Muon's mean lifetime is more than  $2.2 \mu\text{s}$  or (ii) its velocity is more than  $c$ , the velocity of light. The second possibility is impossible as it violates the principle of theory of relativity. Since the cosmic rays Muons are moving violently towards earth with a high speed,  $0.998c$ , their lifetimes are extended in our frame of reference (rest frame) by time dilation to

$$t = \left[ \frac{t_0}{\sqrt{1 - \left(\frac{v}{c}\right)^2}} \right] = \left[ \frac{2.2 \times 10^{-6}}{\sqrt{1 - \left(\frac{0.998c}{c}\right)^2}} \right] = 34.8 \times 10^{-6} \text{ s} = 34.8 \mu\text{s}.$$

We note that the moving muon's mean lifetime almost 16 times longer than the proper lifetime, i.e. that at rest. Thus, the Muon will traverse the distance

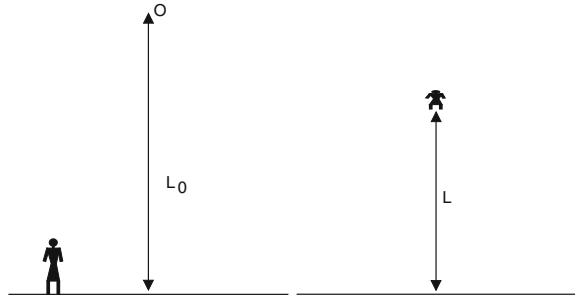
$$vt = .998 \times 3 \times 10^8 \times 34.8 \times 10^{-6} = 10.4 \text{ km}$$

Therefore, we conclude that a muon can reach the ground from altitudes about 10.4 km in spite its lifetime is only  $2.2 \mu\text{s}$ .

**Note 14:** The lifetime of Muon is only  $2.2 \mu\text{s}$  in its own frame of reference, i.e. moving frame of reference. The lifetime of Muon is  $34.8 \mu\text{s}$  in the frame of reference where these altitudes are measured, i.e. with respect to the rest frame of reference.

**Note 15:** Let an observer moving with the Muon. This means the observer and Muon are now in the same reference frame. Since the observer is moving relative to earth, he will find the distance 10.4 km to be contracted as (see Fig. 4.3)

**Fig. 4.3** Distance is contracted for muons



$$L = L_0 \sqrt{1 - \left(\frac{v}{c}\right)^2} = (10.4 \text{ km}) \sqrt{1 - \left(\frac{0.998c}{c}\right)^2} = 0.66 \text{ km}$$

Therefore, a Muon will travel with the speed of about  $0.998c$  in  $2.2 \mu \text{ s}$ .

### 4.7 Terrell Effects

In 1959 James Terrell observed that if one takes a snapshot of a rapidly moving object with respect to the film, then the photograph may give a distorted picture, i.e. the object has gone a rotation rather than contraction. This phenomena is known as Terrell effects. Terrell proved that the shape of the moving body would remain unchanged to the observer, however, it would appear only to be rotated a little. This phenomena is happened due to combination of two different effects: (i) the length contraction effect (ii) the retardation effect. Second effect will arise because the light ray from different parts of the object take different times in reaching the observer's eyes. Hence, when a body is in motion, a distorted image is formed due to the combination of the above two effects.

Let us suppose that a square body flies past an observer with a constant velocity  $v$  at a large distance from the observer. Let  $l$  be length of each side. By length contraction effect, the image of front side will have a length  $l = \sqrt{1 - \frac{v^2}{c^2}}$ . By retardation effect, light rays from the far side must originate  $\frac{l}{c}$  time earlier than the light rays originating from the front face so as to reach the observer at the same observer time. If one takes a photo of this moving body, the snapshot shows the near side and far side as it was  $\frac{l}{c}$  earlier when it was a distance  $\frac{vl}{c}$  to the left. This means side face of the square is seen inclined at an angle  $\theta$  with the front side where  $\tan \theta = \frac{vl}{l} = \frac{v}{c}$ .

*Example 4.1* A particle with mean proper lifetime of  $2 \mu \text{ s}$ , moves through the laboratory with a speed of  $0.9c$ . Calculate its lifetime as measured by an observer in laboratory. What will be the distance traversed by it before decaying? Also find the distance traversed without taking relativistic effect.

**Hint:** We use the result of time dilation,  $\Delta t^1 = \frac{\Delta t}{\sqrt{1 - \frac{v^2}{c^2}}}$ .

Given:  $\Delta t = 2 \mu s = 2 \times 10^{-6}$ ,  $v = 0.9c = 0.9 \times 3 \times 10^8$ . Therefore,  $\Delta t^1 = 4.58 \mu s$ .

Distance traversed  $= 0.9 \times 3 \times 10^8 \times 4.58 \times 10^{-6}$

Distance traversed without relativistic effect  $= 0.9 \times 3 \times 10^8 \times 2 \times 10^{-6}$

*Example 4.2* The proper lifetime of  $\pi^+$  mesons is  $\alpha \mu s$ . If a beam of these mesons of velocity  $f c$  produced ( $f < 1$ ), calculate the distance, the beam can travel before the flux of the mesons beam is reduced to  $e^{-\beta}$  times the initial flux,  $\alpha$  being a constant.

**Hint:** Given:  $\Delta t = \alpha$ ,  $v = f c$ . Therefore,  $\Delta t^1 = \frac{\alpha}{\sqrt{1 - \frac{f^2 c^2}{c^2}}}$  etc.

If  $N_0$  is the initial flux and  $N$  is the flux after time  $t$ , then we have  $N = N_0 e^{-\frac{t}{\tau}}$ ,  $\tau$  being the mean lifetime.

Here,  $N = e^{-\beta} N_0$ , therefore  $e^{-\beta} N_0 = N_0 e^{-\frac{t}{\tau}} \implies t = \beta \tau = \beta \Delta t^1 = \frac{\alpha \beta}{\sqrt{1 - f^2}}$ . Hence, the distance traversed by the beam before the flux of mesons beam is reduced to  $e^{-\beta}$  times the initial flux is  $d = t \times f c$  etc.

*Example 4.3* At what speed should a clock be moved, so that it may appear to loose  $\alpha$  seconds in each minute.

**Hint:** If the clock is to loose  $\alpha$  seconds in each minute, then the moving clock with velocity  $v$  records  $60 - \alpha$  s for each 60 s recorded by stationary clock.

We use the result of time dilation,  $\Delta t^1 = \frac{\Delta t}{\sqrt{1 - \frac{v^2}{c^2}}}$ .

Given:  $\Delta t = 60$  s and  $\Delta t^1 = 60 - \alpha$  s, etc.

*Example 4.4* Calculate the length and orientation of a rod of length 1 metre in a frame of reference which is moving with velocity  $f c$  ( $f < 1$ ) in a direction making an angle  $\theta$  with the rod.

**Hint:** Let the rod is moving with velocity  $f c$  along  $x$ -axis and the rod is inclined at the  $\theta$  angle with  $x$ -axis. We know contraction will take place only in  $x$  direction and not in  $y$  direction. Here,  $l_x = l \cos \theta$  and  $l_y = l \sin \theta$ . Now, using the result of length contraction ( $L = L_0 \sqrt{1 - \frac{v^2}{c^2}}$ ), we have  $l_x^1 = l_x \sqrt{1 - \frac{f^2 c^2}{c^2}} = l \cos \theta \sqrt{1 - f^2}$ ,  $l_y^1 = l_y = l \sin \theta$ . Therefore, the length of the rod as observed from moving frame,  $l^1 = \sqrt{(l_x^1)^2 + (l_y^1)^2}$  etc.

If  $\theta^1$  is the angle made by the rod with  $x$ -axis in moving frame, then  $\tan \theta^1 = \frac{l_y^1}{l_x^1}$ , etc.

*Example 4.5* An astronaut wants to go to a star  $\alpha$  light year away. The rocket accelerates quickly and then moves at a uniform velocity. Calculate with what velocity the rocket must move relative to earth if the astronaut is to reach therein  $\beta$  year as measured by clock at rest on the rocket.

**Hint:** The distance of the star from earth is  $= \alpha y c$  where  $y = 365 \times 24 \times 3,600$ s. In the frame of reference of the astronaut ( i.e.  $S^1$  ), the time required is given by

$$t_2^1 - t_1^1 = \frac{(t_2 - t_1) - \frac{v}{c^2}(x_2 - x_1)}{\sqrt{1 - \frac{v^2}{c^2}}}$$

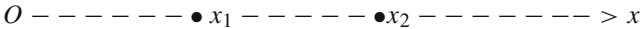
Here,  $(t_2 - t_1)$  is the time difference measured by the clock in earth and  $t_2^1 - t_1^1$  is the time difference measured by the clock in the rocket, i.e. moving frame. Let  $v = fc$  ( $f < 1$ ) be the required velocity of rocket to reach the distance  $x_2 - x_1 = \alpha y c$  in  $t_2^1 - t_1^1 = \beta y$  year. Therefore,  $x_2 - x_1 = \alpha y c = v(t_2 - t_1)$ . Now using these, we get,

$$\beta y = \frac{\frac{\alpha y c}{v} - \frac{v}{c^2}(\alpha y c)}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{\frac{\alpha y c}{fc} - \frac{fc}{c^2}(\alpha y c)}{\sqrt{1 - \frac{f^2 c^2}{c^2}}}, \text{ etc.}$$

*Example 4.6* A flash of light is emitted at the position  $x_1$  on  $x$ -axis and is absorbed at position  $x_2 = x_1 + l$ . In a reference frame  $S^1$  moving with velocity  $v = \beta c$  along  $x$ -axis, find the spatial separation between the point of emission and the point of absorption of light.

**Hint:**

S frame



Here, the frame  $S^1$  is moving with velocity  $v = \beta c$ .

In S frame, let the light is emitted on the position  $x_1$  at time  $t_1 = 0$  and will absorb on the position  $x_2 = x_1 + l$  at time  $t_2 = 0$  where,  $t_2 = \frac{(x_1+l)-x_1}{c} = \frac{l}{c}$ .

In  $S^1$  frame, the points  $x_1$  and  $x_2$  will be

$$x_1^1 = \frac{x_1 - vt_1}{\sqrt{1 - \beta^2}} = \frac{x_1}{\sqrt{1 - \beta^2}}, \quad x_2^1 = \frac{x_2 - vt_2}{\sqrt{1 - \beta^2}} = \frac{(x_1 + l) - \frac{vl}{c}}{\sqrt{1 - \beta^2}}$$

Therefore, the spatial separation is  $l^1 = x_2^1 - x_1^1$  etc.

*Example 4.7* Consider that a frame  $S^1$  is moving with velocity  $v$  relative to the frame S. A rod in  $S^1$  frame makes an angle  $\theta^1$  with respect to the forward direction of motion. What is the angle  $\theta$  as measured in S frame?

**Hint:** Here, x component of the length will be contracted relative to an observer in S frame etc. see Example 4.4.

*Example 4.8* Two rods which are parallel to each other move relative to each other in their length directions. Explain the apparent paradox that either rod may appear longer than the other depending on the state of motion of the observer.



**Hint:** Let us place a rod of length  $l_0$  at rest on the  $x$ -axis of S frame. Consider two frames of reference  $S^1$  and  $S^2$  which are moving with velocities  $v_1$  and  $v_2$ , respectively relative to S frame. To the observer in  $S^1$  frame, the rod appears as

$$l^1 = l_0 \sqrt{1 - \frac{v_1^2}{c^2}} \quad (1)$$

Similarly, to the observer in  $S^2$  frame, the rod appears as

$$l^2 = l_0 \sqrt{1 - \frac{v_2^2}{c^2}} \quad (2)$$

Case 1: If  $v_2 > v_1$ , then  $\frac{v_2^2}{c^2} > \frac{v_1^2}{c^2}$ . Since  $\frac{v_2^2}{c^2} = 1 - \frac{l_2^2}{l_0^2}$  and  $\frac{v_1^2}{c^2} = 1 - \frac{l_1^2}{l_0^2}$ , therefore  $l_1 > l_2$ .

Case 2: If  $v_1 > v_2$ , then  $\frac{v_1^2}{c^2} > \frac{v_2^2}{c^2}$ . As a result,  $l_2 > l_1$ .  
Thus, either rod may appear longer than the other depending on the state of motion of the observer.

*Example 4.9* Four-dimensional volume element 'dxdydzdt' is invariant under Lorentz Transformation.

**Hint:** According to Lorentz-Fitzgerald contraction, we have

$$dx^1 = dx \sqrt{1 - \frac{v^2}{c^2}}, \quad dy^1 = dy, \quad dz^1 = dz \quad (1)$$

According to time dilation,

$$dt^1 = \frac{dt}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (2)$$

The four dimensional volume element in  $S^1$  frame is  $dx^1 dy^1 dz^1 dt^1$  etc.

*Example 4.10* Two events are simultaneous, though not coincide in some inertial frame. Prove that there is no limit on time separation assigned to these events in different frames but the space separation varies from  $\infty$  to a minimum which is the measurement in that inertial frame.

**Hint:** From Lorentz Transformation, we have

$$x_2 - x_1 = \frac{(x_2^1 - x_1^1) + v(t_2^1 - t_1^1)}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (1)$$

$$t_2 - t_1 = \frac{(t_2^1 - t_1^1) + \frac{v}{c^2}(x_2^1 - x_1^1)}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (2)$$

Two events are simultaneous but not coincide, therefore,  $t_2^1 - t_1^1 = 0$  and  $x_2^1 - x_1^1 \neq 0$ . Thus (1) implies,

$$x_2 - x_1 = \frac{(x_2^1 - x_1^1)}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (3)$$

Now, if  $v = 0$ , then  $x_2 - x_1 = x_2^1 - x_1^1$  and  $v = c$ ,  $x_2 - t_1 = \infty$ .

Thus, space separation varies from  $\infty$  to a minimum which is the measurement in the rest frame.

Also, from (2), we get,

$$t_2 - t_1 = \frac{\frac{v}{c^2}(x_2^1 - x_1^1)}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (4)$$

If  $v = 0$ , then  $t_2 - t_1 = 0$  and  $v = c$ ,  $t_2 - t_1 = \infty$ .

Thus, there is no limit for time separation.

*Example 4.11* A circular ring moves parallel to its plane relative to an inertial frame  $S$ . Show that the shape of the ring relative to  $S$  is an ellipse.

**Hint:** Let us consider a frame  $S^1$  containing the ring is moving with velocity  $v$  relative to  $S$  frame in  $x$ -direction. Obviously,  $S^1$  frame is the rest frame with respect to the ring. Let  $(X, Y, 0, T)$  and  $(X^1, Y^1, 0, T^1)$ , respectively, be the coordinates of the centre of the ring relative to  $S$  and  $S^1$ . Consider  $(x, y, 0, T)$ ,  $(x^1, y^1, 0, T^1)$ , respectively, be the coordinates of any point on the circumference of the ring relative to  $S$  and  $S^1$  at the same observer time  $T$ . (here, ring lies in  $xy$  plane.)

From Lorentz Transformation, we have

$$X^1 = \frac{X - vT}{\sqrt{1 - \frac{v^2}{c^2}}}, \quad x^1 = \frac{x - vT}{\sqrt{1 - \frac{v^2}{c^2}}}, \quad Y^1 = Y, \quad y^1 = y \quad (1)$$

Equation of the ring is in  $S^1$  frame is

$$(x^1 - X^1)^2 + (y^1 - Y^1)^2 = a^2 \quad (2)$$

$a$  being the radius of the circular ring.

Using the transformation (1), we get

$$\frac{(x - X)^2}{1 - \frac{v^2}{c^2}} + (y - Y)^2 = a^2 \quad (3)$$

This is an ellipse as observed in S frame. The centre of the ellipse moves along  $x$ -axis with velocity  $v$ .

*Example 4.12* A reference frame  $S^1$  containing a body has the dimensions represented by  $(a\hat{i} + b\hat{j} + c\hat{k})$  is moving with velocity  $f c$  ( $0 < f < 1$ ) along  $x$ -axis relative to a reference system S. How these dimensions will be represented in system S. ( $\hat{i}, \hat{j}, \hat{k}$ ) are unit vectors along respective axes.

**Hint:** The dimension of the body along  $x$ -axis as viewed from S system is  $a^1 = a\sqrt{1 - f^2}$ . However, there are no contractions in the directions of  $y$  and  $z$  axes, etc.

*Example 4.13* In the usual set-up of frames  $S^1$  and S having relative velocity  $v$  along  $x - x^1$ , the origins coinciding at  $t = t^1 = 0$ , we shall find a later time,  $t$ , there is only one plane S on which the clocks agree with those of  $S^1$ . Show that this plane is given by  $x = \frac{c^2}{v} \left[ 1 - \sqrt{1 - \frac{v^2}{c^2}} \right] t$  and moves with velocity  $u = \frac{c^2}{v} \left[ 1 - \sqrt{1 - \frac{v^2}{c^2}} \right]$ . Show also, this velocity is less than  $v$ .

**Hint:** Here the frame  $S^1$  is moving with velocity  $v$  relative to S frame in  $x$  direction. Let  $(x, y, z, t)$  and  $(x^1, y^1, z^1, t^1)$ , respectively, be the coordinates of an event relative to S and  $S^1$ . Then from Lorentz Transformation we have,

$$t = \frac{t^1 + \frac{vx^1}{c^2}}{\sqrt{1 - \frac{v^2}{c^2}}}, \quad x = \frac{x^1 + vt^1}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (1)$$

Considering the situation that clock in S frame agrees with the clock in  $S^1$  frame, i.e. when  $t = t^1$ , then from the second expression in (i), we get,

$$x^1 = x\sqrt{1 - \frac{v^2}{c^2}} - vt \quad (2)$$

Using (2), we get from the first expression in (1) as

$$t\sqrt{1 - \frac{v^2}{c^2}} = t + \frac{v}{c^2} \left[ x\sqrt{1 - \frac{v^2}{c^2}} - vt \right]$$

This gives the equation of the plane  $x = \frac{c^2}{v} \left[ 1 - \sqrt{1 - \frac{v^2}{c^2}} \right] t$  and moves with velocity  $u = \frac{c^2}{v} \left[ 1 - \sqrt{1 - \frac{v^2}{c^2}} \right]$ .

Expanding binomially and neglecting higher power of  $\frac{v^2}{c^2}$ , we get

$$u = \frac{c^2}{v} \left[ 1 - \left( 1 - \frac{v^2}{2c^2} \right) \right], \text{ etc}$$

*Example 4.14* Is it true that two events which occur at the same place and at the same time for one observer will be simultaneous for all observers?

**Hint:** Use

$$t_2^1 - t_1^1 = \frac{(t_2 - t_1) - \frac{v}{c^2}(x_2 - x_1)}{\sqrt{1 - \frac{v^2}{c^2}}} \text{ etc}$$

*Example 4.15* The S and  $S^1$  are two inertial frames of reference. Further, at any instant of time, the origin of  $S^1$  is situated at the position  $(vt, 0, 0)$ . A particle describes a parabolic trajectory,  $x = ut, y = \frac{1}{2}ft^2$  in S frame. Find the trajectory of the particle in  $S^1$  frame. Show also that its acceleration in  $S^1$  frame along  $y^1$  axis is  $\frac{f(1 - \frac{v^2}{c^2})}{(1 - \frac{uv}{c^2})^2}$ .

**Hint:** Since S and  $S^1$  are inertial frames, the relative velocity between them is constant. Also, at  $t$ , the origin of  $S^1$  is situated at the position  $(vt, 0, 0)$ , therefore  $S^1$  must move with a velocity  $v$  along  $x$ -axis. From, Lorentz Transformation, we have

$$t = \frac{t^1 + \frac{vx^1}{c^2}}{\sqrt{1 - \frac{v^2}{c^2}}}, \quad x = \frac{x^1 + vt^1}{\sqrt{1 - \frac{v^2}{c^2}}}$$

Now,  $x = ut$  implies

$$\frac{x^1 + vt^1}{\sqrt{1 - \frac{v^2}{c^2}}} = u \left[ \frac{t^1 + \frac{vx^1}{c^2}}{\sqrt{1 - \frac{v^2}{c^2}}} \right]$$

This gives,

$$x^1 = \left[ \frac{u - v}{1 - \frac{uv}{c^2}} \right] t^1$$

Hence, the particle moves with a uniform velocity  $\frac{u-v}{1-\frac{uv}{c^2}}$  along  $x^1$  axis.

Again,

$$y = y^1 = \frac{1}{2}ft^2 = \frac{1}{2}f \left[ \frac{t^1 + \frac{vx^1}{c^2}}{\sqrt{1 - \frac{v^2}{c^2}}} \right]^2 = \frac{1}{2}f \left[ \frac{t^1 + \frac{v \left( \frac{u-v}{1-\frac{uv}{c^2}} t^1 \right)}{c^2}}{\sqrt{1 - \frac{v^2}{c^2}}} \right]^2$$

This gives

$$y^1 = \left[ \frac{f \left( 1 - \frac{v^2}{c^2} \right)}{2 \left( 1 - \frac{uv}{c^2} \right)^2} \right] t^{12}$$

Hence, the acceleration along  $y^1$  axis  $f_y^1 = \frac{d^2 y^1}{d^2 t^1}$ , etc.

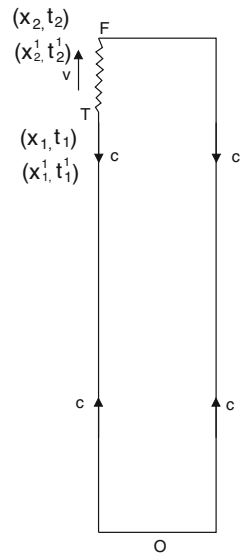
*Example 4.16* The  $S$  and  $S^1$  are two inertial frames of reference. Further, at the instant of time, the origin of  $S^1$  is situated at the position  $(vt, 0, 0)$ . Trajectory of a particle describes a straight line through the origin in  $S$  frame. Find the trajectory of the particle in  $S^1$  frame. Show also that it has no acceleration in  $S^1$  frame.

**Hint:** The parametric equations of the straight line passing through origin is,  $x = t, y = mt$ . See Example 4.15.

*Example 4.17* A rocket of proper length  $L$  metres moves vertically away from Earth with uniform velocity. A light pulse sent from Earth and is reflected from the mirrors at the tail end and the front end of the rocket. If the first pulse is received at the ground  $2T$  s after emission and second pulse is received  $2f \mu\text{s}$  later. Calculate (i) the distance of the rocket from Earth (ii) the velocity of the rocket relative to Earth and (iii) apparent length of the rocket.

**Hint:** (i) The pulse sent from Earth reaches the back part of the rocket  $T$  s after its emission, hence at this instant, the distance of the back part of the rocket from Earth =  $Tc$  metre (see Fig. 4.4).

**Fig. 4.4** Distance of the back part of the rocket from Earth is  $Tc$  metre



- (ii) Let the tail end and the front end of the rocket be characterized by  $(x_1, t_1)$ ,  $(x_2, t_2)$  and  $(x_1^1, t_1^1)$ ,  $(x_2^1, t_2^1)$  in Earth frame S and in the rocket frame  $S^1$ , respectively. From Lorentz Transformation, we have

$$t_2 - t_1 = \frac{(t_2^1 - t_1^1) + \frac{v}{c^2}(x_2^1 - x_1^1)}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (1)$$

Here,  $x_2^1 - x_1^1 = L$ ,  $t_2^1 - t_1^1 = \frac{x_2^1 - x_1^1}{c} = \frac{L}{c}$ ,  $t_2 - t_1 = f \times 10^{-6}$ .  
Therefore, from (1), one can get

$$f \times 10^{-6} = \frac{(\frac{L}{c}) + \frac{v}{c^2}(L)}{\sqrt{1 - \frac{v^2}{c^2}}} \implies v = \text{etc.} \quad (2)$$

- (iii) The apparent length of the rocket  $l = L\sqrt{1 - \frac{v^2}{c^2}}$ , using the value of  $v$  from (2), one can get the required apparent length.

*Example 4.18* Let,  $V_0$  is the volume of a cube in reference system S. Find the volume of the cube as viewed from  $S^1$  frame which is moving with velocity  $v$  along one edge of the cube.

**Hint:** Let  $L_0$  be the length of one edge of the cube in reference system S. Then,  $V_0 = L_0^3$ . Let  $S^1$  is moving along one edge of the cube parallel to  $x$ -axis. Therefore, the length of the edge is contracted by a factor  $\sqrt{1 - \frac{v^2}{c^2}}$  in the direction of motion, i.e. parallel to the  $x$ -axis. The other edges remain unaffected. So, volume of the cube as viewed from  $S^1$  frame is  $V = L_0 L_0 L_0 \sqrt{1 - \frac{v^2}{c^2}} = V_0 \sqrt{1 - \frac{v^2}{c^2}}$

# Chapter 5

## More Mathematical Properties of Lorentz Transformations

### 5.1 Interval

Any event is described by two different things: One is the place where it occurred and another is time when it occurred. Therefore, an event is expressed by three spatial coordinates and one time coordinate. Such a space (four axes of which represent three space coordinates and one time coordinate) is called *Minkowski's Four-dimensional world*. Any event is represented by a point, called *World point* in this space and there corresponds to each particle a certain line called *World line*, i.e. A curve in the four-dimensional space is called *World line*. Let us consider two events which are represented by the coordinates  $(x_1, y_1, z_1, t_1)$  and  $(x_2, y_2, z_2, t_2)$  w.r.t. a frame of reference which is at rest. Then the quantity

$$\rho_{12} = [c^2(t_2 - t_1)^2 - (x_2 - x_1)^2 - (y_2 - y_1)^2 - (z_2 - z_1)^2]^{\frac{1}{2}} \quad (5.1)$$

is called the *Interval* between two events.

### 5.2 The Interval Between Two Events Is Invariant Under Lorentz Transformation

*Proof* Let consider two events which are represented by the coordinates  $(x_1, y_1, z_1, t_1)$  and  $(x_2, y_2, z_2, t_2)$  w.r.t. a frame of reference which is at rest, i.e. in S frame. Suppose, the corresponding coordinates for these events in another frame of reference  $S^1$ , which is moving with uniform velocity along x axis relative to the previous frame, be  $(x_1^1, y_1^1, z_1^1, t_1^1)$  and  $(x_2^1, y_2^1, z_2^1, t_2^1)$ . Then the interval in  $S^1$  frame,

$$\rho_{12}^1 = [c^2(t_2^1 - t_1^1)^2 - (x_2^1 - x_1^1)^2 - (y_2^1 - y_1^1)^2 - (z_2^1 - z_1^1)^2]^{\frac{1}{2}} \quad (5.2)$$

Using Lorentz Transformation, we have

$$(\rho_{12}^1)^2 = c^2 \left[ a \left\{ t_2 - \left( \frac{vx_2}{c^2} \right) \right\} - a \left\{ t_1 - \left( \frac{vx_1}{c^2} \right) \right\} \right]^2 - [a(x_2 - vt_2) - a(x_1 - vt_1)]^2 - (y_2 - y_1)^2 - (z_2 - z_1)^2 = (\rho_{12})^2$$

Thus, the interval between two events is invariant under Lorentz Transformation. We have shown that the interval between two events are invariant, i.e.  $(\rho_{12}^1)^2 = (\rho_{12})^2$ . This gives,

$$\begin{aligned} & [c^2(t_2^1 - t_1^1)^2 - (x_2^1 - x_1^1)^2 - (y_2^1 - y_1^1)^2 - (z_2^1 - z_1^1)^2] \\ & = [c^2(t_2 - t_1)^2 - (x_2 - x_1)^2 - (y_2 - y_1)^2 - (z_2 - z_1)^2] \end{aligned}$$

Now, if the two events occur at the same point in the system  $S^1$ , i.e.

$$(x_2^1 - x_1^1)^2 + (y_2^1 - y_1^1)^2 + (z_2^1 - z_1^1)^2 = 0$$

Then

$$(\rho_{12})^2 = [c^2(t_2 - t_1)^2 - (x_2 - x_1)^2 - (y_2 - y_1)^2 - (z_2 - z_1)^2] = [c^2(t_2^1 - t_1^1)^2] > 0$$

This indicates the interval between two events is a real number.

Real intervals are said to *Time like*. Thus, if the interval between two events is time like, there exists a system of reference in which the two events occur at one and the same place.

Now consider a case when the two events occur at one the same time ,i.e.  $(t_2^1 - t_1^1) = 0$ . Then,

$$\begin{aligned} (\rho_{12})^2 & = [c^2(t_2 - t_1)^2 - (x_2 - x_1)^2 - (y_2 - y_1)^2 - (z_2 - z_1)^2] \\ & = -[(x_2^1 - x_1^1)^2 + (y_2^1 - y_1^1)^2 + (z_2^1 - z_1^1)^2] < 0 \end{aligned}$$

Therefore,  $\rho_{12}$  is imaginary.

Imaginary intervals are said to be *space like*. Here, the space part dominates time part. Hence, if the interval between two events is space like, then there exists a system of reference in which the two events occur simultaneously.

If  $\rho_{12} = 0$ , then we call the interval to be *light like*. Then,

$$c^2(t_2 - t_1)^2 = (x_2 - x_1)^2 + (y_2 - y_1)^2 + (z_2 - z_1)^2,$$

so that the events can be connected to the light signal. Obviously, this is a limiting case dividing space like from time like.

**Note 1:** The expression  $x^2 + y^2 + z^2 - c^2t^2$  remains invariant under L.T. In terms of differentials, the quantity



$$ds^2 = dx^2 + dy^2 + dz^2 - c^2 dt^2 \quad (5.3)$$

remains also invariant under Lorentz Transformation. Here,  $ds$  is known as line element.

Remember that in Euclidean geometry the square of the distance between two infinitely close points in three dimension (i.e. in Cartesian coordinates) can be written as

$$ds^2 = dx_1^2 + dx_2^2 + dx_3^2 \quad (5.4)$$

For a four-dimensional Euclidean space, square of the distance between two infinitely close points will be written (generalizing the above notion)

$$ds^2 = dx_1^2 + dx_2^2 + dx_3^2 + dx_4^2 \quad (5.5)$$

Now, comparing (5.3) and (5.5), one can write

$$x_1 = x, \quad x_2 = y, \quad x_3 = z, \quad x_4 = ict \quad (5.6)$$

Here, we have introduced the imaginary fourth coordinate for mathematical convenience and for giving our equation a more symmetrical form.  $ds$  is the invariant spacetime interval between two events which are infinitely close to each other.

**Note 2:** Suppose a point moves from  $(x, y, z, t)$  to any neighbouring point, then we know the spacetime interval,

$$ds^2 = c^2 dt^2 - dx^2 - dy^2 - dz^2$$

or,

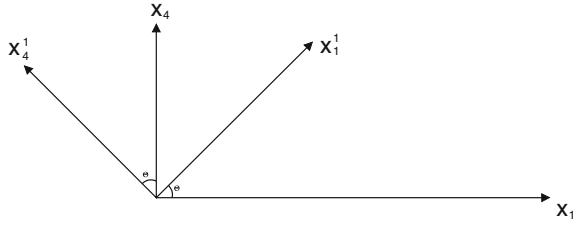
$$\left(\frac{ds}{dt}\right)^2 = c^2 - \left[\left(\frac{dx}{dt}\right)^2 + \left(\frac{dy}{dt}\right)^2 + \left(\frac{dz}{dt}\right)^2\right] = c^2 - v^2$$

If  $c^2 > v^2$ , then  $ds^2 > 0$ , then the distance or interval is positive. It is time-like interval. If  $c^2 < v^2$ , then  $ds^2 < 0$ , then the distance or interval is negative. It is space-like interval. For photon particle  $c^2 = v^2$ ,  $ds^2 = 0$ , then it is light-like or null interval. If  $c^2 < v^2$ , then particle does not obey the law of causality, i.e. if the particle is space-like then, the particle is causally disconnected.

**Note 3:** Suppose two events occur at the same place with respect to  $S^1$  frame, then  $\sum(x_i^1 - y_i^1)^2 = 0$ . Since interval is invariant, we have  $\rho_{12}^2 = (\rho_{12}^1)^2$ . This implies

$$c^2(t_1^1 - t_2^1)^2 = c^2(t_1 - t_2)^2 - \sum(x_i - y_i)^2$$

**Fig. 5.1**  $x_1$  and  $x_4$  axes are rotated through an angle  $\theta$



Writing differential notations, we have

$$c^2(dt^1)^2 = c^2dt^2 - dx^2 - dy^2 - dz^2$$

This yields

$$dt^1 = dt \sqrt{1 - \frac{v^2}{c^2}}$$

Thus we can re-derive time dilation from the invariant property of intervals.

**Theorem 5.1** *Lorentz Transformation may be regarded as a rotation of axes through an imaginary angle.*

*Proof* Let us consider four-dimensional space with coordinates  $x_1, x_2, x_3$  and  $x_4$  where  $x_4 = ict$ . Suppose  $x_1$  and  $x_4$  axes are rotated through an angle  $\theta$  so that new axes are  $x_1^1$  and  $x_4^1$ ;  $x_2, x_3$  axes remain unchanged (see Fig. 5.1). Then obviously  $x_2^1 = x_2, x_3^1 = x_3$  and

$$x_1^1 = x_1 \cos \theta + x_4 \sin \theta \quad (5.7)$$

$$x_4^1 = -x_1 \sin \theta + x_4 \cos \theta \quad (5.8)$$

Let us take angle to be imaginary such that

$$\tan \theta = \frac{iv}{c} = i\beta \quad (5.9)$$

where  $\beta = \frac{v}{c}$ . From (5.9), we have

$$\cos \theta = \frac{1}{a}, \quad \sin \theta = \frac{i\beta}{a} \quad (5.10)$$

where  $a = \sqrt{1 - \beta^2}$ . Now, putting the values of  $\sin \theta, \cos \theta$  and  $x_4$  in (5.7) and (5.8), we get

$$x_1^1 = x_1 \left( \frac{1}{a} \right) + ict \left( \frac{i\beta}{a} \right) = \left( \frac{1}{a} \right) [x_1 - vt]$$

$$ict^1 = -x_1 \left( \frac{i\beta}{a} \right) + ict \left( \frac{1}{a} \right)$$

or,

$$t^1 = \left( \frac{1}{a} \right) \left[ t - \frac{vx_1}{c^2} \right]$$

Thus Eqs. (5.7) and (5.8) give L.T. Hence, we can conclude that L.T. are equivalent to rotation of axes in four-dimensional space through an imaginary angle  $\tan^{-1}(i\beta)$ .

**Theorem 5.2** Lorentz Transformation may be expressed in hyperbolic form.

*Proof* Let us consider,

$$\tanh \phi = \beta = \frac{v}{c}.$$

Then,

$$\cosh \phi = a = \frac{1}{\sqrt{1 - \beta^2}}, \quad \sinh \phi = a\beta$$

Let construct a matrix

$$A = \begin{bmatrix} \cosh \phi & 0 & 0 & -\sinh \phi \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\sinh \phi & 0 & 0 & \cosh \phi \end{bmatrix}$$

We will show that this matrix represents the Lorentz Transformation of  $(x, y, z, t)$ .

Using  $X^1 = AX$ , where,  $X = \begin{bmatrix} x \\ y \\ z \\ ct \end{bmatrix}$ , one can write

$$\begin{bmatrix} x^1 \\ y^1 \\ z^1 \\ t^1 \end{bmatrix} = \begin{bmatrix} \cosh \phi & 0 & 0 & -\sinh \phi \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\sinh \phi & 0 & 0 & \cosh \phi \end{bmatrix} \begin{bmatrix} x \\ y \\ z \\ t \end{bmatrix}$$

From these, we get

$$x^1 = x \cosh \phi - ct \sinh \phi, \quad y^1 = y, \quad z^1 = z, \quad ct^1 = -x \sinh \phi + ct \cosh \phi$$

Now, putting the values of  $\cosh \phi$  and  $\sinh \phi$ , we get

$$x^1 = a[x - vt], \quad y^1 = y, \quad z^1 = z, \quad t^1 = a \left[ t - \frac{xv}{c^2} \right]$$

Hence, the matrix  $A$  expressed in hyperbolic form represents the Lorentz Transformation of  $(x, y, z, t)$ .

**Note 4:** Lorentz Transformation of  $(x, y, z, t)$  can be written in the matrix form as

$$A = \{A_{ij}\} = \begin{bmatrix} a & 0 & 0 & ia\beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -ia\beta & 0 & 0 & a \end{bmatrix}$$

where  $\beta = \frac{v}{c}$  and  $a = \frac{1}{\sqrt{1-\beta^2}}$ . Here,  $v$  is the parameter and  $|A_{ij}| = 1$ .

**Theorem 5.3** *Lorentz Transformations form a group.*

*Proof* We know Lorentz Transformation of  $(x, y, z, t)$  can be written in the matrix form as

$$L = \{A_{ij}\} = \begin{bmatrix} a & 0 & 0 & ia\beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -ia\beta & 0 & 0 & a \end{bmatrix}$$

where  $\beta = \frac{v}{c}$  and  $a = \frac{1}{\sqrt{1-\beta^2}}$ . Here,  $v$  is the parameter and  $|A_{ij}| = 1$ .

Since  $v$  is a parameter, therefore, for two different velocities,  $v_1$  and  $v_2$ , we have  $L_1$  and  $L_2 \in L$  where,

$$L_1 = \begin{bmatrix} \frac{1}{\sqrt{1-(\frac{v_1}{c})^2}} & 0 & 0 & \frac{i(\frac{v_1}{c})}{\sqrt{1-(\frac{v_1}{c})^2}} \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\frac{i(\frac{v_1}{c})}{\sqrt{1-(\frac{v_1}{c})^2}} & 0 & 0 & \frac{1}{\sqrt{1-(\frac{v_1}{c})^2}} \end{bmatrix}$$

$$L_2 = \begin{bmatrix} \frac{1}{\sqrt{1-(\frac{v_2}{c})^2}} & 0 & 0 & \frac{i(\frac{v_2}{c})}{\sqrt{1-(\frac{v_2}{c})^2}} \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\frac{i(\frac{v_2}{c})}{\sqrt{1-(\frac{v_2}{c})^2}} & 0 & 0 & \frac{1}{\sqrt{1-(\frac{v_2}{c})^2}} \end{bmatrix}$$

Now,

$$L_1 L_2 = \begin{bmatrix} \frac{1}{\sqrt{1-(\frac{w}{c})^2}} & 0 & 0 & \frac{i(\frac{w}{c})}{\sqrt{1-(\frac{w}{c})^2}} \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\frac{i(\frac{w}{c})}{\sqrt{1-(\frac{w}{c})^2}} & 0 & 0 & \frac{1}{\sqrt{1-(\frac{w}{c})^2}} \end{bmatrix} \in L$$

where,

$$w = \frac{v_1 + v_2}{1 + \frac{v_1 v_2}{c^2}}$$

Hence, closure property holds good.

We know matrix multiplication is associative.

For  $v=0$ , we get,

$$L_0 = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} \in L$$

Thus  $L_0$  is an identity element.

Now, consider  $L_2$  is the inverse of  $L_1$ . Then  $L_1 L_2 = L_0$  implies

$$\begin{bmatrix} \frac{1}{\sqrt{1-(\frac{w}{c})^2}} & 0 & 0 & \frac{i(\frac{w}{c})}{\sqrt{1-(\frac{w}{c})^2}} \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\frac{i(\frac{w}{c})}{\sqrt{1-(\frac{w}{c})^2}} & 0 & 0 & \frac{1}{\sqrt{1-(\frac{w}{c})^2}} \end{bmatrix} = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

Hence, we get  $w = 0$ , i.e.

$$\frac{v_1 + v_2}{1 + \frac{v_1 v_2}{c^2}} = 0 \Rightarrow v_2 = -v_1$$

In other words,

$$L_2 = \begin{bmatrix} \frac{1}{\sqrt{1-(\frac{v_1}{c})^2}} & 0 & 0 & -\frac{i(\frac{v_1}{c})}{\sqrt{1-(\frac{v_1}{c})^2}} \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ \frac{i(\frac{v_1}{c})}{\sqrt{1-(\frac{v_1}{c})^2}} & 0 & 0 & \frac{1}{\sqrt{1-(\frac{v_1}{c})^2}} \end{bmatrix} \in L$$

Thus, inverse element exists in L [here,  $v_1$  is replaced by  $-v_1$ ].

Therefore, Lorentz Transformations form a group.

*Example 5.1* Show two successive Lorentz Transformation with velocity parameter  $\beta_i = \frac{1}{\sqrt{1-v_i^2/c^2}}$ ,  $i = 1, 2$  equivalent to a single Lorentz Transformation with velocity parameter  $\beta = \beta_1\beta_2(1 + v_1v_2/c^2)$ , where  $v_1$  and  $v_2$  are two velocities with respect to which the two Lorentz Transformation are applied in the same direction.

**Hint:** The Lorentz Transformations with velocity parameter  $v_1$  are

$$x^1 = \frac{x - v_1 t}{\sqrt{1 - \frac{v_1^2}{c^2}}}, \quad y^1 = y, \quad z^1 = z, \quad t^1 = \frac{t - \frac{v_1 x}{c^2}}{\sqrt{1 - \frac{v_1^2}{c^2}}}$$

Thus, the transformation matrix is

$$L(v_1) = \begin{bmatrix} \beta_1 & 0 & 0 & -v_1\beta_1 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\frac{\beta_1 v_1}{c^2} & 0 & 0 & \beta_1 \end{bmatrix}$$

where  $\beta_1 = \frac{1}{\sqrt{1 - \frac{v_1^2}{c^2}}}$ .

Similarly, Lorentz Transformation matrix with velocity parameter  $v_2$

$$L(v_2) = \begin{bmatrix} \beta_2 & 0 & 0 & v_2\beta_2 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\frac{\beta_2 v_2}{c^2} & 0 & 0 & \beta_2 \end{bmatrix}$$

where  $\beta_2 = \frac{1}{\sqrt{1 - \frac{v_2^2}{c^2}}}$ .

Two successive Lorentz Transformation is obtained by multiplying the Transformation matrices of these transformation with each other. Hence

$$\begin{aligned} L(v_1)L(v_2) &= \begin{bmatrix} \beta_1 & 0 & 0 & -v_1\beta_1 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\frac{\beta_1 v_1}{c^2} & 0 & 0 & \beta_1 \end{bmatrix} \begin{bmatrix} \beta_2 & 0 & 0 & -v_2\beta_2 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\frac{\beta_2 v_2}{c^2} & 0 & 0 & \beta_2 \end{bmatrix} \\ &= \begin{bmatrix} \beta_1\beta_2\left(1 + \frac{v_1v_2}{c^2}\right) & 0 & 0 & -(v_1 + v_2)\beta_1\beta_2 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\frac{(v_1+v_2)\beta_1\beta_2}{c^2} & 0 & 0 & \beta_1\beta_2\left(1 + \frac{v_1v_2}{c^2}\right) \end{bmatrix} \end{aligned}$$

$$= \begin{bmatrix} \beta & 0 & 0 & -V\beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\frac{V\beta}{c^2} & 0 & 0 & \beta \end{bmatrix} = L(V)$$

where,  $\beta = \beta_1\beta_2 \left(1 + \frac{v_1v_2}{c^2}\right)$  and  $V = \frac{v_1+v_2}{1+\frac{v_1v_2}{c^2}}$

*Example 5.2* Prove that the Lorentz Transformation of  $x$  and  $t$  can be written in the following forms:

$$ct^1 = -x \sinh \phi + ct \cosh \phi, \quad x^1 = x \cosh \phi - ct \sinh \phi$$

where  $\phi$  is defined by  $\phi = \tanh^{-1} \frac{v}{c}$ .

Establish also the following version of these equations

$$ct^1 + x^1 = e^{-\phi}(ct + x), \quad ct^1 - x^1 = e^{-\phi}(ct - x), \quad e^{2\phi} = \frac{1 + \frac{v}{c}}{1 - \frac{v}{c}}$$

**Hint:**

$$\phi = \tanh^{-1} \frac{v}{c} \implies \tanh \phi = \frac{v}{c}$$

Hence,

$$\cosh \phi = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}, \quad \sinh \phi = \frac{\frac{v}{c}}{\sqrt{1 - \frac{v^2}{c^2}}}$$

Now,

$$x^1 = \frac{x - vt}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{x}{\sqrt{1 - \frac{v^2}{c^2}}} - \frac{ct\left(\frac{v}{c}\right)}{\sqrt{1 - \frac{v^2}{c^2}}} = x \cosh \phi - ct \sinh \phi$$

Similarly,

$$ct^1 = \frac{ct - \frac{vx}{c}}{\sqrt{1 - \frac{v^2}{c^2}}} = -x \sinh \phi + ct \cosh \phi$$

Now,

$$\begin{aligned} ct^1 + x^1 &= (ct + x)(\cosh \phi - \sinh \phi) \\ &= (ct + x) \left[ \frac{1}{2}(e^\phi + e^{-\phi}) - \frac{1}{2}(e^\phi - e^{-\phi}) \right] = e^{-\phi}(ct + x) \end{aligned}$$

$$ct^1 - x^1 = (ct - x)(\cosh \phi + \sinh \phi) = e^\phi (ct - x)$$

Also,

$$\tanh \phi = \frac{v}{c} \implies \frac{\sinh \phi}{\cosh \phi} = \frac{v}{c} \implies \frac{e^\phi - e^{-\phi}}{e^\phi + e^{-\phi}} = \frac{v}{c} \text{ etc}$$

*Example 5.3* Show that the transformations

$$x^1 = x \cosh \phi - ct \sinh \phi, \quad y^1 = y, \quad z^1 = z, \quad ct^1 = -x \sinh \phi + ct \cosh \phi$$

preserve the line element.

**Hint:** We know the line element is

$$\begin{aligned} ds^{12} &= -c^2 dt^2 + dx^2 + dy^2 + dz^2 \\ &= -(-dx \sinh \phi + c dt \cosh \phi)^2 + (dx \cosh \phi - c dt \sinh \phi)^2 + (dy)^2 + (dz)^2 \\ &= -c^2 dt^2 + dx^2 + dy^2 + dz^2 = ds^2 \end{aligned}$$

*Example 5.4* Express the law of addition of parallel velocities in terms of the parameter  $\theta$ .

**Hint:** Suppose a particle is moving with speed  $V$  along  $x$  direction. Let  $\frac{V}{c} = \tanh \Theta$ . The velocity of the particle as measured from  $S^1$  frame which is moving with velocity  $v$  relative to  $S$  frame in  $x$  direction is given by

$$V^1 = \frac{V - v}{1 - \frac{vV}{c^2}}$$

Now, we put  $\frac{V^1}{c} = \tanh \Theta^1$  and  $\frac{v}{c} = \tanh \theta$  in the above expression, we find

$$\Theta^1 = \Theta - \theta$$

So, the law of addition of parallel velocities becomes the addition of  $\theta$  parameter.



# Chapter 6

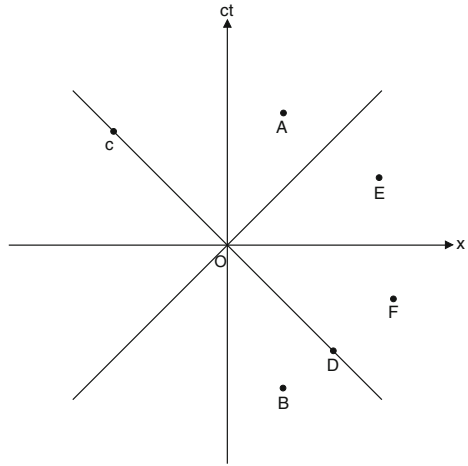
## Geometric Interpretation of Spacetime

### 6.1 Spacetime Diagrams

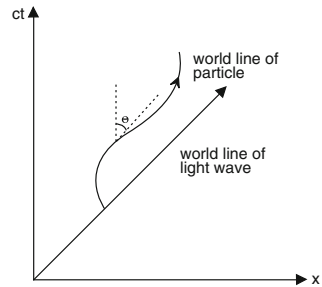
Various types of interval as well as trajectories of the particles in Minkowski's four-dimensional world can be described by some diagrams which are known as spacetime diagrams. To draw the diagram, one has to use some specific inertial frame  $S$  in which one axis indicates the space coordinate  $x$  and other effectively the time axis. However, the time axis is taken as  $ct$  instead of  $t$  so that the coordinates will have the same physical dimensions. We don't need to specify the second and third space coordinates axes, i.e.  $y$  and  $z$  axes. These can be imagined such as both coordinates axes are at right angles to  $ct$  and to  $x$  axis. To visualize the spacetime diagram, we draw the two  $45^\circ$  lines [ $\tan \alpha = \frac{dx}{d(ct)} = \frac{u}{c}$  and we must have  $u < c$  for a material particle; the world line for light signal, i.e.  $u = c$ , is a straight line making an angle  $45^\circ$  with the axes], which are trajectories of light signals through  $O$  ( these light signals pass through the point  $R = 0$  at time  $t = 0$ ;  $R = (x, y, z)$ ). If one includes another axis too, then the set of all such trajectories form a right circular double cone around time axis, with apex at  $O$ , called the *light cone* through  $O$ . However, in general, we just ignore  $y$  and  $z$  axes and the resulting light cone consists of two sloping lines.

One can signify any event by a point on the diagram. Time-like events, i.e. the events which are time-like separated from the apex  $O$  lie inside the light cone, like  $A$  and  $B$  (see Fig. 6.1). Since the sign of  $t$  is invariant, therefore, points in the upper half of the cone possess positive time coordinates and known as inside future light cone. Here, the event  $A$  with  $t_A > 0$  lies inside future light cone. The point  $B$  with  $t_B < 0$  is said to be inside past light cone. Light like events lie on the light cone and as before one can distinguish between future and past; for instance  $C$  lies on the future light cone and  $D$  lies on the past light cone. The region outside the cone consists of space-like events like  $E, F$ . Since time coordinate of space-like event depends on choice of the reference frame, it is usually immaterial whether space-like points lie above or below the  $x$  axis. The world lines on the diagram are the trajectories of the particles. World lines of massless particle (light-like) lie on the light cone and world

**Fig. 6.1** Light cone



**Fig. 6.2** World line of accelerated particle



lines of the massive particles (time-like) lie within the light cone. They are restricted by the speed limit  $c$ .

Let us consider a particle is moving with constant velocity  $u$  in  $x$  direction. Then its world line is a straight line making an angle  $\alpha$  to the time axis where  $|\tan \alpha| = |\frac{u}{c}| < 1$  so that  $|\alpha| < 45^\circ$ . If the particle's velocity  $u(t)$  changes during the motion, i.e. particle is moving with an acceleration, then its world line will be curved rather than a straight line (see Fig. 6.2). Here, the curve makes the angle with the time axis  $\alpha(t) = \tan^{-1} \left( \frac{u(t)}{c} \right)$  that varies but always remains between  $\pm 45^\circ$ . Therefore, it is possible for any signal from  $O$  to reach any event in or on the future light cone (like  $A$  or  $C$ ). Likewise signals will be received by  $O$  from any event in or on the past light cone. Note that World lines passing through  $O$  are confined within the light cone through  $O$ . World lines can not leave the light cone because entry or exit would require  $\alpha(t) > 45^\circ$ , i.e. speed should be greater than  $c$ , which is impossible. Hence, it is not possible to connect between  $O$  and any other event outside the light cone.

### 6.2 Some Possible and Impossible World Lines

Let us consider, Spacetime diagram with the same origin, axes and the light cone as the previous one (see Fig. 6.3). The particle trajectory from O to A is possible because it is within the future light, moreover, its inclination to the time axis is always less than 45°. This indicates that the speed of the particle along the trajectory is always less than c. The trajectory from O to B is impossible as it crosses the light cone.

This means that the speed of the particle should exceed c. Again the trajectory  $E_1E_2E_3O$  is not possible in spite it lies entirely within the light cone. The reason is that near  $E_1$  and  $E_3$  its slope indicates speeds greater than c (in fact, at the points  $E_1$  and  $E_3$  speed of the particle would be infinite).

Remember that in a different inertial frame  $S^1$ , the spacetime diagram looks exactly same as in Fig. 6.3, except that its axes are labelled by  $ct^1$  and  $x^1$ . Two or more events can be entered on both diagrams with different coordinates. However, spacetime separations (intervals)  $\rho_{12}$  between two events remain invariant. In particular, if events lie inside the future light cone of one diagram, then these remain inside of the future light cone of the other.

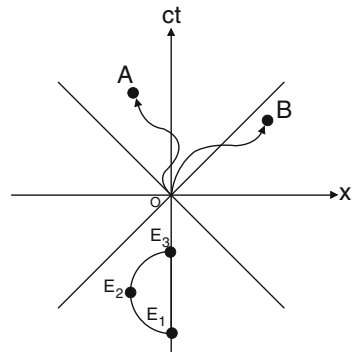
**Note 1:** We know the quantity  $s^2 = x^2 + y^2 + z^2 - c^2t^2$ , called the square of the four-dimensional distance between the event  $(x^\mu)$  and the origin  $(0, 0, 0, 0)$ , remains invariant under Lorentz Transformation. Now, the trajectories of the points whose distance from the origin is zero form a surface described by the equation

$$s^2 = x^2 + y^2 + z^2 - c^2t^2 = 0$$

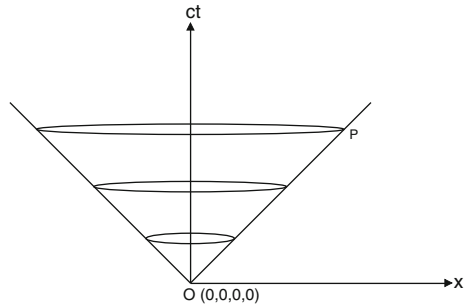
This surface is light cone. This equation describes the propagation of a spherical light wave from the origin  $(0, 0, 0)$  at  $t = 0$ . The spacetime, i.e. (3+1) space is divided in to two invariant separated domains  $S_t$  and  $S_s$  by the light cone (see Fig. 6.4). The region  $S_t$  is defined by  $s^2 = x^2 + y^2 + z^2 - c^2t^2 < 0$ .

From this, we get,

**Fig. 6.3** Possible and impossible events



**Fig. 6.4** The region  $S_t$  is within the light cone and the region  $S_s$  is outside the light cone



$$\frac{s^2}{t^2} = \left[ \frac{\sqrt{x^2 + y^2 + z^2}}{t} \right]^2 - c^2 = v^2 - c^2 < 0$$

This indicates trajectories of the massive particle lie always inside the light cone. Thus, world lines passing through O always be confined within the cone. They cannot enter in the region  $S_s$ , which is defined by  $s^2 = x^2 + y^2 + z^2 - c^2t^2 > 0$ , because entry or exist would require speed greater than  $c$  ( $S_s : \frac{s^2}{t^2} = v^2 - c^2 > 0$ ). Let an event occurs at  $t = 0, x = 0$ , then it can causally connect any point inside the positive light cone.

**Note 2:** Photon has a null trajectory as

$$ds^2 = c^2dt^2 - dx^2 - dy^2 - dz^2$$

or

$$\left( \frac{ds}{dt} \right)^2 = c^2 - \left[ \left( \frac{dx}{dt} \right)^2 + \left( \frac{dy}{dt} \right)^2 + \left( \frac{dz}{dt} \right)^2 \right] = c^2 - v^2$$

For photon,  $v = c$ ,  $\left( \frac{ds}{dt} \right)^2 = 0$ , i.e. *light-like*:  $ds^2 = 0$ .

**Remember** For time-like  $ds^2 > 0$  and space-like:  $ds^2 < 0$ .

**Note 3:** Can an event at  $t = 0, x = 0$  affect the event occurred at P in the future light cone region ?

Yes, because, P is a point which is in future. So an event affects the event occurred at P. But the reverse is not, i.e. event at P can not affect the event at  $t = 0, x = 0$ , since  $t$  is always positive. All the events in the future light cone region occur after the event at  $t = 0, x = 0$ . Light signals emitted at  $t = 0, x = 0$  travel out along the boundary of the light cone and affect the events on the light cone.

**Note 4:** A collision between two particles corresponds to an intersection of their world lines.

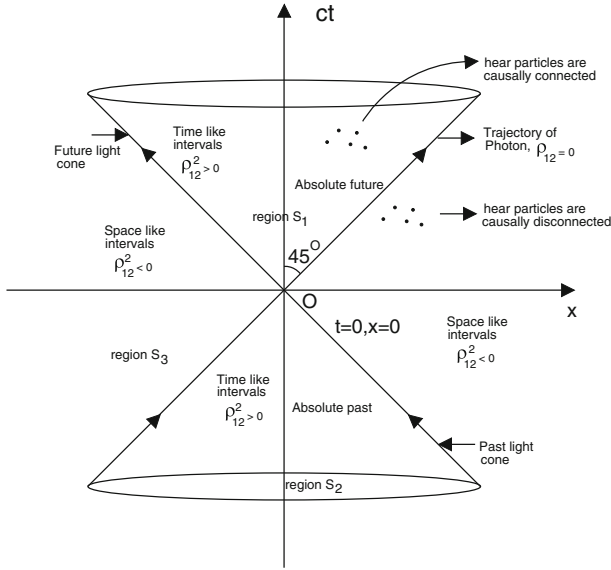


Fig. 6.5 Geometrical representation of spacetime diagram

### 6.3 Importance of Light Cone

The light cone structure is important because inside the light cone particles are causally connected and outside the light cone particles are causally disconnected. Thus, it is very significant in the sense that geometrical representation of spacetime diagram, i.e. light cone is completely consistent with the causality principle. For example, the events within the light cone have a cause-and-effect relationship between events and events outside the light cone have no cause-and-effect relationship between events (Fig. 6.5).

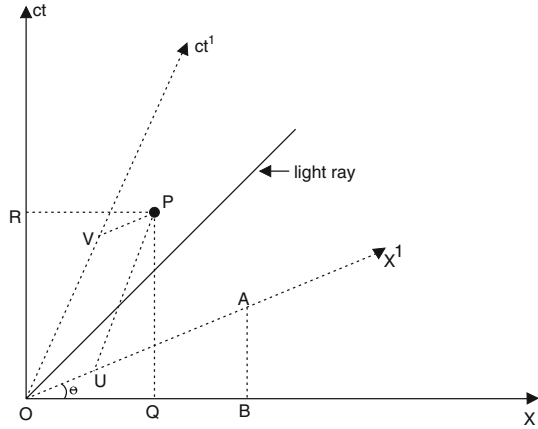
### 6.4 Relationship Between Spacetime Diagrams in S and S<sup>1</sup> Frames

Let us consider two frames of reference S and S<sup>1</sup> and S<sup>1</sup> is moving with constant velocity v along x direction. Let S frame consists of coordinate axes ct and x so that a light ray has the slope  $\frac{\pi}{4}$

$$\left[ \frac{x}{ct} = \tan \theta \implies \tan \theta = \frac{1}{c} \left( \frac{x}{t} \right) = \frac{c}{c} = 1 \implies \theta = \frac{\pi}{4} \right].$$

Now, we try to draw the ct<sup>1</sup> and x<sup>1</sup> axes of S<sup>1</sup>. Here, the equation of x<sup>1</sup> is ct<sup>1</sup> = 0. Using Lorentz Transformation, we get

**Fig. 6.6** Two reference frames have the same origin



$$\frac{c \left( t - \frac{xv}{c^2} \right)}{\sqrt{1 - \frac{v^2}{c^2}}} = 0 \implies ct = \left( \frac{v}{c} \right) x$$

Thus  $x^1$  axis is a straight line,  $ct = \left( \frac{v}{c} \right) x$  with slope  $\left( \frac{v}{c} \right) < 1$ .

Similarly, equation of  $ct^1$  axis,  $x^1 = 0$  gives  $ct = \left( \frac{c}{v} \right) x$ . Hence,  $ct^1$  axis is a straight line,  $ct = \left( \frac{c}{v} \right) x$  having slope  $\left( \frac{c}{v} \right) > 1$ . The fixed points in  $S^1$  frame lie in world line parallel to  $O(ct^1)$  axis. The events with fixed time in  $S^1$  frame lie in world line parallel to  $Ox^1$  axis and are known as lines of simultaneity in  $S^1$  frame. Let P be any event (see Fig. 6.6). Then, the coordinates of P with respect to S system are  $(ct, x) = (OR, OQ)$  and  $(ct^1, x^1) = (OV, OU)$  with respect to  $S^1$  system. Suppose A be any point on  $x^1$  axis having length from O is 1. Therefore, the coordinates of A with respect to  $S^1$  system are  $(ct^1, x^1) = (0, 1)$ . The coordinates of A with respect to S system are  $(ct, x) = \left( \frac{av}{c}, a \right)$  where,  $a = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$

$$\left[ \text{by L.T : } ct^1 = \frac{c \left( t - \frac{xv}{c^2} \right)}{\sqrt{1 - \frac{v^2}{c^2}}} = 0 \quad \text{and} \quad x^1 = \frac{(x - vt)}{\sqrt{1 - \frac{v^2}{c^2}}} = 1 \implies x = a, ct = \frac{c \left( t - \frac{xv}{c^2} \right)}{\sqrt{1 - \frac{v^2}{c^2}}} \right]$$

Hence the distance OA is given by

$$OA = \sqrt{(c^2 t^2 + x^2)} = a \sqrt{1 + \frac{v^2}{c^2}}$$

Thus, scale along the primed axes is greater than the scale along the unprimed by a factor

$$\left[ \frac{1 + \frac{v^2}{c^2}}{\sqrt{1 - \frac{v^2}{c^2}}} \right].$$

## 6.5 Geometrical Representation of Simultaneity, Space Contraction and Time Dilation

Let us consider two frames of reference  $S$  and  $S^1$  and  $S^1$  is moving with constant velocity  $v$  along  $x$  direction. Let  $S$  frame consists of coordinate axes  $ct \equiv M$  &  $x$  and  $S^1$  frame consists of coordinate axes  $ct^1 \equiv M^1$  &  $x^1$ .

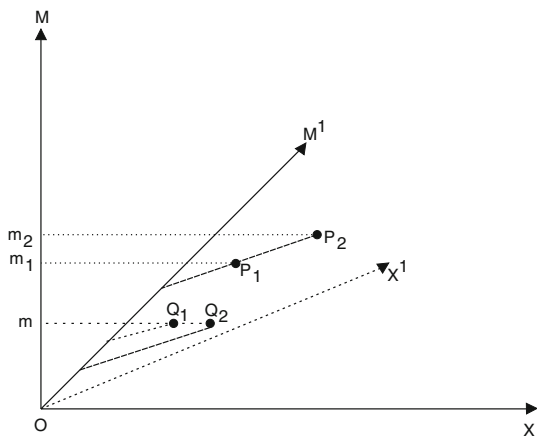
### 6.5.1 Simultaneity

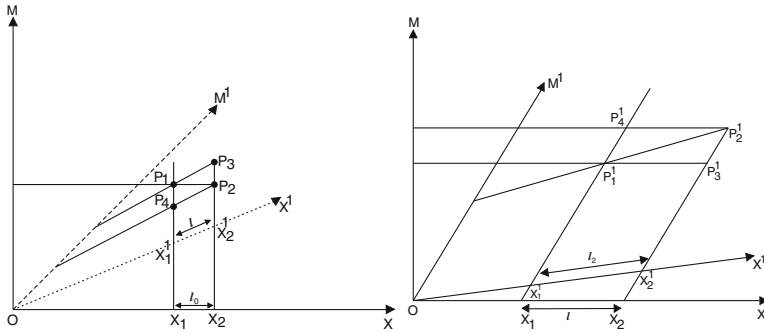
In  $S^1$  frame, two events will be simultaneous if they have the same time coordinate  $m^1$ . Thus, simultaneous events lie on a line parallel to  $x^1$  axis. In the Fig. 6.7,  $P_1$  and  $P_2$  are simultaneous in  $S^1$  system but not simultaneous in  $S$  system as they will be occurring at different times  $m_1$  and  $m_2$  in  $S$  system (i.e. they have different time coordinates). Similarly,  $Q_1$  and  $Q_2$  are two simultaneous events in  $S$  frame but they are separated in times in  $S^1$  frame (i.e. they have different time coordinates  $m_1^1$  and  $m_2^1$  in  $S^1$  frame).

### 6.5.2 Space Contraction

Consider a rod having end points at  $x = x_1$  and  $x = x_2$  rest in  $S$  frame. As time elapses, the world line of each end points traces out a vertical line parallel to the  $M$  axis. The length of the rod ( $l_0$ ) is the distance between its end points measured simultaneously, i.e.  $l_0 = x_2 - x_1$ . In  $S$  system, the length is the distance in  $S$  between the intersections of the world lines (these are parallel to  $M$  axis) with the  $x$  axis or

Fig. 6.7 Simultaneous events





**Fig. 6.8** Lengths are contracted

any line parallel to  $x$  axis. Actually, these intersecting points represent simultaneous events in  $S$ . Like in  $S$  frame, we measure the rod in  $S^1$  frame as the distance between end points measured simultaneously which is the separation of the intersections of the world lines with  $x^1$  axis or any line parallel to the  $x^1$  axis, i.e.  $l = x_2^1 - x_1^1$ . Here, these intersecting points represent simultaneous events in  $S^1$  (see Fig. 6.8).

The length of rod in  $S^1$  system (moving) is  $l = l_0 \sqrt{1 - \frac{v^2}{c^2}}$ , which is clearly less than  $l_0$ . Note that if events are simultaneous in one reference frame, then they are not simultaneous in other frame. To measure the length, observer in  $S$  uses  $P_1$  and  $P_2$ , while the observer in  $S^1$  system uses either  $P_2$  and  $P_4$  or  $P_1$  and  $P_3$ .

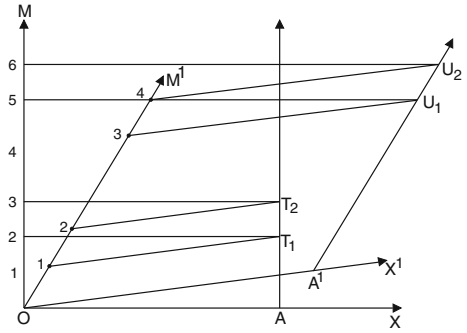
Alternatively, let us consider a rod rest in  $S^1$  frame having end points  $x^1$  and  $x^2$ . The world lines of its end points are parallel to  $M^1$  axis. Its rest length is  $l_0 = x_2^1 - x_1^1$ . Observer in  $S$  system, finds the rod is in motion and to him, the measured length is the distance in  $S$  system between intersections of these world lines with  $x$  axis or any line parallel to  $x$  axis. Clearly, the length of rod in motion is observed by the observer in  $S$  system as  $l = l_0 \sqrt{1 - \frac{v^2}{c^2}}$ , which is less than  $l_0$ .

### 6.5.3 Time Dilation

Let us consider a clock at rest at  $A$  in  $S$  system which ticks off units of time. Since time interval in system  $S$  between ticks being unity, let  $T_1$  and  $T_2$  be two events of ticking at  $m = 2$  and  $m = 3$ . The solid vertical line through  $A$ , i.e.  $T_1 T_2$  is the world line corresponding to the clock in  $S$  system. In  $S^1$  system clock ticks in different places. Therefore, to measure the times interval between  $T_1$  and  $T_2$  in  $S^1$  system, two clocks must be used. The difference in reading of these clocks in  $S^1$  is the difference in times between  $T_1$  and  $T_2$  as measured in  $S^1$  system. Figure 6.9 confirms that the time interval between  $T_1$  and  $T_2$  in  $S^1$  system is greater than unity. Hence, relative to  $S^1$  the clock at rest in system  $S$  appears slowed down. Here,  $S$  clock registered unit time, however,  $S^1$  registered a time greater than one unit.



**Fig. 6.9** Time intervals are increased



Alternatively, a clock at rest at  $A^1$  in  $S^1$  system which ticks off units of time at  $U_1$  and  $U_2$ . Here, the line parallel to  $M^1$  axis through  $A^1$ , i.e. the line  $U_1U_2$  is the world line corresponding to the clock in system  $S^1$ . The time interval to the observer in S frame (relative to which the clock is in motion) is greater than unity.

# Chapter 7

## Relativistic Velocity and Acceleration

### 7.1 Relativistic Velocity Addition

Let us consider three frames of reference  $S$ ,  $S^1$  and  $S^2$ .  $S^1$  is moving with constant velocity  $u$  relative to  $S$  in  $x$  direction and  $S^2$  is moving with constant velocity  $v$  relative to  $S^1$  also in  $x$  direction (see Fig. 7.1). We want to find the transformation equations between  $S^2$  and  $S$ . In Newtonian mechanics, the relative velocity is

$$w = u + v$$

However, we try to see what will happen in special theory of relativity. The transformation equations among  $S$ ,  $S^1$  and  $S^2$  are given by

$$x^1 = \frac{1}{\sqrt{1 - (u/c)^2}} [x - ut], \quad y^1 = y, \quad z^1 = z, \quad t^1 = \frac{1}{\sqrt{1 - (u/c)^2}} \left[ t - \frac{ux}{c^2} \right] \tag{7.1}$$

$$x^2 = \frac{1}{\sqrt{1 - (v/c)^2}} [x^1 - vt^1], \quad y^2 = y^1, \quad z^2 = z^1, \quad t^2 = \frac{1}{\sqrt{1 - (v/c)^2}} \left[ t^1 - \frac{vx^1}{c^2} \right] \tag{7.2}$$

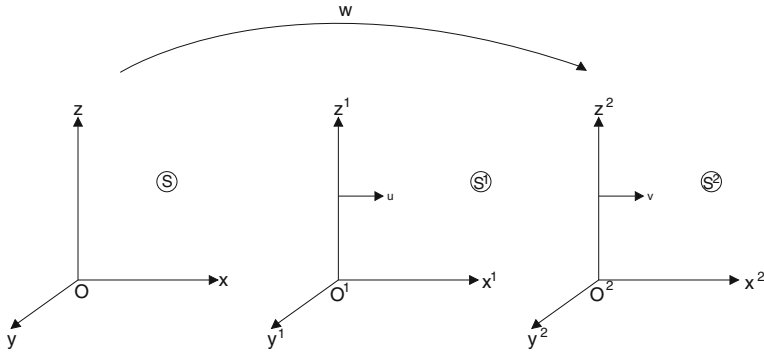
Now, putting the values of  $x^1, t^1$  in the first expression of Eq. (7.2), one gets

$$x^2 = \frac{1}{\sqrt{1 - (v/c)^2}} \frac{1}{\sqrt{1 - (u/c)^2}} \left[ 1 + \frac{uv}{c^2} \right] \left[ x - t \left\{ \frac{u + v}{1 + uv/c^2} \right\} \right] \tag{7.3}$$

Let us choose,

$$w = \frac{u + v}{1 + \frac{uv}{c^2}} \tag{7.4}$$

Using the expression of  $w$  in (7.3), we get



**Fig. 7.1** Addition of velocities

$$x^2 = \frac{x - wt}{\sqrt{1 - (w/c)^2}} \quad (7.5)$$

Obviously,

$$y^2 = y, \quad z^2 = z$$

Again, putting the values of  $x^1, t^1$  in the last expression of Eq. (7.2), one gets as before,

$$t^2 = \frac{t - \frac{wx}{c^2}}{\sqrt{1 - (w/c)^2}} \quad (7.6)$$

The Eq. (7.4) is known as the Einstein law of addition of velocities.

From Eq. (7.4), we have

$$w = \frac{u + v}{1 + \frac{uv}{c^2}} = c \left[ 1 - \frac{(1 - \frac{u}{c})(1 - \frac{v}{c})}{1 + \frac{uv}{c^2}} \right]$$

If  $u < c$  and  $v < c$ , then  $w < c$ . This shows that the composition of velocities which are separately less than  $c$  cannot be added to give a velocity greater than  $c$ . If we put  $v = c$ , i.e. if we consider that a photon travelling with velocity  $c$  along  $x$  direction, then,  $w = c$ . This indicates that if a particle is moving with velocity  $c$ , it will be observed  $c$  in all inertial frames whatever their relative velocity is. This confirms the validity of the second postulate of special theory of relativity that the velocity of light is an absolute quantity independent of the motion of the reference frame. Even when,  $v = u = c$ , then,  $w = c$ .

The relativistic law of addition of velocities reduces to the law of Newtonian mechanics for small values of  $u$  and  $v$  in comparison to the velocity of light, i.e. if  $uv \ll c^2$ , then  $w = u + v$ .

**Note 1:** If one takes,  $u = v = \frac{4c}{5}$ , then relativistic law of addition of velocities gives,  $w = \frac{40c}{41} < c$ . However, according to Newtonian mechanics law of addition of velocities gives,  $w = \frac{8c}{5} > c$ .

**Note 2:** The general Relativistic Addition velocity  $w$  can be written as

$$w = \frac{c\sqrt{c^2(\vec{u} + \vec{v}) \cdot (\vec{u} + \vec{v}) - u^2v^2 + (\vec{u} \cdot \vec{v})^2}}{c^2 - (\vec{u} \cdot \vec{v})}$$

This reduces to the classical formula  $w = |\vec{u} + \vec{v}|$  when  $u, v \ll c$ .

## 7.2 Relativistic Velocity Transformations

Let a frame of reference  $S^1$  is moving with constant velocity  $v$  relative to S frame along x direction. Assume a particle at point P is moving with velocity  $u^1$  relative to  $S^1$  frame and  $u$  its velocity relative to S frame. The space and time coordinates of P in  $S^1$  and S frames are  $(x^1, y^1, z^1, t^1)$  and  $(x, y, z, t)$ , respectively. The component of velocity vector  $\vec{u}$  relative to S frame are

$$u_x = \frac{dx}{dt}, \quad u_y = \frac{dy}{dt}, \quad u_z = \frac{dz}{dt} \quad (7.7)$$

and

$$u = (u_x^2 + u_y^2 + u_z^2)^{\frac{1}{2}}$$

Similarly, the components of velocity vector  $\vec{u}^1$  relative to  $S^1$  frame are

$$u_x^1 = \frac{dx^1}{dt^1}, \quad u_y^1 = \frac{dy^1}{dt^1}, \quad u_z^1 = \frac{dz^1}{dt^1} \quad (7.8)$$

and

$$u^1 = [(u_x^1)^2 + (u_y^1)^2 + (u_z^1)^2]^{\frac{1}{2}}$$

Now, from Lorentz Transformations, we get

$$dx^1 = a(dx - vdt), \quad dy^1 = dy, \quad dz^1 = dz, \quad dt^1 = a\left(dt - \frac{vdx}{c^2}\right) \quad (7.9)$$

where,  $a = \frac{1}{\sqrt{1-v^2/c^2}}$ . Hence,

$$u_x^1 = \frac{dx^1}{dt^1} = \frac{u_x - v}{1 - vu_x/c^2} \quad (7.10)$$

$$u_y^1 = \frac{dy^1}{dt^1} = \frac{u_y \sqrt{1 - v^2/c^2}}{1 - vu_x/c^2} \quad (7.11)$$

$$u_z^1 = \frac{dz^1}{dt^1} = \frac{u_z \sqrt{1 - v^2/c^2}}{1 - vu_x/c^2} \quad (7.12)$$

$$(u^1)^2 = [(u_x^1)^2 + (u_y^1)^2 + (u_z^1)^2] = \frac{(u_x - v)^2 + (u^2 - u_x^2)(1 - v^2/c^2)}{(1 - vu_x/c^2)^2} \quad (7.13)$$

**Note 3:** The general Lorentz Transformation of velocity vector  $\vec{u}^1$  can be written as

$$\vec{u}^1 = \frac{\sqrt{1 - v^2/c^2} \vec{u} + \frac{(\vec{u} \cdot \vec{v}) \vec{v}}{v^2} \left[ 1 - \sqrt{1 - v^2/c^2} \right] - \vec{v}}{1 - \frac{(\vec{u} \cdot \vec{v})}{c^2}}$$

Since,  $\vec{v}$  parallel to x axis, i.e.  $\vec{v} = (v, 0, 0)$ , therefore,  $(\vec{u} \cdot \vec{v}) = u_x v$ .

**Note 4:** The inverse transformations can be obtained by replacing  $v$  by  $-v$  and interchanging primed by unprimed in Eqs. (7.10)–(7.12).

$$u_x = \frac{u_x^1 + v}{1 + \frac{vu_x^1}{c^2}}, \quad u_y = \frac{u_y^1 \sqrt{1 - v^2/c^2}}{1 + \frac{vu_x^1}{c^2}}, \quad u_z = \frac{u_z^1 \sqrt{1 - v^2/c^2}}{1 + \frac{vu_x^1}{c^2}} \quad (7.14)$$

**Note 5:** Einstein's law of addition of velocities is a special case of Relativistic velocity. Putting,  $u_x = u$ ,  $u_y = u_z = 0$  in (7.14), one obtains,

$$u = \frac{u^1 + v}{1 + \frac{vu^1}{c^2}}$$

**Note 6:** The transformation of Lorentz Transformation factor can be found from Eq. (7.13) as

$$\sqrt{1 - (u^1/c)^2} = \frac{\sqrt{1 - (v/c)^2} \sqrt{1 - (u/c)^2}}{\left(1 - \frac{uv}{c^2}\right)}$$

### 7.3 Relativistic Acceleration Transformations

Let acceleration of the particle relative to S and  $S^1$  frames are  $\vec{f}$  and  $\vec{f}^1$ , respectively. The components of the acceleration vector relative to S and  $S^1$  frames are, respectively, as

$$f_x = \frac{du_x}{dt}, \quad f_y = \frac{du_y}{dt}, \quad f_z = \frac{du_z}{dt}$$

and

$$f_x^1 = \frac{du_x^1}{dt^1}, f_y^1 = \frac{du_y^1}{dt^1}, f_z = \frac{du_z^1}{dt^1}$$

Using (7.10), we can get,

$$f_x^1 = \frac{du_x^1}{dt^1} = \frac{f_x \left(1 - \frac{v^2}{c^2}\right)^{\frac{3}{2}}}{\left(1 - \frac{vu_x}{c^2}\right)^3} \quad (7.15)$$

This acceleration is known as *Longitudinal Acceleration*.

Similarly, from (7.11) and (7.12) we get,

$$f_y^1 = \frac{du_y^1}{dt^1} = \left(1 - \frac{v^2}{c^2}\right) \left[ \frac{f_y}{\left(1 - \frac{vu_x}{c^2}\right)^2} + \frac{vf_x u_y}{c^2 \left(1 - \frac{vu_x}{c^2}\right)^3} \right] \quad (7.16)$$

$$f_z^1 = \frac{du_z^1}{dt^1} = \left(1 - \frac{v^2}{c^2}\right) \left[ \frac{f_z}{\left(1 - \frac{vu_x}{c^2}\right)^2} + \frac{vf_x u_z}{c^2 \left(1 - \frac{vu_x}{c^2}\right)^3} \right] \quad (7.17)$$

These are known as *Transverse Acceleration*.

**Note 7:** The general Lorentz Transformation of acceleration vector  $\vec{f}^1$  can be written as

$$\vec{f}^1 = \frac{\vec{f}}{a^2 \left(1 - \frac{\vec{u} \cdot \vec{v}}{c^2}\right)^2} + \frac{(\vec{f} \cdot \vec{v}) \left[ \frac{a\vec{u}}{c^2} - (a-1) \frac{\vec{v}}{v^2} \right]}{a^3 \left(1 - \frac{\vec{u} \cdot \vec{v}}{c^2}\right)^3}$$

**Note 8:** For a particle instantaneously rest in S frame, i.e.  $u_x = u_y = u_z = 0$ , we get,

$$f_x^1 = \frac{du_x^1}{dt^1} = f_x \left(1 - \frac{v^2}{c^2}\right)^{\frac{3}{2}}, f_y^1 = \left(1 - \frac{v^2}{c^2}\right) f_y, f_z^1 = \left(1 - \frac{v^2}{c^2}\right) f_z$$

The transformation formula of acceleration indicates that the acceleration is not invariant in special theory of relativity. In other words, Newton's second law of motion is not invariant under Lorentz Transformation. But, it is evident that acceleration is an absolute quantity i.e. all observers agree whether a body is accelerating or not. In other words, the acceleration of a particle is zero in one frame, it is necessarily zero in all other frames.

## 7.4 Uniform Acceleration

In special theory of relativity, an acceleration is said to be *uniform* if for any co-moving frame its value remains the same. Let S and  $S^1$  be two inertial frames with constant relative velocity  $v$  along x direction. Let velocity of a particle relative to S is  $u$ . If  $S^1$  is a co-moving frame, then  $v = u$ . This means relative velocity of the particle and  $S^1$  frame is zero,  $u^1 = 0$ . Now, if the acceleration is a constant  $f^1 = \frac{du^1}{dt^1} = b$  (say, a constant). From Eq. (7.15), one can write,

$$f = \frac{du}{dt} = \frac{f^1 \left(1 - \frac{u^2}{c^2}\right)^{\frac{3}{2}}}{\left(1 + \frac{vu^1}{c^2}\right)^3} \quad (7.18)$$

Now, using  $u^1 = 0$  and  $\frac{du^1}{dt^1} = b$ , we get

$$\frac{du}{dt} = b \left(1 - \frac{u^2}{c^2}\right)^{\frac{3}{2}} \quad (7.19)$$

Solving this equation, one gets

$$u = \frac{dx}{dt} = \frac{b(t - t_0)}{\left[1 + \frac{b^2(t-t_0)^2}{c^2}\right]} \quad (7.20)$$

[initially, particle starts from rest at  $t = t_0$  ]

However, in Newtonian mechanics, if a particle is moving with uniform acceleration, then,

$$\frac{du}{dt} = \text{constant}$$

The path of this particle is parabola.

Further integration of Eq. (7.20) yields (using  $x = x_0$  at  $t = t_0$ )

$$\frac{(x - x_0 + \frac{c^2}{b})^2}{(\frac{c^2}{b})^2} - \frac{(ct - ct_0)^2}{(\frac{c^2}{b})^2} = 1 \quad (7.21)$$

This is an equation of a hyperbola in (x, ct) space.

## 7.5 Relativistic Transformations of the Direction Cosines

Let us consider the direction cosines of the motion of a particle in S frame be  $(m, n, l)$  and corresponding quantities be  $(m^1, n^1, l^1)$  in  $S^1$  which is moving with constant

velocity  $v$  along  $x$  direction. Then

$$(l, m, n) = \frac{(u_x, u_y, u_z)}{\sqrt{u_x^2 + u_y^2 + u_z^2}} \quad (7.22)$$

Similarly,

$$(l^1, m^1, n^1) = \frac{(u_x^1, u_y^1, u_z^1)}{\sqrt{(u_x^1)^2 + (u_y^1)^2 + (u_z^1)^2}} \quad (7.23)$$

Using Eqs. (7.4)–(7.7), one will get,

$$(l^1, m^1, n^1) = \frac{(l - \frac{v}{u}), m\sqrt{1 - (v/c)^2}, n\sqrt{1 - (v/c)^2}}{D} \quad (7.24)$$

where,

$$D = \frac{u^1}{u} \left(1 - \frac{u_x v}{c^2}\right) = \sqrt{\left[1 - 2lv/u + \frac{v^2}{u^2} + (1 - l^2)\frac{v^2}{c^2}\right]}$$

## 7.6 Application of Relativistic Velocity and Velocity Addition Law

### 7.6.1 The Fizeau Effect: The Fresnel's Coefficient of Drag

Now, we will try to give explanation of index of refraction of moving bodies (Fizeau effect). Fizeau performed the experiment by employing an interferometer to measure the velocity of light propagating in water moving at the velocity  $v$ . Here, we consider two inertial frames: one is  $S$  as laboratory and other is  $S^1$ , the moving water. Since, the water has the velocity  $v$ , therefore, one can think,  $S^1$  frame is moving with velocity  $v$  relative to  $S$  frame which is at rest. Suppose  $\mu$  be the refractive index of water; then,  $\frac{c}{\mu}$  is the velocity of light in water relative to  $S^1$ . If  $V$  is the velocity of a light relative to  $S$  (i.e. laboratory), then using Relativistic Velocity Addition law we get

$$V = \frac{u_x + v}{1 + \frac{u_x v}{c^2}} = \frac{\frac{c}{\mu} + v}{1 + \frac{\frac{c}{\mu} v}{c^2}} = \frac{c}{\mu} \left(1 + \frac{v\mu}{c}\right) \left(1 + \frac{v}{c\mu}\right)^{-1}$$

Expanding binomially and neglecting higher powers of  $\frac{v}{c}$ , we get

$$V = \frac{c}{\mu} + v \left(1 - \frac{1}{\mu^2}\right)$$



Coefficient of drag  $\alpha$  is given by

$$\alpha = \frac{v \left(1 - \frac{1}{\mu^2}\right)}{v} = \left(1 - \frac{1}{\mu^2}\right)$$

At first, Fresnel obtained theoretically this expression and is known as Fresnel's coefficient of drag. Fizeau verified this result experimentally.

### 7.6.2 Aberration of Light

Bradley discovered the phenomena of aberration in 1927. The phenomena of aberration results: The speed of light does not depend on the medium of transmission. However, direction of rays depends on the relative motion of the source and observer. Now, we measure the directions of the light ray emitted from a star from two reference frames  $S$  and  $S^1$  where  $S^1$  frame is moving with velocity  $v$  relative to  $S$  frame in  $x$  direction. Since Earth revolves around the sun about its orbit, therefore, we assume, the frame  $S$  fixed in the sun and  $S^1$  in Earth. Let a light from a star  $P$  be observed by the observers  $O$  and  $O^1$  in systems  $S$  and  $S^1$ , respectively. Suppose the observers notice that the directions of light make an angle  $\alpha$  and  $\alpha^1$  with  $x$  axis, respectively (see Fig. 7.2). Here, the light ray lies on the  $xy$ -plane.

Then,

$$u_x = c \cos \alpha, \quad u_y = c \sin \alpha, \quad u_z = 0$$

$$u_x^1 = c \cos \alpha^1, \quad u_y = c \sin \alpha^1, \quad u_z^1 = 0$$

Using, relativistic velocity transformation law (7.10) and (7.11), we get the *relativistic equation of aberration* as,

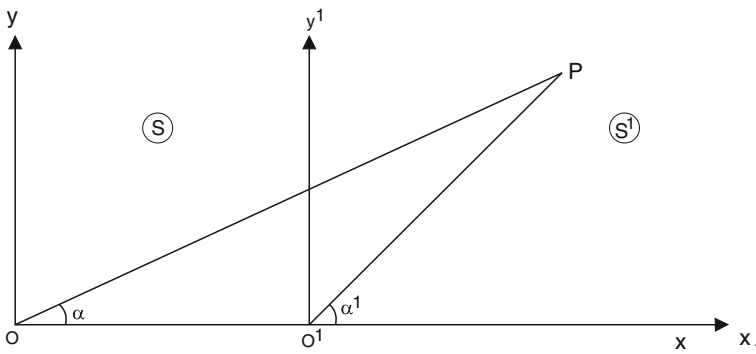


Fig. 7.2 Directions of light make angle  $\alpha$  and  $\alpha^1$  with  $x$  axis

$$\tan \alpha^1 = \frac{u_y}{a(u_x - v)} = \frac{c \sin \alpha}{a(c \cos \alpha - v)} = \frac{\tan \alpha \sqrt{1 - (v/c)^2}}{1 - \frac{v}{c} \sec \alpha}$$

Classical formula from relativistic result:

For,  $v \ll c$  and using  $\alpha^1 = \alpha + \delta\alpha$ , where  $\delta\alpha$  is very small, we get

$$\tan(\alpha + \delta\alpha) = \tan \alpha \left(1 - \frac{v}{c} \sec \alpha\right)^{-1} = \tan \alpha \left(1 + \frac{v}{c} \sec \alpha\right)$$

or

$$\tan(\alpha + \delta\alpha) - \tan \alpha = \frac{\sin \delta\alpha}{\cos(\alpha + \delta\alpha)} = \frac{v \sin \alpha}{c \cos \alpha}$$

Since  $\delta\alpha$  is very, we finally get the *classical value of aberration* as

$$\delta\alpha = \frac{v}{c} \sin \alpha$$

### 7.6.3 Relativistic Doppler Effect

Let a plane wave travelling in the direction  $(l, m, n)$  in S frame is given by

$$\psi = A \exp 2\pi i \left( vt - \frac{lx + my + nz}{\lambda} \right) \quad (7.25)$$

This  $\psi$  satisfies the invariant wave equation

$$\nabla^2 \psi - \frac{1}{c^2} \frac{\partial^2 \psi}{\partial t^2} = 0$$

Suppose another frame of reference  $S^1$  moving with velocity  $U$  in x direction. This plane wave in  $S^1$  is given by

$$\psi^1 = A \exp 2\pi i \left( v^1 t^1 - \frac{l^1 x^1 + m^1 y^1 + n^1 z^1}{\lambda^1} \right) \quad (7.26)$$

Now, using the transformation of  $(x^1, y^1, z^1, t^1)$ ,  $\psi^1$  takes the following form,

$$\psi^1 = A \exp 2\pi i \left[ v^1 a \left( t - \frac{Ux}{c^2} \right) - \frac{l^1 a(x - Ut) + m^1 y + n^1 z}{\lambda^1} \right] \quad (7.27)$$

where  $a = \frac{1}{\sqrt{1 - v^2/c^2}}$ .

Comparing the coefficients of  $x$ ,  $y$ ,  $z$  and  $t$  between Eqs. (7.25) and (7.27), we get,

$$v = a \left( v^1 + \frac{U l^1}{\lambda^1} \right), \quad \frac{l}{\lambda} = a \left( \frac{l^1}{\lambda^1} + \frac{U v^1}{c^2} \right), \quad \frac{m}{\lambda} = \frac{m^1}{\lambda^1}, \quad \frac{n}{\lambda} = \frac{n^1}{\lambda^1} \quad (7.28)$$

For electromagnetic waves, we have

$$v \lambda = v^1 \lambda^1 = c \quad (7.29)$$

From Eqs. (7.28) and (7.29), one can get,

$$\frac{l}{c} = \frac{l^1 + \frac{U}{c}}{c + U l^1} \quad (7.30)$$

If the direction of propagation of the wave makes angles  $\theta$  and  $\theta^1$  with the  $x$  and  $x^1$  axes, therefore,

$$l^1 = \cos \theta^1, \quad l = \cos \theta \quad (7.31)$$

Now, expressing it in terms of angles, we get,

$$\cos \theta = \frac{\cos \theta^1 + \frac{U}{c}}{1 + \frac{U \cos \theta^1}{c}} \quad (7.32)$$

Again, we get

$$\frac{\lambda^1}{\lambda} = \frac{v}{v^1} = \frac{1 + \frac{U \cos \theta^1}{c}}{\sqrt{1 - U^2/c^2}} \quad (7.33)$$

Equation (7.32) gives the *relativistic aberration* and Eq. (7.33) gives *relativistic Doppler effect of light*.

**Note 9:** If the velocity of the source is along the  $x$  axis, then  $\theta = \theta^1 = 0$ , then one gets,

$$v = v^1 \sqrt{\frac{1 + U/c}{1 - U/c}}, \quad \lambda^1 = \lambda \sqrt{\frac{1 + U/c}{1 - U/c}}$$

This is known as *relativistic Longitudinal Doppler effect*.

**Note 10:** If the velocity of the source is perpendicular to the direction of propagation, then  $\theta^1 = 90^\circ$ , then one gets,

$$\frac{\lambda^1}{\lambda} = \frac{v}{v^1} = \frac{1}{\sqrt{1 - U^2/c^2}}$$

This is known as *relativistic Transverse Doppler effect*. This indicates that a change of frequency occurs even when the observer is perpendicular to the direction of motion of the source. This effect is absent in the classical physics.

**Note 11:** Using the relativistic aberration formula, we can also obtain the transformation formula for the *solid angle*  $d\Omega$  of a pencil of rays. From Eq. (7.32), we can write,

$$1 + \frac{U \cos \theta^1}{c} = \frac{1 - \frac{U^2}{c^2}}{1 - \frac{U \cos \theta}{c}}$$

We know,

$$\frac{d\Omega^1}{d\Omega} = \frac{\sin \theta^1 d\theta^1}{\sin \theta d\theta} = \frac{d(\cos \theta^1)}{d(\cos \theta)} = \frac{1 - \frac{U^2}{c^2}}{\left(1 - \frac{U \cos \theta}{c}\right)^2} d\Omega$$

*Example 7.1* An electron is moving with speed  $fc$  in a direction opposite to that of a moving photon. Find the relative velocity of the electron and photon.

**Hint:** We know the speed of photon is  $c$ . Let speed of electron is  $fc$  ( $0 < f < 1$ ). Let the photon and electron are moving along positive and negative directions of  $x$  axis, respectively. Let the electron moving with velocity  $-fc$  be at rest in system  $S$ . Then we assume that system  $S^1$  or laboratory is moving with velocity  $fc$  relative to system  $S$  (electron). Therefore, we may write,  $v = fc, u_x^1 = c$ . Using the theorem of composition of velocities, we get the relative velocity  $u_x$  of the electron and photon as

$$u_x = \frac{u_x^1 + v}{1 + \frac{u_x^1 v}{c^2}} = \frac{c + fc}{1 + \frac{cfc}{c^2}} = c$$

*Example 7.2* An observer notices two particles moving away from him in two opposite directions with velocity  $fc$ . What is the velocity of one particle with respect to other?

**Hint:** We may regard one particle as  $S$  frame, the observer as the  $S^1$  frame and the other particle as the object whose speed is to be determined in the  $S$  frame. Then,  $u^1 = fc$  and  $v = fc$ . Using the theorem of composition of velocities, we get the relative velocity  $u$  of one particle with respect to other as

$$u = \frac{u^1 + v}{1 + \frac{u^1 v}{c^2}} = \frac{fc + fc}{1 + \frac{fcfc}{c^2}} = \text{etc.}$$

*Example 7.3* Two particles come towards each other with speed  $fc$  with respect to laboratory. What is their relative speed?

**Hint:** Consider a system  $S$  in which the particle having velocity  $-fc$  is a rest. Then, the system  $S^1$  i.e. laboratory is moving with velocity  $fc$  relative to the system  $S$ , i.e.  $v = fc$  (see Fig. 7.3). Now, we are to find the velocity of the particle  $u$  in system  $S$  which is moving with velocity  $u^1 = fc$  relative to  $S^1$ , i.e. laboratory. Using the theorem of composition of velocities, we get the relative velocity  $u$  of the particles with respect to other as

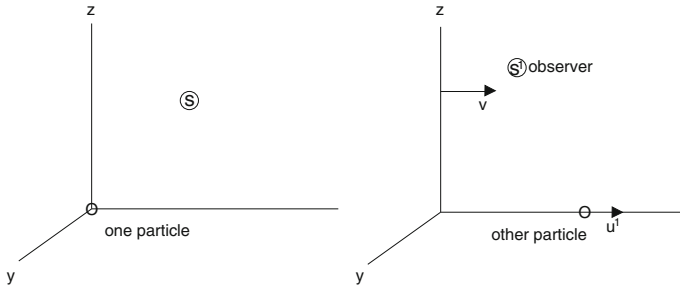


Fig. 7.3 The system  $S^1$  is moving with velocity  $fc$  relative to the system S

$$u = \frac{u^1 + v}{1 + \frac{u^1 v}{c^2}} = \frac{fc + fc}{1 + \frac{fcfc}{c^2}} = \text{etc.}$$

Example 7.4 If  $u$  and  $v$  are velocities in the same direction and  $V$  is their resultant velocity given by  $\tanh^{-1} \frac{V}{c} = \tanh^{-1} \frac{u}{c} + \tanh^{-1} \frac{v}{c}$ , then deduce the law of composition of velocities.

Hint: We know,

$$\tanh^{-1} x = \frac{1}{2} \ln \left( \frac{1+x}{1-x} \right)$$

Let,  $\frac{V}{c} = z, \frac{u}{c} = x, \frac{v}{c} = y$ , then the given condition implies

$$\ln \left( \frac{1+x}{1-x} \right)^{\frac{1}{2}} + \ln \left( \frac{1+y}{1-y} \right)^{\frac{1}{2}} = \ln \left( \frac{1+z}{1-z} \right)^{\frac{1}{2}}$$

or

$$\left( \frac{(1+x)(1+y)}{(1-x)(1-y)} \right) = \left( \frac{1+z}{1-z} \right)$$

or

$$z = \frac{x+y}{1+xy}$$

or

$$\frac{V}{c} = \frac{\frac{u}{c} + \frac{v}{c}}{1 + \frac{u}{c} \cdot \frac{v}{c}}$$

or

$$V = \frac{u+v}{1 + \frac{uv}{c^2}}$$

This is the law of composition of velocities or Einstein’s law of addition of velocities.

*Example 7.5* If  $v$  and  $v^1$  be the velocities of a particle in two inertial frames  $S$  and  $S^1$ . Initially, two system coincided. If  $S^1$  is moving with velocity  $V$  relative to  $S$  in  $x$  direction. Show that  $v^2 = \frac{v^{12} + V^2 + 2v^1 V \cos \theta^1 - (\frac{v^1 v \sin \theta^1}{c})^2}{(1 + \frac{v v^1 \cos \theta^1}{c^2})^2}$ , where  $\theta^1$  is the angle which  $v^1$  makes with  $x^1$  axis. Find also the direction of the velocity in  $S$  system.

**Hint:** According to the problem (see Fig. 7.4),

$$v^2 = v_x^2 + v_y^2, \quad v^{12} = v_x^{12} + v_y^{12} \quad \text{and} \quad v_x^1 = v^1 \cos \theta^1, \quad v_y^1 = v^1 \sin \theta^1$$

Also, we know

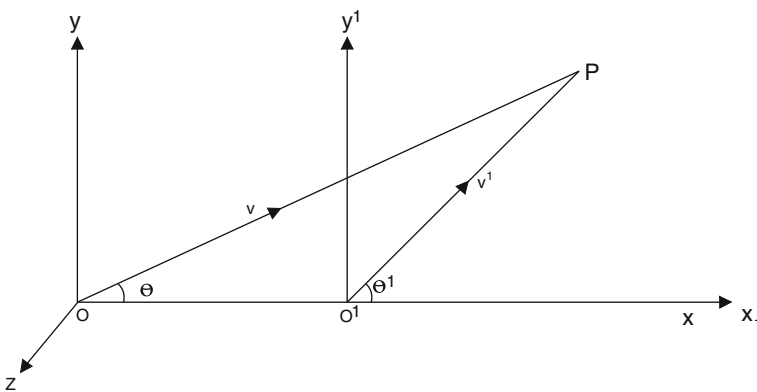
$$v_x^1 = \frac{v_x - V}{1 - \frac{v_x V}{c^2}}, \quad \text{and} \quad v_y^1 = v_y \frac{\sqrt{1 - V^2/c^2}}{1 - \frac{v_x V}{c^2}}$$

Thus

$$v_x = \frac{v_x^1 + V}{1 + \frac{v_x^1 V}{c^2}}, \quad \text{and} \quad v_y = v_y^1 \frac{\sqrt{1 - V^2/c^2}}{1 + \frac{v_x^1 V}{c^2}}$$

Hence,

$$v^2 = \frac{(v_x^1 + V)^2 + v_y^{12}(1 - \frac{V^2}{c^2})}{(1 + \frac{v_x^1 V}{c^2})^2} = \frac{(v^1 \cos \theta^1 + V)^2 + (v^1 \sin \theta^1)^2(1 - \frac{V^2}{c^2})}{(1 + \frac{v^1 \cos \theta^1 V}{c^2})^2} \text{ etc.}$$



**Fig. 7.4** Direction of the motion of the particle makes  $\theta$  and  $\theta^1$  angles with  $x$  and  $x^1$  axes, respectively

Now, the direction of the velocity in S system is given by

$$\tan \theta = \frac{v_y}{v_x} = \frac{\left( \frac{v_y^1 \sqrt{1-V^2/c^2}}{1 + \frac{v_x^1 V}{c^2}} \right)}{\left( \frac{v_x^1 + V}{1 + \frac{v_x^1 V}{c^2}} \right)} = \frac{\left( v^1 \sin \theta^1 \frac{\sqrt{1-V^2/c^2}}{1 + \frac{v^1 \cos \theta^1 V}{c^2}} \right)}{\left( \frac{v^1 \cos \theta^1 + V}{1 + \frac{v^1 \cos \theta^1 V}{c^2}} \right)} \text{ etc.}$$

*Example 7.6* If a photon travels the path in such a way that it moves in  $x^1 - y^1$  plane and makes an angle  $\phi$  with x axis of the system  $S^1$ , then prove that for the frame S,  $u_x^2 + u_y^2 = c^2$ . Here S and  $S^1$  are two inertial frames.

**Hint:** We know the velocity of photon is  $c$ . Its components in  $S^1$  system along x and y axes, respectively are (see Fig. 7.5)

$$u_x^1 = c \cos \phi \text{ and } u_y^1 = c \sin \phi$$

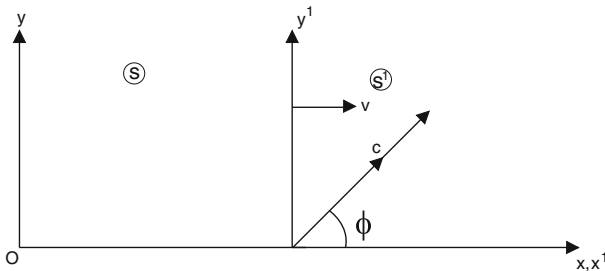
Let the frame  $S^1$  is moving with velocity  $v$  relative to S frame, then from the theorem of composition of velocities, we have

$$u_x = \frac{u_x^1 + v}{1 + \frac{u_x^1 v}{c^2}}, \quad u_y = \frac{u_y^1 \sqrt{1 - v^2/c^2}}{1 + \frac{u_x^1 v}{c^2}}$$

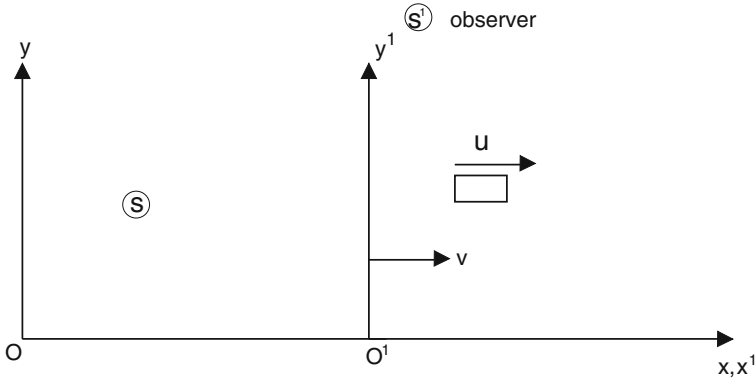
Now,

$$u_x^2 + u_y^2 = \left[ \frac{c \cos \phi + v}{1 + \frac{c \cos \phi v}{c^2}} \right]^2 + \left[ \frac{c \sin \phi \sqrt{1 - v^2/c^2}}{1 + \frac{c \cos \phi v}{c^2}} \right]^2 = c^2$$

*Example 7.7* An observer moving along x axis of S frame at a velocity  $v$  observes a body of proper volume  $V_0$  moving at a velocity  $u$  along the x axis of S. Show that the observer measures the volume to be equal to  $V_0 \frac{\sqrt{c^2 - u^2} \sqrt{c^2 - v^2}}{c^2 - uv}$ .



**Fig. 7.5** The path of the photon makes  $\phi$  angle with x axis



**Fig. 7.6** A body is moving with velocity  $u$  along  $x$  axis

**Hint:** Let  $S^1$  be the frame containing the observer (see Fig. 7.6). The relative velocity between  $S^1$  and  $S$  is  $v$  along  $x$  direction. The velocity of the body relative to  $S$  is  $u$  along  $x$  direction. Therefore, its velocity relative to  $S^1$  is

$$u^1 = \frac{u - v}{1 - \frac{uv}{c^2}}$$

Therefore, an observer measures the volume of the body relative to him as

$$V = V_0 \sqrt{1 - \frac{u^2}{c^2}} = V_0 \sqrt{1 - \frac{\left(\frac{u-v}{1-uv/c^2}\right)^2}{c^2}}$$

**Example 7.8** A particle has velocity  $\vec{u} = (u_1, u_2, u_3)$  in  $S$  and  $\vec{u}^1 = (u_1^1, u_2^1, u_3^1)$  in  $S^1$ . Show that  $c^2 - u^2 = \frac{c^2(c^2 - u_1^2)(c^2 - v^2)}{(c^2 + u_1^1 v)^2}$ .

**Hint:** We know, inverse transformations are obtained from by replacing  $v$  by  $-v$  and subscripted quantities by unsubscripted ones and vice versa. Therefore, from note 6, we can write

$$\sqrt{1 - (u/c)^2} = \frac{\sqrt{1 - (v/c)^2} \sqrt{1 - (u^1/c)^2}}{\left(1 + \frac{u_1^1 v}{c^2}\right)}$$

or,

$$c^2 - u^2 = \frac{c^2(c^2 - u_1^2)(c^2 - v^2)}{(c^2 + u_1^1 v)^2}$$

**Example 7.9** How fast would we need to drive towards a red traffic light for the light to appear green?



**Hint:** We know that the wavelengths of red light and green light are  $\lambda_{\text{red}} \approx 7 \times 10^{-5}$  cm and  $\lambda_{\text{green}} \approx 5 \times 10^{-5}$  cm, respectively. As velocity  $U$  of the car towards the signal, we use the relativistic formula for Doppler effect as  $\lambda = \lambda^1 \sqrt{\frac{1+U/c}{1-U/c}}$ . Thus we have

$$5 \times 10^{-5} = 7 \times 10^{-5} \sqrt{\frac{1-U/c}{1+U/c}} \implies U = \text{etc.}$$

*Example 7.10* Calculate the Doppler shift in wavelength for light of wavelength  $a$  Angstrom (i) when the source approaches to the observer at a velocity  $fc$ . (ii) when the source moves transversely to the line of sight at the same velocity  $fc$ .

**Hint:**

$$(i) \lambda = a \sqrt{\frac{1-fc/c}{1+fc/c}} \text{ etc. } (ii) \lambda = a \frac{1}{\sqrt{1-(fc)^2/c^2}} \text{ etc.}$$

*Example 7.11* If  $u$  and  $u^1$  are the velocities of a particle in the frames  $S$  and  $S^1$ , which are moving with velocity relative to each other. Prove that  $\sqrt{1-(u^1/c)^2} = \frac{\sqrt{1-(v/c)^2} \sqrt{1-(u/c)^2}}{(1-\frac{uv}{c^2})}$ .

**Hint:** Use Eq. (7.13).

*Example 7.12* Show that velocity of light is the same in all inertial frame.

**Hint:** Let velocity in  $S^1$  system is  $c^1$ , then

$$c^1 = \frac{dx^1}{dt^1} = \frac{a(dx - vdt)}{a\left(dt - \frac{vdx}{c^2}\right)} = \frac{\frac{dx}{dt} - v}{1 - \frac{v}{c^2} \frac{dx}{dt}}$$

Now, velocity in  $S$  system is  $\frac{dx}{dt} = c$ , then using  $\frac{dx}{dt} = c$ , we get

$$c^1 = \frac{c - v}{1 - \frac{v}{c^2}c} = c$$

# Chapter 8

## Four Dimensional World

### 8.1 Four Dimensional Space–Time

If any physical law may be expressed in a invariant four dimensional form, then the law will be invariant under Lorentz Transformation. Minkowski said that the external world is formed by four dimensional space–time continuum known as *Minkowski or World space*. The event in world space must be represented by three spatial coordinates and one time coordinate. The events in the world space are represented by points known as *world points*. In this world space there corresponds to each particle a certain line known as *world line*. In other words, a curve in the four dimensional space is called world line. Newton's laws retain their form under Galilean Transformation which is found to be incorrect and hence Newton's laws are inaccurate representation of experimental phenomena. We want to extend the 3-vector law

$$F_i = \frac{d}{dt} \left[ m \left( \frac{dx_i}{dt} \right) \right]$$

to 4-vector form because above 3-vector form does not satisfy the principle of relativity i.e. its form changes under Lorentz Transformation.

The basic representation for such an extension are that:

1. A fourth component of force must be introduced.
2. The time is no more a scalar invariant and changes under Lorentz Transformation.

**Note 1:** In Newtonian mechanics, the physical parameters like position, velocity, momentum, acceleration, force have 3-vector forms. We want to extend the 3-vector form to 4-vector form in the relativistic mechanics in order to satisfy the principle of relativity. Therefore, a fourth component of all physical parameters should be introduced.

## 8.2 Proper Time

The time recorded by a clock moving with a given system is called **proper time**.

Let  $S$  be a reference system which is at rest. Let a clock is moving with velocity  $v$  relative to  $S$ . Then at each instant we can introduce a coordinate system, rigidly linked to the moving clock, which constitutes an inertial reference system  $S^1$ . Suppose at any instant, a clock in  $S$  system indicates a time  $dt$  and at this time, the clock travels a distance  $\sqrt{(dx^2 + dy^2 + dz^2)}$ . Obviously, the relative velocity between the clock and  $S^1$  system is zero i.e. the clock is at rest in the system linked with them i.e. in reference system  $S^1$ . Thus,  $dx^1 = dy^1 = dz^1 = 0$ . The interval

$$d\rho^2 = c^2 dt^2 - dx^2 - dy^2 - dz^2 = c^2 dt^{1^2} - dx^{1^2} - dy^{1^2} - dz^{1^2} = c^2 dt^{1^2}$$

Let the moving clock in  $S^1$  system reads  $dt^1$  corresponding to the time  $dt$ . Then

$$dt^{1^2} = \frac{d\rho^2}{c^2} = \left(\frac{1}{c^2}\right) [dt^2 - dx^2 - dy^2 - dz^2]$$

or

$$dt^1 = dt \left[ 1 - \frac{1}{c^2} \left\{ \left(\frac{dy}{dt}\right)^2 + \left(\frac{dx}{dt}\right)^2 + \left(\frac{dz}{dt}\right)^2 \right\} \right]^{\frac{1}{2}} = dt \sqrt{1 - \frac{v^2}{c^2}} \quad (8.1)$$

[ $v$  is the velocity of the moving clock]

Here,  $dt^1$  is the proper read by the moving clock corresponding to  $dt$  read by the clock at rest. Important to note that  $dt^1 < dt$ . Thus proper time of a moving objects is always smaller than the corresponding time in the clock at rest.

### Alternative:

Let  $S$  be a reference system which is at rest. Let a rocket constitutes another reference system  $S^1$  which is moving with velocity  $v$  relative to  $S$  along +ve direction of  $x$ -axis. Suppose at any instant, a clock in  $S$  system indicates a time  $t$ . Then an observer in rocket i.e. in  $S^1$  frame measures this time as  $t^1$  which is given according to Lorentz Transformation as

$$t^1 = a \left[ t - \frac{vx}{c^2} \right]$$

where  $a = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$ .

Since  $t$  is the time recorded by an observer at rest and  $v$  is the velocity of the rocket, therefore, at this time, the clock i.e. the rocket travels a distance  $x = vt$ . Putting this value in the above equation, we get,

$$t^1 = a \left[ t - \frac{vx}{c^2} \right] = at \left[ 1 - \frac{v^2}{c^2} \right] = t \sqrt{1 - \frac{v^2}{c^2}} \quad (8.2)$$

So that  $t > t^1$ . Thus the time recorded by the clock in rocket i.e. in moving frame differs the time recorded by the clock at rest for the same journey. Hence, two clocks in the two systems  $S$  and  $S^1$  run at different rate. Also we see that the time recorded in the rest frame to be greater than that the moving system. Thus a moving clock runs more slow than a stationary one.

**Theorem 8.1** *Proper time is invariant under Lorentz Transformation.*

*Proof* We know the proper time,

$$d\tau = dt \sqrt{1 - \frac{v^2}{c^2}}$$

or,

$$d\tau = dt \left[ 1 - \frac{1}{c^2} \left\{ \left( \frac{dy}{dt} \right)^2 + \left( \frac{dx}{dt} \right)^2 + \left( \frac{dz}{dt} \right)^2 \right\} \right]^{\frac{1}{2}}$$

This implies

$$d\tau^2 = \frac{1}{c^2} [c^2 dt^2 - dx^2 - dy^2 - dz^2]$$

Now,

$$d\tau^{12} = \frac{1}{c^2} [c^2 dt^{12} - dx^{12} - dy^{12} - dz^{12}]$$

Using the Lorentz Transformation

$$x^1 = a(x - vt); \quad y^1 = y; \quad z^1 = z; \quad t^1 = a \left[ t - \left( \frac{xv}{c^2} \right) \right],$$

we get,

$$\begin{aligned} d\tau^{12} &= \frac{1}{c^2} \left[ c^2 a^2 \left\{ dt - \left( \frac{v dx}{c^2} \right) \right\}^2 - a^2 (dx - v dt)^2 - dy^2 - dz^2 \right] \\ &= \frac{1}{c^2} [c^2 dt^2 - dx^2 - dy^2 - dz^2] = d\tau^2 \end{aligned}$$

Hence proper time is invariant under Lorentz Transformation.

**Alternative:**

We know,

$$\begin{aligned} ds^2 &= c^2 dt^2 - dx^2 - dy^2 - dz^2 \\ &= c^2 dt^2 \left[ 1 - \frac{1}{c^2} \left\{ \left( \frac{dy}{dt} \right)^2 + \left( \frac{dx}{dt} \right)^2 + \left( \frac{dz}{dt} \right)^2 \right\} \right] \\ &= c^2 dt^2 \left[ 1 - \frac{v^2}{c^2} \right] \end{aligned}$$

Therefore,

$$\left( \frac{ds}{c} \right) = dt \sqrt{1 - \frac{v^2}{c^2}} = d\tau$$

Since  $ds$  is invariant under Lorentz Transformation and  $c$  is also an invariant quantity, therefore,  $d\tau$  is invariant under Lorentz Transformation.

**8.3 World Velocity or Four Velocity**

We know the expression

$$dx^2 + dy^2 + dz^2 - c^2 dt^2$$

remains invariant under Lorentz Transformation i.e.

$$dx^2 + dy^2 + dz^2 - c^2 dt^2 = dx^{1^2} + dy^{1^2} + dz^{1^2} - c^2 dt^{1^2}.$$

In terms of differential, the quantity

$$ds^2 = dx^2 + dy^2 + dz^2 - c^2 dt^2$$

remains invariant.

It can be written as

$$\begin{aligned} ds^2 &= dx_1^2 + dx_2^2 + dx_3^2 + dx_4^2, \\ \text{where, } x_1 &= x, \quad x_2 = y, \quad x_3 = z \quad \text{and} \quad x_4 = ict. \end{aligned}$$

Here  $ds$  is called *invariant space-time interval*.

The position four vector i.e. *world point* takes the form

$$x_\mu = [x_1, x_2, x_3, ict] \tag{8.3}$$

World velocity  $u_\mu$  is defined as the rate of change of the position vector of a particle with respect to its proper time i.e.

$$u_\mu = \frac{dx_\mu}{d\tau}. \tag{8.4}$$

The space components of  $u_\mu$  are

$$\frac{dx_j}{d\tau} = \frac{dx_j}{dt\sqrt{1-\frac{v^2}{c^2}}} = \frac{v_j}{\sqrt{1-\frac{v^2}{c^2}}} \tag{8.5}$$

( $j = 1, 2, 3$  and  $v^2 = v_1^2 + v_2^2 + v_3^2$  where  $v_j = \frac{dx_j}{dt}$  is the three dimensional velocity)

The time component of  $u_\mu$  is

$$u_4 = \frac{icdt}{d\tau} = \frac{ic}{\sqrt{1-\frac{v^2}{c^2}}} \tag{8.6}$$

Thus world velocity is given by

$$u_\mu = \frac{dx_\mu}{d\tau} = a \left[ \frac{dx_1}{dt}, \frac{dx_2}{dt}, \frac{dx_3}{dt}, ic \right] = a [v_1, v_2, v_3, ic] \tag{8.7}$$

where  $a = \frac{1}{\sqrt{1-\frac{v^2}{c^2}}}$ .

In Euclidean geometry a vector has three components. In relativity, a vector has four components. Let us consider the transformations of coordinates  $x \rightarrow x^1$ . Then we have

$$dx^1 = \left( \frac{\partial x^1}{\partial x} \right) dx.$$

[we use summation convention, therefore ignore  $\Sigma$  sign]

If the transformation of any vector  $A^\mu$  is given by

$$A^{\mu 1} = \left( \frac{\partial x^{\mu 1}}{\partial x^v} \right) A^v,$$

then  $A^\mu$  is called *contra-variant vector*.

If the transformation of any vector  $B_\mu$  is given by

$$B_\mu^1 = \left( \frac{\partial x^v}{\partial x^{\mu 1}} \right) B_v,$$

then  $B_\mu$  is called *co-variant vector*.

In Minkowski four dimensional space, the distance between two neighboring points is defined by the metric

$$ds^2 = c^2 dt^2 - dx^2 - dy^2 - dz^2 = g_{\mu\nu} dx^\mu dx^\nu$$

Here  $g_{\mu\nu}$  is known as fundamental tensor,

$$g_{11} = g_{22} = g_{33} = -1; \quad g_{44} = 1$$

If  $A^\mu$  i.e. contra-variant vector is known, then co-variant component can be found with the help of fundamental tensor as  $A_\mu = g_{\mu\nu} A^\nu$ .

If  $A_\mu$  i.e. co-variant vector is known, then contra-variant component is given by  $A^\mu = g^{\mu\nu} A_\nu$ . Here  $g^{\mu\nu}$  is the inverse of  $g_{\mu\nu}$ .

$A^\mu A_\mu$  is called *norm of the vector*  $A^\mu$ .

**Theorem 8.2** *Norm is an invariant quantity.*

*Proof*

$$\begin{aligned} A^{\mu 1} A_\mu^1 &= \left[ \left( \frac{\partial x^{\mu 1}}{\partial x^\nu} \right) A^\nu \right] \left[ \left( \frac{\partial x^\sigma}{\partial x^{\mu 1}} \right) A_\sigma \right] \\ &= \left( \frac{\partial x^{\mu 1}}{\partial x^\nu} \right) \left( \frac{\partial x^\sigma}{\partial x^{\mu 1}} \right) A^\nu A_\sigma \\ &= \delta_\nu^\sigma A^\nu A_\sigma = A^\sigma A_\sigma = A^\mu A_\mu \end{aligned}$$

Thus norm is invariant and it is a scalar.

$A^\mu A_\mu = g_{\mu\nu} A^\mu A^\nu$  if  $A^\mu$  is known.

$A^\mu A_\mu = g^{\mu\nu} A_\mu A_\nu$  if  $A_\mu$  is known.

If  $A^\mu A_\mu > 0$ , then  $A^\mu$  is called *time like vector*.

If  $A^\mu A_\mu < 0$ , then  $A^\mu$  is called *space like vector*.

If  $A^\mu A_\mu = 0$ , then  $A^\mu$  is called *null or light like vector*.

**Theorem 8.3** *Any four vector orthogonal to a time like vector is space like.*

*Proof* Let  $A^\mu$  is a time like vector, therefore,

$$A^\mu A_\mu = [(A^0)^2 - (A^1)^2 - (A^2)^2 - (A^3)^2]$$

[ $A^0$  = time component and  $A^1, A^2, A^3$  are spatial components]

Now we transform coordinate axis in such a way that  $A^1 = A^2 = A^3 = 0$ .  
Then  $A^\mu A_\mu = (A^0)^2 > 0$  i.e.  $A^0 \neq 0$ .

If  $B^\mu$  be orthogonal vector to  $A^\mu$ , then

$$A^\mu B_\mu = [A^0 B_0 + A^1 B_1 + A^2 B_2 + A^3 B_3] = 0$$

i.e.

$$A^0 B^0 - (A^1 B^1 + A^2 B^2 + A^3 B^3) = 0$$

Therefore,  $A^0 B^0 = 0$  [as  $A^1 = A^2 = A^3 = 0$ ]

Hence,  $B^0 = 0$  [as  $A^0 \neq 0$ ]

Now,

$$B^\mu B_\mu = [(B^0)^2 - (B^1)^2 - (B^2)^2 - (B^3)^2] = -[(B^1)^2 + (B^2)^2 + (B^3)^2] < 0.$$

Therefore,  $B^\mu$  is a space like vector.

**Theorem 8.4** *Two time like vectors cannot be orthogonal to each other.*

*Proof* Let  $A^\mu$  and  $B^\mu$  be two time like vectors, therefore,

$$A^\mu A_\mu = [(A^0)^2 - (A^1)^2 - (A^2)^2 - (A^3)^2] > 0 \quad (8.8)$$

and

$$B^\mu B_\mu = [(B^0)^2 - (B^1)^2 - (B^2)^2 - (B^3)^2] > 0 \quad (8.9)$$

[ $A^0, B_0$  are time components and  $A^1, A^2, A^3$  &  $B^1, B^2, B^3$  are spatial components]

If possible, let two vectors are orthogonal, then

$$A^0 B^0 - (A^1 B^1 + A^2 B^2 + A^3 B^3) = 0$$

or

$$(A^0 B^0)^2 = (A^1 B^1 + A^2 B^2 + A^3 B^3)^2 \quad (8.10)$$

Multiplying (8.8) and (8.9), we get

$$(A^0 B^0)^2 > [(A^1)^2 + (A^2)^2 + (A^3)^2][(B^1)^2 + (B^2)^2 + (B^3)^2] \quad (8.11)$$

Using (8.10), we get

$$(A^1 B^1 + A^2 B^2 + A^3 B^3)^2 > [(A^1)^2 + (A^2)^2 + (A^3)^2][(B^1)^2 + (B^2)^2 + (B^3)^2]$$



This implies

$$0 > (A^1 B^2 - A^2 B^1)^2 + (A^1 B^3 - A^3 B^1)^2 + (A^2 B^3 - A^3 B^2)^2$$

Since sum of three squares of real quantities in the right hand side is always positive, therefore, the above result is impossible. Hence two time like vectors cannot be orthogonal to each other.

**Theorem 8.5** *Any four vector orthogonal to a space like vector may be time like, space like or null.*

*Proof* Let  $A^\mu$  is a space like vector with  $A^0$  = time component and  $A^1, A^2, A^3$  are spatial components.

Now we choose the inertial frame in such a way that  $A^0 = A^2 = A^3 = 0$  i.e.  $A^\mu = (0, A^1, 0, 0)$ . Then  $A^\mu A_\mu = -(A^1)^2 < 0$  i.e.  $A^1 \neq 0$ .

If  $B^\mu$  be orthogonal vector to  $A^\mu$ , then

$$A^\mu B_\mu = [A^0 B_0 + A^1 B_1 + A^2 B_2 + A^3 B_3] = 0$$

i.e.

$$A^0 B^0 - (A^1 B^1 + A^2 B^2 + A^3 B^3) = 0$$

Therefore,  $A^1 B^1 = 0$  [as  $A^0 = A^2 = A^3 = 0$ ]

Hence,  $B^1 = 0$  [as  $A^1 \neq 0$ ]

Now,

$$B^\mu B_\mu = [(B^0)^2 - (B^2)^2 - (B^3)^2]$$

The right hand side may be positive, negative or zero. Therefore,  $B^\mu$  may be time like, space like or null.

**Theorem 8.6** *Any four vector orthogonal to a null vector is either space like or null.*

*Proof* Let  $A^\mu$  is a null vector with  $A^0$  = time component and  $A^1, A^2, A^3$  are spatial components.

Now we transform the space axes in such a way that  $A^2 = A^3 = 0$ . i.e.  $A^\mu = (A^0, A^1, 0, 0)$ . Then  $A^\mu A_\mu = (A^0)^2 - (A^1)^2 = 0$  implies  $(A^1)^2 = (A^0)^2$ .

If  $B^\mu$  be orthogonal vector to  $A^\mu$ , then

$$A^\mu B_\mu = [A^0 B_0 + A^1 B_1 + A^2 B_2 + A^3 B_3] = 0$$

i.e.

$$A^0 B^0 - (A^1 B^1) = 0$$

Therefore,  $(B^1)^2 = (B^0)^2$  [as  $(A^1)^2 = (A^0)^2$ ]

Now,

$$B^\mu B_\mu = [(B^0)^2 - (B^1)^2 - (B^2)^2 - (B^3)^2] = -[(B^2)^2 + (B^3)^2] \leq 0$$

We use equality sign since, both  $B^2$  and  $B^3$  happen to vanish. Therefore,  $B^\mu$  is either space like or null.

The *magnitude or norm* of the world velocity is given by

$$\begin{aligned} u_\mu u_\mu &\equiv u_1 u_1 + u_2 u_2 + u_3 u_3 + u_4 u_4 \\ &= [av_1 + av_2 + av_3 + aic][av_1 + av_2 + av_3 + aic] \\ &= a^2(v_1^2 + v_2^2 + v_3^2 - c^2) = a^2(v^2 - c^2) = -c^2 \end{aligned}$$

[This is scalar product of given four vectors]

The square of world velocity i.e. norm of the world velocity has a constant magnitude. So it is invariant under Lorentz Transformation. Here we see the usefulness of the four dimensional language in relativity that all equations expressed in four vector language will automatically retain their mathematical form under Lorentz Transformation and hence satisfy the postulates of special theory of relativity.

## 8.4 Lorentz Transformation of Space and Time in Four Vector Form

Consider two systems  $S$  and  $S^1$ , the latter moving with velocity  $v$  relative to former along +ve direction of  $x$ -axis. In Minkowski space, let the coordinates of an event be represented by  $(x, y, z, ict)$  or  $x^\mu$  ( $\mu = 1, 2, 3, 4$ ) where  $x_1 = x$ ,  $x_2 = y$ ,  $x_3 = z$  and  $x_4 = ict$ . The Lorentz Transformation

$$x^1 = a(x - vt), \quad y^1 = y, \quad z^1 = z, \quad t^1 = a \left[ t - \left( \frac{vx}{c^2} \right) \right]$$

can be written as

$$x_1^1 = ax_1 + 0.x_2 + 0.x_3 + ia\beta x_4$$

$$x_2^1 = 0.x_1 + 1.x_2 + 0.x_3 + 0.x_4$$

$$x_3^1 = 0.x_1 + 0.x_2 + 1.x_3 + 0.x_4$$

$$x_4^1 = -ia\beta x_1 + 0.x_2 + 0.x_3 + ax_4$$

where  $a = \frac{1}{\sqrt{1-\beta^2}}$  and  $\beta = \frac{v}{c}$ .

This implies

$$x_\mu^1 = a_{\mu\nu}x_\nu \quad (8.12)$$

where,  $x_\mu^1 = [x_1^1, x_2^1, x_3^1, x_4^1]^T$ ,  $x_\mu = [x_1, x_2, x_3, x_4]^T$  and

$$a_{\mu\nu} = \begin{bmatrix} a & 0 & 0 & ia\beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -ia\beta & 0 & 0 & a \end{bmatrix} \quad (8.13)$$

Equation (8.8) represents the Lorentz Transformation of space and time in four vector form.

Similarly, *Lorentz Transformation of any four vector*  $A_\mu$  ( $\mu = 1, 2, 3, 4$ ) may be expressed as

$$A_\mu^1 = a_{\mu\nu}A_\nu \quad (8.14)$$

where  $a_{\mu\nu}$  is given in (8.9). One can verify  $|a_{\mu\nu}| = 1$ . Also, the Transformation matrix  $a_{\mu\nu}$  is orthogonal since  $[a_{\mu\nu}]^{-1} = [a_{\mu\nu}]^T$ . Thus the Transformation matrix  $a_{\mu\nu}$  is a unitary orthogonal matrix. The inverse Lorentz Transformation matrix  $a_{\mu\nu}^1$  is given by

$$a_{\mu\nu}^1 = [a_{\mu\nu}]^{-1} = \begin{bmatrix} a & 0 & 0 & -ia\beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ ia\beta & 0 & 0 & a \end{bmatrix} \quad (8.15)$$

The inverse Lorentz Transformation of any four vector can be obtained as

$$A_\mu = a_{\mu\nu}^1 A_\nu^1 \quad (8.16)$$

Four vectors are four tensors of rank 1. A four tensor of rank 2 has  $4^2$  components  $T_{\mu\nu}$ . Here, we can also use the Lorentz Transformation matrix to get the required Lorentz transformation of the tensor  $T_{\mu\nu}$  as

$$T_{\mu\nu}^1 = a_{\mu\sigma}a_{\nu\lambda}T_{\sigma\lambda} \quad (8.17)$$

In general, a four tensor of rank  $n$  has  $4^n$  components, which are written with  $n$  indices and they transform as above

$$T_{\mu\nu\rho\dots\omega}^1 = a_{\mu\sigma}a_{\nu\lambda}a_{\rho\alpha}\dots a_{\omega\beta}T_{\sigma\lambda\alpha\dots\beta} \quad (8.18)$$

*Example 8.1* Consider three inertial frames  $S$ ,  $S^1$  and  $S^2$ . Let  $S^1$  is moving with a constant velocity  $v_1$  along  $x$ -axis of  $S$  and  $S^2$  is moving with a constant velocity  $v_2$

along  $y$ -axis of  $S^1$ . Find the transformation between  $S$  and  $S^2$ . Again,  $S^1$  is moving with a constant velocity  $v_2$  along  $y$ -axis of  $S$  and  $S^2$  is moving with a constant velocity  $v_1$  along  $x$ -axis of  $S^1$ . Find the transformation between  $S$  and  $S^2$ . Show also these two transformations are not equal.

**Hint:** The Transformation matrix from  $S$  to  $S^1$  is

$$L_1 = \begin{bmatrix} a_1 & 0 & 0 & ia_1\beta_1 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -ia_1\beta_1 & 0 & 0 & a_1 \end{bmatrix}$$

where  $a_1 = \frac{1}{\sqrt{1-\beta_1^2}}$  and  $\beta_1 = \frac{v_1}{c}$ .

The Transformation matrix from  $S^1$  to  $S^2$  is

$$L_2 = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & a_2 & 0 & ia_2\beta_1 \\ 0 & 0 & 1 & 0 \\ 0 & -ia_2\beta_2 & 0 & a_2 \end{bmatrix}$$

where  $a_2 = \frac{1}{\sqrt{1-\beta_2^2}}$  and  $\beta_2 = \frac{v_2}{c}$ .

Hence the Transformation matrix from  $S$  to  $S^2$  is

$$L_1L_2 = \begin{bmatrix} a_1 & 0 & 0 & ia_1\beta_1 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -ia_1\beta_1 & 0 & 0 & a_1 \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & a_2 & 0 & ia_2\beta_1 \\ 0 & 0 & 1 & 0 \\ 0 & -ia_2\beta_2 & 0 & a_2 \end{bmatrix} = \text{etc.}$$

When the order is reversed then the Transformation matrix from  $S$  to  $S^2$  is

$$L_2L_1 = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & a_2 & 0 & ia_2\beta_1 \\ 0 & 0 & 1 & 0 \\ 0 & -ia_2\beta_2 & 0 & a_2 \end{bmatrix} \begin{bmatrix} a_1 & 0 & 0 & ia_1\beta_1 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -ia_1\beta_1 & 0 & 0 & a_1 \end{bmatrix} = \text{etc.}$$

Obviously,

$$L_1L_2 \neq L_2L_1.$$

*Example 8.2* Find the Lorentz Transformation of four velocity. Using it obtain Einstein's law of addition of velocities.

**Hint:** We know four velocity takes the form

$$U_\mu = [U_1, U_2, U_3, U_4] = \frac{1}{\sqrt{1 - \frac{u^2}{c^2}}} [u_1, u_2, u_3, ic]$$

where  $u^2 = u_1^2 + u_2^2 + u_3^2$ . Now, we use the Lorentz Transformation matrix to get the required Lorentz transformation of  $U_\mu$  as

$$\begin{bmatrix} U_1^1 \\ U_2^1 \\ U_3^1 \\ U_4^1 \end{bmatrix} = \begin{bmatrix} a & 0 & 0 & ia\beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -ia\beta & 0 & 0 & a \end{bmatrix} \begin{bmatrix} U_1 \\ U_2 \\ U_3 \\ U_4 \end{bmatrix}$$

where  $\beta = \frac{v}{c}$  and  $a = \frac{1}{\sqrt{1 - \beta^2}}$ .

Writing explicitly, we have

$$U_1^1 = aU_1 + ia\beta U_4$$

$$U_2^1 = U_2$$

$$U_3^1 = U_3$$

$$U_4^1 = aU_4 - ia\beta U_1$$

Thus we get

$$\frac{u_1^1}{\sqrt{1 - \frac{u_1^2}{c^2}}} = \frac{au_1}{\sqrt{1 - \frac{u^2}{c^2}}} + iva \frac{ic}{c\sqrt{1 - \frac{u^2}{c^2}}} = \frac{a(u_1 - v)}{\sqrt{1 - \frac{u^2}{c^2}}}$$

$$\frac{u_2^1}{\sqrt{1 - \frac{u_2^2}{c^2}}} = \frac{u_2}{\sqrt{1 - \frac{u^2}{c^2}}}, \quad \frac{u_3^1}{\sqrt{1 - \frac{u_3^2}{c^2}}} = \frac{u_3}{\sqrt{1 - \frac{u^2}{c^2}}}$$

$$\frac{ic}{\sqrt{1 - \frac{u_4^2}{c^2}}} = \frac{aic}{\sqrt{1 - \frac{u^2}{c^2}}} - iva \frac{u_1}{c\sqrt{1 - \frac{u^2}{c^2}}}$$

Last equation gives

$$a \left( 1 - \frac{u_1 v}{c^2} \right) = \frac{\sqrt{1 - \frac{u^2}{c^2}}}{\sqrt{1 - \frac{u_1^2}{c^2}}}$$

Using this result, we get the required transformation as

$$u_1^1 = \frac{u_1 - v}{1 - \frac{vu_1}{c^2}}, \quad u_2^1 = \frac{u_2 \sqrt{1 - \frac{v^2}{c^2}}}{1 - \frac{vu_1}{c^2}}, \quad u_3^1 = \frac{u_3 \sqrt{1 - \frac{v^2}{c^2}}}{1 - \frac{vu_1}{c^2}}$$

The inverse transformations are given by

$$u_1 = \frac{u_1^1 + v}{1 + \frac{vu_1^1}{c^2}}, \quad u_2 = \frac{u_2^1 \sqrt{1 - \frac{v^2}{c^2}}}{1 + \frac{vu_1^1}{c^2}}, \quad u_3 = \frac{u_3^1 \sqrt{1 - \frac{v^2}{c^2}}}{1 + \frac{vu_1^1}{c^2}}$$

If we take  $u_2^1 = u_3^1 = 0$ ,  $u_1^1 = u$  and  $u_1 = w$ , then we get

$$w = \frac{u + v}{1 + \frac{uv}{c^2}}$$

This is Einstein's law of addition of velocities.

*Example 8.3* Show that derivative of four velocity is perpendicular to it.

**Hint:** We know the magnitude of four velocity is constant. Hence

$$\frac{d}{d\tau}(u_\mu u_\mu) = \frac{d}{d\tau}(-c^2) = 0$$

Hence,

$$u_\mu \frac{du_\mu}{d\tau} + \frac{du_\mu}{d\tau} u_\mu = 0 \text{ etc.}$$

*Example 8.4* Show that four acceleration is perpendicular to four velocity.

**Hint:** Same as Example 8.3.

*Example 8.5* Show that derivative of four vector with constant magnitude is perpendicular to it.

**Hint:** Same as Example 8.3.

*Example 8.6* Consider the four vectors  $a^\mu$ ,  $b^\mu$ ,  $c^\mu$  whose components are given by

$$a^\mu = (-3, 0, 0, 2), \quad b^\mu = (10, 0, 6, 8), \quad c^\mu = (2, 0, 0, 3)$$

Check the nature of the four vectors.

**Hint:** Here,

$$a^\mu a_\mu = a_0^2 - a_1^2 - a_2^2 - a_3^2 = (-3)^2 - 2^2 > 0, \quad \text{hence, } a^\mu \text{ is time like etc.}$$

*Example 8.7* Consider a particle moving along  $x$ -axis whose velocity as a function of time is  $\frac{dx}{dt} = \frac{at}{\sqrt{b^2 + a^2 t^2}}$ , where  $a$  and  $b$  are constants and velocity of light  $c$  is assumed to be unity. (i) Does the particle's speed exceed the speed of light? (ii) Calculate the component of particle's four velocity. (iii) Express  $x$  and  $t$  in terms of proper time along the trajectory.

**Hint:** (i)

$$\text{Here, } v = \frac{dx}{dt} = \frac{at}{\sqrt{b^2 + a^2 t^2}} < 1$$

since,  $b^2 + a^2 t^2 > a^2 t^2$ , therefore less than the velocity of light.

(ii) The particle's four velocity is given by

$$u_\mu = a[u_x, u_y, u_z, ic], \quad \text{where, } a = \frac{1}{1 - \frac{v^2}{c^2}}$$

Here,  $c = 1$  and  $u_x = v = \frac{at}{\sqrt{b^2 + a^2 t^2}}$ ,  $u_y = 0$ ,  $u_z = 0$  etc.

(iii) The clock of an observer sitting on the particle reads proper time. The proper time elapsed from  $t = 0$  to  $t$  is

$$\tau = \int_0^t dt \sqrt{1 - v^2} = \int_0^t \frac{b dt}{\sqrt{b^2 + a^2 t^2}} = \text{etc.}$$

Particle's trajectory is

$$x(t) - x_0 = \int_0^t \frac{b at dt}{\sqrt{b^2 + a^2 t^2}} = \text{etc.}$$

Now, express  $x$  and  $t$  in terms of proper time  $\tau$ .

*Example 8.8* A  $\pi$  meson is moving with a speed  $v = fc$  ( $0 < f < 1$ ) in a direction  $45^\circ$  to the  $x$ -axis. Calculate the component of  $\pi$  meson's four velocity.

**Hint:** Suppose the  $\pi$  meson is moving in  $x$ - $y$  plane. Therefore,

$$v_x = fc \cos 45^\circ, \quad v_y = fc \sin 45^\circ, \quad v_z = 0$$

The  $\pi$  meson's four velocity is given by

$$u_\mu = a[v_x, v_y, v_z, ic], \quad \text{where, } a = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{1}{\sqrt{1 - f^2}}$$

*Example 8.9* What do you mean by co-moving frame? What should be the components of a 4-velocity vector in its co-moving frame?

**Hint:** A particle is moving in a frame. The observer sat on the particle. So it is also moving with the particle's velocity. This type of reference frame is known as co-moving frame.

The particle is not undergoing any spatial separation with the observer. That is  $dx = dy = dz = 0$ . However, the fourth component which is time cannot be stopped. Thus  $dt \neq 0$ .

Hence space components of  $u_\mu$  are  $\frac{dx_j}{dt} = 0$ .

Now,  $u_4 = \frac{dx_4}{dt} = \frac{d(ict)}{dt} = ic$ . Therefore,  $u_\mu = (0, 0, 0, ic)$ .

*Example 8.10* A particle of rest mass  $m_0$  describes the trajectory  $x = f(t)$ ,  $y = g(t)$ ,  $z = 0$  in an inertial frame S. Find the four velocity components. Also show that norm of the four velocity is  $-c^2$ .

**Hint:** Four velocity is given by

$$u_\mu = \frac{dx_\mu}{d\tau} = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \frac{dx_\mu}{dt}$$

For spatial components are

$$u_i = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \frac{dx_i}{dt}$$

where,

$$v_1 = \dot{x} = \dot{f}(t), \quad v_2 = \dot{y} = \dot{g}(t), \quad v_3 = \dot{z} = 0, \quad v^2 = v_1^2 + v_2^2 + v_3^2 = \dot{f}^2 + \dot{g}^2$$

Here,

$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{1}{\sqrt{1 - \frac{(\dot{f}^2 + \dot{g}^2)}{c^2}}}$$

Now,

$$u_1 = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \frac{dx}{dt} = \frac{\dot{f}}{\sqrt{1 - \frac{(\dot{f}^2 + \dot{g}^2)}{c^2}}}$$

$$u_2 = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \frac{dy}{dt} = \frac{\dot{g}}{\sqrt{1 - \frac{(\dot{f}^2 + \dot{g}^2)}{c^2}}}$$

$$u_3 = 0$$

$$u_4 = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \frac{d(ict)}{dt} = \frac{ic}{\sqrt{1 - \frac{(\dot{f}^2 + \dot{g}^2)}{c^2}}}$$



Now norm of the four velocity is given by

$$\begin{aligned}
 u_\mu u_\mu &= u_1 u_1 + u_2 u_2 + u_3 u_3 + u_4 u_4 \\
 &= \frac{\dot{f}^2}{1 - \frac{(\dot{f}^2 + \dot{g}^2)}{c^2}} + \frac{\dot{g}^2}{1 - \frac{(\dot{f}^2 + \dot{g}^2)}{c^2}} - \frac{c^2}{1 - \frac{(\dot{f}^2 + \dot{g}^2)}{c^2}} \\
 &= -c^2
 \end{aligned}$$

*Example 8.11* A particle of rest mass  $m_0$  describes the circular path  $x = a \cos t$ ,  $y = a \sin t$ ,  $z = 0$  in an inertial frame  $S$ . Find the four velocity components. Also show that norm of the four velocity is  $-c^2$ .

**Hint:** Use  $f(t) = a \cos t$  and  $g(t) = a \sin t$  in Example 8.10.

*Example 8.12* A particle of rest mass  $m_0$  describes the parabolic trajectory  $x = at$ ,  $y = bt^2$ ,  $z = 0$  in an inertial frame  $S$ . Find the four velocity components. Also show that norm of the four velocity is  $-c^2$ .

**Hint:** Use  $f(t) = at$  and  $g(t) = bt^2$  in Example 8.10.

# Chapter 9

## Mass in Relativity

### 9.1 Relativistic Mass

We know the fundamental quantities, length and time are dependent on the observer. Therefore, it is expected that mass would be an observer dependent quantity.

#### 9.1.1 First Method Based on Hypothetical Experiment of Tolman and Lews

Let us consider an elastic collision between two identical perfectly elastic particles P and Q as seen by different inertial frame S and  $S^1$ . Let  $m_0$  be mass of each particle at rest. Here, P is in S system and Q is in  $S^1$  system. We assume that the relative velocity between S and  $S^1$  is  $v$  in which  $S^1$  is approaching to S. Let P and Q have initial velocities along y-axis that are equal in magnitude but opposite in direction, i.e.  $u_y^1 = -u_y$ . Since the collision is elastic, the final velocities have the same magnitude at the initial velocities but in opposite directions. Now, the velocity components as measured by observer in S according to law of transformation of velocities are

$$\left[ u_x = \frac{u_x^1 + v}{1 + \frac{vu_x^1}{c^2}}, \quad u_y = \frac{u_y^1 \sqrt{1 - \frac{v^2}{c^2}}}{1 + \frac{vu_x^1}{c^2}}, \quad u_z = \frac{u_z^1 \sqrt{1 - \frac{v^2}{c^2}}}{1 + \frac{vu_x^1}{c^2}} \right]$$

For particle P, we have

$$u_{xP} = 0, \quad u_{yP} = u_y, \quad u_{zP} = 0$$

For particle Q, we have

$$u_{xQ} = v, \quad u_{yQ} = -u_y \sqrt{1 - \frac{v^2}{c^2}}, \quad u_{zQ} = 0$$

[since,  $u_x^1 = u_y^1 = -u_y$ ,  $u_z^1 = 0$ ]

Thus, the resultant momentum of the whole system before collision as seen from S system is

$$= m_0 u_y \vec{j} + m v \vec{i} - m u_y \sqrt{1 - \frac{v^2}{c^2}} \vec{j} \quad (9.1)$$

where  $m$  is the mass of the particle moving with velocity  $v$  and  $\vec{i}$  &  $\vec{j}$  are unit vectors along  $x$ - and  $y$ -axes, respectively.

Now after elastic collision, the velocities of the particles P and Q as observed from S system are given by

$$w_{xP} = 0, \quad w_{yP} = -u_y, \quad w_{zP} = 0 \quad \text{and} \quad w_{xQ} = v, \quad w_{yQ} = u_y \sqrt{1 - \frac{v^2}{c^2}}, \quad w_{zQ} = 0$$

Thus, the resultant momentum of the whole system after the collision as seen from S system is

$$= -m_0 u_y \vec{j} + m v \vec{i} + m u_y \sqrt{1 - \frac{v^2}{c^2}} \vec{j} \quad (9.2)$$

Using the principle of conservation of momentum, i.e. momentum before impact = momentum after impact, we have

$$m_0 u_y \vec{j} + m v \vec{i} - m u_y \sqrt{1 - \frac{v^2}{c^2}} \vec{j} = -m_0 u_y \vec{j} + m v \vec{i} + m u_y \sqrt{1 - \frac{v^2}{c^2}} \vec{j}$$

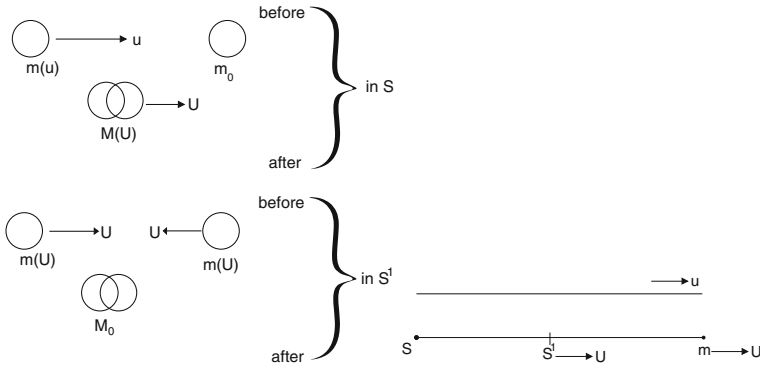
This implies

$$m = \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (9.3)$$

This mass  $m$  of the particle moving with velocity  $v$  is known as *relativistic mass* of the particle whose mass was  $m_0$  at rest. Note that mass of the particle increases with increase of velocity.

### 9.1.2 Second Method Based on D'Inverno's Thought Experiment

As length and time are changed with the motion of the observer, we assume that mass of a particle which is moving with a velocity  $u$  with respect to an inertial frame



**Fig. 9.1** Collisions of two particles as seen from two reference frames

must be a function of  $u$ , i.e.

$$m = m(u). \tag{9.4}$$

Let two particles, one be at rest and other is moving with velocity  $u$  collide (inelastically) and coalesce and combined objects move with velocity  $U$ . The masses of two particles are denoted by  $m_0$  (rest mass) and  $m(u)$  (relativistic mass). The mass of the combined particles be  $M(U)$ . Let us consider the reference system  $S^1$  to be the centre of mass frame. Therefore, it is evident that relative to  $S^1$  frame, collision takes place for the two equal particles coming with same speeds with opposite directions leaving the combined particle with mass  $M_0$  at rest. It is obvious that  $S^1$  has the velocity  $U$  relative to  $S$  (Fig. 9.1).

Now, using the principles of conservation of mass and linear momentum in  $S$  frame, we get,

$$m(u) + m_0 = M(U) \tag{9.5}$$

$$u m(u) + 0 = U M(U) \tag{9.6}$$

These give

$$m(u) = \frac{m_0 U}{u - U} \tag{9.7}$$

The left-hand particle has a velocity  $U$  relative to  $S$ . Now, using Einstein’s law of addition of velocities, we get

$$u = \frac{U + U}{1 + \frac{UU}{c^2}}$$

This equation yields the solution of  $U$  in terms of  $u$  as

$$U = \frac{c^2}{u} \left[ 1 \pm \sqrt{1 - \frac{u^2}{c^2}} \right] \quad (9.8)$$

We should take negative sign as positive sign gives  $U$  to be greater than  $c$ . Using the value of  $U$  in (9.7), we get

$$m(u) = \frac{m_0}{\sqrt{1 - \frac{u^2}{c^2}}} \quad (9.9)$$

### 9.1.3 Third Method

Let a particle of mass  $m$  is moving with velocity  $u$  in an inertial frame  $S$ . Then conservation of linear momentum in  $S$  frame implies

$$m \frac{dx_\mu}{dt} = \text{constant} \quad (9.10)$$

We rewrite this equation in terms of proper time as

$$m \frac{dx_\mu}{d\tau} \sqrt{1 - \frac{u^2}{c^2}} = \text{constant} \quad (9.11)$$

where, velocity of the particle,  $u = \left[ \left( \frac{dx_1}{dt} \right)^2 + \left( \frac{dx_2}{dt} \right)^2 + \left( \frac{dx_3}{dt} \right)^2 \right]^{\frac{1}{2}}$ . Let us consider another reference system  $S^1$  which is moving with velocity  $v$  relative to  $S$  frame. Now, we write the Eq. (9.11) with respect to  $S^1$  frame as

$$m a_{\mu\nu}^1 \frac{dx_\nu^1}{d\tau} \sqrt{1 - \frac{u^2}{c^2}} = \text{constant} \quad (9.12)$$

where  $a_{\mu\nu}^1$  is the inverse Lorentz Transformation matrix whose determinant is non-zero constant. Therefore, Eq. (9.12) implies

$$m \frac{dx_\nu^1}{d\tau} \sqrt{1 - \frac{u^2}{c^2}} = \text{constant} \quad (9.13)$$

Since the principle of conservation of linear momentum holds good in  $S^1$  frame also, we have

$$m^1 \frac{dx_\mu^1}{d\tau} \sqrt{1 - \frac{u^{12}}{c^2}} = \text{constant} \quad (9.14)$$

here,  $m^1$  is the value of mass in  $S^1$  system. Since the constants appearing in right-hand sides of Eqs. (9.13) and (9.14) may not be equal, so we multiply (9.14) by a constant  $k$  to make them equal. Hence

$$\left[ m\sqrt{1 - \frac{u^2}{c^2}} - km^1\sqrt{1 - \frac{u^{12}}{c^2}} \right] \frac{dx_\mu^1}{d\tau} = 0$$

This equation is valid for any arbitrary value of  $\frac{dx_\mu^1}{d\tau}$ , we can write

$$m\sqrt{1 - \frac{u^2}{c^2}} = km^1\sqrt{1 - \frac{u^{12}}{c^2}} = m_0 \quad (9.15)$$

Here,  $m_0$  is the mass of the particle in rest frame. In  $S$  frame, rest mass of the particle is  $m_0$  whereas in  $S^1$  frame rest mass of the particle is  $\frac{m_0}{k}$ . However, rest mass in all inertial frame remains the same. So we must have  $k = 1$ . So a moving particle with velocity  $u$  will have a mass

$$m(u) = \frac{m_0}{\sqrt{1 - \frac{u^2}{c^2}}} \quad (9.16)$$

## 9.2 Experimental Verification of Relativistic Mass

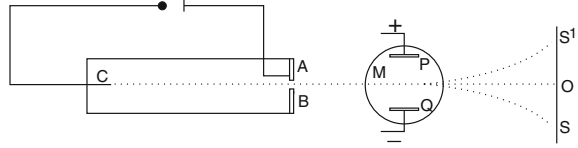
The relativistic mass formula has been verified by many scientists, however, we shall discuss the experiment done by Guey and Lavanchy.

### The Experiment Set-up is as Follows

A cathode–anode system CA is arranged in a vacuum tube where cathode and anode are connected together to the high potential terminal which produces several thousand volts. The cathode rays are collimated into fine beam by the hole B in the anode and then continue a straight line path to strike a photographic plate at O (see Fig. 9.2). Another arrangement is done to produce electric by applying a potential difference between the plates P and Q and magnetic fields by using an electromagnet M indicated by a circle in the path of the electron beam. Both of these fields are successively applied and adjusted in such a way that the deflection of the electron beam would be the same.

The force experienced by electron due to electric field  $X$  as  $Xe$  and that due to magnetic field  $H$  as  $Hev$  where  $v$  be the velocity of the beam of low speed electrons and  $e$  is the charge of the electron. Due to the above assumption,  $Hev = Xe$ , i.e.

**Fig. 9.2** The experiment set-up of Guye and Lavanchy experiment



$$v = \frac{X}{H} \tag{9.17}$$

Since the beam follows a circular path of radius, say  $a$  due to magnetic field, we have  $Hev = \frac{mv^2}{a}$  where  $m$  is mass of the beam of electrons. This implies

$$m = \frac{Hea}{v} \tag{9.18}$$

If  $X^1, H^1, v^1, m^1$  are corresponding quantities to produce the same deflections (here, beam of high speed electrons), then

$$v^1 = \frac{X^1}{H^1} \tag{9.19}$$

$$m^1 = \frac{H^1 ea}{v^1} \tag{9.20}$$

Using the above equations, we get

$$\frac{m^1}{m} = \frac{H^{12} X}{H^2 X^1} \tag{9.21}$$

Guye and Lavanchy made nearly 2,000 determinations of  $\frac{m^1}{m}$  for electrons with velocity ranging from 26 to 48 % that of light and showed that the relativistic formula of mass is almost correct with an accuracy of 1 part in 2,000.

### 9.3 Lorentz Transformation of Relativistic Mass

Let us consider the relativistic mass of a moving particle with velocity  $u$  in  $S$  frame be  $m$  with rest mass  $m_0$  and corresponding quantity be  $m^1$  in  $S^1$  which is moving with constant velocity  $v$  along  $x$ -direction. Here,

$$m = \frac{m_0}{\sqrt{1 - \frac{u^2}{c^2}}} \tag{9.22}$$

Therefore, the Lorentz Transformation of relativistic mass is given by

$$m^1 = \frac{m_0}{\sqrt{1 - \frac{u^2}{c^2}}} \tag{9.23}$$

We know from note-6 of chapter seven that

$$\sqrt{1 - \left(\frac{u^1}{c}\right)^2} = \frac{\sqrt{1 - \left(\frac{v}{c}\right)^2} \sqrt{1 - \left(\frac{u}{c}\right)^2}}{\left(1 - \frac{u_x v}{c^2}\right)} \tag{9.24}$$

Using the result of Eq.(9.24) in (9.23), we get

$$m^1 = \frac{m_0 \left(1 - \frac{u_x v}{c^2}\right)}{\sqrt{1 - \left(\frac{v}{c}\right)^2} \sqrt{1 - \left(\frac{u}{c}\right)^2}} = \frac{m \left(1 - \frac{u_x v}{c^2}\right)}{\sqrt{1 - \left(\frac{v}{c}\right)^2}} \tag{9.25}$$

The inverse transformation is given by

$$m = \frac{m^1 \left(1 + \frac{u_x^1 v}{c^2}\right)}{\sqrt{1 - \left(\frac{v}{c}\right)^2}} \tag{9.26}$$

**Note 1:** The Fig. 9.3 indicates that the relativistic mass increases without bound as  $v$  approaches to  $c$ .

*Example 9.1* A particle of rest mass  $m_0$  and velocity  $fc$  ( $0 < f < 1$ ) collides with another particle of rest mass  $m_0$  at rest and sticks. What will be the velocity and mass of the resulting particle?

**Hint** Let  $M_0$  be the rest mass and  $V$  be the velocity of the resulting particle.

$$\text{Conservation of momentum : } \frac{m_0 f c}{\sqrt{1 - \frac{f^2 c^2}{c^2}}} + m_0 \times 0 = \frac{M_0 V}{\sqrt{1 - \frac{V^2}{c^2}}}$$

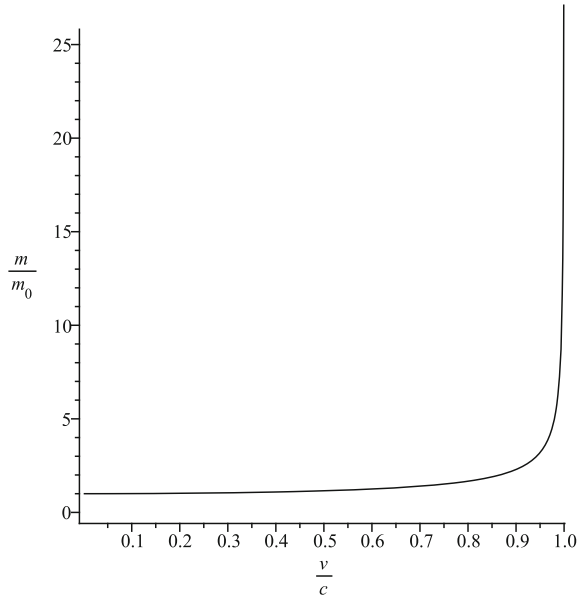
$$\text{Conservation of mass : } m_0 + \frac{m_0}{\sqrt{1 - \frac{f^2 c^2}{c^2}}} = \frac{M_0}{\sqrt{1 - \frac{V^2}{c^2}}}$$

Solving these equations, one can get  $V$  and  $M_0$ .

*Example 9.2* A particle of rest mass  $m_1$  and velocity  $v$  collides with another particle of rest mass  $m_2$  at rest and sticks. Show that the velocity and mass of the resulting particle are given by



**Fig. 9.3** The variation of mass is shown with respect to the velocity



$$M^2 = m_1^2 + m_2^2 + \frac{2m_1m_2}{\sqrt{1 - \frac{v^2}{c^2}}}, \quad V = \frac{m_1v}{m_1 + m_2\sqrt{1 - \frac{v^2}{c^2}}}$$

**Hint**  $M$  be the rest mass and  $V$  be the velocity of the resulting particle.

$$\text{Conservation of momentum : } \frac{m_1v}{\sqrt{1 - \frac{v^2}{c^2}}} + m_2 \times 0 = \frac{MV}{\sqrt{1 - \frac{V^2}{c^2}}}$$

$$\text{Conservation of mass : } m_2 + \frac{m_1}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{M}{\sqrt{1 - \frac{V^2}{c^2}}}$$

Solving these equations, one can get  $V$  and  $M$ .

**Example 9.3** Find the velocity of the particle for which the relativistic mass of the particle exceeds its rest mass by a given fraction  $f$ .

**Hint** Let  $v$  be the velocity of the particle. Now,

$$f = \frac{m - m_0}{m_0} = \frac{m}{m_0} - 1 = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} - 1 \implies v = \text{etc}$$

**Example 9.4** Find the velocity of the particle for which the relativistic mass of the particle  $a$  times its rest mass.

**Hint** Let  $v$  be the velocity of the particle. Now,

$$\frac{m}{m_0} = a \implies \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}} = am_0 \implies v = \text{etc}$$

*Example 9.5* Show that for an inelastic collisions of two particles, the rest mass of the combined particles is greater than the original rest masses.

**Hint** In an inelastic collision, a particle of rest mass  $m_1$  and velocity  $v$  collides with another particle of rest mass  $m_2$  at rest and sticks. Let  $M$  be the rest mass and  $V$  be the velocity of the resulting particle.

$$\text{Conservation of momentum : } \frac{m_1 v}{\sqrt{1 - \frac{v^2}{c^2}}} + m_2 \times 0 = \frac{MV}{\sqrt{1 - \frac{V^2}{c^2}}}$$

$$\text{Conservation of mass : } m_2 + \frac{m_1}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{M}{\sqrt{1 - \frac{V^2}{c^2}}}$$

Solving these equations, one can get  $M$  as

$$M^2 = m_1^2 + m_2^2 + \frac{2m_1 m_2}{\sqrt{1 - \frac{v^2}{c^2}}}$$

Since  $\frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} > 1$ , we have

$$M^2 > m_1^2 + m_2^2 + 2m_1 m_2 = (m_1 + m_2)^2, \text{ etc}$$

*Example 9.6* A particle of rest mass  $m_0$  and velocity  $v$  collides with another particle of rest mass  $m_0$  at rest and sticks (i.e. inelastic collision). Show that the rest mass of the combined particles is greater than the original rest masses by an amount  $\frac{m_0 v^2}{4c^2}$  (neglect  $\frac{v^4}{c^4}$  and higher order).

**Hint** In an inelastic collision, a particle of rest mass  $m_0$  and velocity  $v$  collides with another particle of rest mass  $m_0$  at rest and sticks. Let  $M$  be the rest mass and  $V$  be the velocity of the resulting particle.

$$\text{Conservation of momentum : } \frac{m_0 v}{\sqrt{1 - \frac{v^2}{c^2}}} + m_0 \times 0 = \frac{MV}{\sqrt{1 - \frac{V^2}{c^2}}}$$

$$\text{Conservation of mass : } m_0 + \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{M}{\sqrt{1 - \frac{V^2}{c^2}}}$$

Solving these equations, one can get  $M$  as

$$M^2 = 2m_0^2 + \frac{2m_0^2}{\sqrt{1 - \frac{v^2}{c^2}}} = 2m_0^2 + 2m_0^2 \left(1 - \frac{v^2}{c^2}\right)^{-\frac{1}{2}}$$

Expanding binomially and neglecting  $\frac{v^4}{c^4}$  and higher order, we get

$$M = 2m_0 \left(1 + \frac{v^2}{4c^2}\right)^{\frac{1}{2}} \implies M - 2m_0 = \frac{m_0 v^2}{4c^2}$$

[again, expanding binomially and neglecting  $\frac{v^4}{c^4}$  and higher order]

*Example 9.7* The average lifetime of  $\pi$ -mesons at rest is  $t$  s. A laboratory measurement on  $\pi$ -mesons yields an average lifetime of  $nt$  s. What is the speed of the  $\pi$ -mesons in the laboratory.

**Hint** Here,  $t$  be the lifetime of the rest  $\pi$ -mesons. Let  $v$  be its velocity. According to time dilation,

$$t^1 = nt = \frac{t}{\sqrt{1 - \frac{v^2}{c^2}}} \implies v = \text{etc}$$

*Example 9.8* Let a constant force  $F$  applied on an object with rest mass  $m_0$  at a rest position. Prove that its velocity after a time  $t$  is  $v = \frac{cFt}{\sqrt{m_0^2 c^2 + F^2 t^2}}$ . Prove also that the above result is in agreement with classical result. Further find  $v$  after a very long time.

**Hint** We know

$$F = \frac{d}{dt}(mv) = \frac{d}{dt} \left( \frac{m_0 v}{\sqrt{1 - \frac{v^2}{c^2}}} \right)$$

Now, integrating we get,

$$\int F dt = \int d \left( \frac{m_0 v}{\sqrt{1 - \frac{v^2}{c^2}}} \right)$$

This implies

$$Ft = \frac{m_0 v}{\sqrt{1 - \frac{v^2}{c^2}}} \text{ etc}$$

If  $t$  is small, then  $t^2 + \frac{m_0^2 c^2}{F^2} \approx \frac{m_0^2 c^2}{F^2}$ . Thus

$$v = \frac{cFt}{\sqrt{m_0^2c^2 + F^2t^2}} = \frac{cFt}{F\sqrt{t^2 + \frac{m_0^2c^2}{F^2}}} \approx \frac{cFt}{F\sqrt{\frac{m_0^2c^2}{F^2}}} = \frac{Ftc}{m_0c}$$

This implies

$$F \approx \frac{m_0v}{t} = m_0f$$

[ $f = \frac{v}{t}$  = acceleration]

If  $t$  is large enough, then

$$v = \frac{cFt}{\sqrt{m_0^2c^2 + F^2t^2}} = \frac{cF}{\sqrt{\frac{m_0^2c^2}{t^2} + F^2}} \approx \frac{Fc}{\sqrt{F^2}} = c$$

[since  $\frac{m_0^2c^2}{t^2} \approx 0$  for large  $t$ ]

# Chapter 10

## Relativistic Dynamics

### 10.1 Four Force or Minkowski Force

We proceed to generalize the Newton's equation of motion

$$F_j = \frac{d(mv_j)}{dt}, \quad j = 1, 2, 3 \tag{10.1}$$

which is not invariant under Lorentz Transformation. Its relativistic generalization should be a four vector equation, the spatial part of which would reduce to equation (10.1) in the limit  $\beta = \frac{v}{c} \rightarrow 0$ .

We shall bring the following changes:

1. Since  $t$  is not Lorentz invariant, it should be replaced by proper time  $\tau$ .
2. The rest mass  $m_0$  can be taken as an invariant property of the particle.
3. In place of  $v_j$ , world velocity  $u_\mu$  should be substituted.
4. The left-hand side force  $F_i$  should be replaced by some four vector  $K_\mu$  (called *Minkowski force*).

Thus taking into account all these changes, the relativistic generalization of equation (10.1) is

$$K_\mu = \frac{d(m_0 u_\mu)}{d\tau} \quad \mu = 1, 2, 3, 4 \tag{10.2}$$

The spatial part of the above equation can be written as

$$\frac{ad(m_0 av_j)}{dt} = K_j, \quad a = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$$

or,

$$\frac{d(m_0 av_j)}{dt} = \frac{K_j}{a} \tag{10.3}$$

If we continue to use the classical definition of force and define force as being the time rate of change of momentum in all Lorentz systems, then the classical force should be defined as

$$\vec{F} = \frac{d(m\vec{v})}{dt} \implies F_j = \frac{d(mv_j)}{dt} \quad (10.4)$$

This is called *Relativistic force*.

Actually, it is measured from S frame of a moving particle of rest mass  $m_0$ .

Hence, the space components of the force  $K_\mu$  is related to  $F_j$  as

$$K_j = aF_j \quad (10.5)$$

[these are Relativistic force, i.e. space components of the Minkowski forces]

**Note 1:** Minkowski force and four velocity vector are orthogonal to each other.

*Proof* Here,

$$\begin{aligned} K_\mu u_\mu &= \frac{d(m_0 u_\mu)}{d\tau} u_\mu \\ &= \frac{1}{2} \frac{d(m_0 u_\mu u_\mu)}{d\tau} \\ &= \frac{1}{2} \frac{d(-m_0 c^2)}{d\tau} = 0 \end{aligned}$$

Now, we try to find the time-like part of the force vector  $K_\mu$ , i.e. the fourth component of Minkowski force, with the help of the above theorem.

We write the result  $K_\mu u_\mu = 0$  explicitly which yields

$$K_\mu u_\mu = K_1 u_1 + K_2 u_2 + K_3 u_3 + K_4 u_4 = 0$$

or

$$a^2 F_j v_j + K_4 a i c = 0$$

or

$$a^2 \vec{F} \cdot \vec{v} + K_4 a i c = 0$$

or

$$K_4 = \left( \frac{ia}{c} \right) \vec{F} \cdot \vec{v} \quad (10.6)$$

Therefore, the Minkowski force is given by

$$K_\mu = \left[ aF_1, aF_2, aF_3, \left( \frac{ia}{c} \right) \vec{F} \cdot \vec{v} \right]; \quad a = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (10.7)$$

## 10.2 Four Momentum

In Newtonian mechanics, the momentum vector  $\mathbf{p}$  is obtained by multiplying the velocity vector  $\mathbf{v}$  by a mass  $m$  which is independent of the reference frame. The four momentum vector  $p_\mu$  is obtained similarly from the four velocity vector  $u_\mu = [av_1, av_2, av_3, ic]$  by multiplication with mass factor independent of the frame of reference. This mass factor is called the rest mass,  $m_0$ . Hence we get the *four momentum vector* as,

$$p_\mu = m_0[av_1, av_2, av_3, ic] = am_0[v_1, v_2, v_3, ic] \quad (10.8)$$

The quantity in front of the bracket is generally called the relativistic mass. We therefore have,

$$m = am_0 = \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}}$$

The spatial components are

$$p_j = \frac{m_0 v_j}{\sqrt{1 - \frac{v^2}{c^2}}} = m v_j, \quad j = 1, 2, 3$$

This equation gives the relativistic definition of classical linear momentum of the particle.

The four or time component is

$$p_4 = m_0 u_4 = \frac{m_0 ic}{\sqrt{1 - \frac{v^2}{c^2}}} = icm \quad (10.9)$$

The norm of the four momentum vector is

$$p_\mu p_\mu = -m_0^2 c^2 \quad (10.10)$$

## 10.3 Relativistic Kinetic Energy

Minkowski force is defined by the rate of change of momentum and here, time  $t$  is replaced by proper time  $\tau$ . Therefore,

$$K_\mu = \frac{dp_\mu}{d\tau} = \frac{d(m_0 u_\mu)}{d\tau} \quad (10.11)$$

Thus,

$$K_j = \frac{d(am_0v_j)}{d\tau}$$

From Eq. (10.4), we can write,

$$F_j = \frac{d(am_0v_j)}{dt} \quad (10.12)$$

Again, we know that Minkowski force and four velocity vector are orthogonal to each other, then we have,

$$u_j K_j + u_4 K_4 = 0$$

This implies

$$av_j \frac{d(am_0v_j)}{d\tau} = -aic \frac{d(am_0ic)}{d\tau}$$

or,

$$v_j \frac{d(am_0v_j)}{dt} = v_j F_j \equiv \vec{v} \cdot \vec{F} = \frac{d(am_0c^2)}{dt} \quad (10.13)$$

[using equation (10.12)]

We know that the rate of change of kinetic energy ( $T$ ) of a particle is given by the rate at which the force does work on it which is  $\vec{v} \cdot \vec{F}$ . Using the above definition of the rate of change of kinetic energy, we get from Eq. (10.13)

$$\vec{v} \cdot \vec{F} = \frac{dT}{dt} = \frac{d(am_0c^2)}{dt} \quad (10.14)$$

This implies

$$T = am_0c^2 + T_0, \quad T_0 \text{ is an integration constant}$$

When,  $v = 0$ , then  $T = 0$ , therefore,  $T_0 = -m_0c^2$ , hence

$$T = am_0c^2 - m_0c^2 = (m - m_0)c^2 \quad (10.15)$$

This is the *relativistic kinetic energy* of a freely moving particle.

Actually, it is measured from S frame of a moving particle of rest mass  $m_0$ , i.e. kinetic energy of a moving particle with relativistic effects.

**Note 2:** Relativistic kinetic energy of a moving particle is equal to  $c^2$  times its gain in mass due to motion.

*Alternative derivation of relativistic kinetic energy:*

We have,

$$m = \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}}, \text{ i.e. } m^2c^2 - m^2v^2 = m_0^2c^2$$



From this, we can get

$$2mc^2 dm - 2mv^2 dm - 2m^2 v dv = 0$$

i.e.

$$v^2 dm + mvdv = c^2 dm \quad (10.16)$$

When a force  $\vec{F}$  acts on the mass  $m$  so as to move it through a distance  $d\vec{x}$ , the work done is  $\vec{F} \cdot d\vec{x}$  which is converted into kinetic energy  $T$  as

$$\begin{aligned} T &= \int_{v=0}^v \vec{F} \cdot d\vec{x} \\ &= \int_{v=0}^v \frac{d\vec{p}}{dt} \cdot d\vec{x} \\ &= \int_{v=0}^v \frac{d(m\vec{v})}{dt} \cdot d\vec{x} \\ &= \int_{v=0}^v d(m\vec{v}) \cdot \frac{d\vec{x}}{dt} \\ &= \int_{v=0}^v (m d\vec{v} + \vec{v} dm) \cdot \vec{v} \\ &= \int_{v=0}^v (m\vec{v} \cdot d\vec{v} + v^2 dm) \end{aligned}$$

Since  $\vec{v} \cdot d\vec{v} = v dv$ , therefore, using (10.16), we get

$$T = \int_{m=m_0}^m c^2 dm = mc^2 - m_0 c^2 = (m - m_0)c^2 = m_0 c^2 (a - 1) \quad (10.17)$$

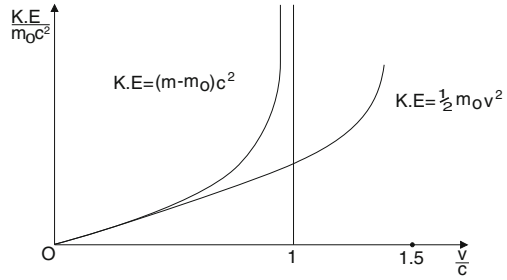
[when,  $v = 0$ ,  $m = m_0 = \text{rest mass}$ ,  $a = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$ ]

**Classical result:** The above relativistic expression for kinetic energy  $T$  must reduce to classical result,  $\frac{1}{2}m_0 v^2$ ,  $\frac{v}{c} \ll 1$ .

**To check:**

Now,

**Fig. 10.1** Variation of classical and relativistic kinetic energy with respect to the velocity



$$\begin{aligned}
 T &= m_0 c^2 (a - 1) \\
 &= m_0 c^2 \left( \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} - 1 \right) \\
 &= m_0 c^2 \left[ \left( 1 - \frac{v^2}{c^2} \right)^{-\frac{1}{2}} - 1 \right]
 \end{aligned}$$

Expanding binomially and neglecting  $\frac{v^4}{c^4}$  and higher order, we get

$$T \approx m_0 c^2 \left[ \left( 1 + \frac{v^2}{2c^2} \right) - 1 \right] = \frac{1}{2} m_0 v^2$$

Hence the result.

The Figure 10.1 shows how the kinetic energy of a moving body varies with its speed according to both classical and relativistic mechanics. At low speeds, i.e.  $\frac{v}{c} \ll 1$ , the formulae give the same results, but diverge at speeds approaching that of light. Relativity theory indicates that a body would need infinite kinetic energy to travel with speed of light, whereas in classical mechanics it would need only a kinetic energy of half its rest energy to have this speed. This again confirms that  $c$  plays the role of limiting velocity.

## 10.4 Mass–Energy Relation

The expression of kinetic energy  $T = (m - m_0)c^2$  suggests that the kinetic energy of a moving body is equal to the increase in mass times the square of the speed of light. Therefore,  $m_0 c^2$  may be regarded as *rest energy* of a body of rest mass  $m_0$ . It is argued that the rest mass has energy  $m_0 c^2$  due to the presence of an internal store of energy. Thus, total energy  $E$  of the body would be the sum of rest energy and kinetic energy (relativistic due to motion). Hence,

$$E = m_0c^2 + (m - m_0)c^2 = mc^2 = \frac{m_0c^2}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (10.18)$$

It is the famous principle of mass–energy equivalence which states a universal equivalence of mass and energy. Here,  $E$  is the relativistic energy of a particle. Einstein argued that mass can be regarded as the source of energy in other word, mass is transformed into energy and vice versa. Thus, law of conservation of mass automatically implies the so called law of conservation of energy.

Recall the four or time component of momentum as

$$p_4 = icm = \frac{iE}{c} \quad (10.19)$$

Hence the *four momentum vector* takes the form as,

$$p_\mu = \left[ p_1, p_2, p_3, \frac{iE}{c} \right] \quad (10.20)$$

## 10.5 Relation Between Momentum and Energy

Total energy and momentum are conserved quantities and the rest energy of a particle is invariant quantity. Now, we try to find how the total energy, rest energy and momentum of a particle are related. We know, total energy is

$$E = mc^2 = \frac{m_0c^2}{\sqrt{1 - \frac{v^2}{c^2}}}$$

This implies

$$E^2 = \frac{m_0^2c^4}{1 - \frac{v^2}{c^2}} \quad (10.21)$$

The momentum is given by

$$p = mv = \frac{m_0v}{\sqrt{1 - \frac{v^2}{c^2}}}$$

This gives

$$p^2c^2 = \frac{m_0^2v^2c^2}{1 - \frac{v^2}{c^2}} \quad (10.22)$$

Subtracting (10.21) from (10.20), we get,

$$E^2 - p^2c^2 = m_0^2c^4 \quad (10.23)$$

The above relation holds for a single particle. For particles in a system that are moving with respect to other, the sum of their individual rest energies may not equal the rest energy of the system.

**Alternative**

We know,

$$p_\mu p_\mu = [p_4p_4 + p_1p_1 + p_2p_2 + p_3p_3] = -\frac{E^2}{c^2} + p^2 \quad (10.24)$$

Also,

$$p_\mu p_\mu = m_0^2 u_\mu u_\mu = -m_0^2 c^2 \quad (10.25)$$

Equating (10.24) and (10.25), we get,

$$-\frac{E^2}{c^2} + p^2 = -m_0c^2 \implies E^2 - p^2c^2 = m_0^2c^4$$

**Note 3:** Since  $m_0c^2$  is an invariant quantity, therefore,  $E^2 - p^2c^2$  is also an invariant quantity.

**Note 4:** Differentiating Eq.(10.23) with respect to  $p$ , we get

$$\frac{dE}{dp} = \frac{pc}{\sqrt{p^2 + m_0^2c^2}} = \frac{pc^2}{E} = v$$

This is another useful relation between momentum and energy.

## 10.6 Evidence in Support of Mass–Energy Relation

In fusion process, when two neutrons and protons make helium  ${}_2\text{He}^4$  nucleus, enormous energy is found. The explanation is as follows:

The mass of two protons =  $2 \times$  mass of hydrogen nucleus

$$= 2 \times m({}_1\text{H}^1) = 2 \times (1.00815) \text{ a.m.u.}$$

Mass of two neutrons =  $2 \times m({}_0n^1) = 2 \times (1.00898) \text{ a.m.u.}$

Therefore,

$$2 \times m({}_1\text{H}^1) + 2 \times m({}_0n^1) = 2 \times (1.00815 + 1.00898) \text{ a.m.u.} = 4.03426 \text{ a.m.u.}$$

The difference in mass of the  ${}^2\text{He}^4$  nucleus and total mass of the constituent nucleus in free state is

$$(4.03426 - 4.00387) = 0.03039 \text{ a.m.u.} = .03039 \times 931.1 \text{ MeV} = 28.3 \text{ MeV}$$

This explains why during fusion process of two neutrons and two protons, tremendous amount of energy is found.

Scientists believe that this fusion process occurs in the sun to provide solar energy.

## 10.7 Force in Special Theory of Relativity

A force is not in general proportional to acceleration in case of special theory of relativity.

In general, force is given by

$$\vec{F} = \frac{d\vec{p}}{dt} = \frac{d(m\vec{v})}{dt}$$

This implies

$$\vec{F} = m \frac{d\vec{v}}{dt} + \vec{v} \frac{dm}{dt} \quad (10.26)$$

Using mass–energy relation,  $E = mc^2$ , we have

$$\frac{dm}{dt} = \frac{1}{c^2} \frac{dE}{dt} = \frac{1}{c^2} \frac{d(T + m_0c^2)}{dt} = \frac{1}{c^2} \frac{dT}{dt}$$

[ $T$  = kinetic energy and  $m_0c^2$  = rest energy]

Again we know,

$$\frac{dT}{dt} = \frac{d(\vec{F} \cdot d\vec{x})}{dt} = \vec{F} \cdot \frac{d\vec{v}}{dt} = \vec{F} \cdot \vec{v}$$

Hence,

$$\frac{dm}{dt} = \frac{1}{c^2} (\vec{F} \cdot \vec{v}) \quad (10.27)$$

Using (10.26) and (10.27), we get,

$$\vec{F} = m \frac{d\vec{v}}{dt} + \left( \frac{1}{mc^2} \right) (\vec{v} \vec{F} \cdot \vec{v}) \quad (10.28)$$

The acceleration  $\vec{f}$  is defined by  $\vec{f} = \frac{d\vec{v}}{dt}$ . Therefore,

$$\vec{f} = \frac{\vec{F}}{m} - \left( \frac{1}{mc^2} \right) (\vec{v} \vec{F} \cdot \vec{v}) \quad (10.29)$$

Thus force is not in general, proportional to acceleration.

Also Eq. (10.29) indicates that in general, acceleration  $\vec{f}$  is not parallel to the force in relativity, since the last term is in the direction of the velocity  $\vec{v}$ . If the velocity is perpendicular to the force, then acceleration  $\vec{f}$  is parallel to the force.

## 10.8 Covariant Formulation of the Newton's Law

The covariant formulation of the Newton's law implies that the force is parallel to the acceleration only when the velocity is either parallel or perpendicular to the acceleration.

Following Newton's law, we write the force as

$$K_\mu = \frac{dp_\mu}{d\tau} = \frac{d(m_0 u_\mu)}{d\tau}$$

The spatial components are

$$K_j = \frac{d(m_0 u_j)}{d\tau} = a \frac{d(am_0 v_j)}{dt}$$

[where  $a = \frac{1}{\sqrt{1-\frac{v^2}{c^2}}}$ ,  $v_j v_j = v^2$ ,  $u_j = av_j$ ]

Thus

$$K_j = m_0 b_j a^2 + \frac{(a^4 m_0 v_j v_k b_k)}{c^2} \quad (10.30)$$

where,  $b_j = \frac{d(v_j)}{dt} = \text{acceleration}$ .

$$\left[ \frac{d(v^2)}{dt} = \frac{d(v_k v_k)}{dt} = 2v_k b_k \right]$$

Note that spatial components of the force will be parallel to the acceleration if second term in (10.30) vanishes. Now, if the velocity vector is perpendicular to the acceleration vector, then  $u_j b_j = 0$  and the second term vanishes. Hence,

$$K_j = m_0 b_j a^2$$

We know the space components of the force  $K_\mu$  is related to  $F_j$  (Newtonian force) as  $K_j = aF_j$ . Therefore,

$$\left(\frac{1}{a}\right) K_j = m_0 b_j a = \frac{m_0 b_j}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (10.31)$$

The coefficient of  $b_j$  of Eq.(10.31) is referred as the *Transverse mass* of the body.

On the other hand, if the velocity is parallel to the acceleration, then  $\vec{b} = h \vec{v}$  [ $h = \text{constant}$ ].

Thus from (10.30), we get,

$$\begin{aligned} K_j &= m_0 b_j a^2 + \frac{(a^4 m_0 v_j v_k h v_k)}{c^2} \\ &= m_0 b_j a^2 + \frac{(a^4 m_0 h v_j v^2)}{c^2} \\ &= m_0 b_j a^4 \end{aligned}$$

[ $v_k v_k = v^2$ ,  $b_j = h v_j$ ]

The space components of the force  $K_\mu$  is related to  $F_j$  (Newtonian force) as  $K_j = a F_j$ . Therefore,

$$F_j = \left(\frac{1}{a}\right) K_j = m_0 b_j a^3 = m_0 b_j \left(1 - \frac{v^2}{c^2}\right)^{-\frac{3}{2}} \quad (10.32)$$

The coefficient of  $b_j$  of Eq.(10.32) is referred as the *Longitudinal mass* of the body.

## 10.9 Examples of Longitudinal Mass and Transverse Mass

The case in which the force  $\vec{F}$  is parallel to velocity, then the acceleration  $\vec{b}$  is parallel to both  $\vec{v}$  and  $\vec{F}$ . The particle moves in a straight line such as a charged particle starts from rest in a uniform electric field is the example of longitudinal mass. The case in which the force  $\vec{F}$  is perpendicular to the velocity  $\vec{v}$ , for then  $\vec{F} \cdot \vec{v} = 0$ . The force on a charged particle moving with velocity  $\vec{v}$  in a magnetic field is the example of transverse mass.

## 10.10 The Lorentz Transformation of Momentum

The four momentum can be written as

$$p_\mu = [p_1, p_2, p_3, p_4] \quad (10.33)$$

where  $p_4 = (\frac{iE}{c})$ . We know Lorentz Transformation of any four vector  $A_\mu$  ( $\mu = 1, 2, 3, 4$ ) may be expressed as

$$A_\mu^1 = a_{\mu\nu}A_\nu$$

where  $a_{\mu\nu}$  is given by

$$a_{\mu\nu} = \begin{bmatrix} a & 0 & 0 & ia\beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -ia\beta & 0 & 0 & a \end{bmatrix}$$

$$[\beta = \frac{v}{c} \text{ and } a = \frac{1}{\sqrt{1-\beta^2}}]$$

Hence,

$$\begin{bmatrix} p_1^1 \\ p_2^1 \\ p_3^1 \\ p_4^1 \end{bmatrix} = \begin{bmatrix} a & 0 & 0 & ia\beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -ia\beta & 0 & 0 & a \end{bmatrix} \begin{bmatrix} p_1 \\ p_2 \\ p_3 \\ p_4 \end{bmatrix} \quad (10.34)$$

This implies,

$$p_1^1 = ap_1 + ia\beta p_4 = ap_1 + ia \left(\frac{v}{c}\right) \left(\frac{iE}{c}\right)$$

i.e.

$$p_1^1 = a \left[ p_1 - \left(\frac{vE}{c^2}\right) \right] \quad (10.35)$$

$$p_2^1 = p_2, \quad p_3^1 = p_3 \quad (10.36)$$

$$p_4^1 = -ia\beta p_1 + ap_4$$

or,

$$\left(\frac{iE^1}{c}\right) = -ia \left(\frac{v}{c}\right) p_1 + a \left(\frac{iE}{c}\right)$$

i.e.

$$E^1 = a[E - vp_1] \quad (10.37)$$

The inverse transformations are obtained by replacing  $v$  by  $-v$  as

$$p_1 = a \left[ p_1^1 + \left(\frac{vE^1}{c^2}\right) \right], \quad p_2 = p_2^1, \quad p_3 = p_3^1, \quad E = a[E^1 + vp_1^1] \quad (10.38)$$

Let a body be at rest in the  $S^1$  frame so that its rest energy is  $m_0$ . This means for observer in S frame the body is moving with velocity  $v$ . So, in  $S^1$  frame,



$$p_1^1 = p_2^1 = p_3^1 = 0, \quad E^1 = m_0 c^2$$

Hence,

$$p_1 = \frac{m_0 v}{\sqrt{1 - \frac{v^2}{c^2}}}, \quad p_2 = 0, \quad p_3 = 0, \quad E = \frac{m_0 c^2}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (10.39)$$

Suppose a body be at rest in the  $S^1$  frame and emits radiation, then its energy is  $E^1$ . Therefore, the components of the momentum four vector in S frame are

$$p_1 = \frac{v E^1}{c^2 \sqrt{1 - \frac{v^2}{c^2}}}, \quad p_2 = 0, \quad p_3 = 0, \quad E = \frac{E^1}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (10.40)$$

### 10.11 The Expression $p^2 - \frac{E^2}{c^2}$ Is Invariant Under Lorentz Transformation

We know

$$p_1^1 = a \left[ p_1 - \left( \frac{vE}{c^2} \right) \right], \quad p_2^1 = p_2, \quad p_3^1 = p_3, \quad E^1 = a [E - vp_1]$$

Now,

$$\begin{aligned} p^{12} - \frac{E^{12}}{c^2} &= p_1^{12} + p_2^{12} + p_3^{12} - \frac{E^{12}}{c^2} \\ &= a^2 \left[ p_1 - \left( \frac{vE}{c^2} \right) \right]^2 + p_2^2 + p_3^2 - \frac{a^2 [E - vp_1]^2}{c^2} \\ &= p^2 - \frac{E^2}{c^2} \quad (\text{using, } p^2 = p_1^2 + p_2^2 + p_3^2) \end{aligned}$$

#### Alternative

The relation between momentum and energy is given by

$$E^2 = p^2 c^2 + m_0^2 c^4$$

From this relation, we can write,

$$E^2 - p^2 c^2 = m_0^2 c^4$$

We know, the rest mass and velocity of light are invariant quantities, therefore,  $E^2 - p^2 c^2$  is also an invariant quantity, i.e.

$$E^{12} - p^{12}c^2 = E^2 - p^2c^2$$

This gives

$$p^{12} - \frac{E^{12}}{c^2} = p^2 - \frac{E^2}{c^2}$$

*Example 10.1* Show that  $m_0 = \frac{p^2c^2 - T^2}{2Tc^2}$ , where,  $m_0$  is the rest mass of a particle with momentum,  $p$  and kinetic energy  $T$ .

**Hint:** We know,

$$E^2 = p^2c^2 + m_0^2c^4 \quad (1)$$

$$T = mc^2 - m_0c^2 = E - m_0c^2 \quad (2)$$

Now,

$$T^2 = E^2 - 2Em_0c^2 + m_0^2c^4 = p^2c^2 + 2m_0^2c^4 - 2Em_0c^2$$

or,

$$p^2c^2 - T^2 = 2Tm_0c^2 \text{ etc}$$

*Example 10.2* Show that if a particle is highly relativistic, then the fractional difference between  $c$  and  $v$  is approximately  $\frac{1}{2} \left( \frac{m_0c^2}{E} \right)^2$  where  $E$  is the relativistic energy.

**Hint:** Since the particle is highly relativistic, then  $\frac{v^2}{c^2}$  is very close to unity, therefore,  $1 - \frac{v^2}{c^2}$  is a very small quantity, i.e.

$$1 - \frac{v^2}{c^2} = \epsilon, \text{ say}$$

or

$$\frac{v^2}{c^2} = 1 - \epsilon \implies \frac{v}{c} = 1 - \frac{1}{2}\epsilon$$

[expanding binomially and neglecting higher order of  $\epsilon$ ]

This implies

$$\epsilon = \frac{2(c - v)}{c}$$

Now,

$$E = mc^2 = \frac{m_0c^2}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{m_0c^2}{\sqrt{\epsilon}}$$

Replacing  $\epsilon$ , we get

$$\frac{(c - v)}{c} = \frac{1}{2} \left( \frac{m_0c^2}{E} \right)^2$$

*Example 10.3* Can a massless particle exist?

**Hint:** We know total energy and relativistic momentum

$$E = mc^2 = \frac{m_0c^2}{\sqrt{1 - \frac{v^2}{c^2}}} \text{ and } p = mv = \frac{m_0v}{\sqrt{1 - \frac{v^2}{c^2}}}$$

When  $m_0 = 0$  and  $v < c$ , then it is evident that  $E = p = 0$ . Thus a massless particle with a speed less than  $c$  can have neither energy nor momentum. However, when  $m_0 = 0$  and  $v = c$ , then,  $E = \frac{0}{0}$  and  $p = \frac{0}{0}$ , which are indeterminate. These imply  $E$  and  $p$  can have any values. Thus, massless particles with non-zero energy and momentum must travel with the speed of light.

Examples of that particle are photon and neutrino.

*Example 10.4* Show that integral of the world force over the world line vanishes if the rest mass is assumed not to change.

**Hint:**

$$\begin{aligned} \int K_\mu dx_\mu &= \int K_\mu \frac{dx_\mu}{d\tau} d\tau = \int K_\mu u_\mu d\tau \\ &= \int \frac{d(m_0 u_\mu)}{d\tau} u_\mu d\tau = \int \frac{d(\frac{1}{2} m_0 u_\mu u_\mu)}{d\tau} d\tau \\ &= - \int \frac{d(\frac{1}{2} m_0 c^2)}{d\tau} d\tau = 0 \end{aligned}$$

Since rest mass is not changed.

*Example 10.5* Find the integral of the relativistic force over 3D path.

**Hint:** When a relativistic force  $\vec{F}$  acts on a mass  $m$ , so as to move it through a distance  $d\vec{x}$ , then  $\int \vec{F} \cdot d\vec{x} = I$  is the work done. Thus,

$$\begin{aligned} I &= \int \vec{F} \cdot d\vec{x} = \int \frac{d(m\vec{v})}{dt} \cdot d\vec{x} \\ &= \int \frac{d(am_0\vec{v})}{dt} \cdot d\vec{x} = \int d(am_0\vec{v}) \cdot \frac{d\vec{x}}{dt} \\ &= \int d(am_0\vec{v}) \cdot \vec{v} = \int am_0\vec{v} \cdot d\vec{v} \\ &= \int am_0 v dv = \int \frac{m_0 v dv}{\sqrt{1 - \frac{v^2}{c^2}}} \end{aligned}$$

[since  $v^2 = \vec{v} \cdot \vec{v} \implies v dv = \vec{v} \cdot d\vec{v}$  and  $a = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$ ]

After integration, we get

$$I = -c^2 m_0 \sqrt{1 - \frac{v^2}{c^2}} + D$$

When  $v = 0$ , work done is zero  $\implies D = m_0 c^2$ .

Thus,

$$I = m_0 c^2 - c^2 m_0 \sqrt{1 - \frac{v^2}{c^2}} = T = \text{relativistic kinetic energy}$$

The above relativistic expression for kinetic energy  $T$  must reduce to classical result,  $\frac{1}{2} m_0 v^2$ , when,  $\frac{v}{c} \ll 1$ .

Now,

$$T = m_0 c^2 \left[ 1 - \left( 1 - \frac{v^2}{c^2} \right)^{\frac{1}{2}} \right]$$

Expanding binomially and neglecting  $\frac{v^4}{c^4}$  and higher order, we get

$$T \approx m_0 c^2 \left[ \left( 1 - 1 + \frac{v^2}{2c^2} \right) \right] = \frac{1}{2} m_0 v^2$$

Hence the result.

*Example 10.6* Calculate the amount gain in the mass of earth in a year, if approximately 2 calories of radiant energy are received by each square centimetre of earth surface per minute.

[Given, radius of earth =  $6.4 \times 10^3$  km.]

**Hint:** Total surface area of earth is  $4\pi r^2$ , therefore, energy received in one minute is  $4\pi r^2 \times$  energy received by each square centimetre of earth surface per minute which is given by (since 2 calories =  $2 \times 4.2$  J)

$$= 4 \times 3.14 \times (6.4 \times 10^6)^2 \times 2 \times 4.2 \times 10^4 \text{ J}$$

Energy received in one year

$$= 4 \times 3.14 \times (6.4 \times 10^6)^2 \times 2 \times 4.2 \times 10^4 \times 60 \times 24 \times 365 \text{ J}$$

Hence, annual gain in the mass of earth

$$\begin{aligned} \Delta m &= \frac{\Delta E}{c^2} = \frac{4 \times 3.14 \times (6.4 \times 10^6)^2 \times 2 \times 4.2 \times 10^4 \times 60 \times 24 \times 365}{(3 \times 10^8)^2} \\ &= 2.524 \times 10^8 \text{ kg per year.} \end{aligned}$$

*Example 10.7* A particle of rest mass  $m_0$  moving with relativistic velocity  $v$  has got a momentum  $p$  and kinetic energy  $T$ . Show that  $\frac{pv}{T} = \frac{T+2m_0c^2}{T+m_0c^2}$ .

**Hint:** We know,

$$E = mc^2 = \frac{pc^2}{v}$$

Therefore,

$$p = \frac{vE}{c^2} = \frac{v}{c^2}(T + m_0c^2) \quad (1)$$

Again,

$$E = T + m_0c^2 = \frac{m_0c^2}{\sqrt{1 - \frac{v^2}{c^2}}} \implies v^2 = \frac{c^2(T^2 + 2m_0Tc^2)}{(T + m_0c^2)^2} \quad (2)$$

Hence,

$$\frac{pv}{T} = \text{etc.}$$

*Example 10.8* Find the velocity that one electron must be given so that its momentum is  $a$  times its rest mass times speed of light. What is the energy of this speed?

**Hint:** The momentum of an electron of rest mass  $m_0 = 9 \times 10^{-31}$  kg moving with velocity  $v$  is given by

$$p = a \times m_0c^2 = \frac{m_0v}{\sqrt{1 - \frac{v^2}{c^2}}} \implies v =$$

Now,

$$E = \frac{m_0c^2}{\sqrt{1 - \frac{v^2}{c^2}}} = \text{etc.}$$

*Example 10.9* Calculate the velocity of an electron having a total energy of  $a$  MeV (rest mass of the electron is  $m_0 = 9 \times 10^{-31}$  kg).

**Hint:** Here,

$$E = \frac{m_0c^2}{\sqrt{1 - \frac{v^2}{c^2}}} = a \text{ MeV} = a \times 10^6 \text{ eV} = a \times 10^6 \times 1.6 \times 10^{-19} \text{ J} \implies v = \text{etc}$$

*Example 10.10* A  $\pi$ -meson of rest mass  $m_\pi$  decays into  $\mu$ -meson of mass  $m_\mu$  and a neutrino of mass  $m_\nu$ . Show that the total energy of the  $\mu$ -meson is  $\frac{1}{2m_\pi}[m_\pi^2 + m_\mu^2 - m_\nu^2]c^2$ .

**Hint:** According to principle of conservation of momentum, if the  $\mu$ -meson of mass  $m_\mu$  moves with momentum  $p$ , then neutrino of mass  $m_\nu$  moves with momentum  $-p$ . Therefore, according to the momentum–energy relation, we have

$$E_\mu^2 = p^2 c^2 + m_\mu^2 c^4, \quad E_\nu^2 = p^2 c^2 + m_\nu^2 c^4 \quad (1)$$

From these, we get,

$$E_\mu^2 - E_\nu^2 = (m_\mu^2 - m_\nu^2) c^4 \quad (2)$$

Total energy

$$E_\mu + E_\nu = E = m_\pi c^2 \quad (3)$$

Solving (2) and (3), one can get, total energy of the  $\mu$ -meson as  $E_\mu =$  etc.

A  $\pi$ -meson of rest mass  $m_\pi$  decays into  $\mu$ -meson of mass  $m_\mu$  and a neutrino of mass  $m_\nu$ . Show that the total energy of the  $\mu$ -meson is  $\frac{1}{2m_\pi} [m_\pi^2 + m_\mu^2 - m_\nu^2] c^2$ .

*Example 10.11* A particle of rest mass  $M$  decays spontaneously into two masses  $m_1$  and  $m_2$ . Obtain energies of the daughter masses. Show also that their kinetic energy is given by  $T_i = \Delta M \left(1 - \frac{m_i}{M} - \frac{\Delta M}{2M}\right) c^2$  where  $\Delta M = M - m_1 - m_2$ , i.e. difference between initial and final masses.

**Hint:** According to principle of conservation of momentum, if the  $m_1$  mass moves with momentum  $p$ , then other mass  $m_2$  moves with momentum  $-p$ . Therefore, according to the momentum–energy relation, we have

$$E_1^2 = p^2 c^2 + m_1^2 c^4, \quad E_2^2 = p^2 c^2 + m_2^2 c^4 \quad (1)$$

(here, energies of the daughter masses are  $E_1$  and  $E_2$ , respectively)

From these, we get,

$$E_1^2 - E_2^2 = (m_1^2 - m_2^2) c^4 \quad (2)$$

Total energy

$$E_1 + E_2 = E = M c^2 \quad (3)$$

Solving (2) and (3), one can get, energies  $E_1$  and  $E_2$  of the daughter masses as

$$E_1 = \frac{(M^2 + m_1^2 - m_2^2) c^2}{2M}, \quad E_2 = \frac{(M^2 + m_2^2 - m_1^2) c^2}{2M}$$

The relativistic kinetic energies  $T_1$  and  $T_2$  of the daughter masses are

$$T_1 = E_1 - m_1 c^2, \quad T_2 = E_2 - m_2 c^2$$

which yields

$$T_1 + T_2 = E_1 + E_2 - m_1 c^2 - m_2 c^2 = M c^2 - m_1 c^2 - m_2 c^2 = \Delta M c^2$$

Now,

$$T_1 = \frac{(M^2 + m_1^2 - m_2^2)c^2}{2M} - m_1c^2 = \frac{(M - m_1 - m_2)(M - m_1 + m_2)c^2}{2M}$$

$$= \Delta M \left( 1 - \frac{m_1}{M} - \frac{\Delta M}{2M} \right) c^2, \text{ etc}$$

*Example 10.12* Calculate the velocity of an electron accelerated by a potential of  $a$  MV (rest mass of the electron is  $m_0 = 9 \times 10^{-31}$  kg).

**Hint:** Here,

$$T = (m - m_0)c^2 = \frac{m_0c^2}{\sqrt{1 - \frac{v^2}{c^2}}} - m_0c^2 = a \times 10^6 \text{ eV} = a \times 10^6 \times 1.6 \times 10^{-19} \text{ J} \implies v = \text{etc}$$

*Example 10.13* By what fraction does the mass of water increase due to increase in its thermal energy, when it is heated from  $a^0$  C to  $b^0$  C?

**Hint:** Let mass of the water to be heated be  $M$  kg. Energy needed in heating the water from  $a^0$  C to  $b^0$  C is

$$\Delta E = M \times 1 \times (b - a) \text{ kilo calories} = M \times 1 \times (b - a) \times 1000 \times 4.2 \text{ J}$$

Mass-energy equivalent relation  $\Delta M = \frac{\Delta E}{c^2}$  gives the required  $\Delta M$ .

The fractional increase in mass of water due to heating is  $\frac{\Delta M}{M}$ , etc.

*Example 10.14* A particle of rest mass  $m_0$  and kinetic energy  $2m_0c^2$  strikes and sticks to a stationary particle of rest mass  $2m_0$ . Find the rest mass  $M_0$  of the composite particle.

**Hint:**

$$\begin{array}{ccc} \odot \longrightarrow v & & \otimes \\ m_0, T = 2m_0c^2 & & T = 0, 2m_0 \end{array}$$

Total energy of the particle of rest mass  $m_0$  moving with velocity  $v$  is

$$E = mc^2 = \frac{m_0c^2}{\sqrt{1 - \frac{v^2}{c^2}}} = 2m_0c^2 + m_0c^2 \tag{1}$$

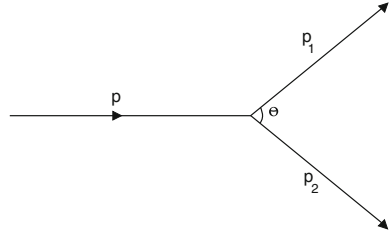
From energy conservation, we have

$$(2m_0c^2 + m_0c^2) + 2m_0c^2 = \frac{M_0c^2}{\sqrt{1 - \frac{V^2}{c^2}}} \tag{2}$$

(Velocity of the composite particle is  $V$ )

Conservation of momentum

**Fig. 10.2** The angle between two decay particles is  $\theta$



$$\frac{m_0 v}{\sqrt{1 - \frac{v^2}{c^2}}} + 0 = \frac{M_0 V}{\sqrt{1 - \frac{V^2}{c^2}}} \quad (3)$$

Solving from (1)–(3), one can get the solutions for  $V$  and  $M_0$  as  $V = \frac{2\sqrt{2}c}{5}$  and  $M_0 = \sqrt{17}m_0$ .

*Example 10.15* An unstable particle of rest mass  $m_0$  and momentum  $p$  decays into two particles of rest masses  $m_1$  and  $m_2$ , momentum  $p_1$  and  $p_2$  and total energies  $E_1$  and  $E_2$ , respectively. Show that

$$m_0 c^4 = (m_1 + m_2) c^4 + 2E_1 E_2 - 2m_1 m_2 c^4 - 2p_1 p_2 c^2 \cos \theta$$

where  $\theta$  is the angle between the two decay particles (Fig. 10.2).

**Hint:** From principle of conservation of energy  $E_0 = E_1 + E_2$ , we have

$$\sqrt{m_0 c^4 + p_0^2 c^2} = \sqrt{m_1 c^4 + p_1^2 c^2} + \sqrt{m_2 c^4 + p_2^2 c^2} \quad (1)$$

From principle of conservation of momentum, we have

$$p^2 = p_1^2 + p_2^2 + 2p_1 p_2 \cos \theta \quad (2)$$

Squaring (1) and using (2), one can obtain the desired result.

*Example 10.16* A meson of rest mass  $\pi$  comes to rest and disintegrates into a muon of rest mass  $\mu$  and a neutrino of zero rest mass. Show that the kinetic energy of motion the muon is  $T = \frac{(\pi - \mu)^2 c^2}{2\pi}$ .

**Hint:**

$$\pi \longrightarrow \mu + \nu$$

According to principle of conservation of momentum, if muon moves with momentum  $p$ , then neutrino moves with momentum  $-p$ . Therefore, according to the momentum–energy relation, we have

$$E_\mu^2 = p^2 c^2 + \mu^2 c^4, \quad E_\nu^2 = p^2 c^2 + \nu^2 c^4 \quad (1)$$



(here, energies of the muon and neutrino are  $E_\mu$  and  $E_\nu$ , respectively and rest mass of the neutrino  $\nu = 0$ )

From these, we get,

$$E_\mu^2 - E_\nu^2 = \mu^2 c^4 \tag{2}$$

Total energy

$$E_\mu + E_\nu = E = \pi c^2 \tag{3}$$

Solving (2) and (3), one can get, energies of muon and neutrino as

$$E_\mu = \frac{(\mu^2 + \pi^2)c^2}{2\pi}, \quad E_\nu = \frac{(\mu^2 - \pi^2)c^2}{2\pi} \tag{4}$$

The kinetic energy of the meson is

$$T = E_\mu - \mu c^2 = \frac{(\mu^2 + \pi^2)c^2}{2\pi} - \mu c^2 = \text{etc}$$

*Example 10.17* A rocket propels itself rectilinearly through empty space by emitting pure radiation in the direction opposite to its motion. If  $v$  is the final velocity relative to its initial frame, prove that the ratio of the initial to the final rest mass of the rocket is given by  $\frac{M_i}{M_f} = \sqrt{\frac{c+v}{c-v}}$ .

**Hint:** Principle of conservation of energy

$$M_i c^2 = \frac{M_f c^2}{\sqrt{1 - \frac{v^2}{c^2}}} + E_\gamma \tag{1}$$

$E_\gamma$  is the energy of radiation. Since it is massless, therefore, its momentum  $p = \frac{E_\gamma}{c}$ .  
Principle of conservation of momentum

$$p = \frac{E_\gamma}{c} = \frac{M_f v}{\sqrt{1 - \frac{v^2}{c^2}}} \tag{2}$$

Assuming rocket started from rest.

Eliminating  $E_\gamma$  from (1) and (2), one can get the required result.

*Example 10.18* A particle of rest mass  $M$  moving at a velocity  $u$  collides with a stationary particle of rest mass  $m$ . If the particles stick together, show that the speed of the composite ball is equal to  $\frac{u\gamma M}{\gamma M + m}$  where  $\gamma = \frac{1}{\sqrt{1 - \frac{u^2}{c^2}}}$ .

**Hint:** Principle of conservation of energy

$$mc^2 + M\gamma c^2 = \frac{(M + m)c^2}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (1)$$

Principle of conservation of momentum

$$M\gamma u = \frac{(M + m)v}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (2)$$

[ $v$  is the velocity of the combined ball]

Using (1) and (2), one can solve for  $v$ .

*Example 10.19* Prove the following relations:

$$(i) T = c\sqrt{m_0^2 c^2 + p^2} - m_0 c^2 \quad (ii) p = \frac{\sqrt{T^2 + 2m_0 c^2 T}}{c} \quad (iii) m_0 = \frac{\sqrt{E^2 - p^2 c^2}}{c^2}$$

**Hint:** We know,

$$E^2 = p^2 c^2 + m_0^2 c^4, \quad E = T + m_0 c^2$$

Now,

$$(i) T = E - m_0 c^2 = \sqrt{p^2 c^2 + m_0^2 c^4} - m_0 c^2 = c\sqrt{p^2 + m_0^2 c^2} - m_0 c^2$$

$$(ii) E^2 = (T + m_0 c^2)^2 = p^2 c^2 + m_0^2 c^4 \implies p = \frac{\sqrt{T^2 + 2m_0 c^2 T}}{c}$$

$$(iii) E^2 = p^2 c^2 + m_0^2 c^4 \implies m_0 = \frac{\sqrt{E^2 - p^2 c^2}}{c^2}$$

*Example 10.20* What is the speed of an electron whose kinetic energy equals  $a$  times its rest energy?

**Hint:**

$$T = (m - m_0)c^2 = \left( \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}} - m_0 \right) c^2 = am_0 c^2 \text{ etc}$$

*Example 10.21* Two particles of rest masses  $m_1$  and  $m_2$  move with velocities  $u_1$  and  $u_2$  collide with another and sticks (i.e. inelastic collision). Show that the rest mass and the velocity of the combined particles are

$$m_3^2 = m_1^2 + m_2^2 + 2m_1 m_2 \gamma_1 \gamma_2 \left( 1 - \frac{u_1 u_2}{c^2} \right), \quad u_3 = \frac{m_1 \gamma_1 u_1 + m_2 \gamma_2 u_2}{m_1 \gamma_1 + m_2 \gamma_2}$$

$$\gamma_1 = \frac{1}{\sqrt{1 - \frac{u_1^2}{c^2}}}, \quad \gamma_2 = \frac{1}{\sqrt{1 - \frac{u_2^2}{c^2}}}$$

**Hint:** In an inelastic collision, let  $m_3$  be the rest mass and  $u_3$  be the velocity of the resulting particle.

$$\text{Conservation of momentum : } m_1\gamma_1u_1 + m_2\gamma_2u_2 = m_3\gamma_3u_3$$

$$\text{Conservation of mass : } m_1\gamma_1 + m_2\gamma_2 = m_3\gamma_3$$

$$\gamma_3 = \frac{1}{\sqrt{1 - \frac{u_3^2}{c^2}}}$$

Solving these equations, one can get  $m_3$  and  $u_3$ .

*Example 10.22* A particle with kinetic energy  $T_0$  and rest energy  $E_0$  strikes an identical particle at rest and gets scattered at an angle  $\theta$ . Show that its kinetic energy  $T$  after scattering is given by  $T = \frac{T_0 \cos^2 \theta}{1 + \frac{T_0 \sin^2 \theta}{2E_0}}$ .

**Hint:** Consider a particle with mass  $m_1$  and momentum  $p_1$  strikes a particle of identical mass at rest. After collision, the particle and rest particle get scattered at angles  $\theta, \phi$  respectively. We have,

$$E^2 = (T + E_0)^2 = c^2p^2 + m_0^2c^4 = c^2p^2 + E_0^2 \implies p^2c^2 = 2TE_0 + T^2 \quad (1)$$

Principle of conservation of momentum

$$p_1 = p_3 \cos \theta + p_4 \cos \phi \quad (2)$$

$$0 = p_3 \sin \theta - p_4 \sin \phi \quad (3)$$

Principle of conservation of energy

$$E_1 + E_2 = E_3 + E_4 \implies E_0 + T_0 + E_0 = E_0 + T + E_0 + T_4 \implies T_4 = T_0 - T \quad (4)$$

From (2) and (3), we get,

$$p_4c^2 = (p_1^2 + p_3^2 - 2p_1p_3 \cos \theta)c^2 \quad (5)$$

Using the relation (1) in (5), we get,

$$2T_4E_0 + T_4^2 = 2T_0E_0 + T_0^2 + 2TE_0 + T^2 - 2\sqrt{2T_0E_0 + T_0^2}\sqrt{2TE_0 + T^2} \cos \theta$$

Replacing  $T_4$  by  $T_0 - T$  and after some simplification, one will get

$$T = \frac{T_0 \cos^2 \theta}{1 + \frac{T_0 \sin^2 \theta}{2E_0}}$$

*Example 10.23* For Minkowski force  $K_\mu$ , show that  $K_\mu K_\mu = \left[ \frac{1 - \frac{v^2}{c^2} \cos^2 \theta}{1 - \frac{v^2}{c^2}} \right] F^2$ ,

where velocity  $\vec{v}$  makes  $\theta$  angle with relativistic force  $\vec{F}$ .

**Hint:** The Minkowski force is given by

$$K_\mu = \left[ aF_1, aF_2, aF_3, \left( \frac{ia}{c} \right) \vec{F} \cdot \vec{v} \right]; \quad a = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$$

Now,

$$\begin{aligned} K_\mu K_\mu &= K_1 K_1 + K_2 K_2 + K_3 K_3 + K_4 K_4 \\ &= a^2 F_1^2 + a^2 F_2^2 + a^2 F_3^2 + \frac{i^2 a^2}{c^2} (\vec{F} \cdot \vec{v})^2 \\ &= a^2 F^2 - \frac{a^2}{c^2} (Fv \cos \theta)^2 = \left[ \frac{1 - \frac{v^2}{c^2} \cos^2 \theta}{1 - \frac{v^2}{c^2}} \right] F^2 \end{aligned}$$

[here,  $\vec{F} \cdot \vec{v} = vF \cos \theta$ ]

*Example 10.24* Show that if a particle decays spontaneously from rest into two or more components, then the rest mass of the particle must be greater than the sum of the rest masses of the resulting components.

**Hint:** Let a particle of rest mass  $m$  decays spontaneously into a number of components of rest masses  $m_1, m_2, m_3, \dots$  and velocities  $v_1, v_2, v_3, \dots$

Conservation of mass

$$E = mc^2 = E_1 + E_2 + E_3 + \dots = \frac{m_1 c^2}{\sqrt{1 - \frac{v_1^2}{c^2}}} + \frac{m_2 c^2}{\sqrt{1 - \frac{v_2^2}{c^2}}} + \frac{m_3 c^2}{\sqrt{1 - \frac{v_3^2}{c^2}}} = \dots$$

Note that  $E_i > m_i c^2$  for  $i = 1, 2, 3, \dots$ . Hence,

$$m > m_1 + m_2 + m_3 + \dots$$

*Example 10.25* A particle of momentum  $p_1$ , rest mass  $m_1$  is incident upon a stationary particle of rest mass  $m_2$ . Show that the velocity of the centre of the mass system is equal to

$$v = \frac{p_1 c^2}{E_1 + E_2} = \frac{p_1 c^2}{\sqrt{p_1^2 c^2 + m_1^2 c^4} + m_2 c^2}$$

**Hint:** In rest system S, total momentum is  $p = p_1 + 0 = p_1$ . Let the centre of the mass system is  $S^1$  which is moving with velocity  $v$  along x-axis. In  $S^1$  system, the total momentum is zero. The total energy in S system before collision is

$$E = E_1 + E_2 = p_1 c^2 + m_1^2 c^4 + m_2 c^2 \quad (\text{here, } p_2 = 0)$$

In S system,

$$p_{1x} = p_1, \quad p_{1y} = p_{1z} = 0, \quad p_{2x} = p_{2y} = p_{2z} = 0$$

In  $S^1$  system,

$$p_{1x}^1 = \frac{p_1 - \frac{vE_1}{c^2}}{\sqrt{1 - \frac{v^2}{c^2}}}, \quad p_{1y}^1 = p_{1y} = 0, \quad p_{1z}^1 = p_{1z} = 0$$

$$p_{2x}^1 = \frac{p_{2x} - \frac{vE_2}{c^2}}{\sqrt{1 - \frac{v^2}{c^2}}} = -\frac{\frac{vE_2}{c^2}}{\sqrt{1 - \frac{v^2}{c^2}}}$$

$$p_{2y}^1 = p_{2z}^1 = 0$$

Now, in  $S^1$  system,

$$p_{1x}^1 + p_{2x}^1 = 0 \implies \frac{p_1 - \frac{v(E_1 + E_2)}{c^2}}{\sqrt{1 - \frac{v^2}{c^2}}} = 0$$

This will give

$$v = \frac{p_1 c^2}{E_1 + E_2} = \frac{p_1 c^2}{\sqrt{p_1^2 c^2 + m_1^2 c^4} + m_2 c^2}$$

*Example 10.26* A particle of rest mass  $m_0$  describes the trajectory  $x = f(t), y = g(t), z = 0$  in an inertial frame S. Find the Minkowski force and the corresponding Newtonian force on the particle.

**Hint:** Minkowski four force is given by

$$K_\mu = \frac{dp_\mu}{d\tau} = \frac{d(m_0 u_\mu)}{d\tau}$$

For spatial component

$$K_i = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \frac{d}{dt} \left( \frac{m_0 v_i}{\sqrt{1 - \frac{v^2}{c^2}}} \right)$$

where,

$$v_1 = \dot{x} = \dot{f}(t), \quad v_2 = \dot{y} = \dot{g}(t), \quad v_3 = \dot{z} = 0, \quad v^2 = v_1^2 + v_2^2 + v_3^2 = \dot{f}^2 + \dot{g}^2$$

Let

$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}, \quad \text{therefore, } \dot{\gamma} = \frac{\gamma^3}{c^2} (\dot{f}\ddot{f} + \dot{g}\ddot{g})$$

Now,

$$\begin{aligned} K_1 &= \gamma \frac{d}{dt} (m_0 v_1 \gamma) = m_0 \gamma [\gamma \dot{v}_1 + v_1 \dot{\gamma}] \\ &= m_0 \gamma^2 \left[ \ddot{f} + \frac{\gamma^2}{c^2} (\dot{f}\ddot{f} + \dot{g}\ddot{g}) \dot{f} \right] \end{aligned}$$

$$\begin{aligned} K_2 &= \gamma \frac{d}{dt} (m_0 v_2 \gamma) = m_0 \gamma [\gamma \dot{v}_2 + v_2 \dot{\gamma}] \\ &= m_0 \gamma^2 \left[ \ddot{g} + \frac{\gamma^2}{c^2} (\dot{f}\ddot{f} + \dot{g}\ddot{g}) \dot{g} \right] \end{aligned}$$

$$K_3 = 0$$

$$K_4 = \gamma \frac{d}{dt} (i m_0 c \gamma) = \frac{i m_0 \gamma^4}{c} [\dot{f}\ddot{f} + \dot{g}\ddot{g}]$$

Newtonian force

$$F_1 = m_0 \frac{dv_1}{dt} = m_0 \ddot{f}, \quad F_2 = m_0 \frac{dv_2}{dt} = m_0 \ddot{g}, \quad F_3 = 0$$

*Example 10.27* A particle of rest mass  $m_0$  describes the parabolic trajectory  $x = at$ ,  $y = bt^2$ ,  $z = 0$  in an inertial frame S. Find the Minkowski force and the corresponding Newtonian force on the particle.

**Hint:** Use  $f(t) = at$  and  $g(t) = bt^2$  in example (10.26).

*Example 10.28* A particle of rest mass  $m_0$  describes the circular path  $x = a \cos t$ ,  $y = a \sin t$ ,  $z = 0$  in an inertial frame S. Find the Minkowski force and the corresponding Newtonian force on the particle.

**Hint:** Use  $f(t) = a \cos t$  and  $g(t) = a \sin t$  in example (10.26).

*Example 10.29* A particle of rest mass  $m_0$  moves on the x-axis along a world line described by the parametric equations  $t(a) = \frac{1}{A} \sinh a$  and  $x(a) = \frac{1}{A} \cosh a$ ,  $a$  is the parameter and  $A$  is a constant. Find four velocity, momentum and acceleration and their norms. Also find the Minkowski force.

**Hint:** The proper time is defined by

$$d\tau = dt\sqrt{1 - v^2}$$

where  $v = \frac{dx}{dt}$  and we have assumed  $c = 1$ . Here,  $y$  and  $z$  coordinates are unimportant and we will be suppressed. We find that  $v = \frac{dx}{dt} = \tanh a$ . Therefore,

$$\tau = \int dt\sqrt{1 - v^2} = \int \frac{dt}{\cosh a} = \int \frac{da}{A} \implies \tau = \frac{a}{A}$$

Hence,

$$t(\tau) = \frac{\sinh(A\tau)}{A}, \quad x(\tau) = \frac{\cosh(A\tau)}{A}$$

Now, we calculate the four velocity vector  $u_\mu = \frac{dx_\mu}{d\tau}$  as

$$u_1 = \frac{dx}{d\tau} = \sinh(A\tau), \quad u_4 = \frac{d(it)}{d\tau} = i \cosh(A\tau)$$

Here, norm of four velocity is given by

$$u_\mu u_\mu = u_1 u_1 + u_4 u_4 = \sinh^2(A\tau) - \cosh^2(A\tau) = -1$$

Note that particle three velocity is  $v = \frac{dx}{dt} = \tanh(A\tau)$ . This velocity never exceeds the speed of light, however, when  $\tau = \pm\infty$ , then  $v \rightarrow 1$ , i.e. approaches to the velocity of light.

The four momentum is given by

$$p_\mu = m_0 u_\mu$$

where  $m_0$  is the rest mass of the particle. Here, norm of four momentum is given by

$$p_\mu p_\mu = -m_0^2$$

The four acceleration  $f_\mu = \frac{du_\mu}{d\tau}$  is given by

$$f_1 = \frac{du_1}{d\tau} = A \cosh(A\tau), \quad f_4 = \frac{du_4}{d\tau} = iA \sinh(A\tau)$$

Here, norm of four acceleration is given by

$$f_\mu f_\mu = f_1 f_1 + f_4 f_4 = A^2 \cosh^2(A\tau) - A^2 \sinh^2(A\tau) = A^2$$

The Minkowski force is given by

$$K_\mu = m_0 f_\mu$$

*Example 10.30* A particle of rest mass  $m_0$  moves on the x-axis along a world line described by the parametric equations  $t = f(\sigma)$  and  $x = g(\sigma)$ ,  $\sigma$  is the parameter. Find four velocity, momentum and acceleration and their norms. Also find the Minkowski force.

**Hint:** The proper time is defined by

$$d\tau = dt\sqrt{1 - v^2}$$

where  $v = \frac{dx}{dt}$  and we have assumed  $c = 1$ . Here,  $y$  and  $z$  coordinates are unimportant and we will be suppressed. We find that  $v = \frac{dx}{dt} = \frac{\dot{g}}{\dot{f}}$  (here,  $\dot{g} = \frac{dg}{d\sigma}$ ). Therefore,

$$\tau = \int dt\sqrt{1 - v^2} = \int \sqrt{\dot{f}^2 - \dot{g}^2} d\sigma$$

The four velocity vector  $u_\mu = \frac{dx_\mu}{d\tau}$  can be obtained as

$$u_1 = \frac{dx}{d\tau} = \frac{\dot{g}}{\sqrt{\dot{f}^2 - \dot{g}^2}}, \quad u_4 = \frac{d(it)}{d\tau} = \frac{i\dot{f}}{\sqrt{\dot{f}^2 - \dot{g}^2}} \text{ etc}$$

*Example 10.31* A particle is moving along x-axis. It is uniformly accelerated in the sense that the acceleration measured in its rest frame is always  $a$ , a constant. Find  $x$  and  $t$  in terms of proper time assuming the particle passes through  $x_0$  at time  $t = 0$  with zero velocity.

**Hint:** Here,  $y$  and  $z$  coordinates are unimportant and we will be suppressed and we have assumed  $c = 1$ . We know four velocity  $u_\mu$  satisfies

$$u_\mu u_\mu = u_1^2 + u_4^2 = -1 \tag{1}$$

Differentiating (1), we get

$$u_1 f_1 + u_4 f_4 = 0 \tag{2}$$

where  $f_\mu = \frac{du_\mu}{d\tau}$  is the four acceleration.

Given:

$$f_\mu f_\mu = f_1^2 + f_4^2 = a^2 \tag{3}$$



Solving these equations, we get

$$f_4 = \frac{du_4}{d\tau} = iau_1 \quad (4)$$

$$f_1 = \frac{du_1}{d\tau} = -iau_4 \quad (5)$$

From these equations, we obtain,

$$\frac{d^2u_1}{d\tau^2} = a^2u_1$$

The solution is obtained as

$$u_1 = A \sinh(a\tau) \quad (6)$$

$A$  is one of the integration constants. The other integration constant is zero as for  $\tau = 0, u_1 = 0$ . Also we obtain the solution for  $u_4$

$$u_4 = Ai \cosh(a\tau) \quad (7)$$

Equation (1) implies  $A = 1$ .

Again,

$$u_1 = \frac{dx_1}{d\tau} = \sinh(a\tau), \quad u_4 = \frac{dx_4}{d\tau} = \frac{d(it)}{d\tau} = i \cosh(a\tau)$$

Using the boundary condition, the solutions of these equations are obtained as

$$x(\tau) = x_0 + \frac{1}{a} [\cosh(a\tau) - 1], \quad t(\tau) = \frac{1}{a} \sinh(a\tau)$$

The world line of the particle is the hyperbola given by

$$t^2 = (x - x_0)^2 + \frac{2}{a} [x - x_0]$$

*Example 10.32* A particle of rest mass  $M$  decays spontaneously into two particles with mass deficit  $\Delta M$ . Show that the kinetic energy of the particle of rest mass  $m_i$  ( $i = 1, 2$ ) is given by  $T_i = \Delta M \left(1 - \frac{m_i}{M} - \frac{\Delta M}{2M}\right) c^2$ .

**Hint:** Let the decayed particles have rest masses  $m_1$  and  $m_2$ . According to principle of conservation of momentum, if the  $m_1$  mass moves with momentum  $p$ , then other mass  $m_2$  moves with momentum  $-p$ . Follow Example 10.11.

# Chapter 11

## Photon in Relativity

### 11.1 Photon

In the beginning of twentieth century, Max Planck argued that light and other electromagnetic radiation consisted of individual packets of energy known as quanta. He proposed that the energy of each quanta is proportional to its frequency  $\nu$ . This hypothesis is known as Planck's hypothesis and is given by

$$E = h\nu$$

[ $h$  is known as Planck's constant]

In fact, after further development of Planck's hypothesis by Einstein, it is verified that light has a dual wave-particle nature. Diffraction, interference, etc. are of wave nature and photoelectric effect, interaction of light with atoms, etc. are of particle nature of light and are called as photons.

The relativistic mass of the photon is given by

$$m = \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}}$$

Here  $m_0$  is the rest mass of the photon. From above equation, we can get

$$m_0 = m\sqrt{1 - \frac{v^2}{c^2}}$$

Using the velocity of light as  $v = c$ , we get,

$$m_0 = 0$$

i.e. the rest mass of the photon is zero. Actually, in all inertial frame, light would never rest. Its velocity always is  $c$ .

Let the direction cosines of the direction of travel of the photon are  $(l, m, n)$ , then

$$\vec{p} = (lp_1 + mp_2 + np_3) = p\hat{n}, \quad p = \sqrt{p_1^2 + p_2^2 + p_3^2}$$

The momentum energy relation becomes

$$E^2 - p^2c^2 = 0 \implies E = pc$$

Using mass energy relation  $E = mc^2$ , we get

$$m = \frac{p}{c}$$

Hence, we have the following important relations of the photon as

$$E = hv, \quad m = \frac{hv}{c^2}, \quad \vec{p} = \frac{hv}{c}\hat{n}$$

## 11.2 Compton Effect

A photon may be described as a particle of zero rest mass with momentum  $\frac{h}{\lambda} = \frac{hv}{c}$  and energy  $hv$ . In 1921, A H Compton proposed the experiment of the scattering of a photon by an electron that if the photon collides with an electron of rest mass  $m$ , it will be scattered at some angle  $\theta$  with a new energy  $hv^1$ . He has shown that the change in energy is related to the scattering angle by the formula  $\lambda^1 - \lambda = 2\lambda_c \sin^2 \frac{\theta}{2}$  where  $\lambda_c = \frac{h}{mc}$  is the *Compton wave length*.

This observed change in frequency of wave length of scattered radiation is known as *Compton effect*.

Let a photon of energy  $hv$  is incident on an electron of rest mass  $m$  and be scattered at an angle  $\theta$  with energy  $hv^1$ . Here the electron which scatters the photon suffers a momentum recoil (see Fig. 11.1).

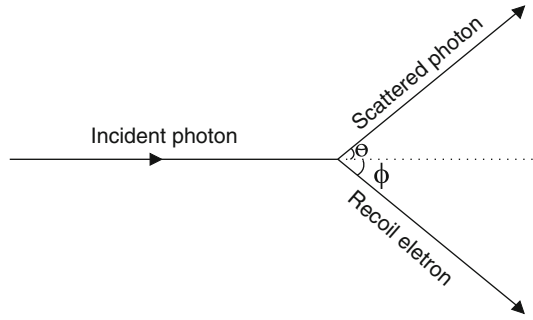
Principle of conservation of energy

$$hv + mc^2 = hv^1 + \frac{mc^2}{\sqrt{1 - \frac{v^2}{c^2}}} \implies \frac{m^2c^4}{1 - \frac{v^2}{c^2}} = [mc^2 + h(v - v^1)]^2 \quad (11.1)$$

[ $v$  is the recoil velocity of the electron]

Principle of conservation of momentum (along and perpendicular to the direction of motion of  $\pi$  meson)

**Fig. 11.1** Scattering of a photon by an electron



$$-\frac{h\nu}{c} + 0 = \frac{h\nu^1}{c} \cos \theta + \frac{mv}{\sqrt{1 - \frac{v^2}{c^2}}} \cos \phi \tag{11.2}$$

$$0 = \frac{h\nu^1}{c} \cos \theta - \frac{mv}{\sqrt{1 - \frac{v^2}{c^2}}} \sin \phi \tag{11.3}$$

Eliminating  $\phi$  from (11.1) and (11.2), we get,

$$\frac{m^2 v^2 c^2}{1 - \frac{v^2}{c^2}} = h^2 (\nu^2 - 2\nu^1 \nu \cos \theta + \nu^{1^2}) \tag{11.4}$$

Subtracting (11.4) from (11.1), we get,

$$\frac{\nu - \nu^1}{\nu \nu^1} = \frac{h}{mc^2} (1 - \cos \theta)$$

Using the result  $\nu \lambda = c$ , we get

$$\lambda^1 - \lambda = \frac{2h}{mc} \sin^2 \frac{\theta}{2} = 2\lambda_c \sin^2 \frac{\theta}{2}$$

### 11.3 The Lorentz Transformation of Momentum of Photon

The four momentum of photon  $p_\mu = [p_1, p_2, p_3, p_4]$  can be written as

$$p_1 = \frac{h\nu l}{c}, \quad p_2 = \frac{h\nu m}{c}, \quad p_3 = \frac{h\nu n}{c}, \quad p_4 = \left( \frac{iE}{c} \right) \tag{11.5}$$

where  $(l, m, n)$  are the direction cosines of the direction of propagation of photon and energy of the photon,  $E = h\nu$ . Here, the direction of the propagation of photon makes  $\theta$  angle with x-axis (see Fig. 11.2).

Using the above transformation formula, we get

$$\begin{bmatrix} \frac{h\nu^1 l^1}{c} \\ \frac{h\nu^1 m^1}{c} \\ \frac{h\nu^1 n^1}{c} \\ \frac{i h\nu^1}{c} \end{bmatrix} = \begin{bmatrix} a & 0 & 0 & i a \beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -i a \beta & 0 & 0 & a \end{bmatrix} \begin{bmatrix} \frac{h\nu l}{c} \\ \frac{h\nu m}{c} \\ \frac{h\nu n}{c} \\ \frac{i h\nu}{c} \end{bmatrix} \quad (11.6)$$

This implies

$$\frac{h\nu^1 l^1}{c} = a \left( \frac{h\nu l}{c} + i \frac{v}{c} \frac{i h\nu}{c} \right)$$

or

$$\nu^1 l^1 = \frac{\nu(l - \frac{v}{c})}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (11.7)$$

$$\frac{h\nu^1 m^1}{c} = \frac{h\nu m}{c} \implies \nu^1 m^1 = \nu m \quad (11.8)$$

$$\frac{h\nu^1 n^1}{c} = \frac{h\nu n}{c} \implies \nu^1 n^1 = \nu n \quad (11.9)$$

$$\frac{i h\nu^1}{c} = a \left( -i \frac{v}{c} \frac{h\nu}{c} + \frac{i h\nu}{c} \right)$$

or

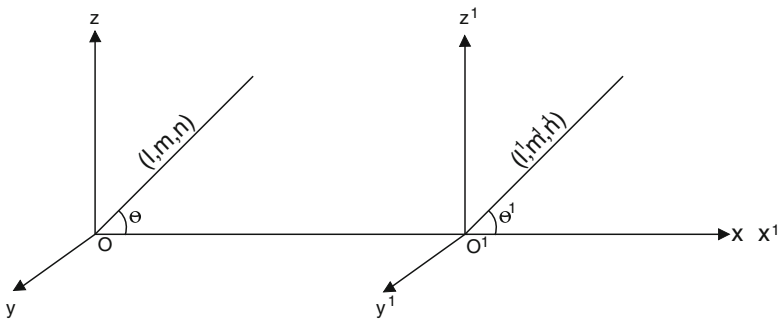


Fig. 11.2 Direction of the propagation of photon makes  $\theta$  angle with x-axis

$$v^1 = \frac{v(l - \frac{lv}{c})}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (11.10)$$

This change of the frequency is due to Doppler effects.

Substituting  $v^1$  in Eq. (11.7), we get,

$$l^1 = \frac{l - \frac{v}{c}}{1 - \frac{lv}{c}} \quad (11.11)$$

Similarly Eqs. (11.8) and (11.9) give,

$$m^1 = \frac{m\sqrt{1 - \frac{v^2}{c^2}}}{1 - \frac{lv}{c}} \quad (11.12)$$

$$n^1 = \frac{n\sqrt{1 - \frac{v^2}{c^2}}}{1 - \frac{lv}{c}} \quad (11.13)$$

These changes in the direction are due to aberration.

If the light propagates in the  $xy$ -plane, then,  $n = n^1 = 0$  and  $l = \cos \theta$ ,  $m = \sin \theta$  where light ray makes  $\theta$  angle with  $x$ -axis. Then Eqs. (11.11) and (11.12) yield

$$\cos \theta^1 = \frac{\cos \theta - \frac{v}{c}}{1 - \frac{v}{c} \cos \theta} \quad (11.14)$$

$$\sin \theta^1 = \frac{\sin \theta \sqrt{1 - \frac{v^2}{c^2}}}{1 - \frac{v}{c} \cos \theta} \quad (11.15)$$

The last two equations give the usual formula of aberration as

$$\cot \theta^1 = \frac{\cot \theta - \frac{v}{c} \csc \theta}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (11.16)$$

## 11.4 Minkowski Force for Photon

Minkowski force for photon can be written in terms of four momentum of photon as

$$K_\mu = \frac{1}{h} p_\mu = \frac{1}{h} \left( p_1 = \frac{h\nu l}{c}, p_2 = \frac{h\nu m}{c}, p_3 = \frac{h\nu n}{c}, \frac{iE}{c} = \frac{i h\nu}{c} \right)$$

Therefore, the Minkowski force for photon can be written in terms of wave length  $\lambda$  of the photon

$$K_\mu = \left( \frac{l}{\lambda}, \frac{m}{\lambda}, \frac{n}{\lambda}, \frac{i}{\lambda} \right) \quad (11.17)$$

[using  $\lambda\nu = c$ ]

If the photon, i.e. electromagnetic wave travels in a direction making the coordinates axes constant, angle of whose direction cosines  $(l, m, n)$  is characterized by

$$\psi = a \exp \left( \frac{mx + my + nz}{\lambda} - \nu t \right)$$

Then, one can write

$$\psi = a \exp(K_\mu x_\mu) \text{ as } x_\mu = (x, y, z, ict)$$

We notice that  $K_\mu x_\mu$  is a scalar quantity. So the phase of the electromagnetic wave  $\psi$  is also a scalar quantity and hence it is invariant under Lorentz Transformation.

*Example 11.1* Show that Minkowski force for photon is a null vector.

**Hint:** We know

$$K_\mu = \left( \frac{l}{\lambda}, \frac{m}{\lambda}, \frac{n}{\lambda}, \frac{i}{\lambda} \right),$$

therefore,

$$\begin{aligned} K_\mu K_\mu &= K_1 K_1 + K_2 K_2 + K_3 K_3 + K_4 K_4 \\ &= \frac{l^2}{\lambda^2} + \frac{m^2}{\lambda^2} + \frac{n^2}{\lambda^2} + \frac{i^2}{\lambda^2} \\ &= \frac{l^2 + m^2 + n^2}{\lambda^2} - \frac{1}{\lambda^2} = 0 \end{aligned}$$

*Example 11.2* Show that four momentum of photon is a null vector.

**Hint:** We know

$$p_\mu = \left[ p_1 = \frac{h\nu l}{c}, p_2 = \frac{h\nu m}{c}, p_3 = \frac{h\nu n}{c}, p_4 = \left( \frac{iE}{c} \right) = \left( \frac{i h\nu}{c} \right) \right],$$

i.e.

$$p_\mu = \left( \frac{lh}{\lambda}, \frac{mh}{\lambda}, \frac{nh}{\lambda}, \frac{ih}{\lambda} \right), \text{ using } \lambda\nu = c$$

therefore,

$$\begin{aligned}
 p_\mu p_\mu &= p_1 p_1 + p_2 p_2 + p_3 p_3 + p_4 p_4 \\
 &= \frac{h^2 l^2}{\lambda^2} + \frac{h^2 m^2}{\lambda^2} + \frac{h^2 n^2}{\lambda^2} + \frac{i^2}{\lambda^2} \\
 &= \frac{h^2 (l^2 + m^2 + n^2)}{\lambda^2} - \frac{h^2}{\lambda^2} = 0
 \end{aligned}$$

*Example 11.3* A meson of rest mass  $\pi$  decays in flight into two photons. If one of the photons is emitted at an angle  $\theta$  to the direction of motion of the meson. Show that the energy of motion of the photon is  $h\nu = \frac{\pi c^2}{2\gamma(1 - \frac{u \cos \theta}{c})}$  (where,  $\gamma = \frac{1}{\sqrt{1 - \frac{u^2}{c^2}}}$ ).

**Hint:** Let  $\nu$  is the frequency of the decay photon. Therefore, both photons have the energy  $E = h\nu$  and momentum  $p = \frac{E}{c} = \frac{h\nu}{c}$ . The direction of the resultant is the direction of motion of the meson. Here, two photons make equal angle  $\theta$  with the direction of motion of the meson, therefore, angle between two photons is  $2\theta$  (see Fig. 11.3).

Principle of conservation of energy

$$\frac{\pi c^2}{\sqrt{1 - \frac{u^2}{c^2}}} = \gamma \pi c^2 = h\nu + h\nu = 2h\nu \tag{1}$$

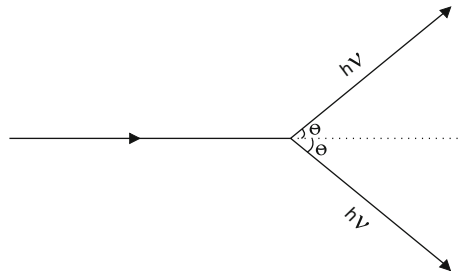
Principle of conservation of momentum

$$\gamma \pi u = \sqrt{p^2 + p^2 + 2pp \cos 2\theta}$$

or

$$\gamma^2 \pi^2 u^2 = 2p^2 + 2p^2 \cos 2\theta = \frac{4h^2 \nu^2}{c^2} \cos^2 \theta \tag{2}$$

**Fig. 11.3** Two photons make equal angle  $\theta$  with the direction of meson





Putting the value of  $h\nu$  from (i) in (ii), we get

$$\cos \theta = \frac{u}{c} \implies \gamma = \frac{1}{\sqrt{1 - \frac{u^2}{c^2}}} = \csc \theta$$

Now,

$$h\nu = \frac{\gamma \pi c^2}{2} = \frac{\csc \theta \pi c^2}{2} = \frac{\pi c^2}{2 \sin \theta} = \frac{\pi c^2}{2 \csc \theta (1 - \cos^2 \theta)} = \frac{\pi c^2}{2\gamma (1 - \frac{u \cos \theta}{c})}$$

*Example 11.4* Show that it is impossible for a photon to transfer all its energy to a free electron.

or show that an electron of finite mass cannot disintegrate into a single photon.

**Hint:** Suppose if possible, a photon containing energy  $E$  and momentum  $p = \frac{E}{c}$  transfers all its energy to an electron of rest mass  $m_0$  and velocity  $v = fc$  where

$$0 < f < 1 \quad (1)$$

Principle of conservation of energy

$$E = \frac{m_0 c^2}{\sqrt{1 - \frac{(fc)^2}{c^2}}} - m_0 c^2 \quad (2)$$

Principle of conservation of momentum

$$p = \frac{E}{c} = \frac{m_0 f c}{\sqrt{1 - \frac{(fc)^2}{c^2}}} \quad (3)$$

From (ii) and (iii), we get,

$$\frac{m_0 f c^2}{\sqrt{1 - f^2}} = \frac{m_0 c^2}{\sqrt{1 - f^2}} - m_0 c^2 \implies f(1 - f) = 0$$

This gives either  $f = 0$  or  $f = 1$  which are contradiction to condition (i). Hence the assumption was wrong, i.e. it is impossible for a photon to transfer all its energy to a free electron.

*Example 11.5* An excited atom of mass  $m_0$ , initially at rest in frame  $S$ , emits a photon and recoils. The internal energy of the atom decreases by  $\Delta E$  and the energy of the photon is  $h\nu$ . Show that  $h\nu = \Delta E (1 - \frac{\Delta E}{2m_0 c^2})$

**Hint:** Let  $M_0$  be the rest mass of the atom after emission and  $v$  be the recoil velocity. Principle of conservation of energy

$$m_0c^2 = \frac{M_0c^2}{\sqrt{1 - \frac{v^2}{c^2}}} + hv = \frac{M_0c^2}{a} + hv \tag{1}$$

Principle of conservation of momentum

$$p = \frac{hv}{c} = \frac{M_0v}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{M_0v}{a} \tag{2}$$

Assuming the atom is excited from rest and  $a = \sqrt{1 - \frac{v^2}{c^2}}$ .

Above two equations can be written as

$$\left(m_0c - \frac{hv}{c}\right)^2 = \left(\frac{M_0^2v^2}{a^2}\right) \tag{3}$$

$$\left(\frac{hv}{c}\right)^2 = \left(\frac{M_0^2c^2}{a^2}\right) \tag{4}$$

Subtracting (4) from (3), we get,

$$hv = \frac{(m_0 + M_0)(m_0 - M_0)c^2}{2m_0}$$

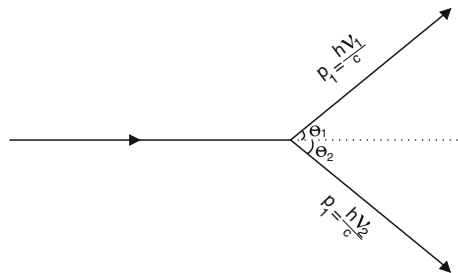
Now,  $\Delta E = (m_0 - M_0)c^2$ , therefore we get,

$$hv = \Delta E \left(1 - \frac{m_0 - M_0}{2m_0}\right) = \Delta E \left(1 - \frac{\Delta E}{2m_0c^2}\right)$$

*Example 11.6* A  $\pi$  meson of rest mass  $m_0$  moving with velocity  $v$  disintegrates into two  $\gamma$  rays. Calculate the energy distribution of  $\gamma$  rays from  $\pi$  meson.

**Hint:** Let  $\nu_1$  and  $\nu_2$  be the frequencies of the two decay  $\gamma$  rays, which make angles  $\theta_1$  and  $\theta_2$  with initial direction of motion of  $\pi$  meson (see Fig. 11.4). Principle of conservation of energy

**Fig. 11.4** Two decay  $\gamma$  rays make  $\theta_1$  and  $\theta_2$  angle with the initial direction of motion of  $\pi$  meson



$$\frac{M_0 c^2}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{M_0 c^2}{a} = h(v_1 + v_2) \quad (1)$$

Principle of conservation of momentum (along and perpendicular to the direction of motion of  $\pi$  meson)

$$\frac{M_0 v}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{M_0 v}{a} = \frac{h}{c}(v_1 \cos \theta_1 + v_2 \cos \theta_2) \quad (2)$$

$$0 = \frac{h}{c}(v_1 \sin \theta_1 - v_2 \sin \theta_2) \quad (3)$$

where  $a = \sqrt{1 - \frac{v^2}{c^2}}$ .

Squaring both sides of (1) and (2) and subtracting with other yields

$$\begin{aligned} M_0 c^4 &= h^2(v_1 + v_2)^2 - h^2(v_1 \cos \theta_1 + v_2 \cos \theta_2)^2 \\ &= h^2[v_1^2 \sin^2 \theta_1 + v_2^2 \sin^2 \theta_2 + 2v_1 v_2(1 - \cos \theta_1 \cos \theta_2)] \end{aligned}$$

Equation (3) implies

$$v_1 \sin \theta_1 = v_2 \sin \theta_2$$

From these two equations, one can solve for  $v_1$  and  $v_2$  as

$$v_1^2 = \frac{M_0^2 c^4 \sin \theta_2}{2h^2 \sin \theta_1 [1 - \cos(\theta_1 + \theta_2)]}, \quad v_2^2 = \frac{M_0^2 c^4 \sin \theta_1}{2h^2 \sin \theta_2 [1 - \cos(\theta_1 + \theta_2)]}$$

here,  $E_1 = hv_1$  and  $E_2 = hv_2$ . Multiplying above two values  $v_1$  and  $v_2$ , we get,

$$\sin\left(\frac{\theta_1 + \theta_2}{2}\right) = \frac{M_0 c^2}{2\sqrt{E_1 E_2}}$$

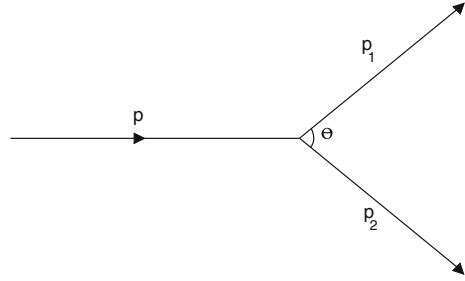
*Example 11.7* A  $\pi$  meson of rest mass  $m_0$  moving with velocity  $v$  disintegrates into two equal  $\gamma$  rays. Show that the angle between two decayed  $\gamma$  rays is  $2\theta$  which is given by  $\sin \theta = \frac{m_0 c^2}{2h\nu}$  where  $\nu$  is the frequency of the decayed  $\gamma$  ray.

**Hint:** In the example 25,  $\theta_1 = \theta_2 = \theta$  and  $v_1 = v_2 = \nu$ .

*Example 11.8* Show that a photon cannot give rise to an electro positron pair in free space.

**Hint:** Suppose, if possible a photon of momentum  $p = \frac{h\nu}{c}$  produces an electron of momentum  $p_1$  and a positron of momentum  $p_2$ . Both have same rest mass  $m_0$ . Let the produced electron and positron makes  $\theta$  angle with each other (see Fig. 11.5).

**Fig. 11.5** Angle between electron and positron is  $\theta$



Principle of conservation of momentum gives

$$p^2 = \frac{h^2 v^2}{c^2} = p_1^2 + p_2^2 + 2p_1 p_2 \cos \theta \quad (1)$$

Principle of conservation of energy implies  $h\nu = E_1 + E_2$ . Now using relation between momentum and energy  $E^2 = p^2 c^2 + m_0^2 c^4$ , we obtain

$$h\nu = \sqrt{p_1^2 c^2 + m_0^2 c^4} + \sqrt{p_2^2 c^2 + m_0^2 c^4} \quad (2)$$

Eliminating  $h\nu$  from these equations, we obtain after some simplification

$$-p_1^2 p_2^2 \sin^2 \theta = m_0^2 c^2 (p_1^2 + p_2^2 + 2p_1 p_2 \cos \theta) \quad (3)$$

Since minimum value of  $\cos \theta$  is  $-1$ . Therefore, right-hand side is always positive. Since left-hand side is either zero or negative, therefore, we conclude that both sides must be zero. Left-hand side will be zero when  $\theta = 0$  or  $\pi$ . For,  $\theta = 0$ , right-hand side is never zero. For,  $\theta = \pi$ , right-hand side gives  $p_1 = p_2$ . But, for  $p_1 = p_2$ , Eq. (11.1) gives  $\nu = 0$ , i.e. the photon is not existed. Therefore, the above process is not possible.

*Example 11.9* Show that the outcome of the collision between two particles cannot be a single photon.

**Hint:** Let a particle of rest mass  $m_2$  moving with velocity  $v_2$  strikes a particle of rest mass  $m_1$  at rest. If possible let a photon of energy  $h\nu$  is produced as an outcome of this collision.

Principle of conservation of momentum:

$$\frac{m_2 v_2}{\sqrt{1 - \frac{v_2^2}{c^2}}} = \frac{h\nu}{c}$$

Principle of conservation of energy:

$$m_1c^2 + \frac{m_2c^2}{\sqrt{1 - \frac{v_2^2}{c^2}}} = hv$$

Eliminating  $m_2$  from these equations, we get

$$\frac{v_2}{c} = \frac{hv}{hv - m_1c^2} > 1$$

Definitely this is impossible. Hence the process cannot be possible.

*Example 11.10* Show that the de Broglie wave length for a material particle rest mass  $m_0$  and a charge  $q$  accelerated from rest through a potential difference  $V$  volt relativistically is given by  $\lambda = \frac{h}{\sqrt{2m_0qV(1 + \frac{qV}{2m_0c^2})}}$ . Calculate the wave length of an

electron having a kinetic energy of 1 MeV.

**Hint:** Here, the kinetic energy of the electron is  $T = qV$ . Therefore, the total energy is  $E = T + m_0c^2 = qV + m_0c^2$ . Now, using relation between momentum and energy  $E^2 = p^2c^2 + m_0^2c^4$ , we obtain,

$$(qV + m_0c^2)^2 = p^2c^2 + m_0^2c^4 \implies p = \sqrt{2m_0qV \left(1 + \frac{qV}{2m_0c^2}\right)}$$

Therefore, the de Broglie wave length is given by

$$\lambda = \frac{h}{p} = \frac{h}{\sqrt{2m_0qV \left(1 + \frac{qV}{2m_0c^2}\right)}}$$

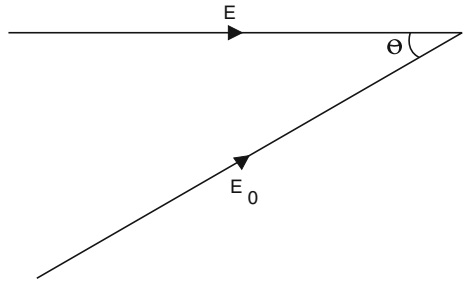
For an electron,  $q = e = 1.6 \times 10^{-19}$  coulomb,  $m_0 = 9.1 \times 10^{-31}$  kg,  $c = 3 \times 10^8$  km and  $h = 6.62 \times 10^{-34}$ . Here,  $eV = 1$  MeV.  $V = 10^6$  volt,  $m_0c^2 = .51$  MeV. Putting all these, one can get  $\lambda = 8.6 \times 10^{-3}$  Angstrom.

*Example 11.11* A photon of energy  $E_0$  collides at an angle  $\theta$  with another photon of energy  $E$ . Prove that the minimum value of  $E_0$  permitting formation of a pair of particles of mass  $m$  is  $E_{th} = \frac{2m^2c^4}{E(1 - \cos\theta)}$ .

**Hint:** Here  $\frac{E}{c}$  and  $\frac{E_0}{c}$  are momenta of the photons with energies  $E$  and  $E_0$  respectively. Before collision takes place, the total momenta was

$$\sqrt{\frac{E^2}{c^2} + \frac{E_0^2}{c^2} + \frac{EE_0 \cos \theta}{c^2}}$$

**Fig. 11.6**  $\theta$  be the angle between two photons



[Here,  $\theta$  be the angle between two photons (see Fig. 11.6)]

Since  $E_{th}$  is the threshold energy, therefore, the momentum of the produced particles will be zero, i.e.

$$\sqrt{\frac{E^2}{c^2} + \frac{E_{th}^2}{c^2} + \frac{E E_{th} \cos \theta}{c^2}} = 0 \tag{1}$$

Principle of conservation of energy

$$E + E_{th} = 2mc^2 \tag{2}$$

These two equations yield

$$E_{th} = \frac{2m^2c^4}{E(1 - \cos \theta)}$$

# Chapter 12

## Relativistic Lagrangian and Hamiltonian

### 12.1 Relativistic Lagrangian

According to classical mechanics, the Lagrangian is given by

$$L = T - V \tag{12.1}$$

where  $T$  is the kinetic energy of the system and is a function of generalized momenta  $p_r$  ( or of the generalized velocity  $\dot{q}_r$ ); while the potential energy  $V$  is the function of generalized coordinates  $q_r$  only.

Again from classical mechanics, we get the canonical momenta as

$$p_r = \frac{\partial L}{\partial \dot{q}_r} = \frac{\partial}{\partial \dot{q}_r}(T - V) = \frac{\partial T}{\partial \dot{q}_r} \tag{12.2}$$

Now we are to find relativistic kinetic energy  $T^*$  such that the relativistic Lagrangian function be  $L^* = T^* - V$ .

We know the relativistic momentum

$$\begin{aligned} p_x &= m_0 \dot{x} [1 - \beta^2]^{-\frac{1}{2}} = \frac{\partial T^*}{\partial \dot{x}} \\ p_y &= m_0 \dot{y} [1 - \beta^2]^{-\frac{1}{2}} = \frac{\partial T^*}{\partial \dot{y}} \\ p_z &= m_0 \dot{z} [1 - \beta^2]^{-\frac{1}{2}} = \frac{\partial T^*}{\partial \dot{z}} \end{aligned} \tag{12.3}$$

where  $\beta = \frac{v}{c}$  and  $v^2 = \dot{x}^2 + \dot{y}^2 + \dot{z}^2$ .

Since  $T^*$  is a function of  $\dot{x}$ ,  $\dot{y}$ ,  $\dot{z}$ , so

$$dT^* = \frac{\partial T^*}{\partial \dot{x}} d\dot{x} + \frac{\partial T^*}{\partial \dot{y}} d\dot{y} + \frac{\partial T^*}{\partial \dot{z}} d\dot{z}$$

Using the relations given in (12.3) we obtain

$$dT^* = m_0[1 - \beta^2]^{-\frac{1}{2}}[\dot{x}d\dot{x} + \dot{y}d\dot{y} + \dot{z}d\dot{z}] \quad (12.4)$$

Also  $v^2 = \dot{x}^2 + \dot{y}^2 + \dot{z}^2$ , so

$$v dv = \dot{x}d\dot{x} + \dot{y}d\dot{y} + \dot{z}d\dot{z}$$

Therefore from (12.4), we get,

$$dT^* = m_0 \left[ 1 - \left( \frac{v}{c} \right)^2 \right]^{-\frac{1}{2}} v dv$$

After integrating, this yields

$$T^* = -m_0 c^2 \left[ 1 - \left( \frac{v}{c} \right)^2 \right]^{\frac{1}{2}} + A \quad (12.5)$$

(A = integration constant)

We know that when  $v \ll c$ ,  $T^* = \frac{1}{2}m_0v^2$  (classical kinetic energy)

Hence, from (12.5), one gets

$$\frac{1}{2}m_0v^2 = -m_0c^2 + A$$

$$\Rightarrow A = \frac{1}{2}m_0v^2 + m_0c^2 = m_0c^2 \quad (\text{since } v \ll c)$$

Therefore

$$\begin{aligned} T^* &= -m_0c^2 \left[ 1 - \left( \frac{v}{c} \right)^2 \right]^{\frac{1}{2}} + m_0c^2 \\ &= m_0c^2 \left[ 1 - \left( 1 - \beta^2 \right)^{\frac{1}{2}} \right] \end{aligned} \quad (12.6)$$

Here V purely depends upon position, i.e. it is independent of velocity. Therefore, it remains unchanged relativistically. So the relativistic Lagrangian function is given by

$$L^* = T^* - V = m_0c^2 \left[ 1 - \left( 1 - \beta^2 \right)^{\frac{1}{2}} \right] - V \quad (12.7)$$

To check whether relativistic Lagrangian function gives relativistic force through Lagrangian's equation analogue to classical mechanics.



Assume the Lagrangian's equation holds similar to classical mechanics in the following form:

$$\frac{d}{dt} \left[ \frac{\partial L^*}{\partial v_i} \right] - \frac{\partial L^*}{\partial x_i} = 0 \quad (12.8)$$

Here,

$$\frac{\partial L^*}{\partial v_i} = m_0 c^2 \left[ 1 - \left( \frac{v}{c} \right)^2 \right]^{-\frac{1}{2}} \left( \frac{v_i}{c^2} \right) = m_0 v_i (1 - \beta^2)^{-\frac{1}{2}}$$

Substituting this in Eq. (12.8), gives,

$$\begin{aligned} \frac{d}{dt} \left[ m_0 v_i (1 - \beta^2)^{-\frac{1}{2}} \right] &= \frac{\partial V}{\partial x_i} = F_i \\ \Rightarrow (1 - \beta^2)^{-\frac{1}{2}} \frac{d}{dt} \left[ m_0 v_i (1 - \beta^2)^{-\frac{1}{2}} \right] &= (1 - \beta^2)^{-\frac{1}{2}} F_i \\ \Rightarrow \frac{d}{d\tau} \left[ m_0 v_i (1 - \beta^2)^{-\frac{1}{2}} \right] &= K_i \end{aligned}$$

[ $\tau$  = proper time and  $K_i$  = space components of Minkowski force]

This is the relativistic equation of motion.

Thus we have shown that the Lagrangian (12.7) is correct.

Now,

$$\frac{\partial L^*}{\partial v_i} = m_0 v_i (1 - \beta^2)^{-\frac{1}{2}} = p_i,$$

i.e. partial derivative of  $L^*$  with respect to the velocity is still the momentum (relativistic).

## 12.2 Relativistic Hamiltonian Function

Analogue to the classical mechanics, the relativistic Hamiltonian function is defined as

$$H^* = \sum p_x \dot{x} - L^*$$

where the canonical momenta are given by

$$p_r = \frac{\partial L^*}{\partial \dot{q}_r} = \frac{\partial T^*}{\partial \dot{q}_r}$$

Now,

$$\begin{aligned}
 H^* &= \sum \frac{\partial T^*}{\partial \dot{x}} \dot{x} - m_0 c^2 \left[ 1 - (1 - \beta^2)^{\frac{1}{2}} \right] + V \\
 &= \sum \frac{\partial}{\partial \dot{x}^*} \left[ m_0 c^2 \left[ 1 - (1 - \beta^2)^{\frac{1}{2}} \right] \right] \dot{x} - m_0 c^2 \left[ 1 - (1 - \beta^2)^{\frac{1}{2}} \right] + V \\
 &= \sum \left[ m_0 c^2 \left[ 1 - \left( \frac{v}{c} \right)^2 \right]^{-\frac{1}{2}} \frac{v}{c^2} \frac{\partial v}{\partial \dot{x}} \right] \dot{x} - m_0 c^2 \left[ 1 - (1 - \beta^2)^{\frac{1}{2}} \right] + V
 \end{aligned}$$

Since  $v^2 = \dot{x}^2 + \dot{y}^2 + \dot{z}^2$ , therefore,  $v \frac{\partial v}{\partial \dot{x}} = \dot{x}$ . It follows that,

$$\begin{aligned}
 H^* &= \sum \left[ m_0 \dot{x}^2 \left[ 1 - \left( \frac{v}{c} \right)^2 \right]^{-\frac{1}{2}} \right] - m_0 c^2 \left[ 1 - (1 - \beta^2)^{\frac{1}{2}} \right] + V \\
 &= m_0 \left[ \dot{x}^2 + \dot{y}^2 + \dot{z}^2 \right] \left[ 1 - \left( \frac{v}{c} \right)^2 \right]^{-\frac{1}{2}} - m_0 c^2 \left[ 1 - (1 - \beta^2)^{\frac{1}{2}} \right] + V \\
 &= m_0 v^2 \left[ 1 - \left( \frac{v}{c} \right)^2 \right]^{-\frac{1}{2}} - m_0 c^2 \left[ 1 - (1 - \beta^2)^{\frac{1}{2}} \right] + V \\
 &= m_0 c^2 \left[ \left( \frac{v}{c} \right)^2 + 1 - \left( \frac{v}{c} \right)^2 \right] \left[ 1 - \left( \frac{v}{c} \right)^2 \right]^{-\frac{1}{2}} - m_0 c^2 + V
 \end{aligned}$$

Thus we obtain the relativistic Hamiltonian as

$$H^* = m_0 c^2 \left[ 1 - \left( \frac{v}{c} \right)^2 \right]^{-\frac{1}{2}} - m_0 c^2 + V \quad (12.9)$$

Using the definition of relativistic mass,  $m = m_0 \left[ 1 - \left( \frac{v}{c} \right)^2 \right]^{-\frac{1}{2}}$ , one can write

$$\begin{aligned}
 H^* &= m c^2 - m_0 c^2 + V \\
 &= T^* + V
 \end{aligned}$$

where  $T^*$  is the relativistic kinetic energy.

Interestingly, we note that analogous to classical Hamiltonian, the relativistic Hamiltonian is the total relativistic energy of the system.

Also,

$$p_x^2 + p_y^2 + p_z^2 = \frac{m_0^2}{1 - \frac{v^2}{c^2}} (\dot{x}^2 + \dot{y}^2 + \dot{z}^2)$$

or,

$$p_x^2 + p_y^2 + p_z^2 = \frac{m_0^2 v^2}{1 - \frac{v^2}{c^2}} = -m_0^2 c^2 + \frac{m_0^2 c^2}{1 - \frac{v^2}{c^2}}$$

This implies

$$\frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} = \sqrt{\frac{p_x^2 + p_y^2 + p_z^2 + m_0^2 c^2}{m_0^2 c^2}} \quad (12.10)$$

Substituting this result into Eq.(12.9), we get another expression for relativistic Hamiltonian as

$$H^* = c\sqrt{p_x^2 + p_y^2 + p_z^2 + m_0^2 c^2} - m_0^2 c^2 + V \quad (12.11)$$

Hamiltonian canonical equations are

$$\dot{x} = \frac{\partial H^*}{\partial p_x} \quad \text{and} \quad \dot{p}_x = -\frac{\partial H^*}{\partial x}$$

From (12.11), we obtain,

$$\dot{x} = \frac{cp_x}{\sqrt{p_x^2 + p_y^2 + p_z^2 + m_0^2 c^2}} = \frac{cp_x \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}}{m_0 c}$$

This implies

$$p_x = \frac{m_0 \dot{x}}{\sqrt{1 - \frac{v^2}{c^2}}}$$

It is relativistic momentum.

Again,

$$\dot{p}_x = \frac{d}{dt} \left[ \frac{m_0 \dot{x}}{\sqrt{1 - \frac{v^2}{c^2}}} \right] = -\frac{\partial H^*}{\partial x} = -\frac{\partial V}{\partial x} = F_x$$

This is the relativistic force.

## 12.3 Covariant Lagrangian and Hamiltonian Formulation

The above expressions for Lagrangian and Hamiltonian predict the correct relativistic equations of motion in a particular Lorentz frame. Here, time coordinate  $t$  has been used as a parameter completely distinct from the spatial coordinates. For a covariant

formulation, all the four coordinates have the similar meaning in world space. Here one should replace  $t$  by the invariant parameter, the proper time  $\tau$ . The covariant form of the Lagrangian should be a function of the coordinates  $x_\mu$  in Minkowski space and their derivatives with respect to  $\tau$ . Here,

$$dt = \frac{dt}{d\tau} d\tau = t^1 d\tau \quad (12.12)$$

[ $^1$  means derivative with respect to  $\tau$ ]

To find the connection between old Lagrangian  $L$  and new Lagrangian  $L^*$ , we use Hamilton's variational principle as

$$0 = \delta I = \delta \int_{\tau_1}^{\tau_2} L^*(x_j, x_j^1, x_4, x_4^1) d\tau = \delta \int_{t_1}^{t_2} L dt = \delta \int_{\tau_1}^{\tau_2} L t^1 d\tau, \quad (12.13)$$

which yields

$$L^*(x_\mu, x_\mu^1) = t^1 L(x_j, \dot{x}_j, t) = t^1 L\left(x_j, \frac{x_4}{ic}, \frac{icx_j^1}{x_4^1}\right) = \frac{x_4^1}{ic} L\left(x_j, \frac{x_4}{ic}, \frac{icx_j^1}{x_4^1}\right) \quad (12.14)$$

[over dot means derivative with respect to  $t$ ,  $x_4 = ict$ ,  $\dot{x}_j = \frac{dx_j}{dt} = \frac{icx_j^1}{x_4^1}$  and  $\mu = 1, 2, 3, 4$ ;  $j = 1, 2, 3$ .]

Here, one can note that the new Lagrangian  $L^*(x_\mu, x_\mu^1)$  is a homogenous function of generalized velocities in the first degree whatever the functional form of old Lagrangian  $L(x_j, \dot{x}_j, t)$ . Then by Euler's theorem

$$L^* = x_\mu^1 \frac{\partial L^*}{\partial x_\mu^1} \quad (12.15)$$

Therefore,

$$\frac{dL^*}{d\tau} = x_\mu^{11} \frac{\partial L^*}{\partial x_\mu^1} + x_\mu^1 \frac{d}{d\tau} \left( \frac{\partial L^*}{\partial x_\mu^1} \right) \quad (12.16)$$

Again since  $L^* = L^*(x_\mu, x_\mu^1)$ , we get

$$\frac{dL^*}{d\tau} = x_\mu^{11} \frac{\partial L^*}{\partial x_\mu^1} + x_\mu^1 \frac{\partial L^*}{\partial x_\mu} \quad (12.17)$$

Subtracting (12.17) from (12.16), we get

$$\left[ \frac{d}{d\tau} \left( \frac{\partial L^*}{\partial x_\mu^1} \right) - \frac{\partial L^*}{\partial x_\mu} \right] x_\mu^1 = 0 \quad (12.18)$$

Thus the homogenous property of new Lagrangian  $L^*(x_\mu, x_\mu^1)$  in generalized velocities gives covariant Lagrange equations.

For spatial part, we have

$$\frac{\partial L^*}{\partial x_j} = \frac{\partial}{\partial x_j} \left[ t^1 L \left( x_j, \frac{x_4}{ic}, \frac{icx_j^1}{x_4^1} \right) \right] = \frac{x_4^1}{ic} \frac{\partial}{\partial x_j} \left[ L \left( x_j, \frac{x_4}{ic}, \frac{icx_j^1}{x_4^1} \right) \right] = \frac{x_4^1}{ic} \frac{\partial L}{\partial x_j} \quad (12.19)$$

$$\begin{aligned} \frac{\partial L^*}{\partial x_j^1} &= \frac{\partial}{\partial x_j^1} \left[ t^1 L \left( x_j, \frac{x_4}{ic}, \frac{icx_j^1}{x_4^1} \right) \right] = \frac{x_4^1}{ic} \frac{\partial}{\partial x_j^1} \left[ L \left( x_j, \frac{x_4}{ic}, \frac{icx_j^1}{x_4^1} \right) \right] \\ &= \frac{x_4^1}{ic} \frac{\partial}{\partial \left( \frac{icx_j^1}{x_4^1} \right)} \left[ L \left( x_j, \frac{x_4}{ic}, \frac{icx_j^1}{x_4^1} \right) \right] \frac{\partial \left( \frac{icx_j^1}{x_4^1} \right)}{\partial x_j^1} = \frac{x_4^1}{ic} \frac{\partial L}{\partial \dot{x}_j} \frac{ic}{x_4^1} = \frac{\partial L}{\partial \dot{x}_j} = p_j \end{aligned} \quad (12.20)$$

Thus the definition of momentum remains the same.

For the fourth component, we have

$$\begin{aligned} \frac{\partial L^*}{\partial x_4} &= \frac{\partial}{\partial x_4} \left[ t^1 L \left( x_j, \frac{x_4}{ic}, \frac{icx_j^1}{x_4^1} \right) \right] = \frac{x_4^1}{ic} \frac{\partial}{\partial x_4} \left[ L \left( x_j, \frac{x_4}{ic}, \frac{icx_j^1}{x_4^1} \right) \right] \\ &= \frac{x_4^1}{ic} \frac{\partial}{\partial \left( \frac{x_4}{ic} \right)} \left[ L \left( x_j, \frac{x_4}{ic}, \frac{icx_j^1}{x_4^1} \right) \right] \frac{\partial \left( \frac{x_4}{ic} \right)}{\partial x_4} = \frac{x_4^1}{ic} \frac{\partial L}{\partial t} \frac{1}{ic} = -\frac{x_4^1}{c^2} \frac{\partial L}{\partial t} \end{aligned} \quad (12.21)$$

$$\begin{aligned} \frac{\partial L^*}{\partial x_4^1} &= \frac{\partial}{\partial x_4^1} \left[ t^1 L \left( x_j, \frac{x_4}{ic}, \frac{icx_j^1}{x_4^1} \right) \right] = \frac{\partial}{\partial x_4^1} \left[ \frac{x_4^1}{ic} L \left( x_j, \frac{x_4}{ic}, \frac{icx_j^1}{x_4^1} \right) \right] \\ &= \frac{L}{ic} + \frac{x_4^1}{ic} \frac{\partial}{\partial \left( \frac{icx_j^1}{x_4^1} \right)} \left[ L \left( x_j, \frac{x_4}{ic}, \frac{icx_j^1}{x_4^1} \right) \right] \frac{\partial \left( \frac{icx_j^1}{x_4^1} \right)}{\partial x_4^1} \\ &= \frac{L}{ic} + \frac{x_4^1}{ic} \left( -\frac{icx_j^1}{x_4^1{}^2} \right) \frac{\partial L}{\partial \dot{x}_j} = \frac{L}{ic} - \frac{\dot{x}_j}{ic} \frac{\partial L}{\partial \dot{x}_j} = \frac{i}{c} \left( \dot{x}_j \frac{\partial L}{\partial \dot{x}_j} - L \right) = \frac{iH}{c} = p_4 \end{aligned} \quad (12.22)$$

Thus we have obtained the fourth component of momentum  $p_4$ . But we know  $p_4 = \frac{iE}{c}$ , therefore, relativistic Hamiltonian is the total relativistic energy of the system.

Using the results given in (12.19) and (12.20), we obtain from (12.18)

$$\frac{x_4^1}{ic} \frac{\partial L}{\partial \dot{x}_j} - \frac{x_4^1}{ic} \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{x}_j} \right) = 0$$

[here,  $\frac{d}{d\tau} = \frac{dt}{d\tau} \frac{d}{dt} = \gamma^1 \frac{d}{dt} = \frac{x_4^1}{ic} \frac{d}{dt}$ ]

This yields the usual Lagrange's equation

$$\frac{\partial L}{\partial \dot{x}_j} - \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{x}_j} \right) = 0$$

Using the result given in (12.21) and (12.22) the fourth component, we get as above

$$-\frac{x_4^1}{c^2} \frac{\partial L}{\partial t} - \frac{x_4^1}{ic} \frac{d}{dt} \left( \frac{iH}{c} \right) = 0 \implies \frac{\partial L}{\partial t} + \frac{dH}{dt} = 0$$

This is a very known result in classical mechanics.

The covariant Hamiltonian function is given by

$$H^* = p_j x_j^1 + p_4 x_4^1 - L^*$$

The result

$$\dot{x}_j = \frac{dx_j}{dt} = \frac{icx_j^1}{x_4^1},$$

implies

$$x_j^1 = \frac{\dot{x}_j x_4^1}{ic}$$

Thus we get

$$H^* = \frac{x_4^1}{ic} (p_j \dot{x}_j + p_4 ic - L) = \frac{x_4^1}{ic} (H - H) = 0$$

Hence covariant Hamiltonian function identically vanishes. However, its derivatives are not, in general, zero as  $H^* = p_\mu x_\mu^1 - L^*$  gives

$$x_\mu^1 = \frac{\partial H^*}{\partial p_\mu} \quad \text{and} \quad p_\mu^1 = \frac{\partial H^*}{\partial x_\mu}.$$

## 12.4 Lorentz Transformation of Force

Let a frame  $S^1$  is moving with velocity  $v$  along  $x$  axis relative to a rest frame  $S$ . Let a particle of mass  $m$  is moving with velocity  $\vec{u}$  with respect to  $S$  frame. In  $S^1$  frame let its mass and velocity be  $m^1$  and  $\vec{u}^1$ , respectively. Let us also suppose that particle is moving under the action of the force  $\vec{F}$  and  $\vec{F}^1$  with respect to  $S$  and  $S^1$  frames, respectively. Then

$$\vec{F} = \frac{d}{dt}(m\vec{u}) \quad ; \quad \vec{F}^1 = \frac{d}{dt^1}(m^1\vec{u}^1)$$

Here,

$$\vec{F} = (F_x, F_y, F_z) \quad , \quad \vec{F}^1 = (F_x^1, F_y^1, F_z^1) \quad \text{and} \quad \vec{u} = (u_x, u_y, u_z) \quad \vec{u}^1 = (u_x^1, u_y^1, u_z^1)$$

We know the following results:

$$\frac{dt^1}{dt} = \gamma \left(1 - \frac{v}{c^2}u_x\right) \quad , \quad m = \frac{m_0}{\sqrt{1 - \frac{u^2}{c^2}}} \quad , \quad u_x^1 = \frac{u_x - v}{1 - \frac{v}{c^2}u_x} \quad , \quad m^1 = \gamma m \left(1 - \frac{v}{c^2}u_x\right)$$

where,  $\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$ ,  $u^2 = u_x^2 + u_y^2 + u_z^2$  and  $m_0$  is the rest mass.

Now,

$$\begin{aligned} F_x^1 &= \frac{d}{dt^1}(m^1 u_x^1) = \frac{d}{dt}(m^1 u_x^1) \frac{dt}{dt^1} = \frac{d}{dt} \left[ \left( \frac{u_x - v}{1 - \frac{v}{c^2}u_x} \right) \gamma m \left(1 - \frac{v}{c^2}u_x\right) \right] \frac{1}{\gamma \left(1 - \frac{v}{c^2}u_x\right)} \\ &= \frac{1}{\left(1 - \frac{v}{c^2}u_x\right)} \frac{d}{dt} [m(u_x - v)] = \frac{1}{\left(1 - \frac{v}{c^2}u_x\right)} \left[ m \frac{du_x}{dt} + u_x \frac{dm}{dt} - vm_0 \frac{d}{dt} \left\{ \left(1 - \frac{u^2}{c^2}\right)^{-\frac{1}{2}} \right\} \right] \\ &= \frac{1}{\left(1 - \frac{v}{c^2}u_x\right)} \left[ m \frac{du_x}{dt} + u_x \frac{dm}{dt} - \frac{vm_0}{c^2} \left(1 - \frac{u^2}{c^2}\right)^{-\frac{3}{2}} u \frac{du}{dt} \right] \\ &= \frac{1}{\left(1 - \frac{v}{c^2}u_x\right)} \left[ m \frac{du_x}{dt} + u_x \frac{dm}{dt} - \frac{vm}{c^2} \left(1 - \frac{u^2}{c^2}\right)^{-1} \left( u_x \frac{du_x}{dt} + u_y \frac{du_y}{dt} + u_z \frac{du_z}{dt} \right) \right] \end{aligned}$$

Taking logarithm of both sides of  $m = \frac{m_0}{\sqrt{1 - \frac{u^2}{c^2}}}$  and differentiating with respect to  $t$ ,

we get,

$$\frac{1}{m} \frac{dm}{dt} = \frac{1}{c^2} \left(1 - \frac{u^2}{c^2}\right)^{-1} u \frac{du}{dt}$$

Using this result, the above expression takes the form

$$\begin{aligned}
 F_x^1 &= \frac{1}{\left(1 - \frac{v}{c^2}u_x\right)} \left[ m \frac{du_x}{dt} + u_x \frac{dm}{dt} - \frac{vm}{c^2} \left( u_x \frac{du_x}{dt} + u_y \frac{du_y}{dt} + u_z \frac{du_z}{dt} \right) \right] \\
 &\quad - \frac{1}{\left(1 - \frac{v}{c^2}u_x\right)} \left[ \frac{v}{c^2} \frac{dm}{dt} (u_x^2 + u_y^2 + u_z^2) \right] \\
 &= \frac{1}{\left(1 - \frac{v}{c^2}u_x\right)} \left[ m \frac{du_x}{dt} \left(1 - \frac{v}{c^2}u_x\right) + u_x \frac{dm}{dt} \left(1 - \frac{v}{c^2}u_x\right) \right] \\
 &\quad - \frac{1}{\left(1 - \frac{v}{c^2}u_x\right)} \left[ mu_y \frac{du_y}{dt} \frac{v}{c^2} + mu_z \frac{du_z}{dt} \frac{v}{c^2} + u_y \frac{dm}{dt} \frac{vu_y}{c^2} + u_z \frac{dm}{dt} \frac{vu_z}{c^2} \right] \\
 &= \frac{d(u_x m)}{dt} - \frac{1}{\left(1 - \frac{v}{c^2}u_x\right)} \left[ \frac{vu_y}{c^2} \frac{d}{dt}(mu_y) + \frac{vu_z}{c^2} \frac{d}{dt}(mu_z) \right]
 \end{aligned}$$

Hence,

$$F_x^1 = F_x - \frac{v}{(c^2 - vu_x)} [u_y F_y + u_z F_z] = \frac{F_x - \frac{v}{c^2} [\vec{u} \cdot \vec{F}]}{\left(1 - \frac{v}{c^2}u_x\right)} \quad (12.23)$$

Now,

$$\begin{aligned}
 F_y^1 &= \frac{d}{dt^1} (m^1 u_y^1) = \frac{d}{dt} (m^1 u_y^1) \frac{dt}{dt^1} \\
 &= \frac{d}{dt} \left[ \left( \frac{u_y}{\gamma \left(1 - \frac{v}{c^2}u_x\right)} \right) \gamma m \left(1 - \frac{v}{c^2}u_x\right) \right] \frac{1}{\gamma \left(1 - \frac{v}{c^2}u_x\right)} \\
 &= \frac{1}{\gamma \left(1 - \frac{v}{c^2}u_x\right)} \frac{d}{dt} (mu_y)
 \end{aligned}$$

Hence,

$$F_y^1 = \frac{\sqrt{1 - \frac{v^2}{c^2}}}{\left(1 - \frac{v}{c^2}u_x\right)} F_y \quad (12.24)$$

Similarly,

$$F_z^1 = \frac{\sqrt{1 - \frac{v^2}{c^2}}}{\left(1 - \frac{v}{c^2}u_x\right)} F_z \quad (12.25)$$



The inverse transformation can be found as

$$F_x = \frac{F_x^1 + \frac{v}{c^2} [\vec{u}^1 \cdot \vec{F}^1]}{\left(1 + \frac{v}{c^2} u_x^1\right)}, \quad F_y = \frac{\sqrt{1 - \frac{v^2}{c^2}}}{\left(1 + \frac{v}{c^2} u_x^1\right)} F_y^1, \quad F_z = \frac{\sqrt{1 - \frac{v^2}{c^2}}}{\left(1 + \frac{v}{c^2} u_x^1\right)} F_z^1 \quad (12.26)$$

Note that in the Newtonian limit,  $v \ll c$ ,  $\vec{F} = \vec{F}^1$ .

Transformation formula of force has an interesting consequence: the force,  $F_x$  in one frame is related to the power developed by the force in another frame,  $\vec{u}^1 \cdot \vec{F}^1$ . Thus in special theory of relativity power and force are related. The transformations (12.26) yield the transformation of power as

$$\vec{u} \cdot \vec{F} = \frac{\vec{u}^1 \cdot \vec{F}^1 + v F_x^1}{\left(1 + \frac{v}{c^2} u_x^1\right)}$$

or inverse relation

$$\vec{u}^1 \cdot \vec{F}^1 = \frac{\vec{u} \cdot \vec{F} - v F_x}{\left(1 - \frac{v}{c^2} u_x\right)}$$

[(12.26) implies

$$\begin{aligned} u_x F_x + u_y F_y + u_z F_z &= u_x F_x^1 + \frac{v}{c^2} u_x (\vec{u}^1 \cdot \vec{F}^1) \\ &+ u_y F_y^1 \sqrt{1 - \frac{v^2}{c^2}} + u_z F_z^1 \sqrt{1 - \frac{v^2}{c^2}} \end{aligned}$$

Using

$$u_x = \frac{u_x^1 + v}{1 + \frac{v}{c^2} u_x^1}, \quad u_y = \frac{u_y^1 \sqrt{1 - \frac{v^2}{c^2}}}{1 + \frac{v}{c^2} u_x^1}, \quad u_z = \frac{u_z^1 \sqrt{1 - \frac{v^2}{c^2}}}{1 + \frac{v}{c^2} u_x^1},$$

and after some simplification, one can get

$$\vec{u} \cdot \vec{F} = \frac{\vec{u}^1 \cdot \vec{F}^1 + v F_x^1}{\left(1 + \frac{v}{c^2} u_x^1\right)}$$

**Particular case:** If the particle is at rest in the  $S^1$  frame, then  $u^1 = 0$  and  $u_x = v$ ,  $u_y = u_z = 0$ . The above transformations assume the following form

$$F_x^1 = F_x, \quad F_y^1 = \frac{F_y}{\sqrt{1 - \frac{v^2}{c^2}}}, \quad F_z^1 = \frac{F_z}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (12.27)$$

### Alternative Proof of the Particular Case

Let a frame  $S^1$  is moving with velocity  $v$  relative to  $S$  along (positive)  $x$  axis and assume a particle of rest mass  $m_0$  is at rest in  $S^1$  frame. This means an observer in  $S$  frame observes the particle is moving with velocity  $v$ .

In  $S$  frame, the force is defined as

$$F = \frac{dp}{dt} = \frac{d}{dt}(mv) = \frac{d}{dt} \left( \frac{m_0 v}{\sqrt{1 - \beta^2}} \right) = m_0 \frac{d}{dt} \left( \frac{v}{\sqrt{1 - \beta^2}} \right) \quad (12.28)$$

where  $\beta = \frac{v}{c}$ ,  $m$  is the relativistic mass of the particle. In  $S^1$  frame, force is defined as

$$F^1 = \frac{dp^1}{dt^1}$$

Since the particle be at rest in  $S^1$  frame, the inverse Lorentz Transformations for momentum and energy give,

$$p_x = \frac{p_x^1 + \frac{v}{c^2} E^1}{\sqrt{1 - \beta^2}} ; \quad p_y = p_y^1 ; \quad p_z = p_z^1 \quad (12.29)$$

From the definition of time dilation, we have

$$\partial t = \frac{\partial t^1}{\sqrt{1 - \beta^2}} \quad \text{or} \quad \partial t^1 = \partial t \sqrt{1 - \beta^2} \quad (12.30)$$

From Eq. (12.29), we have,

$$\partial p_x = \frac{\partial p_x^1 + \frac{v}{c^2} \partial E^1}{\sqrt{1 - \beta^2}} ; \quad \partial p_y = \partial p_y^1 ; \quad \partial p_z = \partial p_z^1 \quad (12.31)$$

From energy and momentum relation  $E^2 = p^2 c^2 + m_0^2 c^4$ , we obtain

$$\partial E = \frac{1}{2} (p^2 c^2 + m_0^2 c^4)^{-\frac{1}{2}} (2pc^2 \partial p) = \frac{pc^2 \partial p}{\sqrt{p^2 c^2 + m_0^2 c^4}}$$

The inverse relation yields

$$\partial E^1 = \frac{p^1 c^2 \partial p^1}{\sqrt{p^{12} c^2 + m_0^2 c^4}} \quad (12.32)$$

We note that the body is at rest in frame  $S^1$ , therefore  $p^1 = 0$  and hence from (12.32)  $\partial E^1 = 0$ .

Consequently, equation (12.31) yields

$$\partial p_x = \frac{\partial p_x^1}{\sqrt{1 - \beta^2}} \quad (12.33)$$

From (12.30) and (12.33), we have

$$\frac{\partial p_x}{\partial t} = \frac{\partial p_x^1}{\partial t^1}$$

Since  $\partial t \rightarrow 0$  implies  $\partial t^1 \rightarrow 0$ , therefore taking these limits, we have

$$\frac{dp_x}{dt} = \frac{dp_x^1}{dt^1} \quad (12.34)$$

Again from (12.30) and (12.31), we have

$$\frac{\partial p_y}{\partial t} = \frac{\partial p_y^1}{\partial t^1} \sqrt{1 - \beta^2}$$

In the limit  $\partial t \rightarrow 0$ , we have

$$\frac{dp_y}{dt} = \sqrt{1 - \beta^2} \frac{dp_y^1}{dt^1} \quad (12.35)$$

Similarly,

$$\frac{dp_z}{dt} = \sqrt{1 - \beta^2} \frac{dp_z^1}{dt^1} \quad (12.36)$$

Hence from (12.34), (12.35), and (12.36), we have

$$\begin{aligned} \frac{dp_x}{dt} &= \frac{dp_x^1}{dt^1} \quad \text{or} \quad F_x = F_x^1 \\ \frac{dp_y}{dt} &= \sqrt{1 - \beta^2} \frac{dp_y^1}{dt^1} \quad \text{or} \quad F_y = F_y^1 \sqrt{1 - \beta^2} \\ \frac{dp_z}{dt} &= \sqrt{1 - \beta^2} \frac{dp_z^1}{dt^1} \quad \text{or} \quad F_z = F_z^1 \sqrt{1 - \beta^2} \end{aligned}$$

## 12.5 Relativistic Transformation Formula for Density

Let us consider two frames S and  $S^1$ ,  $S^1$  is moving with velocity  $v$  relative to S along x direction. Let a parallelepiped of mass  $m$  with volume  $V$  and density  $\rho$  is moving with velocity  $\vec{u}$  with respect to S frame. In  $S^1$  frame its mass, volume, density and velocity be  $m^1$ ,  $V^1$ ,  $\rho^1$  and  $\vec{u}^1$  respectively. Let also  $m_0$  and  $V_0$  be the rest volume and mass respectively and  $l_x, l_y, l_z$  be the lengths of the edge of the body when it is at rest in S frame, then  $V_0 = l_x l_y l_z$ . Since the body is in motion, therefore, lengths of the edges of the body according to length contraction are given by

$$l_x \sqrt{1 - \frac{u_x^2}{c^2}}, l_y \sqrt{1 - \frac{u_y^2}{c^2}}, l_z \sqrt{1 - \frac{u_z^2}{c^2}}$$

Then,

$$V = l_x l_y l_z \sqrt{1 - \frac{u_x^2}{c^2}} \sqrt{1 - \frac{u_y^2}{c^2}} \sqrt{1 - \frac{u_z^2}{c^2}} = AV_0 \quad (12.37)$$

where

$$A = \sqrt{1 - \frac{u_x^2}{c^2}} \sqrt{1 - \frac{u_y^2}{c^2}} \sqrt{1 - \frac{u_z^2}{c^2}} \quad (12.38)$$

Now, the expression for the density in S frame is

$$\rho = \frac{m}{V} = \frac{m_0}{\sqrt{1 - \frac{u^2}{c^2}}} \frac{1}{V_0 A} = \frac{\rho_0}{A \sqrt{1 - \frac{u^2}{c^2}}} \quad (12.39)$$

[here,  $\rho_0$  is the rest density and  $u^2 = u_x^2 + u_y^2 + u_z^2$ ]

The above quantities in  $S^1$  frame

$$u_x^1 = \frac{u_x - v}{1 - \frac{vu_x}{c^2}}, u_y^1 = \frac{u_y \sqrt{1 - \frac{v^2}{c^2}}}{1 - \frac{vu_x}{c^2}}, u_z^1 = \frac{u_z \sqrt{1 - \frac{v^2}{c^2}}}{1 - \frac{vu_x}{c^2}}, m^1 = \frac{m_0}{\sqrt{1 - \frac{u^2}{c^2}}}$$

Lengths of the edges in  $S^1$  frame

$$l_x \sqrt{1 - \frac{u_x^1{}^2}{c^2}}, l_y \sqrt{1 - \frac{u_y^1{}^2}{c^2}}, l_z \sqrt{1 - \frac{u_z^1{}^2}{c^2}}$$

Thus the volume in  $S^1$  frame

$$V^1 = l_x l_y l_z \sqrt{1 - \frac{u_x^2}{c^2}} \sqrt{1 - \frac{u_y^2}{c^2}} \sqrt{1 - \frac{u_z^2}{c^2}} = A^1 V_0 \quad (12.40)$$

where

$$A^1 = \sqrt{1 - \frac{u_x^2}{c^2}} \sqrt{1 - \frac{u_y^2}{c^2}} \sqrt{1 - \frac{u_z^2}{c^2}} \quad (12.41)$$

Therefore, the density of the body as observed from  $S^1$  frame

$$\rho^1 = \frac{m^1}{V^1} = \frac{m_0}{\sqrt{1 - \frac{u^2}{c^2}}} \frac{1}{V_0 A^1} = \frac{\rho_0}{A^1 \sqrt{1 - \frac{u^2}{c^2}}} \quad (12.42)$$

[here,  $u^2 = u_x^2 + u_y^2 + u_z^2$ .]

We know the transformation of Lorentz Transformation factor as (see Chap. 7, note - 6)

$$\sqrt{1 - \left(\frac{u^1}{c}\right)^2} = \frac{\sqrt{1 - \left(\frac{v}{c}\right)^2} \sqrt{1 - \left(\frac{u}{c}\right)^2}}{\left(1 - \frac{u_x v}{c^2}\right)}$$

Using this value,  $\rho^1$  takes the form

$$\rho^1 = \frac{\rho_0}{A^1} \frac{\left(1 - \frac{u_x v}{c^2}\right)}{\sqrt{1 - \left(\frac{v}{c}\right)^2} \sqrt{1 - \left(\frac{u}{c}\right)^2}} = \frac{\rho_0 A \left(1 - \frac{u_x v}{c^2}\right)}{A^1 \sqrt{1 - \left(\frac{v}{c}\right)^2}}$$

Now substituting the values of A and  $A^1$ , we obtain

$$\rho^1 = \frac{\rho_0 \sqrt{1 - \frac{u_x^2}{c^2}} \sqrt{1 - \frac{u_y^2}{c^2}} \sqrt{1 - \frac{u_z^2}{c^2}} \left(1 - \frac{u_x v}{c^2}\right)}{\sqrt{1 - \frac{u_x^2}{c^2}} \sqrt{1 - \frac{u_y^2}{c^2}} \sqrt{1 - \frac{u_z^2}{c^2}} \sqrt{1 - \left(\frac{v}{c}\right)^2}} \quad (12.43)$$

**Particular case:** If the body is at rest in S frame, then

$$u_x = u_y = u_z = 0, \quad u_x^1 = -v, \quad u_y^1 = u_z^1 = 0, \quad \rho = \rho_0$$

Then Eq. (12.42) becomes

$$\rho^1 = \frac{\rho_0}{1 - \left(\frac{v}{c}\right)^2} \quad (12.44)$$

Thus observer sitting in  $S^1$  frame measures the density  $\rho_0$  of the body rest in S frame as  $\rho^1$  given in Eq. (12.44).

**Alternative Proof of the Particular Case**

Let us consider two frames S and  $S^1$ ,  $S^1$  is moving with velocity  $v$  relative to S along x direction. Let a parallelepiped of mass  $m_0$  with volume  $V_0$  and density  $\rho_0$  be at rest in S frame. Then,

$$V_0 = l_x l_y l_z, \quad \rho_0 = \frac{m_0}{V_0}$$

[ $l_x, l_y, l_z$  be the lengths of the edge of the body in S frame]

The volume of parallelepiped measured in  $S^1$  frame will be

$$V^1 = l_x \sqrt{1 - \frac{v^2}{c^2}} l_y l_z = V_0 \sqrt{1 - \frac{v^2}{c^2}} \quad (12.45)$$

[length contraction occurs only along x direction]

Also the mass of parallelepiped measured in  $S^1$  frame will be

$$m^1 = \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (12.46)$$

Hence density in  $S^1$  frame will be

$$\rho^1 = \frac{m^1}{V^1} = \frac{m_0}{\sqrt{1 - \frac{v^2}{c^2}}} \frac{1}{V_0 \sqrt{1 - \frac{v^2}{c^2}}} = \frac{m_0}{V_0} \frac{1}{(1 - \frac{v^2}{c^2})} = \frac{\rho_0}{(1 - \frac{v^2}{c^2})} \quad (12.47)$$

*Example 12.1* A particle of rest mass  $m_0$  moves on the x-axis and attached to the origin by a force  $m_0 k^2 x$ . If it performs oscillation of amplitude  $a$ , then show that the periodic time of this relativistic harmonic oscillation is

$$T = \frac{4}{c} \int_0^a \frac{f dx}{\sqrt{f^2 - m_0^2 c^4}} \quad \text{where} \quad f = m_0 c^2 + \frac{1}{2} k^2 (a^2 - x^2) m_0$$

Also verify that as  $c \rightarrow \infty$ ;  $T = \frac{2\pi}{k}$  and show that if  $\frac{ka}{c}$  is small then  $T = \frac{2\pi}{k} (1 + \frac{3}{16} \frac{k^2 a^2}{c^2})$ .

**Hint:** Let  $v$  be the velocity of Harmonic oscillator along x-axis.

Therefore, the relativistic kinetic energy of harmonic oscillator is  $= (m - m_0)c^2$  where  $m_0 =$  rest mass of harmonic oscillator.

The rest energy of Harmonic Oscillator (H.O.)  $= m_0 c^2$

The potential energy of H.O.  $= \frac{1}{2} k^2 m_0 x^2$

$$\left[ \text{we know, } -\frac{dV}{dx} = F(x); \text{ here, } F(x) = -m_0 k^2 x, \text{ therefore, } V = \frac{1}{2} m_0 k^2 x^2 \right]$$

Total energy of the oscillator is given by

$$E = (m - m_0)c^2 + m_0c^2 + \frac{1}{2}kx^2m_0 = mc^2 + \frac{1}{2}k^2m_0x^2$$

Therefore,

$$E = \frac{m_0c^2}{\sqrt{1 - \frac{v^2}{c^2}}} + \frac{1}{2}k^2m_0x^2 \quad (1)$$

where  $v$  is the velocity of the particle.

If the particle comes to rest at  $x = a$ , i.e.  $v = 0$  at  $x = a$ , then from (1) we have,

$$E = m_0c^2 + \frac{1}{2}k^2m_0a^2 \quad (2)$$

Eq. (1) gives

$$1 - \frac{v^2}{c^2} = \frac{m_0^2c^4}{(E - \frac{1}{2}m_0k^2x^2)^2}$$

or,

$$\frac{v}{c} = \sqrt{1 - \frac{m_0^2c^4}{(E - \frac{1}{2}m_0k^2x^2)^2}} \quad (3)$$

Therefore the period of Harmonic Oscillation is given by

$$\begin{aligned} T &= 4 \int_0^a dt = 4 \int_0^a \frac{dx}{v} \\ &= \frac{4}{c} \int_0^a \frac{dx}{\sqrt{1 - \frac{m_0^2c^4}{(E - \frac{1}{2}m_0k^2x^2)^2}}} \\ &= \frac{4}{c} \int_0^a \frac{E - \frac{1}{2}m_0k^2x^2}{\sqrt{[(E - \frac{1}{2}m_0k^2x^2)^2 - m_0^2c^4]}} dx \end{aligned}$$

Now putting the value of  $E$  from (2), we obtain

$$T = \frac{4}{c} \int_0^a \frac{f dx}{\sqrt{(f^2 - m_0^2c^4)}} \quad (4)$$

where

$$f = m_0 c^2 + \frac{1}{2} m_0 k^2 (a^2 - x^2) \quad (\text{proved})$$

Now substituting  $x = a \sin \phi$ ;  $dx = a \cos \phi d\phi$  in (4), we get,

$$\begin{aligned} T &= \frac{4}{c} \int_0^{\frac{\pi}{2}} \frac{m_0 c^2 + \frac{1}{2} m_0 k^2 (a^2 - a^2 \sin^2 \phi) a \cos \phi}{\sqrt{[m_0 c^2 + \frac{1}{2} m_0 k^2 (a^2 - a^2 \sin^2 \phi)]^2 - m_0^2 c^4}} d\phi \\ &= \frac{4}{c} \int_0^{\frac{\pi}{2}} \frac{(m_0 c^2 + \frac{1}{2} k^2 m_0 a^2 \cos^2 \phi) a \cos \phi}{\sqrt{\frac{1}{4} m_0^2 k^4 a^4 \cos^4 \phi + m_0^2 c^2 k^2 a^2 \cos^2 \phi}} d\phi \\ &= \frac{4a}{c} \int_0^{\frac{\pi}{2}} \frac{m_0 c^2 (1 + \frac{1}{2} \frac{a^2 k^2}{c^2} \cos^2 \phi)}{k a m_0 c \sqrt{1 + \frac{1}{4} \frac{k^2 a^2 \cos^2 \phi}{c^2}}} d\phi \end{aligned}$$

For  $\lambda = \frac{ak}{2c}$ , the period of Harmonic Oscillation is given by

$$T = \frac{4}{k} \int_0^{\frac{\pi}{2}} \frac{1 + 2\lambda^2 \cos^2 \phi}{\sqrt{1 + \lambda^2 \cos^2 \phi}} d\phi \quad (5)$$

Now, as  $c \rightarrow \infty$ ,  $\lambda \rightarrow 0$ , then from (5) we get,  $T = \frac{2\pi}{k}$  (verified).

Again,  $\frac{ka}{c}$  is small, then from (5), we have (i.e.  $\lambda$  is small)

$$\begin{aligned} T &= \frac{4}{k} \int_0^{\frac{\pi}{2}} (1 + 2\lambda^2 \cos^2 \phi) \left(1 - \frac{\lambda^2}{2} \cos^2 \phi + \dots\right) d\phi \\ &\approx \frac{4}{k} \int_0^{\frac{\pi}{2}} \left(1 + \frac{3}{2} \lambda^2 \cos^2 \phi\right) d\phi \\ &= \frac{4}{k} \left[ \frac{\pi}{2} + \frac{3}{2} \lambda^2 \frac{\pi}{4} \right] \\ &= \frac{2\pi}{k} \left[ 1 + \frac{3}{4} \lambda^2 \right] \\ &= \frac{2\pi}{k} \left[ 1 + \frac{3}{16} \frac{a^2 k^2}{c^2} \right] \end{aligned}$$

*Example 12.2* Show that  $\frac{dp_x dy dp_z}{E}$  is invariant under Lorentz Transformation.



**Hint:** We are to show

$$\frac{dp_x^1 dy^1 dp_z^1}{E^1} = \frac{dp_x dy dp_z}{E}$$

We know

$$p_x^1 = a \left[ p_x - \left( \frac{vE}{c^2} \right) \right], \quad p_y = p_y^1, \quad p_z = p_z^1, \quad E^1 = a [E - vp_x], \quad E^2 = p^2 c^2 + m_0^2 c^4$$

$$[\beta = \frac{v}{c}, \quad p^2 = p_x^2 + p_y^2 + p_z^2 \text{ and } a = \frac{1}{\sqrt{1 - \beta^2}}]$$

Now,

$$\frac{dp_x^1 dy^1 dp_z^1}{E^1} = \frac{a \left[ dp_x - \left( \frac{v}{c^2} dE \right) \right] dy dp_z}{a [E - vp_x]}$$

$$E^2 = p^2 c^2 + m_0^2 c^4 = (p_x^2 + p_y^2 + p_z^2) c^2 + m_0^2 c^4 \implies 2E \frac{dE}{dp_x} = c^2 p_x$$

Hence

$$\frac{dp_x^1 dy^1 dp_z^1}{E^1} = \frac{a \left[ 1 - \left( \frac{v}{c^2} \frac{dE}{dp_x} \right) \right] dp_x dy dp_z}{a [E - vp_x]} = \frac{dp_x dy dp_z}{E}$$

# Chapter 13

## Electrodynamics in Relativity

### 13.1 Relativistic Electrodynamics

When charges are in motion, the electric and magnetic fields are associated with this motion, which have space and time variation. This phenomenon is called electromagnetism. The study involves time-dependent electromagnetic fields and the behaviour of which is described by a set of equations called Maxwell's equations.

### 13.2 Equation of Continuity

According to the principle of conservation of charge the net amount of charge in an isolated system remains constant, i.e. the time rate of increase of charge within the volume equal to the net rate of flow of charge into the volume. This statement of conservation of charge can be expressed by the equation of continuity, which can be written as

$$\nabla \cdot \vec{J} + \frac{\partial \rho}{\partial t} = 0 \quad (13.1)$$

where  $\vec{J}$  = current density,  $\rho$  = charge density.

This states that 'total current flowing out of some volume must be equal to the rate of decrease of charge within the volume'.

If

$$\frac{\partial \rho}{\partial t} = 0, \quad (13.2)$$

then

$$\nabla \cdot \vec{J} = 0 \quad (13.3)$$

This expresses the fact that in case of stationary currents there is no net outward flux of current density.

### 13.3 Maxwell's Equations

Prior to Maxwell, there were four fundamental equations of electromagnetism prescribed by different scientists. Maxwell modified these equations and put them together in compact form known as Maxwell's equations in electromagnetism. The Maxwell's equations represent mathematical expression of certain results. These equations cannot be verified directly, however their application to any situation can be verified. The electromagnetic theory of light is based on Maxwell's equations of electromagnetic field. These equations are differential forms of elementary laws of electricity and magnetism given by Gauss, Faraday and Ampere.

[1] Gauss's law in electrostatics

$$\begin{aligned}\nabla \cdot \vec{E} &= \left(\frac{1}{\epsilon_0}\right) \rho \\ &= 0, \quad \text{for a region without charge}\end{aligned}\tag{13.4}$$

[2] Application of Gauss's law to static magnetic fields. It has no name

$$\nabla \cdot \vec{B} = 0\tag{13.5}$$

[3] Faraday's law and Lenz's law of electromagnetic induction

$$\nabla \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}\tag{13.6}$$

[4] Ampere's law with Maxwell's corrections

$$\nabla \times \vec{B} - \mu_0 \epsilon_0 \left(\frac{\partial \vec{E}}{\partial t}\right) = \mu_0 \vec{J}\tag{13.7}$$

Ampere's law:

$$\nabla \times \vec{B} = \mu_0 \vec{J}\tag{13.8}$$

where,  $\vec{E}$  = Electric field strength,  $\vec{B}$  = Magnetic field strength,

$\vec{J}$  = current density,  $\rho$  = charge density.

$\mu_0$  = magnetic permeability of free space =  $4\pi \times 10^{-7}$  Wb/A m,

$\epsilon_0$  = electric permittivity of free space =  $8.86 \times 10^{-12}$  C<sup>2</sup>/m<sup>2</sup>.

Here,

$$\mu_0 \epsilon_0 = \frac{1}{c^2} \quad (13.9)$$

Maxwell's field equations can be solved for  $\vec{B}$  and  $\vec{E}$  in terms of a scalar function  $\phi$  and a vector function  $\vec{A}$  as

$$\vec{B} = \nabla \times \vec{A} \quad (13.10)$$

[it follows from (13.5) as  $\nabla \cdot \nabla \times \vec{A} = 0$ ]

$$\vec{E} = -\left(\frac{\partial \vec{A}}{\partial t}\right) - \nabla \phi \quad (13.11)$$

Faraday's law given in (13.6) implies

$$\nabla \times \vec{E} = -\left(\frac{\partial \vec{B}}{\partial t}\right) = -\left(\frac{\partial}{\partial t}\right)(\nabla \times \vec{A}) = -\nabla \times \left(\frac{\partial \vec{A}}{\partial t}\right)$$

This gives

$$\vec{E} = -\left(\frac{\partial \vec{A}}{\partial t}\right) - \nabla \phi$$

[ $\nabla \phi$  is similar to integration constant as  $\nabla \times \nabla \phi = 0$ ]

where,  $\vec{A}$  is known as *magnetic potential* and  $\phi$  indicates *electric potential*.

### 13.4 Derivation of Equation of Continuity from Maxwell's Equations

If the charges move with velocity  $\vec{u}$ , then current density

$$\vec{j} = \rho \vec{u} \quad (13.12)$$

Taking divergence of equation (13.7), we get

$$\begin{aligned} \nabla \cdot \nabla \times \vec{B} &= \mu_0 \epsilon_0 \nabla \cdot \left(\frac{\partial \vec{E}}{\partial t}\right) + \mu_0 \nabla \cdot \vec{J} \\ &= \mu_0 \epsilon_0 \left[\frac{\partial(\nabla \cdot \vec{E})}{\partial t}\right] + \mu_0 \nabla \cdot \vec{J} \end{aligned}$$

As  $\nabla \cdot \nabla \times \vec{B} = 0$ , we have

$$\mu_0 \varepsilon_0 \left[ \frac{\partial(\nabla \cdot \vec{E})}{\partial t} \right] + \mu_0 \nabla \cdot \vec{J} = 0$$

Substituting,  $\nabla \cdot \vec{E} = (1/\varepsilon_0) \rho$  in the above equation, we get

$$\nabla \cdot \vec{J} + \frac{\partial \rho}{\partial t} = 0$$

which is the equation of continuity and expresses the law of conservation of total charge.

### 13.5 Displacement Current

In the set of Maxwell's equations, the following equation

$$\nabla \times \vec{B} - \mu_0 \varepsilon_0 \left( \frac{\partial \vec{E}}{\partial t} \right) = \mu_0 \vec{J}$$

is the Ampere's law with Maxwell's corrections. The quantity  $(\partial \vec{E} / \partial t)$  was introduced by Maxwell and is called *displacement current*.

Ampere's law was

$$\nabla \times \vec{B} = \mu_0 \vec{J} \quad (13.13)$$

If one applies the divergence to (13.13), then one can obtain

$$\nabla \cdot \nabla \times \vec{B} = \mu_0 (\nabla \cdot \vec{J})$$

This implies

$$\nabla \cdot \vec{J} = 0 \quad (13.14)$$

This equation is valid for steady-state problems, i.e. when charge density is not changing. The complete relation is given by the continuity equation

$$\nabla \cdot \vec{J} + \frac{\partial \rho}{\partial t} = 0 \quad (13.15)$$

Therefore, Maxwell realized that the definition of total current density is incomplete and he suggested that some quantity be added to the right-hand side of (13.13), i.e. to  $\vec{J}$ . Assuming it to be  $\vec{J}^1$ , Eq. (13.13) becomes

$$\nabla \times \vec{B} = \mu_0(\vec{J} + \vec{J}^1) \quad (13.16)$$

Taking divergence of equation (13.16), one finds

$$\nabla \cdot \nabla \times \vec{B} = \mu_0 \nabla \cdot (\vec{J} + \vec{J}^1)$$

This implies

$$\nabla \cdot \vec{J}^1 = -\nabla \cdot \vec{J} = \frac{\partial \rho}{\partial t} \quad (13.17)$$

From Gauss's law, one can see that electric field  $E$  is related to the charge density  $\rho$  by

$$\nabla \cdot \vec{E} = \left(\frac{1}{\epsilon_0}\right) \rho \quad (13.18)$$

Substituting the value of  $\rho$  from (13.18) into (13.17), we obtain

$$\nabla \cdot \vec{J}^1 = \frac{\partial[\epsilon_0(\nabla \cdot \vec{E})]}{\partial t} = \nabla \cdot \frac{\partial(\epsilon_0 \vec{E})}{\partial t}$$

This is true for any arbitrary volume, therefore we have

$$\vec{J}^1 = \epsilon_0 \frac{\partial \vec{E}}{\partial t} \quad (13.19)$$

Thus if we add  $\epsilon_0 \partial \vec{E} / \partial t$  to the right-hand side of (13.13), then its divergence will satisfy the continuity equation (13.15) so that the discrepancy is removed.

With  $\epsilon_0 \partial \vec{E} / \partial t$  included, we have generalized Ampere's law

$$\nabla \times \vec{B} - \mu_0 \epsilon_0 \left( \frac{\partial \vec{E}}{\partial t} \right) = \mu_0 \vec{J} \quad (13.20)$$

The new quantity first introduced by Maxwell and he called the added term  $\epsilon_0 \partial \vec{E} / \partial t$  in (13.20), the *displacement current*. After the inclusion of displacement current by Maxwell, the modified Ampere's law includes time varying fields. According to Maxwell this added term is as effective as the conduction current  $\vec{J}$  for producing magnetic field.

### 13.6 Transformation for Charge Density

Let us consider two frames S and  $S^1$ ,  $S^1$  is moving with velocity  $v$  relative to S along x direction. Let a parallelepiped of volume  $V_0$  containing charge density  $\rho_0$  be at rest in S frame. Then,

$$V_0 = l_x l_y l_z$$

[ $l_x, l_y, l_z$  are the lengths of the edge of the body in S frame]

The volume of parallelepiped measured in  $S^1$  frame is

$$V^1 = l_x \sqrt{1 - \frac{v^2}{c^2}} l_y l_z = V_0 \sqrt{1 - \frac{v^2}{c^2}} \quad (13.21)$$

[length contraction occurs only along x direction]

Let charge density be measured in  $S^1$  frame  $\rho^1$  and charge contained in the volume  $V^1$  is  $\rho^1 V^1$ . We know charge is invariant, which implies

$$\rho^1 V^1 = \rho_0 V_0 \quad (13.22)$$

Using (13.21), we obtain,

$$\rho^1 = \frac{\rho_0}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (13.23)$$

This indicates that when a charged particle is moving then charge density will be increased.

### 13.7 Four Current Vector

Let us consider an elementary volume element

$$dV = dx_1 dx_2 dx_3,$$

containing rest charge density  $\rho_0$  moving with velocity  $v$ . Then charge in this volume is

$$dq = \rho dx_1 dx_2 dx_3 \quad (13.24)$$

[here,  $\rho = \rho_0 / \sqrt{1 - \frac{v^2}{c^2}}$ ]

Now, we multiply both sides of (13.24) by a four vector  $dx_\mu$ , which yields

$$dq dx_\mu = \rho dx_1 dx_2 dx_3 dx_\mu = \rho \frac{dx_\mu}{dt} dx_1 dx_2 dx_3 dt \quad (13.25)$$

We know charge is invariant and this implies the left-hand side of Eq. (13.25) is a four vector. Therefore, the right-hand side of this equation must be a four vector. Again, we know the four-dimensional volume element is an invariant quantity, i.e.

$$dx_1 dx_2 dx_3 dx_4 = \text{invariant}$$

or,

$$dx_1 dx_2 dx_3 d(ict) = \text{invariant}$$

i.e.

$$dx_1 dx_2 dx_3 dt = \text{invariant}$$

Hence, the right-hand side of Eq. (13.25) must be a four vector. In other words,

$$\rho \frac{dx_\mu}{dt}$$

should be a four vector.

This four vector is represented by  $J_\mu$  is known as current four vector, i.e.

$$J_\mu = \rho \frac{dx_\mu}{dt} = \frac{\rho_0}{\sqrt{1 - \frac{v^2}{c^2}}} \frac{dx_\mu}{dt} = \rho_0 \frac{dx_\mu}{d\tau} = \rho_0 u_\mu \quad (13.26)$$

Thus current four vector  $J_\mu$  is defined by multiplying the invariant rest charge density  $\rho_0$  by the four velocity vector  $u_\mu$  so that

$$J_1 = \rho \frac{dx_1}{dt} = \frac{\rho_0}{\sqrt{1 - \frac{v^2}{c^2}}} v_1 = \rho_0 u_1, \quad J_2 = \rho \frac{dx_2}{dt} = \rho v_2 = \frac{\rho_0}{\sqrt{1 - \frac{v^2}{c^2}}} v_2 = \rho_0 u_2,$$

$$J_3 = \rho \frac{dx_3}{dt} = \frac{\rho_0}{\sqrt{1 - \frac{v^2}{c^2}}} v_3 = \rho_0 u_3, \quad J_4 = \rho \frac{dx_4}{dt} = \rho \frac{d(ict)}{dt} = ic\rho = \frac{ic\rho_0}{\sqrt{1 - \frac{v^2}{c^2}}}$$

Thus,

$$J_\mu = [J_1, J_2, J_3, J_4 \equiv ic\rho] = [\rho_0 u_1, \rho_0 u_2, \rho_0 u_3, ic\rho] \quad (13.27)$$



### 13.8 Equation of Continuity in Covariant Form

We write the continuity equation

$$\nabla \cdot \vec{J} + \frac{\partial \rho}{\partial t} = 0$$

explicitly as

$$\frac{\partial J_1}{\partial x_1} + \frac{\partial J_2}{\partial x_2} + \frac{\partial J_3}{\partial x_3} + \frac{\partial \rho}{\partial t} = 0$$

or

$$\frac{\partial J_1}{\partial x_1} + \frac{\partial J_2}{\partial x_2} + \frac{\partial J_3}{\partial x_3} + \frac{\partial(ict\rho)}{\partial(ict)} = 0$$

or

$$\frac{\partial J_1}{\partial x_1} + \frac{\partial J_2}{\partial x_2} + \frac{\partial J_3}{\partial x_3} + \frac{\partial J_4}{\partial x_4} = 0$$

or

$$\frac{\partial J_\mu}{\partial x_\mu} = 0 \tag{13.28}$$

This is the continuity equation in covariant form. Since the left-hand side of Eq. (13.28) is a four-dimensional scalar, therefore it is invariant under Lorentz Transformation. Also, this equation indicates that four divergence of the current four vector  $J_\mu$  vanishes.

#### Proof of Invariance of Continuity Equation

**Hint:** We have to show  $\partial J_\mu^1 / \partial x_\mu^1 = \partial J_\mu / \partial x_\mu$ , i.e.

$$\frac{\partial J_1^1}{\partial x_1^1} + \frac{\partial J_2^1}{\partial x_2^1} + \frac{\partial J_3^1}{\partial x_3^1} + \frac{\partial J_4^1}{\partial x_4^1} = \frac{\partial J_1}{\partial x_1} + \frac{\partial J_2}{\partial x_2} + \frac{\partial J_3}{\partial x_3} + \frac{\partial J_4}{\partial x_4}$$

We know that

$$x_1 = a(x_1^1 + vt^1); \quad x_2 = x_2^1; \quad x_3 = x_3^1; \quad t = a \left[ t^1 + \left( \frac{vx_1^1}{c^2} \right) \right]$$

[ $\beta = v/c$  and  $a = 1/\sqrt{1-\beta^2}$ ]

Applying chain rule, we have

$$\frac{\partial}{\partial x_1^1} = \frac{\partial}{\partial x_1} \frac{\partial x_1}{\partial x_1^1} + \frac{\partial}{\partial x_2} \frac{\partial x_2}{\partial x_1^1} + \frac{\partial}{\partial x_3} \frac{\partial x_3}{\partial x_1^1} + \frac{\partial}{\partial t} \frac{\partial t}{\partial x_1^1}$$

Thus we get,

$$\frac{\partial}{\partial x_1^1} = a \left( \frac{\partial}{\partial x_1} + \frac{v}{c^2} \frac{\partial}{\partial t} \right) \quad (13.29)$$

Similarly,

$$\frac{\partial}{\partial x_2^1} = \frac{\partial}{\partial x_2}, \quad \frac{\partial}{\partial x_3^1} = \frac{\partial}{\partial x_3}, \quad \frac{\partial}{\partial t^1} = a \left( v \frac{\partial}{\partial x_1} + \frac{\partial}{\partial t} \right) \quad (13.30)$$

Now,

$$\begin{aligned} \frac{\partial J_1^1}{\partial x_1^1} + \frac{\partial J_2^1}{\partial x_2^1} + \frac{\partial J_3^1}{\partial x_3^1} + \frac{\partial J_4^1}{\partial x_4^1} &= a \left( \frac{\partial [a(J_1 - v\rho)]}{\partial x_1} + \frac{v}{c^2} \frac{\partial [a(J_1 - v\rho)]}{\partial t} \right) \\ &+ \frac{\partial J_2}{\partial x_2} + \frac{\partial J_3}{\partial x_3} + a \left( v \frac{\partial [a(\rho - \frac{v}{c^2} J_1)]}{\partial x_1} \right. \\ &\quad \left. + \frac{\partial [a(\rho - \frac{v}{c^2} J_1)]}{\partial t} \right) \\ &= \frac{\partial J_1}{\partial x_1} + \frac{\partial J_2}{\partial x_2} + \frac{\partial J_3}{\partial x_3} + \frac{\partial J_4}{\partial x_4} \end{aligned}$$

### 13.9 Transformation of Four-Current Vector

The four current vector can be written as

$$J_\mu = [J_1, J_2, J_3, J_4] \quad (13.31)$$

where  $J_4 = ic\rho$ .

We know Lorentz transformation of any four vector  $A_\mu$  ( $\mu = 1, 2, 3, 4$ ) may be expressed as

$$A_\mu^1 = a_{\mu\nu} A_\nu$$

where  $a_{\mu\nu}$  is given as

$$a_{\mu\nu} = \begin{bmatrix} a & 0 & 0 & ia\beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -ia\beta & 0 & 0 & a \end{bmatrix}$$

$$[\beta = v/c \text{ and } a = 1/\sqrt{1 - \beta^2}]$$

Hence,

$$\begin{bmatrix} J_1^1 \\ J_2^1 \\ J_3^1 \\ J_4^1 \end{bmatrix} = \begin{bmatrix} a & 0 & 0 & ia\beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -ia\beta & 0 & 0 & a \end{bmatrix} \begin{bmatrix} J_1 \\ J_2 \\ J_3 \\ J_4 \end{bmatrix} \quad (13.32)$$

This implies,

$$J_1^1 = aJ_1 + ia\beta J_4 = aJ_1 + ia \left( \frac{v}{c} \right) (ic\rho)$$

i.e.

$$J_1^1 = a [J_1 - v\rho] \quad (13.33)$$

$$J_2^1 = J_2, \quad J_3^1 = J_3 \quad (13.34)$$

$$J_4^1 = -ia\beta J_1 + aJ_4$$

or,

$$(ic\rho^1) = -ia \left( \frac{v}{c} \right) J_1 + a (ic\rho),$$

i.e.

$$\rho^1 = a \left[ \rho - \left( \frac{v}{c^2} \right) J_1 \right] \quad (13.35)$$

The inverse transformations are obtained by replacing  $v$  by  $-v$  as

$$J_1 = a \left[ J_1^1 + v\rho^1 \right], \quad J_2 = J_2^1, \quad J_3 = J_3^1, \quad \rho = a \left[ \rho^1 + \left( \frac{v}{c^2} \right) J_1^1 \right] \quad (13.36)$$

Let a body containing charge density be at rest in the  $S^1$  frame so that its rest charge density is  $\rho_0$ . This means for observer in  $S$  frame the charged body is moving with velocity  $v$ . So, in  $S^1$  frame,

$$J_1^1 = J_2^1 = J_3^1 = 0, \quad \rho^1 = \rho_0$$

Hence,

$$J_1 = av\rho_0, \quad J_2 = 0, \quad J_3 = 0, \quad \rho = a\rho_0 \quad (13.37)$$

### 13.10 Maxwell's Equations in Covariant Form

In terms of magnetic potential  $\vec{A}$  and electric potential  $\phi$ , the Maxwell's equations can be written as

$$\square^2 \vec{A} = -\mu_0 \vec{J} \quad (13.38)$$

$$\square^2 \phi = - \left( \frac{1}{\epsilon_0} \right) \rho \quad (13.39)$$

with

$$\nabla \cdot \vec{A} + \frac{1}{c^2} \frac{\partial \phi}{\partial t} = 0 \quad (13.40)$$

Here,  $\square^2 = \nabla^2 - 1/c^2 \cdot \partial^2/\partial t^2$  is *d'Alembertian operator*.

Equation (13.40) is called *Lorentz-Gauge condition*.

*Proof* From the Maxwell's equation (13.7), we have

$$\nabla \times \vec{B} - \mu_0 \epsilon_0 \left( \frac{\partial \vec{E}}{\partial t} \right) = \mu_0 \vec{J}$$

Using the  $\vec{B}$  and  $\vec{E}$  in terms of  $\vec{A}$  and  $\phi$  from (13.10) and (13.11), we obtain

$$\begin{aligned} \nabla \times \nabla \times \vec{A} &= \mu_0 \epsilon_0 \left( \frac{\partial \left[ - \left( \frac{\partial \vec{A}}{\partial t} \right) - \nabla \phi \right]}{\partial t} \right) + \mu_0 \vec{J} \\ \implies \nabla (\nabla \cdot \vec{A}) - \nabla^2 \vec{A} &= -\mu_0 \epsilon_0 \left( \frac{\partial^2 \vec{A}}{\partial t^2} \right) - \mu_0 \epsilon_0 \left( \frac{\partial (\nabla \phi)}{\partial t} \right) + \mu_0 \vec{J} \\ \implies \nabla^2 \vec{A} - \frac{1}{c^2} \frac{\partial^2 \vec{A}}{\partial t^2} &= \nabla \left[ \nabla \cdot \vec{A} + \frac{1}{c^2} \frac{\partial \phi}{\partial t} \right] - \mu_0 \vec{J} \\ \implies \nabla^2 \vec{A} - \frac{1}{c^2} \frac{\partial^2 \vec{A}}{\partial t^2} &\equiv \square^2 \vec{A} = -\mu_0 \vec{J} \end{aligned}$$

provided

$$\nabla \cdot \vec{A} + \frac{1}{c^2} \frac{\partial \phi}{\partial t} = 0$$

Again from Maxwell equation (13.4), we have

$$\begin{aligned} \left( \frac{1}{\epsilon_0} \right) \rho &= \nabla \cdot \vec{E} \\ &= \nabla \cdot \left[ - \left( \frac{\partial \vec{A}}{\partial t} \right) - \nabla \phi \right] \\ &= -\nabla^2 \phi - \frac{\partial (\nabla \cdot \vec{A})}{\partial t} \end{aligned}$$

or

$$\begin{aligned}\square^2\phi + \frac{1}{c^2}\frac{\partial^2\phi}{\partial t^2} + \frac{\partial(\nabla\cdot\vec{A})}{\partial t} &= -\left(\frac{1}{\varepsilon_0}\right)\rho \\ \implies \square^2\phi + \frac{\partial}{\partial t}\left[\frac{1}{c^2}\frac{\partial\phi}{\partial t} + \nabla\cdot\vec{A}\right] &= -\left(\frac{1}{\varepsilon_0}\right)\rho \\ &\implies \square^2\phi = -\left(\frac{1}{\varepsilon_0}\right)\rho\end{aligned}$$

provided

$$\nabla\cdot\vec{A} + \frac{1}{c^2}\frac{\partial\phi}{\partial t} = 0$$

### 13.10.1 The d'Alembertian Operator Is Invariant Under Lorentz Transformation

*Proof* From (13.29) and (13.30), we have

$$\frac{\partial}{\partial x_1^1} = a\left(\frac{\partial}{\partial x_1} + \frac{v}{c^2}\frac{\partial}{\partial t}\right), \quad \frac{\partial}{\partial x_2^1} = \frac{\partial}{\partial x_2}, \quad \frac{\partial}{\partial x_3^1} = \frac{\partial}{\partial x_3}, \quad \frac{\partial}{\partial t^1} = a\left(v\frac{\partial}{\partial x_1} + \frac{\partial}{\partial t}\right)$$

From these, we can get

$$\frac{\partial^2}{\partial x_1^1{}^2} = a^2\left(\frac{\partial^2}{\partial x_1^2} + \frac{v^2}{c^4}\frac{\partial^2}{\partial t^2} + \frac{2v}{c^2}\frac{\partial^2}{\partial x\partial t}\right) \quad (13.41)$$

$$\frac{\partial^2}{\partial x_2^1{}^2} = \frac{\partial^2}{\partial x_2^2}, \quad \frac{\partial^2}{\partial x_3^1{}^2} = \frac{\partial^2}{\partial x_3^2}, \quad \frac{\partial^2}{\partial t^1{}^2} = a^2\left(v^2\frac{\partial^2}{\partial x_1^2} + \frac{\partial^2}{\partial t^2} + 2v\frac{\partial^2}{\partial x\partial t}\right) \quad (13.42)$$

Therefore,

$$\begin{aligned}\square^{12} &= \frac{\partial^2}{\partial x_1^1{}^2} + \frac{\partial^2}{\partial x_2^1{}^2} + \frac{\partial^2}{\partial x_3^1{}^2} - \frac{1}{c^2}\frac{\partial^2}{\partial t^1{}^2} \\ &= a^2\left(\frac{\partial^2}{\partial x_1^2} + \frac{v^2}{c^4}\frac{\partial^2}{\partial t^2} + \frac{2v}{c^2}\frac{\partial^2}{\partial x\partial t}\right) + \frac{\partial^2}{\partial x_2^2} + \frac{\partial^2}{\partial x_3^2} \\ &\quad - \frac{a^2}{c^2}\left(v^2\frac{\partial^2}{\partial x_1^2} + \frac{\partial^2}{\partial t^2} + 2v\frac{\partial^2}{\partial x\partial t}\right) \\ &= \frac{\partial^2}{\partial x_1^2} + \frac{\partial^2}{\partial x_2^2} + \frac{\partial^2}{\partial x_3^2} - \frac{1}{c^2}\frac{\partial^2}{\partial t^2} = \square^2\end{aligned}$$

It is shown that the d'Alembertian operator is the invariant four-dimensional Laplacian, therefore, the right-hand sides of (13.38) and (13.39) should be the components of a four vector. Hence, Lorentz covariance requires that the potential  $\phi$  and  $A$  form a four-vector potential.

### 13.10.2 Lorentz-Gauge Condition in Covariant Form

The Lorentz-Gauge condition is

$$\nabla \cdot \vec{A} + \frac{1}{c^2} \frac{\partial \phi}{\partial t} = 0$$

This can be written as

$$\nabla \cdot \vec{A} + \frac{\partial \left( \frac{i\phi}{c} \right)}{\partial(ict)} = 0$$

or

$$\frac{\partial A_1}{\partial x_1} + \frac{\partial A_2}{\partial x_2} + \frac{\partial A_3}{\partial x_3} + \frac{\partial A_4}{\partial x_4} = 0$$

i.e.

$$\frac{\partial A_\mu}{\partial x_\mu} = 0$$

This gives a clue about the fourth component of the four potential vector which is

$$A_4 = \frac{i\phi}{c} \quad (13.43)$$

Therefore, the four potential vector can be written as

$$A_\mu = [A_1, A_2, A_3, A_4] \quad (13.44)$$

where  $A_4 = i\phi/c$ .

Now, the Maxwell equation (13.39) takes the following form:

$$\square^2 A_4 = -\mu_0 J_4 \quad (13.45)$$

Hence the total Maxwell's equations with Lorentz-Gauge condition can be written in covariant forms as

$$\square^2 A_\mu = -\mu_0 J_\mu \quad (13.46)$$

$$\frac{\partial A_\mu}{\partial x_\mu} = 0 \quad (13.47)$$

### 13.10.3 Gauge Transformations

From Eq. (13.10), we can construct a new vector potential keeping unchanged the magnetic field as,

$$\vec{A}^1 = \vec{A} + \nabla\psi \quad (13.48)$$

where  $\psi$  is any arbitrary scalar quantity.

$$[B^1 = \nabla \times \vec{A}^1 = \nabla \times (\vec{A} + \nabla\psi) = \nabla \times \vec{A} = \vec{B} \quad \text{as } \nabla \times \nabla\psi = 0]$$

If this addition keeps the electric field given in Eq. (13.11) unchanged, then we have to reconstruct the scalar potential  $\phi$  in term of a new choice of scalar potential  $\phi^1$  as

$$\begin{aligned} \vec{E}^1 &= -\left(\frac{\partial \vec{A}^1}{\partial t}\right) - \nabla\phi^1 = -\left(\frac{\partial(\vec{A} + \nabla\psi)}{\partial t}\right) - \nabla\phi^1 \\ &= -\left(\frac{\partial \vec{A}}{\partial t}\right) - \nabla\left(\frac{\partial\psi}{\partial t}\right) - \nabla\phi^1 \\ &= -\left(\frac{\partial \vec{A}}{\partial t}\right) - \nabla\left(\frac{\partial\psi}{\partial t} + \phi^1\right) = -\left(\frac{\partial \vec{A}}{\partial t}\right) - \nabla\phi = E \end{aligned}$$

provided

$$\phi = \frac{\partial\psi}{\partial t} + \phi^1$$

i.e.

$$\phi^1 = \phi - \frac{\partial\psi}{\partial t} \quad (13.49)$$

Hence, electric field strength  $E$  and magnetic field strength  $B$  remain invariant under the transformations given in (13.48) and (13.49). Therefore, the set of transformations (13.48) and (13.49) leaves the Maxwell's equations unchanged and the transformations are called *Gauge transformations*.

The physical laws which remain invariant under gauge transformations are called *gauge invariants*.

When one applies gauge transformations to the Maxwell's equations, the gauge function  $\psi$  should not be taken arbitrarily; rather Lorentz-Gauge condition (13.40) imposes the restriction on  $\psi$ .

Lorentz-Gauge condition for gauge transformations is

$$\nabla \cdot \vec{A}^1 + \frac{1}{c^2} \frac{\partial\phi^1}{\partial t} = 0$$

Replacing the values from (13.48) and (13.49), we have

$$\nabla \cdot (\vec{A} + \nabla\psi) + \frac{1}{c^2} \frac{\partial \left( \phi - \frac{\partial\psi}{\partial t} \right)}{\partial t} = 0$$

This implies

$$\nabla \cdot \vec{A} + \nabla^2\psi + \frac{1}{c^2} \left( \frac{\partial\phi}{\partial t} \right) - \frac{1}{c^2} \left( \frac{\partial^2\psi}{\partial t^2} \right) = 0$$

Using Lorentz-Gauge condition

$$\nabla \cdot \vec{A} + \frac{1}{c^2} \frac{\partial\phi}{\partial t} = 0,$$

we obtain

$$\nabla^2\psi - \frac{1}{c^2} \left( \frac{\partial^2\psi}{\partial t^2} \right) = \square^2\psi = 0$$

Thus the solution of this equation can be used to construct the gauge transformation functions  $A^1$  and  $\phi^1$  from the old  $A$  and  $\phi$ .

### 13.11 Transformation of Four-Potential Vector

The four-potential vector can be written as

$$A_\mu = [A_1, A_2, A_3, A_4] \quad (13.50)$$

where  $A_4 = i\phi/c$ . The first three components belong to magnetic potential and the fourth corresponds to electric potential.

We know Lorentz transformation of any four vector  $A_\mu$  ( $\mu = 1, 2, 3, 4$ ) may be expressed as

$$A_\mu^1 = a_{\mu\nu} A_\nu$$

where  $a_{\mu\nu}$  is given as

$$a_{\mu\nu} = \begin{bmatrix} a & 0 & 0 & ia\beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -ia\beta & 0 & 0 & a \end{bmatrix}$$

$$[\beta = v/c \text{ and } a = 1/\sqrt{1 - \beta^2}]$$



Hence,

$$\begin{bmatrix} A_1^1 \\ A_2^1 \\ A_3^1 \\ A_4^1 \end{bmatrix} = \begin{bmatrix} a & 0 & 0 & ia\beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -ia\beta & 0 & 0 & a \end{bmatrix} \begin{bmatrix} A_1 \\ A_2 \\ A_3 \\ A_4 \end{bmatrix} \quad (13.51)$$

This implies,

$$A_1^1 = aA_1 + ia\beta A_4 = aA_1 + ia \left(\frac{v}{c}\right) \left(\frac{i\phi}{c}\right)$$

i.e.

$$A_1^1 = a \left[ A_1 - \frac{v}{c^2} \phi \right] \quad (13.52)$$

$$\begin{aligned} A_2^2 &= A_2, & A_3^3 &= A_3 \\ A_4^1 &= -ia\beta A_1 + aA_4 \end{aligned} \quad (13.53)$$

or,

$$\left(\frac{i\phi^1}{c}\right) = -ia \left(\frac{v}{c}\right) A_1 + a \left(\frac{i\phi}{c}\right)$$

i.e.

$$\phi^1 = a [\phi - vA_1] \quad (13.54)$$

The inverse transformations are obtained by replacing  $v$  by  $-v$  as

$$A_1 = a \left[ A_1^1 + \frac{v}{c^2} \phi^1 \right], \quad A_2 = A_2^1, \quad A_3 = A_3^1, \quad \phi = a [\phi^1 + vA_1^1] \quad (13.55)$$

### 13.12 The Electromagnetic Field Tensor

The electric field strength  $\vec{E}$  and magnetic field strength  $\vec{B}$  in the Maxwell's equations are defined in terms of electromagnetic potentials  $\vec{A}$  and  $\phi$  by

$$\vec{B} = \nabla \times \vec{A} \quad \text{and} \quad \vec{E} = - \left( \frac{\partial \vec{A}}{\partial t} \right) - \nabla \phi$$

In view of these equations, we define a second rank anti-symmetric tensor  $F_{\mu\nu}$  which is constructed from the (four-dimensional) curl of the four potential  $A_\mu$  as

$$F_{\mu\nu} = \text{curl}(A_\mu) = \frac{\partial A_\nu}{\partial x_\mu} - \frac{\partial A_\mu}{\partial x_\nu} \quad (13.56)$$

Here,  $F_{\mu\nu} = -F_{\nu\mu}$  and  $F_{\mu\mu} = 0$ . Interestingly, one can note that the components of  $F_{\mu\nu}$  have remarkable correspondence with the components of electric field strength  $\vec{E}$  and magnetic field strength  $\vec{B}$  e.g.,

$$F_{12} = \frac{\partial A_2}{\partial x_1} - \frac{\partial A_1}{\partial x_2} = (\nabla \times \vec{A})_3 = B_3$$

$$(B_1 i + B_2 j + B_3 k) = \begin{bmatrix} i & j & k \\ \frac{\partial}{\partial x_1} & \frac{\partial}{\partial x_2} & \frac{\partial}{\partial x_3} \\ A_1 & A_2 & A_3 \end{bmatrix}$$

$$F_{14} = \frac{\partial A_4}{\partial x_1} - \frac{\partial A_1}{\partial x_4} = \frac{\partial \left( \frac{i\phi}{c} \right)}{\partial x_1} - \frac{\partial A_1}{\partial (ict)}$$

$$= \left( \frac{i}{c} \right) \frac{\partial \phi}{\partial x_1} - \left( \frac{1}{ic} \right) \frac{\partial A_1}{\partial t}$$

$$= \left( \frac{i}{c} \right) \left[ -\frac{\partial \phi}{\partial x_1} - \frac{\partial A_1}{\partial t} \right]$$

$$= \left( \frac{i}{c} \right) E_1 = - \left( \frac{iE_1}{c} \right) \text{ etc.}$$

The explicit values of the components of the above tensor may be depicted in the following form:

$$F_{\mu\nu} = \begin{bmatrix} 0 & B_3 & -B_2 & -\frac{iE_1}{c} \\ -B_3 & 0 & B_1 & -\frac{iE_2}{c} \\ B_2 & -B_1 & 0 & -\frac{iE_3}{c} \\ \frac{iE_1}{c} & \frac{iE_2}{c} & \frac{iE_3}{c} & 0 \end{bmatrix} \quad (13.57)$$

This second rank anti-symmetric tensor  $F_{\mu\nu}$  is called *electromagnetic field strength tensor*.

Again, we notice that the choice of the vector  $A_\mu$  given in (13.10) may not be unique. One can choose the new one as

$$A_\mu^* = A_\mu + \nabla\psi, \quad (\psi \text{ is a scalar function}) \quad (13.58)$$

It is easy to verify that the replacement of  $A_\mu$  by  $A_\mu^*$  does not change the expression given in (13.10). However, Lorentz gauge condition  $\partial A_\mu / \partial x_\mu = 0$  restricts the arbitrary choice of  $\psi$ . The function  $\psi$  should satisfy the

$$\nabla^2 \psi - \frac{1}{c^2} \left( \frac{\partial^2 \psi}{\partial t^2} \right) = \square^2 \psi = 0 \quad (13.59)$$

Maxwell's equations given in (13.4)–(13.7) can be expressed in terms of electromagnetic field strength tensor  $F_{\mu\nu}$  as

$$\frac{\partial F_{\mu\nu}}{\partial x_\nu} = \mu_0 J_\mu \quad (13.60)$$

$$\frac{\partial F_{\mu\nu}}{\partial x_\lambda} + \frac{\partial F_{\nu\lambda}}{\partial x_\mu} + \frac{\partial F_{\lambda\mu}}{\partial x_\nu} = 0 \quad (13.61)$$

The first and fourth equations in a set of Maxwell's equations can be expressed by Eq. (13.60). For example, when  $\mu = 1$ , the Eq. (13.60) gives

$$\frac{\partial F_{12}}{\partial x_2} + \frac{\partial F_{13}}{\partial x_3} + \frac{\partial F_{14}}{\partial x_4} = \mu_0 J_1$$

or

$$\frac{\partial B_3}{\partial x_2} - \frac{\partial B_2}{\partial x_3} - \left(\frac{i}{c}\right) \left(\frac{1}{ic}\right) \frac{\partial E_1}{\partial t} = \mu_0 J_1$$

or

$$\frac{\partial B_3}{\partial x_2} - \frac{\partial B_2}{\partial x_3} - \left(\frac{1}{c^2}\right) \frac{\partial E_1}{\partial t} = \mu_0 J_1$$

or

$$\frac{\partial B_3}{\partial x_2} - \frac{\partial B_2}{\partial x_3} - \mu_0 \varepsilon_0 \frac{\partial E_1}{\partial t} = \mu_0 J_1$$

or

$$(\nabla \times \vec{B})_1 - \mu_0 \varepsilon_0 \frac{\partial E_1}{\partial t} = \mu_0 J_1 \text{ etc.}$$

When  $\mu = 4$ , the Eq. (13.60) gives

$$\frac{\partial F_{42}}{\partial x_2} + \frac{\partial F_{43}}{\partial x_3} + \frac{\partial F_{41}}{\partial x_1} = \mu_0 J_4$$

or

$$\frac{\partial \left(\frac{iE_2}{c}\right)}{\partial x_2} + \frac{\partial \left(\frac{iE_3}{c}\right)}{\partial x_3} + \frac{\partial \left(\frac{iE_1}{c}\right)}{\partial x_1} = \mu_0 ic\rho$$

or

$$\frac{\partial E_2}{\partial x_2} + \frac{\partial E_3}{\partial x_3} + \frac{\partial E_1}{\partial x_1} = \frac{\rho}{\varepsilon_0} \left(\text{as } \varepsilon_0 \mu_0 = \frac{1}{c^2}\right)$$

or

$$\nabla \cdot \vec{E} = \frac{\rho}{\varepsilon_0}$$

The second and third equations in a set of Maxwell's equations can be expressed by Eq. (13.61). For example, if  $\mu$ ,  $\nu$  and  $\lambda$  assume the values from any combination of (1, 2, 3), we always get from Eq. (13.61)

$$\frac{\partial F_{12}}{\partial x_3} + \frac{\partial F_{23}}{\partial x_1} + \frac{\partial F_{31}}{\partial x_2} = 0$$

or

$$\frac{\partial B_3}{\partial x_3} + \frac{\partial B_1}{\partial x_1} + \frac{\partial B_2}{\partial x_2} = 0 \implies \nabla \cdot \vec{B} = 0$$

Again, if we take  $\mu = 2$ ,  $\nu = 4$  and  $\lambda = 1$ , then Eq. (13.61) gives

$$\frac{\partial F_{24}}{\partial x_1} + \frac{\partial F_{41}}{\partial x_2} + \frac{\partial F_{12}}{\partial x_4} = 0$$

or

$$-\frac{\partial \left( \frac{iE_2}{c} \right)}{\partial x_1} + \frac{\partial \left( \frac{iE_1}{c} \right)}{\partial x_2} + \frac{\partial B_3}{\partial (ict)} = 0$$

or

$$(\nabla \times \vec{E})_3 = -\frac{\partial B_3}{\partial t} \text{ etc.}$$

Equation (13.61) can also be written in terms of dual electromagnetic field tensor  $*F_{\mu\nu}$  as

$$\frac{\partial *F_{\mu\nu}}{\partial x_\nu} = 0$$

Here, dual electromagnetic field tensor  $*F_{\mu\nu}$  is defined as

$$*F_{\mu\nu} = \frac{1}{2} \epsilon_{\mu\nu\alpha\beta} F_{\alpha\beta}$$

where  $\epsilon_{\mu\nu\alpha\beta}$  is Levi-Civita tensor defined as

$$\begin{aligned} \epsilon_{\mu\nu\alpha\beta} &= 0, \quad \text{if any two indices are equal} \\ &= +1 \quad \text{if } \mu\nu\alpha\beta \text{ is an even permutation of } 0, 1, 2, 3 \\ &= -1 \quad \text{if } \mu\nu\alpha\beta \text{ is an odd permutation of } 0, 1, 2, 3 \end{aligned}$$

The components of  $*F_{\mu\nu}$  is given by

$$*F_{\mu\nu} = \begin{bmatrix} 0 & -\frac{iE_3}{c} & -\frac{iE_2}{c} & B_1 \\ \frac{iE_3}{c} & 0 & -\frac{iE_1}{c} & B_2 \\ -\frac{iE_2}{c} & \frac{iE_1}{c} & 0 & B_3 \\ -B_1 & -B_2 & -B_3 & 0 \end{bmatrix}$$

The Maxwell's equations given in (13.60) and (13.61) are in tensor form. We know a tensor equation must be the same in all coordinate systems, i.e. invariant. Therefore, the Maxwell equations are invariant under Lorentz Transformation. Thus, it is self-evident the covariance of Maxwell's equations under Lorentz Transformation.

### 13.13 Lorentz Transformation of Electromagnetic Fields

The components in the electromagnetic field strength tensor  $F_{\mu\nu}$  are electric and magnetic fields  $\vec{E}$  and  $\vec{B}$ . Here  $F_{\mu\nu}$  is a second rank anti-symmetric tensor. Therefore, one can use the Lorentz Transformation matrix  $a_{\mu\nu}$  to get the required Lorentz transformation of the electromagnetic field strength tensor  $F_{\mu\nu}$  as

$$F_{\mu\nu}^1 = a_{\mu\alpha} a_{\nu\beta} F_{\alpha\beta} \quad (13.62)$$

where

$$a_{\mu\nu} = \begin{bmatrix} a & 0 & 0 & ia\beta \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -ia\beta & 0 & 0 & a \end{bmatrix}$$

$$[\beta = v/c \text{ and } a = 1/\sqrt{1 - \beta^2}]$$

Using the above transformation equations, one can get the transformation of electric and magnetic field components as

$$E_1^1 = E_1, \quad E_2^1 = a[E_2 - c\beta B_3], \quad E_3^1 = a[E_3 + c\beta B_2] \quad (13.63)$$

$$B_1^1 = B_1, \quad B_2^1 = a \left[ B_2 + \left( \frac{\beta}{c} \right) E_3 \right], \quad B_3^1 = a \left[ B_3 - \left( \frac{\beta}{c} \right) E_2 \right] \quad (13.64)$$

The above relations can be obtained from (13.62) as follows:

[1] For magnetic field components:

Let us choose  $\mu = 3$  and  $\nu = 1$  because  $F_{31} = B_2$ . Writing explicitly the components of (13.62), we get,

$$F_{31}^1 = a_{3\alpha} a_{1\beta} F_{\alpha\beta} = a_{33}[a_{11}F_{31} + a_{14}F_{34}]$$

or

$$B_2^1 = aB_2 + ia\beta \left( \frac{-iE_3}{c} \right) = a \left[ B_2 + \left( \frac{\beta}{c} \right) E_3 \right]$$

Now we choose  $\mu = 2$  and  $\nu = 3$  because  $F_{23} = B_1$ . Equation (13.62) implies

$$F_{23}^1 = a_{2\alpha}a_{3\beta}F_{\alpha\beta} = a_{22}a_{33}F_{23} \implies B_1^1 = B_1$$

The choice  $\mu = 1$  and  $\nu = 2$  gives  $F_{12} = B_3$ . Hence, (13.62) yields

$$F_{12}^1 = a_{1\alpha}a_{2\beta}F_{\alpha\beta} = a_{11}a_{22}F_{12} + a_{14}a_{22}F_{42}]$$

or

$$B_3^1 = aB_3 + ia\beta \left( \frac{iE_2}{c} \right) = a \left[ B_3 - \left( \frac{\beta}{c} \right) E_2 \right]$$

[2] For electric field components:

$$F_{24}^1 = a_{2\alpha}a_{4\beta}F_{\alpha\beta} = a_{22}a_{44}F_{24} + a_{22}a_{41}F_{21}]$$

or

$$\begin{aligned} -\left( \frac{E_2^1}{c} \right) &= a - \left( \frac{E_2}{c} \right) + ia\beta B_3 \\ \implies E_2^1 &= a(E_2 - c\beta B_3) \\ F_{14}^1 &= a_{1\alpha}a_{4\beta}F_{\alpha\beta} = a_{11}a_{44}F_{14} + a_{14}a_{41}F_{41} \\ \implies E_1^1 &= E_1 \\ F_{34}^1 &= a_{3\alpha}a_{4\beta}F_{\alpha\beta} = a_{33}a_{44}F_{34} + a_{33}a_{41}F_{31} \\ \implies E_3^1 &= a(E_3 + c\beta B_2) \end{aligned}$$

### Alternative

The electric field strength  $\vec{E}$  and magnetic field strength  $\vec{B}$  in the Maxwell's equations are defined in terms of electromagnetic potentials  $\vec{A}$  and  $\phi$  by

$$\vec{B} = \nabla \times \vec{A} \text{ and } \vec{E} = -\left( \frac{\partial \vec{A}}{\partial t} \right) - \nabla\phi$$

We know the transformation of electromagnetic potentials  $\vec{A}$  and  $\phi$

$$A_1^1 = a \left[ A_1 - \frac{v}{c^2} \phi \right], \quad A_2^2 = A_2, \quad A_3^1 = A_3, \quad \phi^1 = a [\phi - vA_1]$$

$$[\beta = v/c \text{ and } a = 1/\sqrt{1 - \beta^2}]$$

We also know

$$\frac{\partial}{\partial x_1^1} = a \left( \frac{\partial}{\partial x_1} + \frac{v}{c^2} \frac{\partial}{\partial t} \right), \quad \frac{\partial}{\partial x_2^1} = \frac{\partial}{\partial x_2}, \quad \frac{\partial}{\partial x_3^1} = \frac{\partial}{\partial x_3}, \quad \frac{\partial}{\partial t^1} = a \left( v \frac{\partial}{\partial x_1} + \frac{\partial}{\partial t} \right)$$

Now the Lorentz Transformations of electric field strength  $\vec{E}$  and magnetic field strength  $\vec{B}$  will be

$$B^1 = \nabla A^1 \quad (13.65)$$

$$\vec{E}^1 = - \left( \frac{\partial \vec{A}^1}{\partial t} \right) - \nabla \phi^1 \quad (13.66)$$

In terms of components the transformations of magnetic field strength are written as

$$iB_1^1 + jB_2^1 + kB_3^1 = i \left( \frac{\partial A_3^1}{\partial x_2^1} - \frac{\partial A_2^1}{\partial x_3^1} \right) + j \left( \frac{\partial A_1^1}{\partial x_3^1} - \frac{\partial A_3^1}{\partial x_1^1} \right) + k \left( \frac{\partial A_2^1}{\partial x_1^1} - \frac{\partial A_1^1}{\partial x_2^1} \right)$$

Now,

$$\begin{aligned} B_1^1 &= \left( \frac{\partial A_3^1}{\partial x_2^1} - \frac{\partial A_2^1}{\partial x_3^1} \right) = \left( \frac{\partial A_3}{\partial x_2} - \frac{\partial A_2}{\partial x_3} \right) = B_1 \\ B_2^1 &= \left( \frac{\partial A_1^1}{\partial x_3^1} - \frac{\partial A_3^1}{\partial x_1^1} \right) = \frac{\partial \left[ a \left( A_1 - \frac{v}{c^2} \phi \right) \right]}{\partial x_3} - a \left( \frac{\partial}{\partial x_1} + \frac{v}{c^2} \frac{\partial}{\partial t} \right) A_3 \\ \Rightarrow B_2^1 &= a \left[ B_2 + \left( \frac{\beta}{c} \right) E_3 \right] \\ B_3^1 &= \left( \frac{\partial A_2^1}{\partial x_1^1} - \frac{\partial A_1^1}{\partial x_2^1} \right) = a \left( \frac{\partial}{\partial x_1} + \frac{v}{c^2} \frac{\partial}{\partial t} \right) A_2 - \frac{\partial \left[ a \left( A_1 - \frac{v}{c^2} \phi \right) \right]}{\partial x_2} \\ \Rightarrow B_3^1 &= a \left[ B_3 - \left( \frac{\beta}{c} \right) E_2 \right] \end{aligned}$$

Similarly, the transformation components of electric field strength are given by

$$iE_1^1 + jE_2^1 + kE_3^1 = i \left( \frac{\partial A_1^1}{\partial t^1} - \frac{\partial \phi^1}{\partial x_1^1} \right) + j \left( \frac{\partial A_2^1}{\partial t^1} - \frac{\partial \phi^1}{\partial x_2^1} \right) + k \left( \frac{\partial A_3^1}{\partial t^1} - \frac{\partial \phi^1}{\partial x_3^1} \right)$$

Now,

$$E_1^1 = \left( \frac{\partial A_1^1}{\partial t^1} - \frac{\partial \phi^1}{\partial x_1^1} \right) = a \left( v \frac{\partial}{\partial x_1} + \frac{\partial}{\partial t} \right) \left[ a \left( A_1 - \frac{v}{c^2} \phi \right) \right]$$

$$\begin{aligned}
& -a \left( \frac{\partial}{\partial x_1} + \frac{v}{c^2} \frac{\partial}{\partial t} \right) [a(\phi - vA_1)] \\
\implies E_1^1 &= E_1 \\
E_2^1 &= \left( \frac{\partial A_2^1}{\partial t^1} - \frac{\partial \phi^1}{\partial x_2^1} \right) = a \left( v \frac{\partial}{\partial x_1} + \frac{\partial}{\partial t} \right) A_2 - \frac{\partial [a(\phi - vA_1)]}{\partial x_2} \\
\implies E_2^1 &= a(E_2 - vB_3) \\
E_3^1 &= \left( \frac{\partial A_3^1}{\partial t^1} - \frac{\partial \phi^1}{\partial x_3^1} \right) = a \left( v \frac{\partial}{\partial x_1} + \frac{\partial}{\partial t} \right) A_3 - \frac{\partial [a(\phi - vA_1)]}{\partial x_3} \\
\implies E_3^1 &= a(E_3 + vB_2)
\end{aligned}$$

The inverse transformation of electric and magnetic field components can be obtained by replacing  $v$  by  $-v$  and interchanging primes and unprimes.

$$E_1 = E_1^1, \quad E_2 = a[E_2^1 + c\beta B_3^1], \quad E_3 = a[E_3^1 - c\beta B_2^1] \quad (13.67)$$

$$B_1 = B_1^1, \quad B_2 = a \left[ B_2^1 - \left( \frac{\beta}{c} \right) E_3^1 \right], \quad B_3 = a \left[ B_3^1 + \left( \frac{\beta}{c} \right) E_2^1 \right] \quad (13.68)$$

*It is seen that electromagnetic field components are unaffected in the direction of motion, however, transverse components are modified.*

Here, the transformations of electric and magnetic field components can be written in terms of the components of the fields parallel ( $\parallel$ ) and perpendicular ( $\perp$ ) to the direction of the relative velocity between S and  $S^1$ .

$$E_{\parallel}^1 = E_{\parallel}, \quad E_{\perp}^1 = a \left[ E_{\perp} + (\vec{v} \times \vec{B}_{\perp}) \right] \quad (13.69)$$

$$B_{\parallel}^1 = B_{\parallel}, \quad B_{\perp}^1 = a \left[ B_{\perp} - \frac{1}{c^2} (\vec{v} \times \vec{E}_{\perp}) \right] \quad (13.70)$$

However, one can write the above transformation equations for electric and magnetic field components in more compact forms as

$$\vec{B}^1 = a \left[ \vec{B} - \frac{1}{c^2} (\vec{v} \times \vec{E}) + \frac{(\vec{v} \cdot \vec{B}) \vec{v}}{v^2} \left( \frac{1}{a} - 1 \right) \right] \quad (13.71)$$

$$\vec{E}^1 = a \left[ \vec{E} + \frac{1}{c^2} (\vec{v} \times \vec{B}) + \frac{(\vec{v} \cdot \vec{E}) \vec{v}}{v^2} \left( \frac{1}{a} - 1 \right) \right] \quad (13.72)$$

If one substitutes  $\vec{v} = (v, 0, 0)$  in Eqs. (13.71) and (13.72), then Eqs. (13.69) and (13.70) will recover.



It is interesting to note that a purely electric field or a purely magnetic field in one frame of reference will appear to be a combination of electric and magnetic fields in another frame of reference except the components in the direction of relative motion between the frames. Thus, if one of these fields is absent in one frame of reference, then that field may appear in another reference frame. For example, suppose in  $S$  frame there is no magnetic field. The transformation equation (13.71) implies that in  $S^1$  frame, magnetic field exists which is given as  $\vec{B}^1 = -\frac{1}{c^2}(\vec{v} \times \vec{E})$ .

### 13.14 Maxwell's Equations Are Invariant Under Lorentz Transformations

Maxwell's equations can be expressed in the covariant form with the help of electromagnetic four potentials  $A_\mu$  and current four potentials  $J_\mu$  along with the Lorentz-Gauge condition as

$$\square^2 A_\mu = -\mu_0 J_\mu, \quad \frac{\partial A_\mu}{\partial x_\mu} = 0$$

Now, to prove the invariance of Maxwell's equations under Lorentz Transformations, we will have to show that the above equations must retain the same form in  $S^1$  frame. Thus Maxwell's equations will be invariant if

$$\square^{12} A_\mu^1 = -\mu_0 J_\mu^1, \quad \frac{\partial A_\mu^1}{\partial x_\mu^1} = 0$$

We know that D'Alembertian operator is invariant under Lorentz Transformations i.e.  $\square^{12} = \square^2$ .

(i)  $\mu = 1$ :

$$\begin{aligned} \square^{12} A_1^1 &= \square^2 A_1^1 \\ &= \square^2 \left[ a \left( A_1 - \frac{v}{c^2} \phi \right) \right] \\ &= a \square^2 A_1 - a \frac{v}{c^2} \square^2 \phi \\ &= -a \mu_0 J_1 - a \frac{v}{c^2} \left( -\frac{\rho}{\epsilon_0} \right) \\ &= -a \mu_0 J_1 + a \rho \mu_0 v \\ &= -a \mu_0 (J_1 - v \rho) = -\mu_0 J_1^1 \end{aligned}$$

Thus

$$\square^{12} A_1^1 = -\mu_0 J_1^1$$

(ii)  $\mu = 2$ :

$$\begin{aligned}\square^{12} A_2^1 &= \square^2 A_2^1 \\ &= \square^2 A_2 = -\mu_0 J_2^1\end{aligned}$$

Hence

$$\square^{12} A_2^1 = -\mu_0 J_2^1, \quad \text{as } J_2^1 = J_2$$

(iii)  $\mu = 3$ :

$$\begin{aligned}\square^{12} A_3^1 &= \square^2 A_3^1 \\ &= \square^2 A_3 = -\mu_0 J_3^1\end{aligned}$$

That is

$$\square^{12} A_3^1 = -\mu_0 J_3^1, \quad \text{as } J_3^1 = J_3$$

(iv)  $\mu = 4$ :

$$\begin{aligned}\square^{12} A_4^1 &= \square^2 A_4^1 \\ &= \square^2 [a (A_4 - i\beta A_1)] \\ &= a \square^2 A_4 - i\beta \square^2 A_1 \\ &= -a\mu_0 J_4 - ai\beta(-\mu_0 J_1) \\ &= -a\mu_0 (J_4 - i\beta J_1) = -\mu_0 J_4^1\end{aligned}$$

Therefore

$$\square^{12} A_4^1 = -\mu_0 J_4^1$$

The cases (i)–(iv) indicate that

$$\square^{12} A_\mu^1 = -\mu_0 J_\mu^1$$

To show the invariance of Lorentz Gauge condition  $\partial A_\mu^1 / \partial x_\mu^1 = 0$ , we use the transformation of  $A_\mu^1$  which is given by

$$A_\mu^1 = a_{\mu\nu} A_\nu$$

This implies

$$A_\mu^1 = \frac{\partial x_\mu^1}{\partial x_\nu} A_\nu$$

or

$$\frac{\partial A_{\mu}^1}{\partial x_{\mu}^1} = \frac{\partial}{\partial x_{\mu}^1} \frac{\partial x_{\mu}^1}{\partial x_{\nu}} A_{\nu}$$

or

$$\frac{\partial A_{\mu}^1}{\partial x_{\mu}^1} = \frac{\partial A_{\nu}}{\partial x_{\nu}}$$

We know Lorentz Gauge condition  $\partial A_{\nu}/\partial x_{\nu} = 0$ , hence

$$\frac{\partial A_{\mu}^1}{\partial x_{\mu}^1} = 0$$

Thus Maxwell's equations are invariant under Lorentz Transformations.

### Alternative

We know Maxwell's equations are given by

$$\frac{\partial F^{\mu\nu}}{\partial x_{\nu}} = \mu_0 J^{\mu}, \quad \frac{\partial F_{\mu\nu}}{\partial x_{\lambda}} + \frac{\partial F_{\nu\lambda}}{\partial x_{\mu}} + \frac{\partial F_{\lambda\mu}}{\partial x_{\nu}} = 0$$

Therefore to prove its invariance, we will have to show

$$\frac{\partial F^{1\mu\nu}}{\partial x_{\nu}^1} = \mu_0 J^{1\mu}, \quad \frac{\partial F_{\mu\nu}^1}{\partial x_{\lambda}^1} + \frac{\partial F_{\nu\lambda}^1}{\partial x_{\mu}^1} + \frac{\partial F_{\lambda\mu}^1}{\partial x_{\nu}^1} = 0$$

Note that we have used the first equation in contravariant form of electromagnetic field tensor. This tensor transforms according to the rule given as

$$F^{1\mu\nu} = \frac{\partial x_{\mu}^1}{\partial x_{\beta}} \frac{\partial x_{\nu}^1}{\partial x_{\sigma}} F^{\beta\sigma}$$

Now, we have

$$\begin{aligned} \frac{\partial F^{1\mu\nu}}{\partial x_{\nu}^1} &= \frac{\partial}{\partial x_{\nu}^1} \left[ \frac{\partial x_{\mu}^1}{\partial x_{\beta}} \frac{\partial x_{\nu}^1}{\partial x_{\sigma}} F^{\beta\sigma} \right] \\ &= \frac{\partial x_{\alpha}}{\partial x_{\nu}^1} \frac{\partial}{\partial x_{\alpha}} \left[ \frac{\partial x_{\mu}^1}{\partial x_{\beta}} \frac{\partial x_{\nu}^1}{\partial x_{\sigma}} F^{\beta\sigma} \right] \\ &= \frac{\partial x_{\alpha}}{\partial x_{\nu}^1} \frac{\partial x_{\mu}^1}{\partial x_{\beta}} \frac{\partial x_{\nu}^1}{\partial x_{\sigma}} \frac{\partial}{\partial x_{\alpha}} [F^{\beta\sigma}] \end{aligned}$$

$$\begin{aligned}
&= \frac{\partial x_\mu^1}{\partial x_\beta} \delta_\sigma^\alpha \frac{\partial}{\partial x_\alpha} [F^{\beta\sigma}] \\
&= \frac{\partial x_\mu^1}{\partial x_\beta} \frac{\partial}{\partial x_\alpha} [F^{\beta\alpha}] \\
&= \frac{\partial x_\mu^1}{\partial x_\beta} \mu_0 J^\beta
\end{aligned}$$

Therefore, we get

$$\frac{\partial F^{1\mu\nu}}{\partial x_\nu^1} = \mu_0 J^{1\mu}$$

For the second Maxwell equation we have used the covariant form of electromagnetic field tensor. This tensor transforms according to the following rule

$$F^1{}_{\mu\nu} = \frac{\partial x_\beta}{\partial x_\mu^1} \frac{\partial x_\lambda}{\partial x_\nu^1} F_{\beta\lambda}$$

Now, we have

$$\begin{aligned}
\frac{\partial F^1{}_{\mu\nu}}{\partial x_\sigma^1} &= \frac{\partial}{\partial x_\sigma^1} \left[ \frac{\partial x_\beta}{\partial x_\mu^1} \frac{\partial x_\lambda}{\partial x_\nu^1} F_{\beta\lambda} \right] \\
&= \frac{\partial x_\alpha}{\partial x_\sigma^1} \frac{\partial}{\partial x_\alpha} \left[ \frac{\partial x_\beta}{\partial x_\mu^1} \frac{\partial x_\lambda}{\partial x_\nu^1} F_{\beta\lambda} \right]
\end{aligned}$$

Thus

$$\frac{\partial F^1{}_{\mu\nu}}{\partial x_\sigma^1} = \frac{\partial x_\alpha}{\partial x_\sigma^1} \frac{\partial x_\beta}{\partial x_\mu^1} \frac{\partial x_\lambda}{\partial x_\nu^1} \frac{\partial}{\partial x_\alpha} [F_{\beta\lambda}]$$

Similarly,

$$\begin{aligned}
\frac{\partial F^1{}_{\nu\sigma}}{\partial x_\mu^1} &= \frac{\partial x_\beta}{\partial x_\mu^1} \frac{\partial x_\lambda}{\partial x_\nu^1} \frac{\partial x_\alpha}{\partial x_\sigma^1} \frac{\partial}{\partial x_\beta} [F_{\lambda\alpha}] \\
\frac{\partial F^1{}_{\sigma\mu}}{\partial x_\nu^1} &= \frac{\partial x_\lambda}{\partial x_\nu^1} \frac{\partial x_\alpha}{\partial x_\sigma^1} \frac{\partial x_\beta}{\partial x_\mu^1} \frac{\partial}{\partial x_\lambda} [F_{\alpha\beta}]
\end{aligned}$$

After adding these three results, we get

$$\frac{\partial F^1{}_{\mu\nu}}{\partial x_\sigma^1} + \frac{\partial F^1{}_{\nu\sigma}}{\partial x_\mu^1} + \frac{\partial F^1{}_{\sigma\mu}}{\partial x_\nu^1} = \frac{\partial x_\alpha}{\partial x_\sigma^1} \frac{\partial x_\beta}{\partial x_\mu^1} \frac{\partial x_\lambda}{\partial x_\nu^1} \left[ \frac{\partial F_{\beta\lambda}}{\partial x_\alpha} + \frac{\partial F_{\lambda\alpha}}{\partial x_\beta} + \frac{\partial F_{\alpha\beta}}{\partial x_\lambda} \right] = 0$$

### 13.15 Lorentz Force on a Charged Particle

Now we determine the electromagnetic forces experienced by charged particle moving in an electromagnetic field. Let us consider two frames  $S$  and  $S^1$ ,  $S^1$  is moving with velocity  $v$  relative to  $S$  along  $x$  direction. Let a charged particle of charge  $q$  be rest in system  $S^1$ . This means the charged particle is moving with velocity  $\vec{v} = v\vec{i} = (v, 0, 0)$  relative to system  $S$ . Since the particle is at rest in system  $S^1$ , therefore no magnetic field exists in  $S^1$  i.e. field is purely electrostatic. In other words,

$$B_x^1 = B_y^1 = B_z^1 = 0 \quad (13.73)$$

But the particle is moving relative to  $S$ , therefore to an observer in  $S$  frame the purely electric field appears both as an electric and a magnetic field. We know the Lorentz transformation formulae of force components as

$$F_x^1 = F_x, \quad F_y^1 = \frac{F_y}{\sqrt{1 - \beta^2}}, \quad F_z^1 = \frac{F_z}{\sqrt{1 - \beta^2}} \quad (13.74)$$

[see the results given in Eq. (12.27)]

Note that if  $F$  is the electric field measured in system  $S$  due to charge stationary in it and  $\vec{F}^1$  is the electric field measured in system  $S^1$ , moving with velocity  $v$  relative to  $S$  in any direction, then resolving the electric field along and transverse to velocity  $v$ , are

$$F_{\parallel}^1 = F_{\parallel} \quad \text{and} \quad F_{\perp}^1 = \frac{F_{\perp}}{\sqrt{1 - \beta^2}}$$

where  $\parallel$  and  $\perp$  are along and perpendicular to the direction of velocity.

The force  $F^1$  in  $S^1$  system due to the electric field strength  $\vec{E}^1$  is  $\vec{F}^1 = q\vec{E}^1$  since  $q$  is invariant under Lorentz transformation. Writing explicitly we get

$$F_x^1 = qE_x^1, \quad F_y^1 = qE_y^1, \quad F_z^1 = qE_z^1 \quad (13.75)$$

Using (13.74), we get

$$\begin{aligned} F_x &= F_x^1 = qE_x^1, & F_y &= F_y^1 \sqrt{1 - \beta^2} = qE_y^1 \sqrt{1 - \beta^2}, \\ F_z &= F_z^1 \sqrt{1 - \beta^2} = qE_z^1 \sqrt{1 - \beta^2} \end{aligned} \quad (13.76)$$

However, we know Lorentz transformations of electric field components as

$$E_x^1 = E_x, \quad E_y^1 = \frac{E_y - vB_z}{\sqrt{1 - \beta^2}}, \quad E_z^1 = \frac{E_z + vB_y}{\sqrt{1 - \beta^2}} \quad (13.77)$$

Using (13.77), components of force from (13.76) are

$$F_x = qE_x, \quad F_y = q(E_y - vB_z), \quad F_z = q(E_z + vB_y) \quad (13.78)$$

Therefore we obtain,

$$\vec{F} = F_x \vec{i} + F_y \vec{j} + F_z \vec{k} = qE_x \vec{i} + q(E_y - vB_z) \vec{j} + q(E_z + vB_y) \vec{k}$$

or

$$\begin{aligned} \vec{F} &= q\vec{E} + qv(B_y \vec{k} - B_z \vec{j}) \\ &= q\vec{E} + q[v\vec{i} \times (B_x \vec{i} + B_y \vec{j} + B_z \vec{k})] \\ &= q\vec{E} + q(\vec{v} \times \vec{B}) \end{aligned}$$

i.e.,

$$\vec{F} = q[\vec{E} + \vec{v} \times \vec{B}] \quad (13.79)$$

The forces experienced by charged particle moving in an electromagnetic field is known as *Lorentz force* on a particle of charge  $q$ . Note that this Lorentz force is a direct consequences of the relativity principle.

### Lorentz Force in Covariant Form

We have seen that the equation of motion of a particle of rest mass  $m_0$  and charge  $q$  in an electromagnetic field ( $\vec{E}$  and  $\vec{B}$ ) is given by Lorentz force equation as

$$\vec{F} = q[\vec{E} + \vec{v} \times \vec{B}], \quad \text{where } \vec{v} \text{ is the velocity of the charge}$$

Let the force experienced per unit volume of charge density  $\rho$  is  $\vec{f}$ . Then the above expression gives

$$\vec{f} = \rho[\vec{E} + \vec{v} \times \vec{B}]$$

This implies

$$\vec{f} = \rho\vec{E} + \rho\vec{v} \times \vec{B} = \rho\vec{E} + \vec{J} \times \vec{B} \quad (13.80)$$

Writing explicitly this equation, we get,

$$f_1 = \rho E_1 + J_2 B_3 - J_3 B_2, \quad f_2 = \rho E_2 + J_3 B_1 - J_1 B_3, \quad f_3 = \rho E_3 + J_1 B_2 - J_2 B_1 \quad (13.81)$$

We know electromagnetic field strength tensor  $F_{\mu\nu}$  can be expressed in terms of the components of electric field strength  $\vec{E}$  and magnetic field strength  $\vec{B}$  as

$$F_{\mu\nu} = \begin{bmatrix} 0 & B_3 & -B_2 & -\frac{iE_1}{c} \\ -B_3 & 0 & B_1 & -\frac{iE_2}{c} \\ B_2 & -B_1 & 0 & -\frac{iE_3}{c} \\ \frac{iE_1}{c} & \frac{iE_2}{c} & \frac{iE_3}{c} & 0 \end{bmatrix}$$

Thus, Eq. (13.81) can be written in terms of the components of electromagnetic field strength tensor  $F_{\mu\nu}$  as

$$\begin{aligned} f_1 &= F_{11}J_1 + F_{12}J_2 + F_{13}J_3 + F_{14}J_4 \\ f_2 &= F_{21}J_1 + F_{22}J_2 + F_{23}J_3 + F_{24}J_4 \\ f_3 &= F_{31}J_1 + F_{32}J_2 + F_{33}J_3 + F_{34}J_4 \end{aligned}$$

Here  $J_\mu$  is the current four vector with the fourth component  $J_4 = ic\rho$ . Note that the right-hand sides of the above equations are the components of four vectors of electromagnetic field and current density. Therefore, it is expected that the above equations could be written in tensor form as

$$f_\mu = m_0 \frac{du_\mu}{d\tau} = F_{\mu\nu}J_\nu \quad (13.82)$$

Thus we have the expression of *four-dimensional Lorentz force* or the Lorentz Force in covariant form. Here the first three components in (13.82) represent the ordinary three-dimensional Lorentz force and the fourth component represents the rate of work done on the particle.

Here the fourth component is

$$\begin{aligned} f_4 &= F_{41}J_1 + F_{42}J_2 + F_{43}J_3 + F_{44}J_4 \\ &= \frac{iE_1}{c}J_1 + \frac{iE_2}{c}J_2 + \frac{iE_3}{c}J_3 + 0 = \frac{i}{c}\vec{E} \cdot \vec{J} = \frac{i\rho}{c}\vec{E} \cdot \vec{v} \end{aligned}$$

Thus the fourth component represents the amount of work done by the electric field.

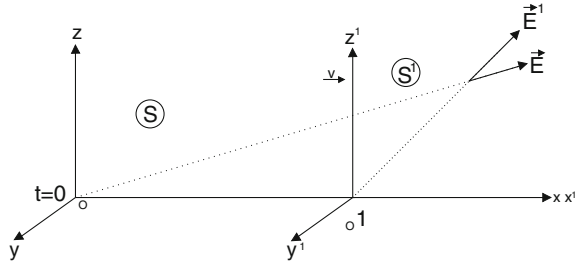
The expression given in (13.82) can also be written using Maxwell's equation in the following form:

$$f_\mu = \frac{1}{\mu_0} F_{\mu\nu} \frac{\partial F_{\nu\lambda}}{\partial x_\lambda} \quad \text{as} \quad \frac{\partial F_{\nu\lambda}}{\partial x_\lambda} = \mu_0 J_\nu \quad (13.83)$$

### 13.16 Electromagnetic Field Produced by a Moving Charge

Consider a particle of charge  $q$  moving with uniform velocity  $v$  in  $S$  system. We choose another system  $S^1$  moving with velocity  $v$  relative to  $S$  system in which this charge is at rest at the origin (see Fig. 13.1). Since the charge is at rest in  $S^1$ ,

**Fig. 13.1** Charge  $q$  and reference frame  $S^1$  are moving with velocity  $v$  along  $x$  direction



therefore it produces only electric field  $\vec{E}^1$  in  $S^1$ . Let P be any point whose position vector is  $\vec{r}^1$  relative to  $S^1$ . Therefore, from Coulomb's law, the electric field at P in  $S^1$  is given as

$$\vec{E}^1 = \frac{q\vec{r}^1}{4\pi\epsilon_0 r^{13}} \quad \text{here, } \vec{B}^1 = 0 \quad r^1 = |\vec{r}^1| \quad (13.84)$$

Writing explicitly, we have

$$E_x^1 = \frac{qx^1}{4\pi\epsilon_0 r^{13}}, \quad E_y^1 = \frac{qy^1}{4\pi\epsilon_0 r^{13}}, \quad E_z^1 = \frac{qz^1}{4\pi\epsilon_0 r^{13}} \quad (13.85)$$

$$B_x^1 = 0, \quad B_y^1 = 0, \quad B_z^1 = 0 \quad (13.86)$$

The reverse transformations relations give the components of electric and magnetic field ( $\vec{E}$  and  $\vec{B}$ ) at point P in S which are given as

$$E_x = E_x^1, \quad E_y = \frac{E_y^1 + vB_z^1}{\sqrt{1 - \beta^2}} = \frac{E_y^1}{\sqrt{1 - \beta^2}}, \quad E_z = \frac{E_z^1 - vB_y^1}{\sqrt{1 - \beta^2}} = \frac{E_z^1}{\sqrt{1 - \beta^2}} \quad (13.87)$$

$$B_x = B_x^1 = 0, \quad B_y = \frac{B_y^1 - \frac{v}{c^2}E_z^1}{\sqrt{1 - \beta^2}} = -\frac{\frac{v}{c^2}E_z^1}{\sqrt{1 - \beta^2}}, \quad B_z = \frac{B_z^1 - \frac{v}{c^2}E_y^1}{\sqrt{1 - \beta^2}} = \frac{\frac{v}{c^2}E_y^1}{\sqrt{1 - \beta^2}} \quad (13.88)$$

Let position vector of P be  $\vec{r}$  relative to S. Then the components of the radius vector  $\vec{r}$  are related to the components of the radius vector  $\vec{r}^1$  as

$$x^1 = \frac{x - vt}{\sqrt{1 - \beta^2}}; \quad y^1 = y; \quad z^1 = z; \quad t^1 = \frac{t - \frac{vx}{c^2}}{\sqrt{1 - \beta^2}}$$



At time  $t = 0$ , when  $q$  is at origin  $O$  of  $S$ , then the above equations give:

$$x^1 = \frac{x}{\sqrt{1-\beta^2}}; \quad y^1 = y; \quad z^1 = z; \quad \text{and} \quad t^1 = -\frac{\frac{vx}{c^2}}{\sqrt{1-\beta^2}}$$

Let  $\theta$  be the angle made by  $\vec{r}$  with  $x$ -axis, then

$$x = r \cos \theta \quad \text{where} \quad r^2 = x^2 + y^2 + z^2$$

This implies

$$y^2 + z^2 = r^2 \sin^2 \theta$$

We know that

$$r^{12} = x^{12} + y^{12} + z^{12}$$

Therefore, at  $t = 0$

$$r^{12} = \frac{x^2}{1-\beta^2} + y^2 + z^2 = \frac{r^2}{1-\beta^2} (1 - \beta^2 \sin^2 \theta)$$

Hence

$$r^1 = \frac{r}{\sqrt{1-\beta^2}} \left(1 - \beta^2 \sin^2 \theta\right)^{\frac{1}{2}}$$

Thus the components of electric field of charge  $q$  at  $t = 0$  as seen in  $S$  are given as

$$E_x = E_x^1 = \frac{1}{4\pi\epsilon_0} \frac{qx^1}{r^{13}} = \frac{1}{4\pi\epsilon_0} \frac{\frac{qx}{\sqrt{1-\beta^2}} (1-\beta^2)^{\frac{3}{2}}}{\left[r^3 (1-\beta^2 \sin^2 \theta)^{\frac{3}{2}}\right]} \quad (13.89)$$

$$E_y = \frac{E_y^1}{\sqrt{1-\beta^2}} = \frac{1}{\sqrt{1-\beta^2}} \frac{qy^1}{4\pi\epsilon_0 r^{13}} = \frac{1}{4\pi\epsilon_0} \frac{qy(1-\beta^2)}{r^3 (1-\beta^2 \sin^2 \theta)^{\frac{3}{2}}} \quad (13.90)$$

$$E_z = \frac{E_z^1}{\sqrt{1-\beta^2}} = \frac{1}{\sqrt{1-\beta^2}} \frac{1}{4\pi\epsilon_0} \frac{qz^1}{r^{13}} = \frac{1}{4\pi\epsilon_0} \frac{qz(1-\beta^2)}{r^3 (1-\beta^2 \sin^2 \theta)^{\frac{3}{2}}} \quad (13.91)$$

Writing in compact form, we get the electric field strength of charge  $q$  at  $t = 0$  as observed in  $S$

$$\vec{E} = iE_x + jE_y + kE_z = \frac{1}{4\pi\epsilon_0} \frac{q(1-\beta^2)\vec{r}}{r^3 (1-\beta^2 \sin^2 \theta)^{\frac{3}{2}}} \quad (13.92)$$

This indicates that  $\vec{E}$  is a radial field directed out along  $\vec{r}$ . We can write Eq. (13.88) in the following form:

$$\vec{E} = iE_x + jE_y + kE_z = \frac{1}{4\pi\epsilon_0} \frac{q(1-\beta^2)}{r^2(1-\beta^2\sin^2\theta)^{\frac{3}{2}}} \frac{\vec{r}}{r} \quad (13.93)$$

This implies the field is an inverse square one.

Similarly, the components of magnetic field  $\vec{B}$  in S follow explicitly from substituting  $\vec{B}^1$  and  $\vec{E}^1$  as

$$B_x = B_x^1 = 0 \quad (13.94)$$

$$\begin{aligned} B_y &= -\frac{\frac{v}{c^2}E_z^1}{\sqrt{1-\beta^2}} = -\frac{1}{4\pi\epsilon_0} \frac{\frac{v}{c^2}}{\sqrt{1-\beta^2}} \frac{qz^1}{r^{1^3}} \\ &= -\frac{1}{4\pi\epsilon_0 c^2} \frac{v}{\sqrt{1-\beta^2}} \frac{qz(1-\beta^2)^{\frac{3}{2}}}{r^3(1-\beta^2\sin^2\theta)^{\frac{3}{2}}} \end{aligned}$$

or

$$B_y = -\frac{\mu_0}{4\pi r^3} \frac{qvz(1-\beta^2)}{r^3(1-\beta^2\sin^2\theta)^{\frac{3}{2}}} \quad (\text{since } \mu_0\epsilon_0 = \frac{1}{c^2}) \quad (13.95)$$

$$\begin{aligned} B_z &= \frac{\frac{v}{c^2}E_y^1}{\sqrt{1-\beta^2}} = \frac{1}{4\pi\epsilon_0} \frac{\frac{v}{c^2}}{\sqrt{1-\beta^2}} \frac{qy^1}{r^{1^3}} \\ &= \frac{1}{4\pi\epsilon_0 c^2} \frac{v}{\sqrt{1-\beta^2}} \frac{qy(1-\beta^2)^{\frac{3}{2}}}{r^3(1-\beta^2\sin^2\theta)^{\frac{3}{2}}} \end{aligned}$$

or

$$B_z = \frac{\mu_0}{4\pi r^3} \frac{qvy(1-\beta^2)}{r^3(1-\beta^2\sin^2\theta)^{\frac{3}{2}}} \quad (13.96)$$

Writing in compact form, we get the magnetic field strength of charge  $q$  at  $t = 0$  as observed in S

$$\vec{B} = \vec{i}B_x + \vec{j}B_y + \vec{k}B_z = \frac{\mu_0}{4\pi} \frac{q(1-\beta^2)(\vec{v} \times \vec{r})}{(1-\beta^2\sin^2\theta)r^3} \quad (13.97)$$

This is *relativistic Biot-Savart Law* for magnetic field of a moving charge.

Note that  $\vec{v} = (v, 0, 0)$ , i.e. the moving direction is x-axis, so that x component of the magnetic field is zero. For classical limit,  $\beta \ll 1$ , Eq. (13.97) implies

$$\vec{B} = \frac{\mu_0}{4\pi} \frac{q(\vec{v} \times \vec{r})}{r^3} \quad (13.98)$$

This is the *classical Biot-Savart Law* for magnetic field of a moving charge.

### 13.17 Relativistic Lagrangian and Hamiltonian Functions of a Charged Particle in an Electromagnetic Field

Assume a particle of mass  $m_0$  and charge  $q$  moving with velocity  $\vec{v}$  in an electromagnetic field ( $\vec{E}$  and  $\vec{B}$ ). The force on the charged particle is given by Lorentz force

$$\vec{F} = q[\vec{E} + \vec{v} \times \vec{B}]$$

Magnetic field  $\vec{B}$  and electric field  $\vec{E}$  can be written in terms of electric potential  $\phi$  and magnetic potential  $\vec{A}$  as

$$\vec{B} = \nabla \times \vec{A}, \quad \vec{E} = -\left(\frac{\partial \vec{A}}{\partial t}\right) - \nabla \phi$$

Using these values, the Lorentz force takes the form

$$\vec{F} = q \left[ -\left(\frac{\partial \vec{A}}{\partial t}\right) - \nabla \phi + \vec{v} \times (\nabla \times \vec{A}) \right]$$

or

$$\vec{F} = q \left[ -\nabla \phi - \left(\frac{\partial \vec{A}}{\partial t} + (\vec{v} \cdot \nabla) \vec{A}\right) + \nabla(\vec{v} \cdot \vec{A}) \right] \quad (13.99)$$

The total derivative of  $\vec{A}$  can be expressed as

$$\frac{d\vec{A}}{dt} = \frac{\partial \vec{A}}{\partial t} + \sum_{i=1}^3 \frac{\partial \vec{A}}{\partial x_i} \frac{dx_i}{dt} = (\vec{v} \cdot \nabla) \vec{A} + \frac{\partial \vec{A}}{\partial t} \quad (13.100)$$

Using (13.100), the Eq. (13.99) yields

$$\vec{F} = q \left[ -\nabla \phi - \frac{d\vec{A}}{dt} + \nabla(\vec{v} \cdot \vec{A}) \right] = q \left[ -\nabla(\phi - \vec{v} \cdot \vec{A}) - \frac{d\vec{A}}{dt} \right] \quad (13.101)$$

Now, we write the x component of both sides of Eq. (13.101) as

$$F_1 = q \left[ -\frac{\partial}{\partial x} (\phi - \vec{v} \cdot \vec{A}) - \frac{d\vec{A}_1}{dt} \right] \quad (13.102)$$

One can write  $d\vec{A}_1/dt$  in the following form:

$$\frac{d\vec{A}_1}{dt} = \frac{d}{dt} \left[ \frac{\partial}{\partial \dot{x}} (\vec{v} \cdot \vec{A}) \right] \text{ as } \frac{\partial}{\partial \dot{x}} (\vec{v} \cdot \vec{A}) = \frac{\partial}{\partial v_1} (\vec{v} \cdot \vec{A}) = A_1$$

Again electric potential  $\phi$  is independent of the velocity, so  $\partial\phi/\partial\dot{x} = 0$ . Hence Eq. (13.102) can be written in the following form:

$$F_1 = -\frac{\partial U}{\partial x} + \frac{d}{dt} \left[ \frac{\partial U}{\partial \dot{x}} \right], \quad (13.103)$$

where

$$U = q(\phi - \vec{v} \cdot \vec{A}) \quad (13.104)$$

$U$  is the generalized potential.

Thus the relativistic Lagrangian  $L_{\text{em}}$  for a charged particle in an electromagnetic field is  $L_{\text{em}} = L_m - U$ , where  $L_m = m_0 c^2 \left( 1 - \sqrt{1 - \frac{v^2}{c^2}} \right)$  is the relativistic Lagrangian for the free particle. The terms in  $U$  describe the interaction of the charge with the field. Hence, we have

$$L_{\text{em}} = m_0 c^2 \left( 1 - \sqrt{1 - \frac{v^2}{c^2}} \right) - q\phi + q\vec{v} \cdot \vec{A} \quad (13.105)$$

For classical limit ( $v \ll c$ ), we have

$$L_{\text{em}} = \frac{1}{2} m_0 v^2 - q\phi + q\vec{v} \cdot \vec{A} \quad (13.106)$$

This Lagrangian indicates that the potential is velocity dependent.

The generalized momentum with components  $(p_1, p_2, p_3)$  of the charged particle in an electromagnetic field is defined as  $p_i = \partial L_{\text{em}} / \partial \dot{x}_i$ . Thus we get

$$p_1 = \frac{\partial L_{\text{em}}}{\partial \dot{x}} \equiv \frac{\partial L_{\text{em}}}{\partial v_1} = \left( \frac{m_0 \dot{x}}{\sqrt{1 - \frac{v^2}{c^2}}} + qA_1 \right)$$

Hence the generalized momentum is given by the vector equation

$$\vec{p} = \left( \frac{m_0 \vec{v}}{\sqrt{1 - \frac{v^2}{c^2}}} + q \vec{A} \right) \quad (13.107)$$

For classical limit ( $v \ll c$ ), we have

$$\vec{p} = (m_0 \vec{v} + q \vec{A}) \quad (13.108)$$

The Hamiltonian for a charged particle in an electromagnetic field is defined as

$$H_{\text{em}} = p_i \dot{x}_i - L_{\text{em}} = \vec{p} \cdot \vec{v} - L_{\text{em}}$$

Putting the values of  $\vec{p}$  and  $L_{\text{em}}$ , we get

$$H_{\text{em}} = \left( \frac{m_0 \vec{v} \cdot \vec{v}}{\sqrt{1 - \frac{v^2}{c^2}}} + q \vec{A} \cdot \vec{v} \right) - m_0 c^2 \left( 1 - \sqrt{1 - \frac{v^2}{c^2}} \right) + q\phi - q\vec{v} \cdot \vec{A}$$

or

$$H_{\text{em}} = m_0 c^2 \left( \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} - 1 \right) + q\phi \quad (13.109)$$

For classical limit ( $v \ll c$ ), we have

$$H_{\text{em}} = \frac{1}{2} m_0 v^2 + q\phi \quad (13.110)$$

One can also write the Hamiltonian in terms of momentum as

$$\left[ (H_{\text{em}} - q\phi) + m_0 c^2 \right]^2 - [(\vec{p} - q\vec{A}) \cdot (\vec{p} - q\vec{A})] c^2 = m_0 c^4$$

or

$$H_{\text{em}} = c \sqrt{(\vec{p} - q\vec{A})^2 + m_0 c^2} + q\phi - m_0 c^2 \quad (13.111)$$

*Example 13.1* Show that the charge of an electron is relativistically invariant.

**Hint:** Let us consider two frames S and  $S^1$ ,  $S^1$  is moving with velocity  $v$  relative to S along x direction. Then the transformation equations for current and charge density are

$$J_1^1 = a [J_1 - v\rho], \quad J_2^1 = J_2, \quad J_3^1 = J_3, \quad \rho^1 = a \left[ \rho - \left( \frac{v}{c^2} \right) J_1 \right],$$

where  $a = 1/\sqrt{1 - \frac{v^2}{c^2}}$ .

Let us consider the charged electron to be at rest in S frame, then obviously current density  $J$  is zero. Hence the above transformations become

$$J_1^1 = -av\rho, \quad J_2^1 = J_2, \quad J_3^1 = J_3, \quad \rho^1 = \frac{\rho}{\sqrt{1 - \frac{v^2}{c^2}}}$$

It is apparent that the charge contained in the elementary volume  $dV^1 = dx_1^1 dx_2^1 dx_3^1$  in  $S^1$  frame is

$$\begin{aligned} dq^1 &= \rho^1 dV^1 = \rho^1 dx_1^1 dx_2^1 dx_3^1 \\ &= \rho^1 \left( dx_1 \sqrt{1 - \frac{v^2}{c^2}} \right) dx_2 dx_3 \\ &= \rho dx_1 dx_2 dx_3 = dq \end{aligned}$$

Thus, both frames of reference S and  $S^1$  measure the same charge. In other words, charge of an electron is relativistically invariant.

*Example 13.2* Show that  $A^2 - \frac{\phi^2}{c^2}$  is invariant under Lorentz Transformation, where  $A^2 = A_1^2 + A_2^2 + A_3^2$ .

**Hint:** One can note that

$$A_\mu A_\mu = A_1 A_1 + A_2 A_2 + A_3 A_3 + A_4 A_4 = A^2 - \frac{\phi^2}{c^2}$$

Now,

$$\begin{aligned} A_\mu^1 A_\mu^1 &= A_1^1 A_1^1 + A_2^1 A_2^1 + A_3^1 A_3^1 + \frac{i\phi^1}{c} \frac{i\phi^1}{c} \\ &= a^2 \left[ A_1 - \frac{v}{c^2} \phi \right]^2 + A_2^2 + A_3^2 - \frac{1}{c^2} a^2 [\phi - vA_1]^2 \\ &= A^2 - \frac{\phi^2}{c^2} = A_\mu A_\mu \quad \left( \text{putting the value, } a = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \right) \end{aligned}$$

*Example 13.3* Show that  $J^2 - \rho^2 c^2$  is invariant under Lorentz Transformation, where  $J^2 = J_1^2 + J_2^2 + J_3^2$ .

**Hint:** As in Example (13.2), we note that

$$J_\mu J_\mu = J_1 J_1 + J_2 J_2 + J_3 J_3 + J_4 J_4 = J^2 - \rho^2 c^2 \text{ etc.}$$

*Example 13.4* Show that  $AJ - \phi\rho$  is invariant under Lorentz Transformation, where  $J^2 = J_1^2 + J_2^2 + J_3^2$ .

**Hint:** Here,

$$A_\mu J_\mu = A_1 J_1 + A_2 J_2 + A_3 J_3 + A_4 J_4 = AJ - \phi\rho \text{ etc.}$$

*Example 13.5* Find the non-relativistic limit of the inverse transformation of four current vector. Interpret the results.

**Hint:** The inverse transformations of four current vector are given as

$$J_1 = a \left[ J_1^1 + v\rho^1 \right], \quad J_2 = J_2^1, \quad J_3 = J_3^1, \quad \rho = a \left[ \rho^1 + \left( \frac{v}{c^2} \right) J_1^1 \right]$$

For non-relativistic limit  $v \ll c$ , the above results become

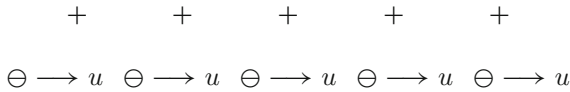
$$J_1 \approx J_1 + v\rho^1, \quad J_2 = J_2^1, \quad J_3 = J_3^1, \quad \rho \approx \rho^1 \text{ etc.}$$

*Example 13.6* Show that D'Alembertian operator  $\square^2$  is invariant under Lorentz Transformations but not Laplacian  $\nabla^2$ .

**Hint:** See the text.

*Example 13.7* If total charge is an invariant quantity, how can it be that a neutral wire in one frame appears to be charged in another frame?

**Hint:** Let us consider a long straight current carrying wire at rest in S frame. In this wire, the free electrons move with velocity  $u$  to the right and the positive ions are at rest.



Here, the number of free electrons and the number of ions per unit volume are the same, say  $n$ , so that the net charge in any volume of the wire is zero. In this consideration the positive charges (ions) are at rest in S frame, whereas electrons are in motion. Now, the negative charge density ( due to electrons ) is  $\rho^- = -ne$ , where  $e$  is the magnitude of an electronic charge and the positive charge density ( due to ions ) is  $\rho^+ = ne$ . Therefore, the net charge density measured in S frame is

$$\rho = \rho^- + \rho^+ = 0$$

For current density,  $J_x^- = -neu$ ,  $J_x^+ = 0$ , so that

$$J_x = J_x^- + J_x^+ = u\rho^-$$

Suppose  $S^1$  frame is moving with velocity  $v$  relative to  $S$  frame along  $x$  direction. Now, observer in  $S^1$  frame measures the current density as

$$\rho^{-1} = \frac{\rho^- - \frac{vJ_x^-}{c^2}}{\sqrt{1 - \frac{v^2}{c^2}}}, \quad \rho^{+1} = \frac{\rho^+ - \frac{vJ_x^+}{c^2}}{\sqrt{1 - \frac{v^2}{c^2}}}$$

Therefore, he obtains the net charge as ( putting the values of  $J_x^-, J_x^+, \rho^-, \rho^+$  )

$$\rho^1 = \rho^{-1} + \rho^{+1} = \frac{\frac{new}{c^2}}{\sqrt{1 - \frac{v^2}{c^2}}},$$

which is positive and non-zero. Thus observer in  $S^1$  frame finds the wire to be positively charged.

*Example 13.8* Show that  $c^2B^2 - E^2$  is invariant under Lorentz Transformation.

**Hint:** We will have to show

$$c^2B^{12} - E^{12} = c^2(B_1^{12} + B_2^{12} + B_3^{12}) - (E_1^{12} + E_2^{12} + E_3^{12}) = c^2B^2 - E^2$$

Use the transformations of electric and magnetic field components as

$$E_1^1 = E_1, \quad E_2^1 = a[E_2 - c\beta B_3], \quad E_3^1 = a[E_3 + c\beta B_2]$$

$$B_1^1 = B_1, \quad B_2^1 = a \left[ B_2 + \left( \frac{\beta}{c} \right) E_3 \right], \quad B_3^1 = a \left[ B_3 - \left( \frac{\beta}{c} \right) E_2 \right] \text{ etc.}$$

*Example 13.9* Show that  $\vec{E} \cdot \vec{B}$  is invariant under Lorentz Transformation.

**Hint:** Here one has to show

$$\vec{E}^1 \cdot \vec{B}^1 = E_1^1 B_1^1 + E_2^1 B_2^1 + E_3^1 B_3^1 = \vec{E} \cdot \vec{B}$$

*Example 13.10* Show that the continuity equation is self-contained in the homogeneous pair of Maxwell's equation.

**Hint:** The Maxwell's equation in terms of electromagnetic tensor  $F_{\mu\nu}$  is

$$\frac{\partial F_{\mu\gamma}}{\partial x_\gamma} = \mu_0 J_\mu \quad (1)$$

Differentiating (i) with respect to  $x_\mu$ , we get

$$\frac{\partial^2 F_{\mu\gamma}}{\partial x_\mu \partial x_\gamma} = \mu_0 \frac{\partial J_\mu}{\partial x_\mu} \quad (2)$$



Since  $F_{\gamma\mu}$  is antisymmetric, i.e.  $F_{\gamma\mu} = -F_{\mu\gamma}$ , Then from Eq. (ii) we get,

$$-\frac{\partial^2 F_{\gamma\mu}}{\partial x_\mu \partial x_\gamma} = \mu_0 \frac{\partial J_\mu}{\partial x_\mu} \quad (3)$$

Interchanging the dummy indices  $\mu$  and  $\gamma$  in the above equation, we get

$$-\frac{\partial^2 F_{\gamma\mu}}{\partial x_\gamma \partial x_\mu} = \mu_0 \frac{\partial J_\mu}{\partial x_\mu} \quad (4)$$

Using the property of perfect differential

$$\text{i.e.} \left( \frac{\partial^2}{\partial x_\mu \partial x_\gamma} = \frac{\partial^2}{\partial x_\gamma \partial x_\mu} \right)$$

$$-\frac{\partial^2 F_{\gamma\mu}}{\partial x_\mu \partial x_\gamma} = \mu_0 \frac{\partial J_\mu}{\partial x_\mu} \quad (5)$$

Equations (1) + (5) give

$$2\mu_0 \frac{\partial J_\mu}{\partial x_\mu} = 0$$

i.e.

$$\frac{\partial J_1}{\partial x_1} + \frac{\partial J_2}{\partial x_2} + \frac{\partial J_3}{\partial x_3} + \frac{\partial J_4}{\partial x_4} = 0$$

or

$$\frac{\partial J_1}{\partial x_1} + \frac{\partial J_2}{\partial x_2} + \frac{\partial J_3}{\partial x_3} + \frac{\partial (ic\rho)}{\partial (ict)} = 0$$

This implies

$$\nabla \cdot \vec{J} + \frac{\partial \rho}{\partial t} = 0,$$

which is the equation of continuity.

*Example 13.11* Show that a purely electric field in one frame appears both as an electric and a magnetic field to an observer moving relative to the first.

**Hint:** use Eqs. (13.69) and (13.70).

*Example 13.12* Show that a purely magnetic field in one frame appears both as an electric and a magnetic field to an observer moving relative to the first.

**Hint:** use Eqs. (13.69) and (13.70).

*Example 13.13* Starting from the four-dimensional form of homogeneous Maxwell's equation  $\partial F^{\mu\nu}/\partial x_\nu = 0$  obtain the wave equation for the field in a vacuum in the four-dimensional form.

**Hint:**

$$\frac{\partial F^{\mu\nu}}{\partial x_\nu} = 0 \implies \frac{\partial \left[ \frac{\partial A_\nu}{\partial x_\mu} - \frac{\partial A_\mu}{\partial x_\nu} \right]}{\partial x_\nu} = 0$$

or

$$\frac{\partial^2 A_\nu}{\partial x_\nu \partial x_\mu} - \frac{\partial^2 A_\mu}{\partial x_\nu \partial x_\nu} = 0$$

We know,

$$\frac{\partial^2 A_\nu}{\partial x_\nu \partial x_\mu} = \frac{\partial}{\partial x_\mu} \frac{\partial A_\nu}{\partial x_\nu} = 0$$

[The Lorentz-Gauge condition gives  $\partial A_\nu / \partial x_\nu = 0$ ]

Thus we get

$$\frac{\partial^2 A_\mu}{\partial x_\nu \partial x_\nu} = 0$$

In other words,

$$\square^2 A_\mu = 0 \quad \text{here, } \frac{\partial^2}{\partial x_\nu \partial x_\nu} \equiv \square^2 = \nabla^2 - \frac{1}{c^2} \frac{\partial^2}{\partial t^2}$$

*Example 13.14* Assuming the Lorentz-Gauge condition  $\partial A_\nu / \partial x_\nu = 0$  show that Maxwell's equation gives  $\partial^2 A_\mu / \partial x_\nu \partial x_\nu = -\mu_0 J_\mu$ .

**Hint:** Maxwell's equation gives

$$\frac{\partial F^{\mu\nu}}{\partial x_\nu} = \mu_0 J_\mu \implies \frac{\partial \left[ \frac{\partial A_\nu}{\partial x_\mu} - \frac{\partial A_\mu}{\partial x_\nu} \right]}{\partial x_\nu} = \mu_0 J_\mu$$

or

$$\frac{\partial^2 A_\nu}{\partial x_\nu \partial x_\mu} - \frac{\partial^2 A_\mu}{\partial x_\nu \partial x_\nu} = \mu_0 J_\mu$$

We know,

$$\frac{\partial^2 A_\nu}{\partial x_\nu \partial x_\mu} = \frac{\partial}{\partial x_\mu} \frac{\partial A_\nu}{\partial x_\nu} = 0$$

[Using the Lorentz-Gauge condition gives  $\partial A_\nu / \partial x_\nu = 0$ ]

Thus we get

$$\frac{\partial^2 A_\mu}{\partial x_\nu \partial x_\nu} = -\mu_0 J_\mu$$

*Example 13.15* Starting with Maxwell's equation show that  $\partial J_\mu / \partial x_\mu = 0$ .

**Hint:** See Example 13.10.

*Example 13.16* Using the transformation laws of electric  $E$  and magnetic  $B$  field components find the transformation laws for force.

**Hint:** The forces experienced by charged particle moving with velocity  $\vec{v}$  in an electromagnetic field is known as *Lorentz force* on a particle of charge  $q$  which is given by  $\vec{F} = q[\vec{E} + \vec{v} \times \vec{B}]$ . Writing explicitly, we have

$$F_1 = qE_1, \quad F_2 = q(E_2 - vB_3), \quad F_3 = q(E_3 + vB_2)$$

The components of the force in  $S^1$  system are due to electrostatic force

$$F_1^1 = qE_1^1, \quad F_2^1 = qE_2^1, \quad F_3^1 = qE_3^1$$

We also know the transformations of electric field components as

$$E_1^1 = E_1, \quad E_2^1 = a[E_2 - c\beta B_3], \quad E_3^1 = a[E_3 + c\beta B_2]$$

Therefore, we have

$$\begin{aligned} F_1^1 &= qE_1^1 = qE_1, & F_2^1 &= qE_2^1 = qa[E_2 - c\beta B_3] = aF_2, \\ F_3^1 &= qE_3^1 = aq[E_3 + c\beta B_2] = aF_3 \end{aligned}$$

*Example 13.17* Show that  $F_{\mu\nu}F_{\mu\nu} = 2[B^2 - \frac{E^2}{c^2}]$ .

**Hint:**

$$\begin{aligned} F_{\mu\nu}F_{\mu\nu} &= F_{1\nu}^2 + F_{2\nu}^2 + F_{3\nu}^2 + F_{4\nu}^2 \\ &= [F_{11}^2 + F_{21}^2 + F_{31}^2 + F_{41}^2] + [F_{12}^2 + F_{22}^2 + F_{32}^2 + F_{42}^2] \\ &\quad + [F_{13}^2 + F_{23}^2 + F_{33}^2 + F_{43}^2] + [F_{14}^2 + F_{24}^2 + F_{34}^2 + F_{44}^2] \end{aligned}$$

We know the explicit values of the components of the electromagnetic Field tensor  $F_{\mu\nu}$  as

$$F_{\mu\nu} = \begin{bmatrix} 0 & B_3 & -B_2 & -\frac{iE_1}{c} \\ -B_3 & 0 & B_1 & -\frac{iE_2}{c} \\ B_2 & -B_1 & 0 & -\frac{iE_3}{c} \\ \frac{iE_1}{c} & \frac{iE_2}{c} & \frac{iE_3}{c} & 0 \end{bmatrix}$$

Putting the values of the components of the electromagnetic field tensor, one can get,

$$F_{\mu\nu}F_{\mu\nu} = 2 \left[ B_1^2 + B_2^2 + B_3^2 - \frac{E_1^2}{c^2} - \frac{E_2^2}{c^2} - \frac{E_3^2}{c^2} \right] = 2 \left[ B^2 - \frac{E^2}{c^2} \right]$$

*Example 13.18* Show that  $1/64\pi^2(E^2+B^2)^2 - \frac{1}{16\pi^2}(\vec{E} \times \vec{B}) \cdot (\vec{E} \times \vec{B})$  is invariant under Lorentz Transformation.

**Hint:** Using this result

$$(\vec{E} \times \vec{B}) \cdot (\vec{E} \times \vec{B}) = E^2B^2 - (\vec{E} \cdot \vec{B})^2,$$

the given expression implies

$$\frac{1}{64\pi^2}(E^2+B^2)^2 - \frac{1}{16\pi^2}(\vec{E} \times \vec{B}) \cdot (\vec{E} \times \vec{B}) = \frac{1}{64\pi^2}[(E^2-B^2)^2 + 4(\vec{E} \cdot \vec{B})^2]$$

Here,  $(E^2 - B^2)$  and  $(\vec{E} \cdot \vec{B})$  are invariant, hence, etc.

*Example 13.19* Show that if  $\vec{E} \geq c\vec{B}$  in one inertial frame, then  $\vec{E}^1 \geq c\vec{B}^1$  in all other inertial frame.

**Hint:** We know  $c^2B^2 - E^2$  is invariant quantity, therefore,

$$c^2B^2 - E^2 = c^2B^2 - E^2 \leq 0 \implies \vec{E}^1 \leq c\vec{B}^1$$

*Example 13.20* Show that if electric field  $\vec{E}$  and magnetic field  $\vec{B}$  are perpendicular in one inertial frame, then they are perpendicular in all other inertial frames.

**Hint:** We know,  $\vec{E} \cdot \vec{B}$  is invariant quantity, therefore,

$$\vec{E}^1 \cdot \vec{B}^1 = \vec{E} \cdot \vec{B} = 0 \text{ since } \vec{E} \text{ and } \vec{B} \text{ are perpendicular}$$

*Example 13.21* Show that for a given electromagnetic field, we can find an inertial frame in which either  $\vec{B} = 0$  if  $\vec{E} \geq c\vec{B}$  or  $\vec{E} = 0$  if  $\vec{E} \leq c\vec{B}$  at a given point if  $\vec{E} \cdot \vec{B} = 0$  at that point.

**Hint:**

**Case I:** Let  $S^1$  be any frame in which no magnetic field is present. Therefore,

$$c^2B^2 - E^2 = c^2B^{1^2} - E^{1^2} = -E^{1^2} < 0 \quad \text{i.e. } \vec{E} \geq c\vec{B}$$

We know from Eq. (13.13)

$$B_{\parallel}^1 = B_{\parallel} = 0, \quad B_{\perp}^1 = a \left[ B_{\perp} - \frac{1}{c^2} (\vec{v} \times \vec{E}_{\perp}) \right] = 0$$

These give

$$\vec{B} = \frac{1}{c^2} (\vec{v} \times \vec{E})$$

From this expression, we see that

$$\vec{v} \cdot \vec{B} = \vec{v} \cdot \frac{1}{c^2} (\vec{v} \times \vec{E}) = 0$$

Therefore, velocity vector of  $S^1$  frame is perpendicular to the magnetic field. Since electric field exists in  $S^1$ , so the velocity vector of  $S^1$  frame is also perpendicular to the electric field. This implies the relative velocity vector between  $\vec{E}$  and  $\vec{B}$  frames is perpendicular to the plane containing S and  $S^1$  frames. Also,

$$\vec{E} \cdot \vec{B} = \vec{E} \cdot \frac{1}{c^2} (\vec{v} \times \vec{E}) = 0$$

Now,

$$\vec{E} \times \vec{B} = \vec{E} \times \frac{1}{c^2} (\vec{v} \times \vec{E}) = \frac{E^2}{c^2} \vec{v}$$

This implies

$$\vec{v} = \frac{c^2 (\vec{E} \times \vec{B})}{E^2}$$

Note that this type of frame is not unique.

**Case II:** Let  $S^1$  be any frame in which no electric field is present. Therefore,

$$c^2 B^2 - E^2 = c^2 B_{\perp}^2 - E^2 = c^2 B_{\perp}^2 > 0 \text{ i.e. } \vec{E} \leq c \vec{B}$$

We know from Eq. (13.12)

$$E_{\parallel}^1 = E_{\parallel} = 0, \quad E_{\perp}^1 = a[E_{\perp} + (\vec{v} \times \vec{B}_{\perp})] = 0$$

These give

$$\vec{E} = -(\vec{v} \times \vec{B})$$

From this expression, we see that

$$\vec{v} \cdot \vec{E} = \vec{v} \cdot (\vec{v} \times \vec{B}) = 0$$

Therefore, velocity vector of  $S^1$  frame is perpendicular to the electric field. Since magnetic field exists in  $S^1$ , therefore the velocity vector of  $S^1$  frame is also perpendicular to the magnetic field. This implies the relative velocity vector between  $\vec{E}$  and  $\vec{B}$  frames is perpendicular to the plane containing S and  $S^1$  frames. Also,

$$\vec{E} \cdot \vec{B} = -(\vec{v} \times \vec{B}) \cdot \vec{B} = 0$$

Now,

$$\vec{B} \times \vec{E} = \vec{B} \times [-(\vec{v} \times \vec{B})] = -B^2 \vec{v}$$

This implies

$$\vec{v} = \frac{c^2(\vec{E} \times \vec{B})}{B^2}$$

Note that this type of frame is not unique.

*Example 13.22* Find the trajectory of a particle of charge  $q$ , rest mass  $m_0$  moving in a uniform electrostatic field  $E$  along x-axis. Given that the initial momentum in the y direction is  $p_0$  and initial momentum in the x direction is zero.

**Hint:** Since the electrostatic field  $E$  is in x direction, we assume the initial velocity of the charged particle to be in the y direction. The the charged particle moves in the xy-plane. Here the equation of motion is

$$\vec{F} = \frac{d\vec{p}}{dt} = q\vec{E}$$

From the given condition,

$$\frac{dp_x}{dt} = qE, \quad \frac{dp_y}{dt} = 0$$

These yield

$$p_x = qEt, \quad p_y = p_0$$

[at  $t = 0$ ,  $p_x = 0$ ,  $p_y = p_0$ ] The total energy of the particle

$$T_E = \sqrt{p^2 c^2 + m_0^2 c^4} = \sqrt{c^2(p_x^2 + p_y^2) + m_0^2 c^4} = \sqrt{(c^2 p_0^2 + c^2 q^2 E^2 t^2) + m_0^2 c^4}$$

Again, we know

$$\vec{p} = m\vec{v} \text{ and } T_E = mc^2, \quad \text{therefore, } \vec{v} = \frac{\vec{p}}{T_E} c^2$$

This implies

$$v_x = \frac{dx}{dt} = \frac{p_x}{T_E} c^2 = \frac{qEt c^2}{\sqrt{(c^2 p_0^2 + c^2 q^2 E^2 t^2) + m_0^2 c^4}}$$

and

$$v_y = \frac{dy}{dt} = \frac{p_y}{T_E} c^2 = \frac{p_0 c^2}{\sqrt{(c^2 p_0^2 + c^2 q^2 E^2 t^2) + m_0^2 c^4}}$$

Solving these equations, we get

$$x = \frac{1}{qE} \sqrt{(c^2 p_0^2 + c^2 q^2 E^2 t^2) + m_0^2 c^4}$$

$$y = \frac{p_0 c}{qE} \sinh^{-1} \left( \frac{cqEt}{\sqrt{c^2 p_0^2 + m_0^2 c^4}} \right)$$

Eliminating  $t$  from these equations, we get the required trajectory of the charged particle as

$$x = \frac{\sqrt{c^2 p_0^2 + m_0^2 c^4}}{qE} \cosh \left( \frac{qEy}{p_0 c} \right)$$

# Chapter 14

## Relativistic Mechanics of Continua

### 14.1 Relativistic Mechanics of Continuous Medium (Continua)

Here we give our attention on the behaviour of the continuous distribution of matter—elastic bodies, liquids and gases. These so called continuous media possesses energy density, momentum density and stresses. In our previous discussion on relativistic electromagnetic theory, we have noticed that these quantities together are expressed in a four vector form and therefore are Lorentz covariant. So we would expect the same for the continuous Medium.

In the mechanics of continuous matter, we will deal with local volume element rather than the individual particle. A volume element contains sufficient number of individual particles that reflect the characteristic of continuous Medium.

Let  $\rho$  be the density and  $\vec{u}$  be the average velocity of matter; then the change of the density in a volume element is given by the non-relativistic continuity equation

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{u}) = 0 \tag{14.1}$$

This equation implies the influx of matter into the volume element. Let a force acting on a unit volume of matter be denoted by  $\vec{g}$ . Then one can write the equation of motion using Newton’s laws as

$$\rho \frac{d\vec{u}}{dt} = \vec{g} \tag{14.2}$$

Here the velocity  $\vec{u}$  of the unit volume of matter depends on space and time coordinates, i.e. on  $x_1, x_2, x_3$  and  $t$ .

Applying chain rule, we have

$$\frac{du_i}{dt} = \frac{\partial u_i}{\partial t} + \sum_{k=1}^{k=3} \frac{\partial u_i}{\partial x_k} \frac{\partial x_k}{\partial t} = \frac{\partial u_i}{\partial t} + u_{i,k} u_k \tag{14.3}$$



Here we have used summation convention and  $u_{i,k} = \frac{\partial u_i}{\partial x_k}$ .

By making use of Eq. (14.3), we get from (14.2)

$$\rho \left[ \frac{\partial u_i}{\partial t} + u_{i,k} u_k \right] = g_i \quad (14.4)$$

We can rearrange the left-hand side expression as

$$\begin{aligned} \rho \left[ \frac{\partial u_i}{\partial t} + u_{i,k} u_k \right] &= \frac{\partial(\rho u_i)}{\partial t} - u_i \frac{\partial \rho}{\partial t} + (\rho u_i u_k)_{,k} - u_i (\rho u_k)_{,k} \\ &= \frac{\partial(\rho u_i)}{\partial t} + (\rho u_i u_k)_{,k} - u_i \left[ \frac{\partial \rho}{\partial t} + (\rho u_k)_{,k} \right] \end{aligned}$$

Using continuity Eq. (14.1), finally we have

$$\rho \left[ \frac{\partial u_i}{\partial t} + u_{i,k} u_k \right] = \frac{\partial(\rho u_i)}{\partial t} + (\rho u_i u_k)_{,k} \quad (14.5)$$

Therefore, Eq. (14.4) yields

$$g_i = \frac{\partial(\rho u_i)}{\partial t} + (\rho u_i u_k)_{,k} \quad (14.6)$$

There are two types of forces acting on a continuous medium (fluid) namely external (volume forces) and internal or elastic forces (stresses). The external forces such as the forces due to gravitational field or electric field are acting on the matter in a given infinitesimal volume element proportional to the size of the volume element. Sometimes external forces are called volume or body forces.

Let  $F_i$  be the  $i$ th component of the external force exerted on the continuous medium (fluid), then

$$F_i = \int_V f_i dV \quad (14.7)$$

where  $V$  is the volume of the continuous medium (fluid). Here,  $f_i$  is the force per unit volume.

The second type of forces (internal) is the stress within the material itself, i.e. the forces due to mutual action of particles in the two adjoining volume elements which are proportional to the area of the surface of separation. These are also known as area forces. The components of the area force,  $dq_i$  depend linearly on the components of the elementary area  $dA_k$ .

$$dq_i = t_{ik} dA_k \quad (14.8)$$

Here  $dA_k$  is the element of vector area perpendicular to  $x_k$  direction. Thus  $dq_i$  is the force acting in the  $x_i$  direction. The stress  $t_{ik}$ , i.e. the force per unit area must be symmetric in its indices

$$t_{ik} = t_{ki} \quad (14.9)$$

Here the first index  $i$  indicates that the direction of the force is along  $x_i$  direction and second index implies the normal to the surface is along  $x_k$ . If the components of the stress  $t_{ik}$  are not symmetric, then the total angular momentum will change without applying any external (volume forces)  $f_i$  at the rate

$$\frac{dM_i}{dt} = \int_V \delta_{ikl} t^{kl} dV$$

[ $V$  is the volume of the whole body and  $\delta_{ikl}$  is the Levi-Civita tensor]

Now the  $i$ th component of the area force exerted on the continuous medium (fluid) inside a closed surface  $A$  by the fluid outside is given by

$$H_i = - \int_A t_{ik} dA_k = - \int_V t_{ik,k} dV \quad (\text{by Gauss law}) \quad (14.10)$$

where  $dA_k$  is an elementary vector surface area and  $V$  is the volume enclosed by  $A$ . Therefore, area force, i.e. internal force acting on a unit volume of the continuous medium is

$$h_i = -t_{ik,k} \quad (14.11)$$

Hence the total force which acts on the matter contained in a volume  $V$  is given as

$$G_i = \int_V f_i dV - \int_V t_{ik,k} dV = \int_V (f_i - t_{ik,k}) dV \quad (14.12)$$

Therefore, the force per unit volume is

$$g_i = f_i - t_{ik,k} \quad (14.13)$$

From (14.6) and (14.13), we obtain

$$f_i = \frac{\partial(\rho u_i)}{\partial t} + (\rho u_i u_k + t_{ik}),_k \quad (14.14)$$

*These three Eqs. (14.12)–(14.14) together with the continuity Eq. (14.1) determine the motion of a continuous medium in non-relativistic mechanics under the influence of internal and external forces.*

Thus if one knows the volume or external force  $f_i$  such as gravitational force or electromagnetic force and the stresses  $t_{ik}$  which depend on the internal deformation, then the a continuum system will be completely determined. For example, in case of perfect fluid,  $t_{ik} = p\delta_{ik}$  where  $p$  is the pressure.

To find the equation of continuity and equation of motion in a continuous medium in special theory of relativity, we use a special coordinate system. In a continuous medium it is convenient to use a special coordinate system  $S^0$  known as the co-moving coordinate system, with respect to which the matter is at rest at a world point  $P^0$ .

At this world point, the formulation of the equations is greatly simplified, because all classical terms containing undifferentiated velocity components vanish. The components of the four velocity  $U_\mu$  can be expressed as

$$U_i = au_i \quad \text{and} \quad U_4 = ic,$$

where  $a = \frac{1}{\sqrt{1-\frac{u^2}{c^2}}}$ .

We note that the first derivatives of the spatial part of the velocity  $U_i$  are equal to those of  $u_i$  where as the first derivative of the temporal part of the velocity  $U_4$  vanishes as  $U_4 = ic$ . In relativistic mechanics, energy flux is  $c^2$  times the momentum density. This follows from mass energy relation. The non-relativistic momentum density is  $\rho u_i$ . However, in addition, we have to consider the flux of energy on account of the stress.

Two same and opposite forces  $t_{ik}dA_k$  and  $-t_{ik}dA_k$  act on either side of the elementary vector surface area  $dA_k$  depending on which way the normal to the surface points. The amount of energy gained on one side is equal to that lost on the other. The amount of energy flux vector is  $u_i t_{ik}$  and the corresponding momentum density is  $\frac{u_i t_{ik}}{c^2}$ .

[ $E = mc^2$ ;  $E = t_{ik}, m = \frac{E}{c^2}$ , therefore, the momentum density  $\rho u_i = \frac{m}{V} u_i = \frac{u_i t_{ik}}{c^2}$  as  $V = \text{unit volume}$ ]

Hence the total relativistic momentum density of the continuous medium will be

$$P_k = \rho u_k + \frac{u_i t_{ik}}{c^2} \tag{14.15}$$

Here, we note that  $P_k$  vanishes in the special coordinate system but not their derivatives. Therefore, the equation of continuity in a continuous medium in special theory of relativity will be

$$\frac{\partial \rho}{\partial t} + P_{k,k} = 0 \tag{14.16}$$

Substituting the expression for  $P_k$  from (14.15), we obtain,

$$\frac{\partial \rho}{\partial t} + \rho u_{k,k} + \frac{u_{i,k} t_{ik}}{c^2} = 0 \quad (14.17)$$

As we know  $t$  is related to  $x_4$  as  $x_4 = ict$ , therefore above equation can be written as

$$ic\rho_{,4} + \rho u_{k,k} + \frac{u_{i,k} t_{ik}}{c^2} = 0 \quad (14.18)$$

*This is the equation of continuity in a continuous medium in special theory of relativity.*

Similarly, we can get the equation of motion by adding the contribution from the energy flux as

$$f_i = \frac{\partial \left( \rho u_i + \frac{u_i t_{ik}}{c^2} \right)}{\partial t} + (\rho u_i u_k + t_{ik})_{,k} \quad (14.19)$$

As before, we replace  $t$  by  $x_4 = ict$  and obtain *the motion of a continuous medium in relativistic mechanics under the influence of internal and external forces.*

$$f_j = ic\rho u_{j,4} + \frac{i}{c} u_{k,4} t_{kj} + (\rho u_j u_k + t_{jk})_{,k} \quad (14.20)$$

# Appendix A

The investigation of relations which remain valid when we change from coordinate system to any other is the chief aim of tensor calculus. The laws of physics can not depend on the frame of reference which the physicists choose for the purpose of description. Accordingly it is convenient to utilize the tensor calculus as the mathematical background in which such laws can be formulated.

## Transformation of Coordinates:

Let there be two reference systems  $S$  and  $S'$ .  $S'$  system depends on  $S$  system i.e,

$$x'^i = \phi^i(x^1, x^2, \dots, x^n) ; i = 1, 2, \dots, n \quad (\text{A.1})$$

Differentiation of (A.1) gives,

$$dx'^i = \frac{\partial \phi^i}{\partial x^r} dx^r = \frac{\partial x'^i}{\partial x^r} dx^r \quad (\text{A.2})$$

or,

$$dx^i = \frac{\partial x^i}{\partial x'^m} dx'^m \quad (\text{A.3})$$

Vectors in  $(S')$  system are related with  $(S)$  system.

There are two possible ways of transformations:

$$\bar{A}^i = \frac{\partial x^i}{\partial x^j} A^j$$

it is called **Contravariant vector**.

$$\bar{A}_i = \frac{\partial x^j}{\partial \bar{x}^i} A_j$$

it is called **Covariant vector**.

The expression  $A^i B_i$  is an invariant i.e.,

$$\bar{A}^i \bar{B}_i = A^i B_i.$$

$A^i B^j = A^{ij}$  form  $n^2$  quantities.

If the transformation of  $A^{ij}$  like

$$\bar{A}^{ij} = \frac{\partial \bar{x}^i}{\partial x^k} \frac{\partial \bar{x}^j}{\partial x^l} A^{kl}$$

then,  $A^{ij}$  is known as **Contravariant tensor of rank two**.

If

$$\bar{A}_{ij} = \frac{\partial x^k}{\partial \bar{x}^i} \frac{\partial x^l}{\partial \bar{x}^j} A_{kl}$$

then,  $A_{ij}$  is known as **Covariant tensor of rank two**.

**Mixed tensor of rank two** :  $A^i_j$

Here the transformation,

$$\bar{A}^i_j = \frac{\partial \bar{x}^i}{\partial x^k} \frac{\partial x^l}{\partial \bar{x}^j} A^k_l$$

$$T_{\alpha\beta}^\gamma = A_{\alpha\beta} B^\gamma \rightarrow \text{Outer Product.}$$

$$T_{\alpha\beta}^\alpha \rightarrow \text{Contraction.}$$

$$A^i_k B^k_{mn}$$

or

$$A^i_k B^m_{ij}$$

are known as **Inner Product**.

$$T_{\alpha\beta} = T_{\beta\alpha} \rightarrow \text{Symmetric}$$

$$T_{\alpha\beta} = -T_{\beta\alpha} \rightarrow \text{Antisymmetric.}$$

**The Line element:**

The distance between two neighbouring points is given by

$$ds^2 = g_{ij} dx^i dx^j. \tag{A.4}$$

Here,  $g_{ij}$  are functions of  $x^i$ . If  $g = |g_{ij}| \neq 0$ , then the space is called **Rimannian Space**.

$g_{ij}$  is called **fundamental tensor** (covariant tensor of order two).

In Euclidean space:

$$ds^2 = dx^2 + dy^2 + dz^2$$

In Minkowski flat space time, the line element

$$ds^2 = dx^0^2 - dx^1^2 - dx^2^2 - dx^3^2.$$

In order that the distance  $ds$  between two neighbouring points be real, the equation (A.4) will be amended to

$$ds^2 = e g_{ij} dx^i dx^j \quad [e = \pm 1]$$

Here  $e$  is called the indicator and takes the value  $+1$  or  $-1$  so that  $ds^2$  is always positive.

The contravariant tensor  $g^{ij}$  is defined by,

$$g^{ij} = \frac{\Delta_{ij}}{g},$$

here  $\Delta_{ij}$  is the cofactor of  $g_{ij}$ .

Obviously,

$$g_{ab} g^{bc} = \delta_a^c.$$

One can raise or lowering the indices of any tensor with the help of  $g^{ab}$  and  $g_{ab}$  as

$$\begin{aligned} g_{ac} T^{ab} &= T_c^b \\ T_{ab} g^{ac} &= T_b^c \\ g_{ab} A^b &= A_a \\ g^{ab} A_a &= A^b \\ g^{ab} g^{cd} g_{bd} &= g^{ac}. \end{aligned}$$

**Theorem:**

The expression  $\sqrt{-g} d^4x$  is an invariant volume element.

**Hints:** We know

$$g'_{ab} = \frac{\partial x^c}{\partial x'^a} \frac{\partial x^d}{\partial x'^b} g_{cd}$$

$\Rightarrow$

$$\det(g'_{ab}) = \left| \frac{\partial x}{\partial x'} \right|^2 \det(g_{cd})$$

$\Rightarrow$

$$g' = \left| \frac{\partial x}{\partial x'} \right|^2 g \quad (\text{A.5})$$

Also the volume element  $d^4x$  transform into  $d^4x'$  as

$$d^4x' = \left| \frac{\partial x'}{\partial x} \right| d^4x \quad (\text{A.6})$$

From (A.5) and (A.6) we get

$$\sqrt{-g'} d^4x' = \sqrt{-g} d^4x.$$

**Magnitude  $l$  of a vector  $A^\mu$**  is defined as

$$l^2 = A^\mu A_\mu = g_{\mu\nu} A^\mu A^\nu = g^{\mu\nu} A_\mu A_\nu$$

**Angle between two vectors  $A^\mu$  and  $B^\mu$ :**

We know in vector algebra, angle between two vectors  $\vec{A}$  and  $\vec{B}$  is defined as

$$\cos \theta = \frac{\vec{A} \cdot \vec{B}}{|\vec{A}| |\vec{B}|}$$

Similarly, the angle between two vectors  $A^\mu$  and  $B^\mu$  is defined as

$$\begin{aligned} \cos \theta &= \frac{\text{scalar product of } A^\mu \text{ and } B^\mu}{\text{length of } A^\mu \times \text{length of } B^\mu} \\ &= \frac{A^\mu B_\mu}{\sqrt{(A^\mu A_\mu)(B^\mu B_\mu)}} \\ &= \frac{g^{\mu\nu} A_\mu B_\nu}{\sqrt{g^{\alpha\beta} A_\alpha B_\beta}(g^{\rho\sigma} A_\rho B_\sigma)} \end{aligned}$$



**Levi Civita Tensor:**

$$\epsilon^{abcd} = +1,$$

if  $a, b, c, d$  is an even permutation of  $0, 1, 2, 3$ . i.e., in cyclic order.

$$= -1,$$

if  $a, b, c, d$  is odd permutation of  $0, 1, 2, 3$ . i.e., not in cyclic order.

$$= 0$$

if any two indices are equal.

The components of  $\epsilon_{abcd}$  are obtained from  $\epsilon^{abcd}$  by lowering the indices in the usual way, and multiplying it by  $(-g)^{-1}$ :

$$\epsilon_{abcd} = g_{a\mu} g_{b\nu} g_{c\gamma} g_{d\sigma} (-g)^{-1} \epsilon^{\mu\nu\gamma\sigma}.$$

For example,

$$\begin{aligned} \epsilon_{0123} &= g_{0\mu} g_{1\nu} g_{2\gamma} g_{3\sigma} (-g)^{-1} \epsilon^{\mu\nu\gamma\sigma} \\ &= (-g)^{-1} \det g_{\mu\nu} \\ &= -1 \end{aligned}$$

In general,

$$\epsilon_{abcd} = 1,$$

if  $a, b, c, d$  is an even permutation of  $0, 1, 2, 3$ .

$$= -1,$$

if  $a, b, c, d$  is odd permutation of  $0, 1, 2, 3$ .

$$= 0 \text{ otherwise.}$$

**Generalized Kronecker Delta:**

$\delta_{\mu\nu}^{\alpha\beta}$  can be constructed as follows:

$$\begin{aligned} \delta_{\mu\nu}^{\alpha\beta} &= \begin{vmatrix} \delta_{\mu}^{\alpha} & \delta_{\mu}^{\beta} \\ \delta_{\nu}^{\alpha} & \delta_{\nu}^{\beta} \end{vmatrix} \\ &= +1, \quad \alpha \neq \beta, \alpha\beta = \mu\nu \end{aligned}$$

$$= -1, \quad \alpha \neq \beta, \quad \alpha\beta = \nu\mu$$

$$= 0, \quad \text{otherwise.}$$

We can define  $\delta_{\mu\nu\xi}^{\alpha\beta\gamma}$  and  $\delta_{\mu\nu\xi\sigma}^{\alpha\beta\gamma\delta}$  as follows:

$$\delta_{\mu\nu\xi}^{\alpha\beta\gamma} = \begin{vmatrix} \delta_{\mu}^{\alpha} & \delta_{\mu}^{\beta} & \delta_{\mu}^{\gamma} \\ \delta_{\nu}^{\alpha} & \delta_{\nu}^{\beta} & \delta_{\nu}^{\gamma} \\ \delta_{\xi}^{\alpha} & \delta_{\xi}^{\beta} & \delta_{\xi}^{\gamma} \end{vmatrix}$$

$$= +1, \quad \alpha \neq \beta \neq \gamma, \quad \alpha\beta\gamma \text{ is an even permutation of } \mu\nu\xi$$

$$= -1, \quad \alpha \neq \beta \neq \gamma, \quad \alpha\beta\gamma \text{ is an odd permutation of } \mu\nu\xi$$

$$= 0, \quad \text{otherwise.}$$

$$\delta_{\mu\nu\xi\sigma}^{\alpha\beta\gamma\delta} = \begin{vmatrix} \delta_{\mu}^{\alpha} & \delta_{\mu}^{\beta} & \delta_{\mu}^{\gamma} & \delta_{\mu}^{\delta} \\ \delta_{\nu}^{\alpha} & \delta_{\nu}^{\beta} & \delta_{\nu}^{\gamma} & \delta_{\nu}^{\delta} \\ \delta_{\xi}^{\alpha} & \delta_{\xi}^{\beta} & \delta_{\xi}^{\gamma} & \delta_{\xi}^{\delta} \\ \delta_{\sigma}^{\alpha} & \delta_{\sigma}^{\beta} & \delta_{\sigma}^{\gamma} & \delta_{\sigma}^{\delta} \end{vmatrix}$$

Likewise the tensor  $\delta_{\mu\nu\xi\sigma}^{\alpha\beta\gamma\delta}$  can be expressed in a similar way.

*Example 1:* Show that

$$\delta_{\mu\beta}^{\alpha\beta} = 3\delta_{\mu}^{\alpha}$$

*Example 2:* Show that

$$\delta_{\alpha}^{\alpha} = 4$$

*Example 3:* Show that

$$\delta_{\mu\gamma\tau}^{\alpha\beta\tau} = 2\delta_{\mu\gamma}^{\alpha\beta}$$

**Hints:**

Here,

$$\delta_{\mu\gamma\tau}^{\alpha\beta\tau} = \begin{vmatrix} \delta_{\mu}^{\alpha} & \delta_{\mu}^{\beta} & \delta_{\mu}^{\tau} \\ \delta_{\gamma}^{\alpha} & \delta_{\gamma}^{\beta} & \delta_{\gamma}^{\tau} \\ \delta_{\tau}^{\alpha} & \delta_{\tau}^{\beta} & \delta_{\tau}^{\tau} \end{vmatrix}$$

Now, expand along third row and use  $\delta_{\tau}^{\tau} = 4$

## Appendix B

### Lagrangian and Euler-Lagrange's Equation:

Let Cartesian coordinate  $x_i$  can be written in terms of generalized coordinate  $q_k$  as

$$x_i = F_i(q_k, t) \quad (\text{B.1})$$

Now, one can get

$$dx_i = \frac{\partial x_i}{\partial q_k} dq_k + \frac{\partial x_i}{\partial t} dt \quad (\text{B.2})$$

The Cartesian velocity can be found as

$$v_i = \frac{dx_i}{dt} = \frac{\partial x_i}{\partial q_k} \dot{q}_k + \frac{\partial x_i}{\partial t} \quad (\text{B.3})$$

Therefore, we have the kinetic energy as

$$T = \frac{1}{2} \sum m_i v_i^2 = \frac{1}{2} m_i \left( \frac{\partial x_i}{\partial q_k} \dot{q}_k + \frac{\partial x_i}{\partial t} \right) \left( \frac{\partial x_i}{\partial q_k} \dot{q}_k + \frac{\partial x_i}{\partial t} \right) \quad (\text{B.4})$$

Thus,  $T = f(q_i, \dot{q}_i, t)$  and hence

$$\frac{\partial T}{\partial \dot{q}_k} = m_i v_i \frac{\partial v_i}{\partial \dot{q}_k} = m_i v_i \frac{\partial x_i}{\partial q_k} \quad (\text{B.5})$$

$$\frac{d}{dt} \left( \frac{\partial T}{\partial \dot{q}_k} \right) = m_i v_i \frac{d}{dt} \left( \frac{\partial x_i}{\partial q_k} \right) + m_i v_i \frac{\partial x_i}{\partial q_k} \quad (\text{B.6})$$

We have the following identity

$$\frac{d}{dt} \left( \frac{\partial x_i}{\partial q_k} \right) = \frac{\partial^2 x_i}{\partial q_l \partial q_k} \dot{q}_l + \frac{\partial}{\partial t} \left( \frac{\partial x_i}{\partial q_k} \right) = \frac{\partial}{\partial q_k} \left[ \frac{\partial x_i}{\partial q_l} \dot{q}_l + \frac{\partial x_i}{\partial t} \right] = \frac{\partial v_i}{\partial q_k} \quad (\text{B.7})$$

Putting the result of (B.7) in (B.6) we get,

$$\begin{aligned} \frac{d}{dt} \left( \frac{\partial T}{\partial \dot{q}_k} \right) &= m_i v_i \frac{\partial v_i}{\partial q_k} + m_i \dot{v}_i \frac{\partial x_i}{\partial q_k} = \frac{\partial T}{\partial q_k} + m_i \dot{v}_i \frac{\partial x_i}{\partial q_k} \frac{d}{dt} \left( \frac{\partial T}{\partial \dot{q}_k} \right) - \frac{\partial T}{\partial q_k} \\ &= m_i \dot{v}_i \frac{\partial x_i}{\partial q_k} \end{aligned} \quad (\text{B.8})$$

If we multiply both side by virtual changes  $\delta q_k$ , then we have

$$\left[ \frac{d}{dt} \left( \frac{\partial T}{\partial \dot{q}_k} \right) - \frac{\partial T}{\partial q_k} \right] \delta q_k = m_i \dot{v}_i \frac{\partial x_i}{\partial q_k} \delta q_k = m_i \ddot{x}_i \delta x_i \quad (\text{B.9})$$

Remind D'Alambert's Principle

$$(m_i \ddot{x}_i - X_i) \delta x_i = 0 \quad (\text{B.10})$$

For the conservative force, the potential energy  $V$  yields

$$\delta V = X_i \delta x_i = X_i \frac{\partial x_i}{\partial q_k} \delta q_k = Q_k \delta q_k = - \frac{\partial V}{\partial q_k} \delta q_k \quad (\text{B.11})$$

$Q_k = - \frac{\partial V}{\partial q_k}$  is called generalized force.

(B.9), (B.10) and (B.11)  $\Rightarrow$

$$\begin{aligned} \left[ \frac{d}{dt} \left( \frac{\partial T}{\partial \dot{q}_k} \right) - \frac{\partial T}{\partial q_k} \right] \delta q_k &= - \frac{\partial V}{\partial q_k} \delta q_k \\ \Rightarrow \left[ \frac{d}{dt} \left\{ \frac{\partial}{\partial \dot{q}_k} (T - V) \right\} - \frac{\partial}{\partial q_k} (T - V) \right] \delta q_k &= 0 \end{aligned}$$

(since  $V$  is not a function of  $\dot{q}_i$ )

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}_k} \right) - \frac{\partial L}{\partial q_k} = 0 \quad (\text{B.12})$$

$L = T - V \rightarrow$  Lagrangian and (B.12) **Euler-Lagrangian** Equations.

$p_i = \frac{\partial L}{\partial \dot{q}_i} \rightarrow$  generalized momentum.

**Note:** If the Lagrangian does not involve a certain coordinate that coordinate is called cyclic and the corresponding momentum will be a constant of motion (or a conserved quantity).

### Action Integral:

Almost all fundamental equations of physics can be obtained with the help of a variational principle, choosing suitable Lagrangian or action in different cases.

Hamilton's variational principle states that  $\delta S = \delta \int_1^2 L dt = 0 \Rightarrow$  Lagrange's equation of motion.

$\delta S = 0 \Rightarrow$  extremization of  $S \Leftrightarrow$  equations of motion.

Choose a point particle with mass  $m$  whose position is given by  $x_i(t)$  at time  $t$ . Let it is moving in a time independent potential  $V(x_i)$ . The action of this event is given by

$$S([x_i]; t_1, t_2) \equiv \int_{t_1}^{t_2} L dt = \int_{t_1}^{t_2} \left[ \frac{1}{2} m \frac{dx_i}{dt} \frac{dx_i}{dt} - V(x_i) \right] \quad (\text{B.13})$$

It is a functional of the path  $x_i(t)$  for  $t_1 < t < t_2$  which depends on initial and final times  $t_1, t_2$ .

(A functional is an operator whose range lies on the real time  $R$  or the complex plane  $C$ ).

Thus for a given path  $x_i(t)$ , we associate a number called the functional (in this case  $S$ ).

Consider the change of  $S$  for a small deformation of path

$$x_i(t) \rightarrow x_i(t) + \delta x_i(t) \quad (\text{B.14})$$

(note that there are no deformation i.e., no variations at the end points,  $\delta x_i(t_1) = \delta x_i(t_2) = 0$ )

Then

$$S[x_i + \delta x_i] = \int_{t_1}^{t_2} dt \left[ \frac{1}{2} m \frac{d}{dt} (x_i + \delta x_i) \frac{d}{dt} (x_i + \delta x_i) - V(x_i + \delta x_i) \right] \quad (\text{B.15})$$

Now

$$\frac{d}{dt} (x_i + \delta x_i) \frac{d}{dt} (x_i + \delta x_i) \approx \frac{dx_i}{dt} \frac{dx_i}{dt} + 2 \left( \frac{d}{dt} \delta x_i \right) \frac{dx_i}{dt}$$

(neglecting  $o(\delta x)^2$ )

$$\approx \frac{dx_i}{dt} \frac{dx_i}{dt} - 2 \frac{d^2 x_i}{dt^2} \delta x_i + 2 \frac{d}{dt} \left( \delta x_i \frac{dx_i}{dt} \right)$$

$$V(x_i + \delta x_i) \approx V(x_i) + \delta x_i \partial_i V \quad (\partial_i \equiv \frac{\partial}{\partial x_i})$$

Thus

$$S[x_i + \delta x_i] = S[x_i] + \int_{t_1}^{t_2} dt \delta x_i \left( -\partial_i V - m \frac{d^2 x_i}{dt^2} \right) + m \int_{t_1}^{t_2} dt \frac{d}{dt} \left( \delta x_i \frac{dx_i}{dt} \right) \quad (\text{B.16})$$

Since at the end points  $\delta x_i(t_1) = \delta x_i(t_2) = 0$ , the last term vanishes.

The equation (B.16) indicates that  $S$  does not change under an arbitrary  $\delta x_i$  if the classical equations of motion are valid.

Thus we have a clue of correspondence between equations of motion and extremization of action integral  $S$ .

**Note:** Here boundary conditions have to be supplied externally. This example dealt with particle mechanics. It can be generalized to classical field theory as in Maxwell's Electrodynamics or Einstein's General relativity.

### Lagrangian Formulation:

#### Newtonian Mechanics:

The action function is given by

$$S[q, \dot{q}] = \int_{t_1}^{t_2} L(q, \dot{q}) dt$$

Here the integration over a specific path  $q(t)$ . No variations,  $\delta q(t)$  of this path at the end points,  $\delta q(t_1) = \delta q(t_2) = 0$ .

Extremization of the action function  $\Rightarrow \delta S = 0$

i.e.,

$$\begin{aligned} 0 = \delta S &= \int_{t_1}^{t_2} \delta L dt = \int_{t_1}^{t_2} \left( \frac{\partial L}{\partial q} \delta q + \frac{\partial L}{\partial \dot{q}} \delta \dot{q} dt \right) \\ &= \left[ \frac{\partial L}{\partial \dot{q}} \delta q \right]_{t_1}^{t_2} + \int_{t_1}^{t_2} \left( \frac{\partial L}{\partial q} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} \right) \delta q dt \\ &\Rightarrow \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} - \frac{\partial L}{\partial q} = 0 \end{aligned}$$

This is Euler-Lagrangian equation for a one dimensional mechanical system.

#### Field Theory:

Here, we are interested the dynamical behavior of a field  $q(x^\alpha)$  in curved space time.

Let  $W$  be an arbitrary region in the space time manifold bounded by a closed hypersurface  $\partial W$ . The Lagrangian  $L(q, q_{,\alpha})$  is a scalar function of the field and its first derivative.

Thus action function is given by

$$S[q] = \int_W L(q, q_{,\alpha}) \sqrt{-g} d^4x$$

Note that the variation of  $q$  is arbitrary within  $W$  but vanishes on  $\partial W$ ,  $[\delta q]_{\partial W} = 0$   
Now,

$$\begin{aligned} \delta S &= \int_W \left[ \frac{\partial L}{\partial q} \delta q + \frac{\partial L}{\partial q_{,\alpha}} \delta q_{,\alpha} \right] \sqrt{-g} d^4x \\ &= \int_W [L' \delta q + (L^\alpha \delta q)_{;\alpha} - L^\alpha_{;\alpha} \delta q] \sqrt{-g} d^4x \end{aligned}$$

$$(L' \equiv \frac{\partial L}{\partial q}; \quad L^\alpha \equiv \frac{\partial L}{\partial q_{,\alpha}})$$

Using, Gauss divergence theorem, one gets

$$\delta S = \int_W (L' - L^\alpha_{;\alpha}) \delta q \sqrt{-g} d^4x + \oint_{\partial W} L^\alpha \delta q d\Sigma_\alpha$$

Since,  $[\delta q]_{\partial W} = 0$ , finally one can get

$$\delta S = 0 \Rightarrow \nabla_\alpha \frac{\partial L}{\partial q_{,\alpha}} - \frac{\partial L}{\partial q} = 0$$

This is Euler-Lagrange equation for a single scalar field  $\phi$  ( $\nabla$  stands for covariant derivative).

### Examples of Lagrangian for some fields:

(a) A scalar field  $\psi$ :

This can represent e.g., the  $\pi^0$  meson. The Lagrangian

$$L = \frac{1}{2} \left( g^{\mu\nu} \psi_{,\nu} \psi_{,\mu} + \frac{m^2}{\hbar^2} \psi^2 \right)$$

Euler Lagrangian equations are

$$g^{\alpha\beta} \psi_{;\alpha\beta} - \frac{m^2}{\hbar^2} \psi = 0$$

$$(L^\alpha = -g^{\alpha\beta} \psi_{;\beta}; \quad L^\alpha_{;\alpha} = -g^{\alpha\beta} \psi_{;\alpha\beta}; \quad L' = -\frac{m^2}{\hbar^2} \psi)$$

This is Klein-Gordon equation.

**(b) The electromagnetic field:**

$$L = \frac{1}{16\pi} F_{ab} F_{cd} g^{ac} g^{bd},$$

where electromagnetic tensor

$$F_{ab} = \partial_a A_b - \partial_b A_a;$$

$A_a \rightarrow$  electromagnetic potential.

Euler-Lagrange equation yields

$$F_{ab};c g^{bc} = 0$$

**(c) General Relativity:**

The Einstein's field equations in general relativity can be derived from the principle of Least Action  $\delta S = 0$  where

$$S = \int_W \sqrt{-g} [L_G + 2kL_F] d^4x$$

( $W$  be an arbitrary region in the space time manifold bounded by a closed hypersurface  $\partial W$ )

Here we choose  $L_G = R$  as the Lagrangian for the gravitational field with  $R = g_{\mu\nu} R^{\mu\nu}$ .  $L_F$  is the Lagrangian for all the other fields ( $\phi$ ).  $k =$  Einstein's gravitational constant  $= \frac{8\pi G}{c^4}$ .

The Einstein field equation in general relativity,  $R_{\alpha\beta} - \frac{1}{2}g_{\alpha\beta}R = kT_{\alpha\beta}$  are recovered by varying  $S[g; \phi]$  with respect to  $g_{\alpha\beta}$ . The variation is subjected to the condition

$$[\delta g_{\alpha\beta}]_{\partial W} = 0.$$



# References

1. Bandopadhyaya, N.: Special Theory of Relativity. Academic Publishers, Dordrecht (1998)
2. Banerjee, S., Banerjee, A.: The Special Theory of Relativity. Prentice Hall of India, New Delhi (2002)
3. Barton, G.: Introduction to the Relativity Principle. Wiley, New York (1999)
4. Bergmann, P.G.: Introduction to the Theory of Relativity. Prentice Hall of India, New Delhi (1992)
5. Carmeli, M.: Cosmological Special Relativity. World Scientific, Singapore (2002)
6. Das, A.: The Special Theory of Relativity, Springer, New York (1993)
7. D'Inverno, R.: Introducing Einstein's Relativity. Clarendon Press, Oxford (1992)
8. Einstein, A.: On the electrodynamics of moving bodies. <http://www.fourmilab.ch/etexts/einstein/specrel/www/> (1905). Accessed 30 June 1905
9. Einstein, A., Lorentz, H.A., Weyl, H., Minkowski, H.: The Principle of Relativity. Dover Publication, New York (1924)
10. French, A.P.: Special Relativity. W W Norton, New York (1968)
11. Gibilisco, S.: Understanding Einstein's Theories of Relativity. Dover publication, New York (1983)
12. Goldstein, H.: Classical Mechanics. Narosa, New Delhi (1998)
13. Griffiths, D.J.: Electrodynamics. Wiley Eastern, New Delhi (1978)
14. Jackson, J.D.: Electrodynamics. Wiley Eastern, New Delhi (1978)
15. Kogut, J.B.: Introduction to Relativity, Academic Press, New York (2000)
16. Landau, L.D., Lifshitz, E.M.: Classical Theory of Fields, Pergamon Press, New York (1975)
17. Lord, E.A.: Tensor Relativity and Cosmology. Tata McGraw-Hill, New Delhi (1976)
18. Misner, C.W., Thorne, K., Wheeler, J.: Gravitation. Freeman, San Francisco (1973)
19. Moller, C.: The Theory of Relativity. Oxford University Press, Oxford (1972)
20. Naber, G.L.: The Geometry of Minkowski Spacetime: An Introduction to the Mathematics of the Special Theory of Relativity, Springer, New York (1992)
21. Ney, E.P.: Electromagnetism and Relativity. John Weatherhill, Tokyo (1965)
22. Pathria, R.K.: The Theory of Relativity. Dover Publications, New York (1974)
23. Pauli, W.: Theory of Relativity, Dover Publications, New York (1958)
24. Poisson, E.: A Relativist's Toolkit. Cambridge University Press, Cambridge (2004)
25. Prakash, S.: Relativistic Mechanics. Pragati Prakashan, Meerut (2000)
26. Puri, S.P.: Special Theory of Relativity. Pearson, India (2013)
27. Resnick, R.: Introduction to Special Relativity. Wiley Eastern Ltd., New Delhi (1985)
28. Rindler, W.: Relativity. Clarendon Press, Oxford (2002)

29. Rosser, W.G.V.: *An Introduction to the Theory of Relativity*. Butterworths, London (1964)
30. Schwartz, H.M.: *Introduction to Special Relativity*. McGraw Hill, New York (1968)
31. Sharipov, R.: *Classical Electrodynamics and Theory of Relativity*. Samizdat Press (2003)
32. Srivastava, M.P.: *Special Theory of Relativity*. Hindustan Publishing Corporation, New Delhi (1992)
33. Synge, J.L., Griffith, B.A.: *Classical Mechanics*, McGraw Hill, New York (1949)
34. Taylor, E., Wheeler, J.: *Spacetime Physics*. Freeman, New York (1992)
35. Taylor, J.G.: *Special Relativity*. Oxford University, Oxford (1965)
36. Tolman, R.: *Relativity, Thermodynamics and Cosmology*. Dover, New York (1934)
37. Tourrence, P.: *Relativity and Gravitation*. Cambridge University Press, Cambridge (1992)
38. Ugare, V.A.: *Special Theory of Relativity*. Mir Publishers, Moscow (1979)
39. Weinberg, S.: *Gravitation and Cosmology*. Wiley, New York (1972)
40. Williams, W.S.C.: *Introducing Special Relativity*. Taylor and Francis, London (2002)
41. Woodhouse, N.M.J.: *Special Relativity*. Springer, Berlin (2003)

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